

## Background-independent condensed matter models for quantum gravity

This article has been downloaded from IOPscience. Please scroll down to see the full text article.

2011 New J. Phys. 13 095006

(<http://iopscience.iop.org/1367-2630/13/9/095006>)

View [the table of contents for this issue](#), or go to the [journal homepage](#) for more

Download details:

IP Address: 194.94.224.254

The article was downloaded on 10/01/2012 at 12:37

Please note that [terms and conditions apply](#).

## Background-independent condensed matter models for quantum gravity

**Alioscia Hamma<sup>1</sup> and Fotini Markopoulou<sup>1,2,3,4,5</sup>**

<sup>1</sup> Perimeter Institute for Theoretical Physics, 31 Caroline St. N, Waterloo, ON, N2L 2Y5, Canada

<sup>2</sup> Department of Physics, University of Waterloo, Waterloo, ON, N2L 3G1, Canada

<sup>3</sup> Max Planck Institute for Gravitational Physics (Albert Einstein Institute), Am Mühlenberg 1, D-14476 Golm, Germany

<sup>4</sup> Santa Fe Institute, 1399 Hyde Park Road, Santa Fe, NM 87501, USA

E-mail: [fotini@perimeterinstitute.ca](mailto:fotini@perimeterinstitute.ca)

*New Journal of Physics* **13** (2011) 095006 (26pp)

Received 26 November 2010

Published 14 September 2011

Online at <http://www.njp.org/>

doi:10.1088/1367-2630/13/9/095006

**Abstract.** A number of recent proposals on a quantum theory of gravity are based on the idea that spacetime geometry and gravity are derivative concepts and only apply at an approximate level. There are two fundamental challenges to any such approach. At the conceptual level, there is a clash between the ‘timelessness’ of general relativity and emergence. Secondly, the lack of a fundamental spacetime renders difficult the straightforward application of well-known methods of statistical physics to the problem. We recently initiated a study of such problems using spin systems based on the evolution of quantum networks with no *a priori* geometric notions as models for emergent geometry and gravity. In this paper, we review two such models. The first model is a model of emergent (flat) space and matter, and we show how to use methods from quantum information theory to derive features such as the speed of light from a non-geometric quantum system. The second model exhibits interacting matter and geometry, with the geometry defined by the behavior of matter. This model has primitive notions of gravitational attraction that we illustrate with a toy black hole, and exhibits entanglement between matter and geometry and thermalization of the quantum geometry.

<sup>5</sup> Author to whom any correspondence should be addressed.

**Contents**

<b>1. Introduction</b>	<b>2</b>
<b>2. Spin systems on a dynamical lattice</b>	<b>5</b>
2.1. Graphs and space . . . . .	6
2.2. Purpose of models . . . . .	7
<b>3. Quantum graphity: a model of emergent geometry and matter</b>	<b>7</b>
3.1. Speed of light via the Lieb–Robinson (LR) bound . . . . .	10
3.2. Transition in the causal structure . . . . .	11
3.3. From a dynamical to a fixed lattice . . . . .	12
<b>4. A unitary model of interacting matter and geometry</b>	<b>13</b>
4.1. A toy black hole . . . . .	16
<b>5. Matter–geometry entanglement and thermalization</b>	<b>17</b>
5.1. Thermalization via subsystem dynamics . . . . .	17
5.2. Thermalization in our model . . . . .	19
<b>6. Summary and outlook</b>	<b>20</b>
<b>Acknowledgments</b>	<b>25</b>
<b>References</b>	<b>25</b>

**1. Introduction**

Research in quantum gravity is currently at a very interesting point. In the past decade, the field has been presented with a great opportunity and challenge: quantum gravity may have observational consequences. In place of the long-standing frustration, that quantum effects of the gravitational field become important only at  $10^{-35}$  m, a tiny number we believed to be beyond observations, we now have investigations on different sources of observational data, ranging from astrophysical (where large distances may amplify the tiny quantum gravity signals) to observational cosmology (where the CMB spectrum may carry signs of quantum gravity effects) to the LHC (for a recent review of quantum gravity phenomenology, see [1]). In all these cases, the theoretical explanation of the observations may require new physics involving quantum gravity. It is exciting and challenging to find a theory that connects to these observations.

In this new environment, the novel direction of emergent gravity is particularly relevant. The central idea is that spacetime geometry and gravity are derivative concepts and only apply at an approximate level. Quantum gravity is one more situation in physics where we have the low-energy theory—general relativity and quantum field theory—and are looking for the high-energy, microscopic one. As an analogy, consider the example of going from thermodynamics to the kinetic theory. What we currently know is the low-energy theory, the analogue of fluid dynamics. We are looking for the microscopic theory, the analogue of the quantum molecular dynamics. Just as there are no waves in the molecular theory, we may not find geometric degrees of freedom in the fundamental theory. Not surprisingly, this significant shift in perspective opens up new routes that may take us out of the old problems.

Results in this direction include long-standing topics such as black hole thermodynamics [2] and matrix models [3], the AdS/CFT correspondence and the appearance of gravitons in string theory, general relativity as thermodynamics [7], Horava gravity [6],

emergence of a Lorentzian metric and aspects of gravitation such as Hawking radiation in analogue models of gravity [4] and condensed matter approaches such as Volovik's work on emergent Lorentz invariance at the Fermi point [5] and Wen's work on emergent matter and gravitons from a bosonic spin system [16], the broader idea of spacetime as a condensate [8] or aspects of emergence in causal dynamical triangulations (CDT) [9] and group field theory and similar ideas in simplicial [43] and causal models [42]. While approaches of this kind are new and therefore less established, this is a very promising and timely direction, especially in the light of the opportunities and challenges to quantum gravity research coming from the recent developments in observations. At the same time, results on emergent gravity are mainly a collection of hints and analogies. We are still lacking a microscopic theory that unifies and provides context to the existing results and that, of course, can be tested against experiment.

We are particularly interested in what the emergence paradigm has to say about the problem of time. The 'problem of time' refers to the clash in the roles that time plays in general relativity and quantum theory, and is considered fundamental in quantum gravity research since many other problems can be traced to it. We can review it very briefly as follows (for classic accounts of the issue in far more depth, see, e.g., [10]).

In general relativity, spacetime, the four-dimensional (4D) curved manifold with metric  $g_{\mu\nu}$  and curvature  $R_{\mu\nu}$ , and matter, the stress–energy tensor  $T_{\mu\nu}$ , are dynamical and affect each other: matter tells spacetime how to curve and spacetime tells matter where to go. This is the content of the Einstein equations

$$R_{\mu\nu} - \frac{1}{2}g_{\mu\nu}R = \frac{8\pi G}{c^4}T_{\mu\nu}, \quad (1)$$

where  $G$  is Newton's constant and  $c$  is the speed of light. The Einstein equations are invariant under diffeomorphisms of the spacetime manifold, operations that map spacetime points to other spacetime points. This means that spacetime coordinates are not physical; instead, events and their relations are physical. A given collection of events, and the order in which they occurred, is physically meaningful but can be represented by several different metrics, all related by diffeomorphisms. The correct physical quantity is the equivalence class of metrics under diffeomorphisms, usually called a *geometry*. Now let us consider a pure gravity scenario, a universe with no matter, where the right-hand side of equation (1) is zero. With no physical objects to mark events, it is very difficult to construct observables that measure local physical properties. In the 3 + 1 canonical decomposition of the Einstein action, evolution is a diffeomorphism and it is impossible to locally disentangle the effects of change from the effects of changing coordinates. The Hamiltonian for the evolution of space is just a constraint. This is what is often called the *timelessness* of general relativity.

Timelessness becomes significant in quantum gravity. Quantum gravity is often seen as the problem of unifying or reconciling general relativity and quantum theory, combining the physics of the very large with that of the very small. Quantum theory always uses a fixed spacetime, and that time is a parameter, not even an operator. The incompatibility of general relativity and quantum theory can be stated in many ways and a classic one is *the problem of time*: adapting the Schrödinger equation to a diffeomorphism invariant context by quantizing equation (1) gives the Wheeler–deWitt equation

$$\hat{H}|\Psi_U\rangle = 0. \quad (2)$$

Here,  $|\Psi_U\rangle$  is the wavefunction of the universe,  $\hat{H}$  is the quantum Hamiltonian constraint and 0 means there is no time.

We believe that an emergence scenario for gravity can open up new potential solutions to the problem of time, mainly in two ways:

1. *No pure gravity.* Timelessness in the argument above arises in a universe of pure gravity. If we allow matter, it is possible to define clocks and time. While we do live in a universe with matter, general relativity allows for pure gravity solutions and there is no way to exclude them in the current setup. The emergence of gravity may provide a mechanism for excluding pure gravity solutions, for example if gravity is simply the thermodynamics of matter and there is no gravity without matter.
2. *Fundamental versus geometric time.* If gravity is emergent, then geometry and diffeomorphisms should also be emergent. It is then possible that one can have a fundamental time, consistent with quantum mechanics, at the microscopic level, and an emergent or geometric time, the  $g_{00}$  component of the emergent metric, macroscopically. Quantum theory and general relativity could be reconciled if we have a reason why they should apply at different levels, and emergence can provide that reason<sup>6</sup>.

Our aim is twofold. First, we wish to study emergent gravity from the perspective of background independence and time in quantum gravity. In addition, we would like to see a unified microscopic picture behind all the various hints of why gravity may be emergent. Emergence, for us, means that as we go to the Planck scale, the Hamiltonian model becomes very complicated and messy. It is at low energies that symmetries and the simplicity of physical laws emerge. This kind of emergence is very anti-reductionist, because the building blocks of Nature can be very complicated (and not fundamental; they are perhaps made of even more complicated building blocks), whereas the emergent structures are simple and beautiful. In the case of gravity, this means, for example, that diffeomorphisms should appear as an emergent symmetry.

We will take a straightforward approach to the problem. Emergence is studied primarily in statistical/condensed matter physics. The classic paradigm is the Ising model: the microscopic physics is described by spins on a lattice, while the emergent physics in the right phase can exhibit ferromagnetism. Ferromagnetism is an emergent phenomenon, as it is not a property of the microscopic degrees of freedom; instead, it is an ordering of the collective. The possibility that gravity may be emergent suggests that quantum gravity ought to be studied as a problem in statistical physics. The question we will be asking is: ‘what is the analogue of the Ising model for gravity?’

In this direction, we propose to study *background-independent spin systems* to model the emergence of geometry and gravity. We will attempt to adapt the methods and paradigms of statistical physics to the analogue of dynamical geometry: we will study spin systems on a dynamical lattice. We initiated this program in [11, 12], where we proposed *quantum graphity*, a spin system on a dynamical lattice whose purpose was to study the emergence of flat space and matter, and developed this further in [14] with another such spin system, modeling the interaction of geometry and matter. While this is a new direction of research, we have studied

<sup>6</sup> An appealing feature of this scenario, for us, is that if diffeomorphism and Lorentz invariance are emergent symmetries, there will be departures from the exact symmetry, providing observational tests for theories of emergent gravity.

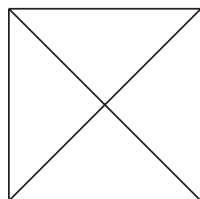
several subjects, including the emergence of (flat) geometry and matter [11, 12], deriving the speed of light from first principles [38], matter/geometry interaction and entanglement and issues in quantum cosmology [14] and the distinction between fundamental and emergent time [15]. Not surprisingly, we find that the change in perspective that emergent gravity brings does indeed raise new opportunities. It also creates new puzzles, which we hope to be able to clarify in the relatively straightforward context of spin systems. The purpose of this paper is to review the models and results and outline future directions of research and the major open issues.

The outline of this paper is as follows. In section 2, we describe the basic idea of graphity models, which is to describe the dynamical lattice using the space of adjacency qubits of the complete graph. We explain what we mean by using a graph as space, and outline the task involved in finding the right Hamiltonian for the spin system. In section 3, we review the first quantum graphity model of [12]. The Hamiltonian for this model is an extension of the string network condensation mechanism of Wen and collaborators for emergent matter, which we review before extending it to dynamical lattices. We then describe a method to derive the emergent speed of light from the microscopic physics using the Lieb–Robinson (LR) bound for information propagation in a local spin system and discuss the extent to which this means that the model has an effective Lorentzian spacetime. We end this section by presenting a new tool for dealing with dynamical graphs, the transformation from the complete to the line graph. In section 4, we review the model of interacting matter and geometry that was proposed in [14]. In this model, we have attempted to implement the idea of ‘geometry tells matter where to go and matter tells geometry how to curve’ in a spin Hamiltonian. The question is to what extent this captures elements of gravitational behavior and we study it by looking at a toy black hole. In addition, the model produces entanglement between matter and geometry, which gives rise to the possibility that each of the two, studied separately, will show thermalization. We present the results that support this. In section 6, we outline the major issues and directions for this program of research, namely the quantum behavior of matter–geometry interaction, time and Lorentz invariance, and gravity.

## 2. Spin systems on a dynamical lattice

Adjacency, an essential aspect of geometry, is fixed in spin systems such as the Ising model. The spin system is given by a Hamiltonian on a lattice and this lattice determines adjacency. A local Hamiltonian then is a sum of the terms of neighbor interactions. For example, in the Ising model, the Hamiltonian is  $H = -\sum_{\langle ij \rangle} J_{ij} \sigma_i \sigma_j$ , where the sum is over all adjacent spins  $i$  and  $j$ . As explained above, we instead want adjacency to be dynamical and so we will turn adjacency into a quantum degree of freedom.

We do this by starting with the lattice of *all possible* adjacencies. For  $N$  systems, this is  $K_N$ , the complete graph on  $N$  vertices.  $K_N$  has  $\binom{N}{2}$  links, corresponding to all possible pairings of its  $N$  vertices. For example, for  $N = 4$ , the graph of all possible pairs  $ij$  with  $i, j = 1, \dots, 4$  is  $K_4$ :





We now give each link  $e_{ij}$  in  $K_N$  a Hilbert space  $\mathcal{H}_{e_{ij}} \simeq \mathbb{C}^2$  with basis elements  $\{|1\rangle, |0\rangle\}$ , and interpret  $|1\rangle_{ij}$  to mean that the edge  $e_{ij}$  is *on* and  $i, j$  are adjacent, and  $|0\rangle$  to mean that the edge is *off*. The total state space then is

$$\mathcal{H}_{\text{graph}} = \bigotimes_{e=1}^{\frac{N(N-1)}{2}} \mathcal{H}_e. \quad (3)$$

By having *on* and *off* degrees of freedom on the edges, a generic state in  $\mathcal{H}_{\text{graph}}$  is a superposition of subgraphs of  $K_N$ , which, for large enough  $N$  and to the extent that graphs can represent geometry, can be thought of as a quantum superposition of quantum geometries.

### 2.1. Graphs and space

We will utilize the graph of adjacencies as a primitive form of geometry. Adjacency of course defines neighborhoods, a basic aspect of topology. Together with a local Hamiltonian, it determines a causal set of events, or interactions. A causal set has long been known to contain the information on the metric, up to the conformal factor.

Graphs are used in many different approaches to quantum gravity, from spin networks in quantum gravity [20, 21] and loop quantum gravity [19], spin foams [33] and group field theory [22], to the partial orders in causal sets [23] and quantum causal histories [24]. There is substantial freedom in the assignment of geometry to a graph. One can consider the graph as the dual to a piecewise linear triangulation of some given dimension or use measures of geometry such as the Hausdorff dimension of the graph. A unified and organized way of assigning a geometry to a general graph does not currently exist in quantum gravity research. For example, even if the evolution of our states defines a (quantum) causal set, recovering the metric when one only has the causal set is a formidable problem.

Fortunately for our purposes, the correspondence between a graph and space is clear in two extremal situations: no space and flat, or nearly flat, space. In our model  $K_N$  represents no space<sup>7</sup>. The neighborhood of a vertex in  $K_N$  is the entire  $K_N$ , meaning that in that state of adjacency there is no notion of subsystems and hence no notion of locality. We interpret this state as no space. On the other hand, if the graph is a regular lattice, say a cubic lattice, there is also a straightforward interpretation of it as flat space in three dimensions.

Back to the state space  $\mathcal{H}_{\text{graph}}$ , we note that, for very large  $N$ , the regular cubic lattice is a subgraph of  $K_N$ , which means a particular state in  $\mathcal{H}_{\text{graph}}$ . Simply using these two extremal cases we can sketch out the basic idea behind our models. We will think of *space as order*, and will consider space to be emergent when geometrical symmetries such as the homogeneity and isotropy of the FRW metric appear.  $K_N$ , in contrast, is permutation invariant. The scenario of *geometrogenesis* proposed in [11, 12] is that at early times, or high energies, the universe is in the  $K_N$  phase, where there is no useful notion of space, and at low energy it freezes in the symmetric state of an FRW metric. The questions regarding emergence of geometry will then be: a generic state in  $\mathcal{H}_{\text{graph}}$  is a quantum superposition of graphs and, therefore, to the extent that the graph represents a geometry, a superposition of geometries. We do not see such superpositions macroscopically. Why? In addition, can the system settle in a regular lattice as its ground state or long-lived metastable state?

<sup>7</sup> One can argue that the configuration in which all links are in the  $|0\rangle$  state is also a representation of ‘no space’.  $K_N$  appears to be a better candidate for the high-energy densities one expects at early times and in quantum gravitational regimes (in our model, a  $|1\rangle$  link has a higher energy than a  $|0\rangle$  link).

## 2.2. Purpose of models

The task in constructing quantum graphity models is to find a Hamiltonian such that the model:

1. shows how a regular, smooth geometry emerges from disordered or *no* geometry;
2. exhibits primitive notions of gravity, e.g. attraction, horizons, negative heat capacity, etc, at the collective level;
3. lets us investigate quantum effects such as quantum interference of geometries and entanglement of matter and geometry;
4. can be used to investigate questions on emergent gravity, such as time versus emergence, and to develop new tools for emergent gravity physics in an explicit context.

In this direction, we have already studied several subjects, including the merging of (flat) geometry and matter [11, 12], deriving the speed of light from first principles [38], matter/geometry interaction and entanglement and issues in quantum cosmology [14], as well as fundamental versus geometric time [15].

We will now review two models of [11, 12] and [14].

## 3. Quantum graphity: a model of emergent geometry and matter

*Quantum graphity*, first proposed in [11] and [12], is a spin system in which locality, geometric symmetries and matter arise at low energy when a spin system is cooled to its ground state. The model is an extension of the string-net condensation mechanism for emergent  $U(1)$  matter, photons and fermions proposed by Wen and collaborators [16]. This model is on a fixed lattice and we extended it to the dynamical quantum lattice defined by equation (3). The result is that the microscopic degrees of freedom that determine the ground state lattice are the same as those that give rise to the  $U(1)$  matter in the appropriate limit.

The state space of this model is  $\mathcal{H}_{\text{graph}} = \bigotimes_e \mathcal{H}_e$ , with the edge labels extended to

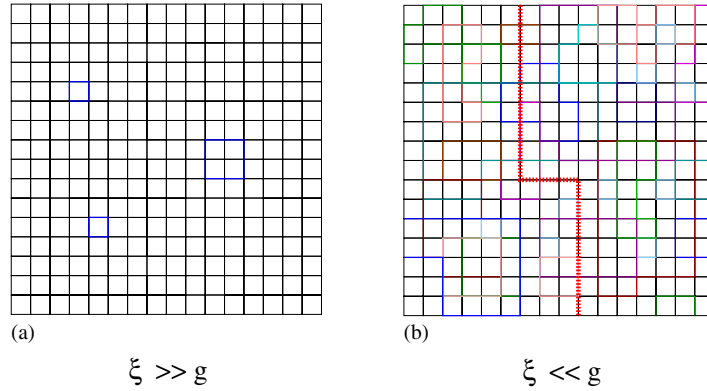
$$\mathcal{H}_e = \text{span}\{|j_{ij}, m_{ij}\rangle\} = \text{span}\{|0, 0\rangle, |1, -1\rangle, |1, 0\rangle, |1, 1\rangle\} \simeq \mathbb{C}^4. \quad (4)$$

As previously,  $j_{ij} = \{|1\rangle, |0\rangle\}$  means that the edge  $e_{ij}$  is *on* or *off*, respectively. With the  $m_{ij}$  labels, we have one *off* state and three *on* states. A finite dimensional Hilbert space at every edge means that this is nothing else than a spin system. In our interpretation, a configuration of *on* states is a graph  $\Gamma_N$ . If the graph has enough order, it will represent space. Now we see that an *on* edge can have three different values, or colors. These colors allow us to draw networks on  $\Gamma_N$ . For instance, if  $\Gamma_N$  is a square lattice, a basis state in  $\mathcal{H}_{\text{graph}}$  can now be a configuration of colored strings, or string-nets, like in figure 1. Following the mechanism of string-net condensation invented by Wen and collaborators [16], the strings of colored edges will give rise to an emergent gauge field or field of matter. A string condensate is a liquid of fluctuating strings, with the collective excitations above such states corresponding to gauge bosons and fermions. In string net condensation, just as phonons are emergent collective excitations of a crystal, gauge bosons and fermions are collective excitations of a pattern of strings. We will outline this mechanism below.

Let us first define the following four operators acting on each  $\mathcal{H}_e$ ,

$$\begin{aligned} J|j, m\rangle &= j|j, m\rangle, \\ M|j, m\rangle &= m|j, m\rangle, \end{aligned} \quad (5)$$





**Figure 1.** Two configurations of colored strings, or string-nets: (a) the black links are in the state  $|1, 0\rangle$  and the blue links in the state  $|1, 1\rangle$ , while all the other edges in  $K_N$  are in the state  $|0, 0\rangle$ . (b) A state with many colors. Again, all the edges in  $K_N$  that are not drawn are in the state  $|0, 0\rangle$ . In both figures, the  $\{j_{ab}\}$  degrees of freedom are frozen in the configuration of a square lattice.

and the raising and lowering operators

$$\begin{aligned} M^+ |j, m\rangle &= \frac{1}{\sqrt{2}} \sqrt{(j-m)(j+m+1)} |j, m+1\rangle, \\ M^- |j, m\rangle &= \frac{1}{\sqrt{2}} \sqrt{(j+m)(j-m+1)} |j, m-1\rangle. \end{aligned} \quad (6)$$

The relevant commutation relations are

$$[M^+, M^-] = M, \quad [M, M^\pm] = \pm M^\pm. \quad (7)$$

It is important that all these operators annihilate the *off* state:

$$J|0, 0\rangle = M|0, 0\rangle = M^\pm|0, 0\rangle = 0. \quad (8)$$

Now, on each edge of  $K_N$  we can define creation and annihilation operators  $a^\dagger$  and  $a$ , with  $a|0\rangle = 0$ ,  $a|1\rangle = |0\rangle$ , and the number operator  $n$ ,  $a^\dagger a|n\rangle = |n\rangle$ . We use these to define the operator  $N$ , a quantum analogue of the adjacency matrix of a graph:  $N_{ij}$  is the number operator acting on edge  $e_{ij}$  in the graph and returning the value of that edge, e.g.  $N_{13}(|e_{12}\rangle \otimes |e_{13}\rangle \otimes \dots) = n_{13}(|e_{12}\rangle \otimes |e_{13}\rangle \otimes \dots)$ . Powers of  $N_{ij}^{(L)}$  tell us how many paths of length  $L$  connect  $i$  and  $j$ . For example, paths of length 2 are counted by  $N_{ij}^{(2)} = \sum_k N_{ik} N_{kj}$ , etc. The important property of these operators that we will use in this model is that normal-ordering  $N$  returns the number of *non-overlapping* paths from  $i$  to  $j$  of length  $L$  (for why this is so, see [12]).

We can now define the Hamiltonian of the model. It has the following terms:

- Vertex term:

$$H_V = g_V \sum_i e^{(v_0 - \sum_j N_{ij})^2}, \quad (9)$$

for  $g_V > 0$  and  $v_0$  a free parameter. This term is minimized when all vertices have degree  $v_0$ .

- Loop (and string) terms: Using the full state space (the  $m$ -values on the edge Hilbert spaces), the Hamiltonian that determines the loop distribution in the ground state graph

is a generalization of Wen's string-net condensation [16] to a dynamical lattice. A string condensate is defined by the types of strings of the model, namely the 'colors' we use, and the way we allow them to branch. To explain the mechanism, let us start by freezing all the edges in some configuration  $\Gamma_N$ . Then we can consider a Hamiltonian that is the sum of a constraint term, a potential energy term and a fluctuation term:

$$H_{\text{strings}} = V H_C + \xi H_{\text{pot}} + g H_{\text{kin}}. \quad (10)$$

The constraint term says that it costs a large energy  $V \gg \xi, g$  to have open strings. So, at low energies, all the allowed string configurations are closed string-nets, like in figures 1(a) and (b). The potential term  $\xi$  is a string tension. When  $\xi \gg g$  the energy of a string configuration is proportional to its length and therefore only a few small strings are allowed in the ground state; see figure 1(a). On the other hand, when  $\xi \ll g$  the kinetic energy makes the string fluctuate a lot and the ground state is a superposition of many closed string configurations with arbitrary length, as in figure 1(b). This is the so-called *string condensed state*. The modes of vibration of these strings are the first category of excitations above the ground state. The second type is made of defects, namely violations of the constraint  $V H_C$ . In the string pictures, such defects are endpoints of open strings. In [16], it was shown that the fluctuations of closed strings correspond to U(1) gauge bosons (photons), while endpoints of open strings correspond to charged fermions (electrons).

To extend this construction to the case in which the  $\{j_{ij}\}$  degrees of freedom are not frozen in a particular lattice  $\Gamma_N$  like the square lattice of figures 1(a) and (b) [11, 12], we modified the Hamiltonian of [16] to the terms

$$H_{\text{loops}} = -g_L \sum_{\text{loops}} \sum_{L=0}^{\infty} \frac{r^L}{L!} \prod_{a=1}^L M_a^{\pm}, \quad (11)$$

with  $\prod_{a=1}^L M_a^{\pm} = M_{ij}^+ M_{jk}^- \dots M_{yz}^+ M_{zi}^-$ , and

$$H_{\text{strings}} = g \sum_i \left( \sum_j M_{ij} \right)^2 + \xi \sum_{ij} M_{ij}^2. \quad (12)$$

$g$  and  $\xi$  are positive couplings. The ground state of string consists of all links having  $m = 0$ . Again, the  $g$  term is the kinetic energy term that makes strings fluctuate, whereas the  $\xi$  term is a potential energy corresponding to a string tension.

At the same time, the term  $H_{\text{loops}}$  also determines the ground state lattice. This can be seen more easily if we consider the term equation (11) acting on the reduced state space  $\mathcal{H}_{|0,1\rangle}$ , i.e. ignoring the  $m$  labels. In that case, the term (11) can be written as

$$H_{\text{loops}} = -g_L \sum_{i,j} \delta_{ij} e^{r N_{abij}} = -g_L \sum_{i,j} \delta_{ij} \sum_{L=0}^{\infty} \frac{r^L}{L!} N_{ij}^{(L)}. \quad (13)$$

When normal ordered,  $H_{\text{loops}}$  counts all non-overlapping paths of different lengths and lowers the energy when the number of loops of length  $L^*$  is maximized, where  $L^*$  depends on the choice of the parameter  $r$ . In [12], we found a choice of parameters ( $v_0 = 3$  and  $r \geq 7.1$ ) for which *the honeycomb lattice is a stable local minimum*.

- Evolution terms: the terms above are eigen-operators of graph states and do not change the linking structure between vertices. The time evolution of the model should include terms

that change the graph. In [12], we suggested local graph evolution moves, much like one does in spin foam models, for example by terms that exchange linking or add/subtract edges:



So far, the focus of our work in this model has been to look for the lowest-energy graph states, for which this part of the Hamiltonian is not required. As a result, we have no insight into the particular properties of some type of evolution over another.

The full Hamiltonian of the model,

$$H = H_V + H_{\text{loops}} + H_{\text{strings}}, \quad (14)$$

(normal ordered) is then a modification of the string-net condensation mechanism of [16] in that the supporting lattice is dynamical. An interesting feature of the model is that the same degrees of freedom give rise to the U(1) gauge bosons and charged fermions and also shape the lattice. This can be thought of as a novel form of unification of the matter and geometry, with the separation between the two appearing only at the emergent level. In the right region of the couplings, the model has a stable energy minimum that exhibits both a lattice configuration (the honeycomb) and string-net condensation [12, 26]<sup>8</sup>. The space, fields and electrons have emerged all together.

### 3.1. Speed of light via the Lieb–Robinson (LR) bound

We have seen that in string-net condensation, in the limit  $\xi \ll g$ , one obtains emergent photons as collective excitations of fluctuating strings. In the continuum limit, these excitations are described by the Lagrangian of electromagnetism and the speed of light is a function of the microscopic couplings,  $c = \sqrt{\xi g}$ . An important question now arises: does the emergence of the Maxwell equations imply the emergence of Lorentz invariance? Or is the fact that the elementary excitations behave like photons not enough to have Lorentz invariance? Indeed, do we even have a light cone? What forbids other signals to travel faster than light?

The system we are considering is not relativistic; after all, we are using ordinary non-relativistic quantum mechanics. It is well known that in principle in non-relativistic quantum mechanics it is possible to send signals arbitrarily far away in an arbitrarily short time. This is easy to understand in perturbation theory. If we perturb some eigenstate of a quantum many-body system with a local perturbation, for example by flipping one spin in a long spin chain, the system will be away from equilibrium and will start evolving in time. For some arbitrarily short time  $t_\epsilon$ , there is an order in perturbation theory such that a spin arbitrarily far from the perturbed spin will be away from equilibrium as well, and will feel the perturbation. We know that ordinary non-relativistic quantum mechanics is not the right theory and one should use quantum field theory instead; still, it seems strange that such a violation of causality is possible. Well, the point is that such violation is not that strong after all. If the spins are very far away and the time is very short, to see the effect one must go to a *very* high order in perturbation theory. In effect, this defines an emergent light cone, outside which the signal is exponentially suppressed.

<sup>8</sup> A similarly motivated model of emergent regular space was recently proposed in [27].

The key for this to happen is locality. The type of system we are considering goes under the name of *local bosonic system*. A local bosonic system is a system in which locality is constrained at the fundamental level. By this we mean that the Hilbert space is the tensor product of small local Hilbert spaces, and that the Hamiltonian is the sum of local operators with support on only a finite (and small) number of local Hilbert spaces. In this case, one can formalize the emergent light cone structure in a rigorous way, using the LR bound, a bound to the violation of causality that one can have [17]. This bound implies the existence of a maximum speed  $v_{\text{LR}}$  of signals in the medium. The speed  $v_{\text{LR}}$  determines an effective light cone such that signals outside the light cone are exponentially suppressed.

To be more precise, consider different-time commutators of observables with support on regions  $P$  and  $Q$  of the system that are spatially separated by a distance  $d_{PQ}$ . Consider a Hamiltonian that is a sum of local terms  $H = \sum_{\langle ab \rangle} h_{ab}$  with  $\langle ab \rangle$  being the edges of a graph. Then we can bound the commutator between two observables  $O_P(t)$ ,  $O_Q(0)$  as

$$\|[O_P(t), O_Q(0)]\| \leq 2\|P_P\|\|O_Q\|C \exp\{a(d_{PQ} - v_{\text{LR}}t)\}, \quad (15)$$

where the constants  $a$  and  $C$  depend on the graph and the speed  $v_{\text{LR}}$  depends on both the graph and the maximum strength of the interactions  $\|h_{ab}\|$  [17]. In [38], we applied the LR theorem to Wen's Hamiltonian for emergent photons to find the LR speed of the emergent U(1) excitations. Usually the problem with the bound is that it is too loose because it overkills by choosing the maximum strength of interactions. This would yield an estimate for the maximum speed of signals  $v_{\text{LR}} \simeq g$ , which is much higher than the speed of emergent light  $c$ . We were able to find a tighter bound in which  $v_{\text{LR}} \simeq \sqrt{\xi g} \equiv c$  (recall that the speed of light in the emergent Maxwell equations is related to the fundamental couplings by  $c = \sqrt{\xi g}$ ). Therefore, emergent light is also the fastest signal in the system.

This is a strong argument in favor of the emergence of Lorentz symmetry from a spin system. However, more work needs to be done in this direction. We also note that the above derivation of an emergent light cone does not address the problem of different species traveling at different speeds, a problem generic to emergent gravity approaches.

### 3.2. Transition in the causal structure

An important feature of the LR speed defined by the bound is that it depends on the degree of the vertices on the lattice:

$$v_{\text{LR}} \propto d. \quad (16)$$

One can understand this intuitively: the transmission of information and correlations through the lattice is a quantum effect due to the sum over all the possible paths that connect two points. Therefore, the higher the number of paths that connect two points  $P$  and  $Q$  on the lattice, the stronger the signal at  $P$  from  $Q$ . Or, equivalently, more paths open up the light cones and increase  $v_{\text{LR}}$ . The vertex degree is a measure of the number of paths between two points in the lattice, which is why we find that the LR speed is proportional to it. In section 4.1, we will use this fact to describe a toy black hole.

For the model defined above,  $v_{\text{LR}} \propto d$  means that at high energy, when more edges are on and therefore the average vertex degree is higher, the speed of light is also higher. If, for example, we approximate the transition from high- to low-energy states by a sequence of

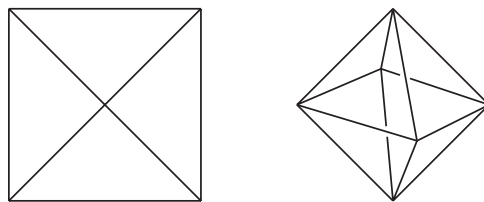
hypercubes of dimension  $D$ , then, since  $D \propto d$ , we will find that  $v_{\text{LR}} \propto D$ , and hence it drops with the dimension.

In [38], we pointed out that this effect can be used to resolve the homogeneity problem in cosmology. This is the question of why we see the sky at a uniform temperature, the cosmic microwave background, including distant parts that have had no causal connection in the past. It is usually solved by assuming an inflationary period in the early universe. Here, the causal mechanism for thermal equilibrium is the opening up of the light cones caused by more *on* edges at higher energies (early times). This is interesting as it illustrates the possibility of a transition in the causal structure and that such a quantum gravity effect can have observable consequences that are not limited to the Planck scale. Of course, further work is needed to check if a variation of the speed of light like the one proposed here is consistent with either observations. We note that, while we can always defined a LR speed of information propagation, for a highly connected lattice the string network condensation mechanism does not apply, and we do not expect  $v_{\text{LR}}$  to be the speed of light as there is no emergent light.

### 3.3. From a dynamical to a fixed lattice

Central to our motivation for studying emergent gravity using spin systems is to gain access to tools to describe emergence used in statistical physics and condensed matter theory. However, to accommodate the possibility of dynamical geometry, we are using spin systems on dynamical lattices. Many of the tools that we would like to use are of course designed for spin systems on a fixed lattice. For example, we would like to use mean field theory analysis to find out if the transition to the flat graph described above is a phase transition. How can we do that if the connectivity of the lattice is a variable?

It turns out that the basic idea of writing the ensemble of graphs as subgraphs of  $K_N$  for some large enough  $N$  also allows us to transform the dynamical lattice in the model above to a spin system on a fixed lattice. This can be done by mapping  $K_N$  to its line graph  $L(K_N)$ , as follows [18]: every edge of  $K_N$  maps to a vertex in  $L(K_N)$ . Thus,  $L(K_N)$  is a graph with  $N(N-1)/2$  vertices. We now connect two vertices in  $L(K_N)$  if the corresponding edges in  $K_N$  shared a vertex. For example, for  $K_4$ , the original graph and its line graph  $L(K_4)$  are



It is now easy to see that the values  $j_e = \{|1\rangle, |0\rangle\}$  on the original system become simple quantum Ising spins on the vertices of the line graph. A particular graph, a state in  $\mathcal{H}_{\text{graph}}$ , is simply an Ising spin configuration on the fixed lattice  $L(K_N)$ .

In [18], we used this mapping in a weak coupling approximation of the model. With a mean-field theory approximation and low-temperature expansion, we studied the properties of the model near-zero temperature. We found that the model is dual to an Ising model with external nonzero magnetic field if we neglect the interaction terms. In particular, we showed that the average vertex degree is naturally a good order parameter for the mean-field theory approximation and we found that the parameter  $v_0$  plays the role of the external magnetic field. Since  $v_0$  is never zero, the model has no phase transition and at  $T = 0$  the system goes to the

ground state as expected. We also confirmed that, as the temperature drops, the speed of the emergent light must drop with the vertex degree.

This method of dealing with dynamical graphs is likely to be of more general interest in background-independent quantum gravity approaches, where dynamical graphs and ensembles of graphs are commonly used. As long as in the theory in question an isolated vertex is physically equivalent to no vertex, the ensemble of graphs can always be written in terms of the state space  $\mathcal{H}_{\text{graph}}$ ; that is, the graphs can be viewed as subgraphs of  $K_N$  for some sufficiently large  $N$ . If the degrees of freedom reside on the edges only, then the above mapping turns a quantum ensemble of graphs into an Ising-like model. For theories with degrees of freedom on the vertices (as in general spin foam models for example), a generalization of the above mapping may be possible.

#### 4. A unitary model of interacting matter and geometry

In the previous section we described a background-independent model for emergent locality, spatial geometry and matter. This model leaves open several questions. An important one is that in the scenario of [12] the universe starts at a high-temperature configuration ( $K_N$ ) and evolves to a low-energy one (a regular local lattice). This has clear limitations when applied to a cosmological context. In effect, what we described above assumes an external heat bath in which we can dump the energy of all the links that we turn off. What we would like instead is a unitary, energy-conserving model. In [14], we constructed such a model, in which graph edges can be deleted and matter created and vice versa. We can turn links off and still conserve energy because we can transfer the energy from the links to the matter. This is essentially an extension of the Hubbard model to a dynamical lattice. The second motivation for the model we are about to present now goes back to the interpretation of general relativity as ‘geometry tells matter where to go and matter tells geometry how to curve’. We wish to know whether it is possible to implement this kind of behavior in a spin system and, if so, study to what extent such a spin system can show aspects of gravitation.

The model of [14] then is motivated by the questions:

1. If gravity is emergent, is there an analogue of the Ising model for gravity? Can we get horizons, attraction, negative heat capacity, etc, from a spin system with a local Hamiltonian?
2. The spin system ultimately should describe a cosmological theory (just as general relativity does at the emergent level). Can we have unitary microscopic evolution that equilibrates to appear approximately thermal for long enough times?
3. In a system with state space  $\mathcal{H}_{\text{geometry}} \otimes \mathcal{H}_{\text{matter}}$ , study quantum effects of the matter–geometry interaction, such as entanglement of matter and geometry and superposition of quantum geometries.

We began addressing the above questions in [14] and [37]. Geometry in this model is described, as before, by basis states  $\{|1\rangle, |0\rangle\}$  on the edges  $e$  of complete graph  $K_N$  on  $N$  vertices, interpreted as the edges being on or off, respectively. That is,  $\mathcal{H}_{\text{graph}} = \bigotimes_{e \in K_N} \mathcal{H}_e$ ;  $\mathcal{H}_e \simeq \mathbb{C}^2$ . We implemented quantum matter in the simplest way possible, by allowing for bosons that live on the vertices of the graph:  $\mathcal{H}_{\text{matter}} = \bigotimes_{i=1}^N \mathcal{H}_i$ , where  $\mathcal{H}_i$  is the state space of a harmonic



oscillator on vertex  $i \in K_N$ . The state space of the model then is

$$\mathcal{H} = \mathcal{H}_{\text{graph}} \otimes \mathcal{H}_{\text{matter}} = \bigotimes_{e \in K_N} \mathcal{H}_e \bigotimes_{i=1}^N \mathcal{H}_i, \quad (17)$$

with basis states of the form

$$|\Psi\rangle = |\Psi_{\text{graph}}\rangle \otimes |\Psi_{\text{matter}}\rangle = |e_1, \dots, e_{\frac{N(N-1)}{2}}\rangle \otimes |n_1, \dots, n_N\rangle. \quad (18)$$

We give the links energy when they are in the *on* state:

$$H_{\text{link}} = -U \sum_{ij} \sigma_{ij}^z, \quad (19)$$

while the boson energy is given by

$$H_v = \sum_{i=1}^N H_i = - \sum_i \mu_i b_i^\dagger b_i, \quad (20)$$

where  $b_i$  and  $b_i^\dagger$  are the annihilation and creation operators for a boson on site  $i$ .

As in the previous model, we can have a superposition of interactions, or graph edges. For example, the state

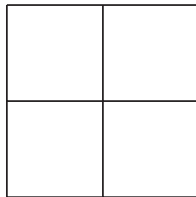
$$|\psi_{ij}\rangle = \frac{|10\rangle \otimes |1\rangle_{ij} + |10\rangle \otimes |0\rangle_{ij}}{\sqrt{2}} \quad (21)$$

describes a particle at  $i$  and a particle at  $j$  and a superposition of them interacting and not interacting or, equivalently, a superposition of geometries. The state

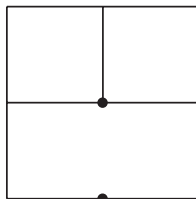
$$|\phi_{ij}\rangle = \frac{|00\rangle \otimes |1\rangle_{ij} + |11\rangle \otimes |0\rangle_{ij}}{\sqrt{2}}, \quad (22)$$

on the other hand, describes *entanglement of matter with geometry* and distinguishes this model from the previous one.

The basic idea in this model is that the energy of a graph edge can be transferred to matter, and vice versa. That is, starting from the configuration



we can erase a link to create two bosons on the vertices of the erased link (the black dots represent bosons):



The above process is reversible, i.e. we can eliminate two bosons and replace them with a link connecting their supporting vertices. This exchange is done using the term in the Hamiltonian<sup>9</sup>

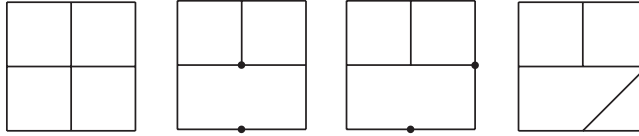
$$H_{\text{ex}} = k \sum_{ij} (|0\rangle\langle 1|_{ij} b_i^\dagger b_j^\dagger + |1\rangle\langle 0|_{ij} b_i b_j). \quad (24)$$

In addition, the bosons can hop around the lattice *but only where a link exists*, i.e. they can hop from  $i$  to  $j$  only if the link  $ij$  is on:

$$H_{\text{hop}} = -t \sum_{ij} |1\rangle\langle 1|_{ij} (b_i^\dagger b_j + b_i b_j^\dagger). \quad (25)$$

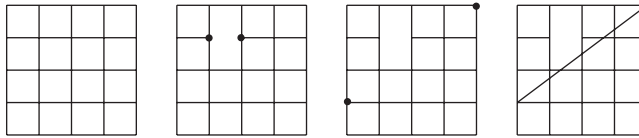
This is an important feature of the model as it means that *it is the behavior of the matter that gives the graph the meaning of geometry*. An edge between two nodes  $i, j$  means that there is a hopping term in the Hamiltonian between  $i$  and  $j$ . Since the meaning of the edges is given by the dynamics of the particles, a pure graph without particles is just a mathematical structure without physical content. To the extent that such a model can capture aspects of gravity, pure gravity will not be meaningful.

The combination of  $H_{\text{ex}}$  and  $H_{\text{hop}}$  causes the lattice to change, as for example in this sequence:



Note that it is more likely that bosons will be created in more highly connected parts of the lattice, as there will be a higher density of links to be turned into matter and, similarly, new links are more likely to be created where there is a higher concentration of matter.

With the term  $H_{\text{ex}}$  as given by equation (24) above, it is possible to destroy bosons and create a new graph link in their place, no matter how far apart, with respect to the lattice, the bosons are. The following sequence is an example:



In [14], we eliminated the possibility of such non-local connections by allowing boson–edge exchange only if there already is a path of length  $L$  connecting the boson sites, where  $L$  is some short distance:

$$H_{\text{ex}} = k \sum_{ij} (|0\rangle\langle 1|_{ij} P_{ij}^L b_i^\dagger b_j^\dagger + P_{ij}^L |1\rangle\langle 0|_{ij} b_i b_j), \quad (26)$$

<sup>9</sup> In [14], we introduced the more general Hamiltonian that created  $R$  pairs of bosons for every link destroyed, and vice versa:

$$H_{\text{ex}} = k \sum_{ij} (|0\rangle\langle 1|_{ij} (b_i^\dagger b_j^\dagger)^R + |1\rangle\langle 0|_{ij} (b_i b_j)^R). \quad (23)$$

Here, for simplicity, we take  $R = 1$ .

where

$$P_{ij}^L = \sum_{k_1, \dots, k_{L-1}} P_{ik_1} P_{k_1 k_2} \cdots P_{k_{L-1} j}; \quad P_{ij} = |1\rangle\langle 1|_{ij}. \quad (27)$$

In what follows, we will take  $L = 2$ . Note that, with the above dynamics, the graph cannot disconnect.

The total Hamiltonian for this model then is

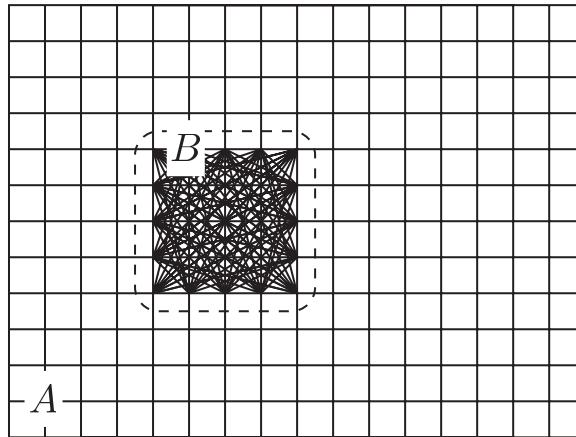
$$H = H_{\text{link}} + H_v + H_{\text{ex}} + H_{\text{hop}}. \quad (28)$$

#### 4.1. A toy black hole

As we discussed in the introduction, our working interpretation of the physical content of general relativity is that geometry tells matter where to go and matter tells geometry how to curve. Can the reverse be true, i.e. in a system where geometry tells matter where to go and matter tells geometry how to curve, do we see aspects of gravity?

The model we just defined is one where geometry (the graph) determines where matter is allowed to go, while matter influences geometry: a particle at vertex  $i$  and one at vertex  $j$  can be replaced by a new edge  $ij$  where one did not previously exist and so change the distance between  $i$  and  $j$ . This is not quite the same as general relativity, but the model does have some features that are reminiscent of gravitational behavior. We will illustrate this by constructing a toy ‘black hole’ in our spin system.

Consider a state of the system in which the graph is a flat lattice but with a region that is highly connected. We call the flat region  $A$  and the highly connected region  $B$ :



To make things simple, if  $B$  contains  $N_B$  vertices, we ask that  $B$  is completely connected so that the degree of vertices in  $B$  is  $\sim N_B$ .

It is intuitively clear that  $B$  will act as a trap: there are many more links to the interior of  $B$  than paths leading to the outside  $A$ . A boson that hops its way to the boundary of  $B$  is far more likely to take its next step into  $B$ 's interior and once inside there are many more edges that will keep it inside than edges leading back to  $A$ . This intuition can be made more quantitative by using the LR speed of light that we defined and calculated in section 3.1 to find out what happens to a ray of light traveling on this lattice<sup>10</sup>. Recall that we found that the LR

<sup>10</sup> We can add Wen's  $U(1)$  Hamiltonian for emergent light [16] in equation (28). In the phase where the couplings of the  $U(1)$  theory are small with respect to the other couplings in the Hamiltonian, we have electromagnetic waves traveling on the lattice.

speed is proportional to the degree of the lattice. Hence, in region  $A$ ,  $c_A \sim 1$ , while in region  $B$ ,  $c_B \sim N_B$ . Now consider a particle traveling in this graph and crossing the boundary between  $A$  and  $B$ . Snell's law states that the critical angle for total internal reflection is

$$\theta_c = \sin^{-1} \frac{c_B}{c_A} = \sin^{-1} N_B, \quad (29)$$

and so the probability of a light ray escaping  $B$  is of the order of  $N_B^{-1}$ . In the large  $N_B$  limit, region  $B$  traps light and matter.

$N_B$  is not infinite and so light and matter can eventually escape. Assuming that  $A$ , even though less dense, is much larger than  $B$ ,  $B$  will evaporate until its edges are of a density comparable to  $A$ . The whole process is of course unitary, but the emitted quanta are entangled with the remnant in  $B$ . That is, the spectrum of the emitted radiation will be mixed even though the underlying dynamics is unitary. To the extent that this configuration can be thought of as a toy black hole, it illustrates the resolution of the information loss paradox via matter–geometry entanglement and no black hole singularity.

Of course, there is no claim here that this is a real black hole. There are important differences between what we just described and black holes in general relativity: this is not a spacetime construction; the horizon can be seen by a local observer crossing it. We really are modeling a black hole using two media of different refractive indices. Nonetheless, it illustrates that in this model there is an entropic form of attraction: highly connected regions attract bosons in the sense that the particles are more likely to be found at vertices of higher degree. In addition, there is an analogue of negative heat capacity: bosons are more likely to be created in the highly connected parts of the lattice as there is a higher density of links that can be turned into matter, and new links are more likely to be created where there is a higher concentration of matter.

## 5. Matter–geometry entanglement and thermalization

One of the advantages of approaching the problem of quantum gravity using dynamical lattice by means of spin systems and hopping particles is that we have at our hand methods of quantum information and many-body theory. One of the features of the model of the previous section is the tensor product structure  $\mathcal{H} = \mathcal{H}_{\text{graph}} \otimes \mathcal{H}_{\text{matter}}$  in the state space. It is natural to ask what the role of entanglement is with respect to such bipartition. In other words, how much can matter and space get entangled? This question is relevant in understanding the sense in which matter and space, taken separately, can thermalize.

### 5.1. Thermalization via subsystem dynamics

The quantum evolution of the whole system is unitary and is given by

$$\rho(t) = U(t)\rho(0)U^\dagger(t), \quad (30)$$

where  $U(t) = e^{-iHt}$ . Now, suppose we start with a pure state  $\rho(0)$  which is a product state with respect to the bipartition  $\mathcal{H} = \mathcal{H}_{\text{graph}} \otimes \mathcal{H}_{\text{matter}}$ . If the unitary  $U(t)$  acts as an entangling gate, the state  $\rho(t)$  will be entangled. The evolution of the matter alone will be given by

$$\rho_{\text{matter}}(t) = \text{Tr}_{\text{graph}} \rho(t) = \text{Tr}_{\text{graph}} [U(t)\rho(0)U^\dagger(t)]. \quad (31)$$

The quantum evolution for the matter only is that of an open system and thus is not unitary. That is, the state  $\rho_{\text{matter}}(t)$  can be mixed. The question then is whether there is enough entanglement

to make  $\rho_{\text{matter}}(t)$  a thermal state. A first thing to note is that the dimension of the matter Hilbert space is  $\dim \mathcal{H}_{\text{matter}} = N \dim \mathcal{H}_i$  while the dimension of the Hilbert space of the graph is  $\dim \mathcal{H}_{\text{graph}} = N^2 \dim \mathcal{H}_e$ . So, as long as  $\dim \mathcal{H}_e \simeq \dim \mathcal{H}_i$ , it is possible for the matter to be in a completely mixed state and therefore have a thermal radiation up to infinite temperature. The details of the model determine what kind of thermal radiation is possible to get.

On the other hand, the quantum evolution of the graph will also be that of an open quantum system and therefore the graph will not evolve unitarily. In the right regime of the couplings, namely a weak coupling between graph and matter, an effective thermal behavior for the graph is possible. For this to happen, first of all, we need the Hamiltonian for the graph to be non-degenerate. Although  $H_{\text{graph}}$  is highly degenerate, such degeneracy is easily lifted by random perturbations. Now, let  $\sigma_m(E)$  be the density of states given by the Hamiltonian  $H_{\text{matter}}$ . We can treat the matter as a heat bath for the graph with inverse temperature

$$\beta(E) = \frac{d \ln \sigma(E)}{dE}. \quad (32)$$

The effective temperature  $\beta(E)$  is set by the energy of the initial state of the system  $\rho(0)$ , which is supposed to be separable  $\rho(0) = \rho_{\text{graph}}^0 \otimes \rho_{\text{matter}}^0$  and such that  $\rho_{\text{matter}}^0$  has energy peaked around a certain value  $E$ . Let us also suppose that the interaction Hamiltonian  $H^I = H_{\text{hop}} + H_{\text{ex}}$  is a perturbation, describing scattering processes that approximatively conserve the unperturbed energies (weak coupling). Now consider any observable  $A$  on the graph. These observables constitute all the information we can extract about geometry in our system. The quantum evolution of the expectation value  $\langle A \rangle(t)$  is of course given by

$$\langle A \rangle(t) = \text{Tr} \{ \rho(t) A \otimes \mathbf{1}_{\text{matter}} \}. \quad (33)$$

Under the hypothesis above, it is possible to prove [28] that the long time value of  $\langle A \rangle(t)$  is given by

$$\lim_{t \rightarrow \infty} \langle A \rangle(t) \simeq \frac{\text{Tr}_{\text{graph}} \{ A e^{-\beta(E) H_{\text{graph}}} \}}{\text{Tr}_{\text{graph}} \{ e^{-\beta(E) H_{\text{graph}}} \}}. \quad (34)$$

This means that, even though the total system is not at the equilibrium and is evolving unitarily, and in particular the particles are not in some thermal equilibrium, nevertheless there exists an effective temperature  $\beta(E)$  such that the large time limit of the expectation value of any observable on the graph gives the canonical expectation value at the temperature  $\beta(E)$ . This is a rather old idea (see, e.g., [29]), which we hope finds a concrete application in our system. It is the subject of our current research to show that a similar result can be valid also for system Hamiltonians that can produce an interesting geometry.

*Equilibration in probability.* A related approach to the issue of equilibration in a closed quantum system is to look at equilibration in probability. This approach can be used to study both the models presented here. Even though the long time limit of the expectation value of an observable  $\hat{A}_L$  does not exist, they would still spend most of their time very close to their time average  $\bar{A}$ . Here the subscript  $L$  reminds us that the system has a finite (linear) size  $L$ . To be more precise, consider the time average  $\bar{\rho} = \lim_{\tau \rightarrow \infty} \tau^{-1} \int_0^\tau \rho(s) ds$ , which always exists and is given by the  $\rho_0$  totally dephased in the eigenbasis  $\Pi_n = |E_n\rangle\langle E_n|$  of the Hamiltonian:  $\bar{\rho} = \sum_n \Pi_n \rho_0 \Pi_n$ . We suppose the spectrum  $\{E_n\}$  of the Hamiltonian to be non-degenerate. This is not a very strict

condition since a random perturbation of the Hamiltonian will lift any degeneracies. The time average of the observable  $\hat{A}$  is given by

$$\overline{A(t)} = \lim_{\tau \rightarrow \infty} \frac{1}{\tau} \int_0^\tau A(t) dt = \text{Tr} \left( \sum_n \Pi_n \rho_0 \Pi_n \hat{A} \right) = \text{Tr} \{ \bar{\rho} \hat{A} \}. \quad (35)$$

Equilibration in probability means then that

$$\lim_{L \rightarrow \infty} \text{Pr} \{ |A_L(t) - \overline{A(t)}| \geq \epsilon \} = 0. \quad (36)$$

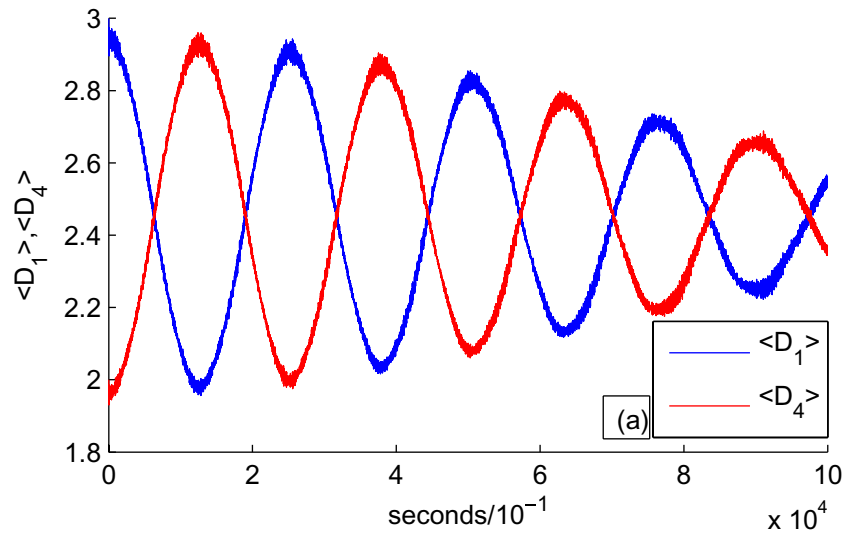
Equilibration in probability is a typical feature of quantum many-body systems away from equilibrium. The equilibrium expectation values  $\bar{A}$  do not need to be those given by the canonical distribution and therefore this is a much weaker feature.

### 5.2. Thermalization in our model

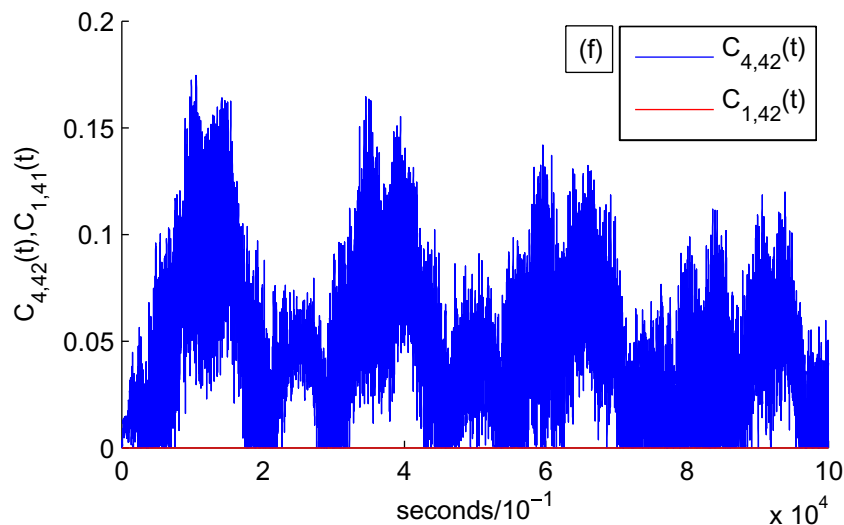
In order to study equilibration and thermalization in this model, we need to resort to numerical simulation. Even to simulate the small graph with four nodes, we need to consider the Hamiltonian equation (28) for hard the core bosons so that the local Hilbert space of a vertex is just 2D in the regime where the hopping and exchange terms are a perturbation of the potential energy terms  $H_{\text{link}}$  and  $H_{\text{hop}}$ . This is the regime that in condensed matter physics is called the Mott insulator when referring to the Hubbard model: particles are localized and there is no transport. Eigenstates of the model are then approximatively product states of the edge/particle configurations, which allows for a direct evaluation of equation (35). The entanglement dynamics and the signs of thermalization for the model equation (28) with just four vertices have been studied numerically in [14], showing that thermalization occurs in the Mott insulating phase of the model. As can be seen in figure 2, the expectation values of the degree operators evolve towards a typical value. In the Mott insulating phase thus, we have at least equilibration in probability. Moreover, we see that there is a nontrivial entanglement dynamics. In figure 3, we show the time evolution of the concurrence between a vertex and an edge attached to it. Also this entanglement is slowly damping, showing a sign of equilibration.

As we have seen, equilibration in probability means that the long time expectation values of observables are given by equation (35). We would like to know whether in our model, equilibration in probability gives *thermal* expectation values for the relevant observables. Let us consider an initial state  $\rho_0 = |\psi_0\rangle\langle\psi_0|$  that is very peaked around a certain energy  $E_0$ . We choose a properly normalized initial state  $|\psi_0\rangle = \alpha \sum_n \exp((E_n - E_0)^2/2\sigma) |E_n\rangle$ . Then we can compute, say, the long time expectation value  $\bar{D}$  of the observable  $D$  that measures the average degree of the nodes in the graph. From equation (35), one can see that this value is a function of the initial state  $\rho_0$  and therefore of the effective temperature  $\beta(E_0)$ . Then we can consider the Gibbs state at the same temperature and compute the thermal expectation value of  $D$ . The two curves are shown in figure 4. For a large range of temperatures the two expectation values are very similar. Therefore the typical, equilibration value of such an observable corresponds to the thermal one at the effective temperature  $\beta(E_0)$ . A systematic study of the thermalization properties of this kind of model is the subject of our current research [37].





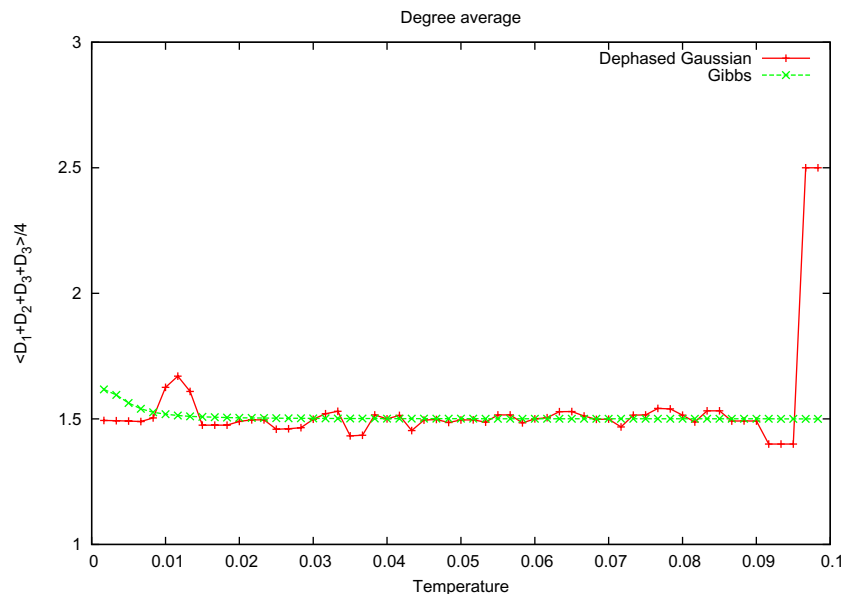
**Figure 2.** The time evolution for the expectation value of the degree number observables  $D_1$ ,  $D_4$  at two nodes 1, 4 for the system in the Mott insulating phase. The two values are oscillating but equilibrating towards a common value independent of the initial state.



**Figure 3.** Entanglement dynamics in the model. Here we plot the concurrence between the number of particles  $n_4$  at node 4 and the insisting edge 24 (blue line). We see a non-trivial evolution, which shows equilibration at large times. The red line shows the same quantity for node 1, showing that an edge only entangles with the nodes it comes from and therefore  $C_{1,42} = 0$  during the whole evolution.

## 6. Summary and outlook

The idea that gravity may be an emergent phenomenon is not new, although in recent years there has been growing interest in this possibility. We presented here a summary of the ideas



**Figure 4.** The long time limit expectation value of  $D$  is the red curve with crosses, while its thermal value is represented by the green curve with  $\times$ . For a large range of temperatures the two expectation values are very similar.

and current status of our program to study emergent gravity using the traditional methods of statistical physics and spin systems, adapted to the new problem. Possibly the most characteristic of the new features of these spin systems is that the lattice is dynamical, reflecting the dynamical nature of geometry in general relativity.

This emphasis on what is often called background independence means that our work has common elements, but also fundamental differences, with other background-independent approaches to quantum gravity. For example, the use of graph states as representations of spatial geometry is shared with loop quantum gravity [31], spin foams [33] and group field theory [22] and indirectly (mapping graphs to triangulations) with CDT [32] or Regge calculus. There are, of course, differences in the precise assignment of geometry among these approaches, and in our case there is freedom to decide on which geometric interpretation of the graph states is most appropriate<sup>11</sup>. Once past the dynamical graph states, however, there are substantial differences, coming from our focus on gravity being emergent. In loop quantum gravity, CDT, spin foams and group field theory, quantum analogs of geometry and gravitational properties such as the quantum Hilbert–Einstein action or Lorentz invariance are built into the high-energy theory. Geometric properties are present at the microscopic level: in loop quantum gravity, for example, the state space is given in terms of spin network basis states, understood as a discretization of space since [20, 21, 34] and, in addition, embedded in 3D space. The dynamics of the microstates is governed by the Hamiltonian constraint, obtained by the canonical quantization of the Hilbert–Einstein action. In spin foams or group field theory, the microscopic states similarly are based on simplicial graphs or complexes carrying algebraic geometric data, including the

<sup>11</sup> For example, our basic trick for dealing with superpositions of graphs by regarding them as subgraphs of  $K_N$  can be used in those approaches too, as can the useful mapping of  $K_N$  to its line graph that we used for the mean field theory analysis in [44].

Lorentz group. Depending on the model, a spin foam may be based on states that are embedded complexes or combinatorial (abstract, non-embedded graphs). In many models, the states carry representations of the Lorentz group, while the dynamics uses the Regge action, a discrete form of the Hilbert–Einstein action. In CDT, the states are simplicial combinatorial objects, again encoding microscopic Lorentzian geometry, while the dynamics contains the Regge action. Of course, the important question is how many of these features survive at the low energy or continuum limit. Is the input of geometric and gravitational properties at the microscopic level necessary for the continuum limit to be gravity? The answer is at present unclear. CDT results indicate that while certain properties such as causality (or topological restrictions) are important, others, including the gravitational action, may not be [35]. In spin foams or group field theory, it is not clear whether the Lorentz group representations on the microstates ensure the reappearance of the Lorentz group in the continuum limit. In contrast, we regard quantum gravity as an emergent phenomenon from the low-energy theory of a condensed matter system. Emergence is concerned with the study of the macroscopic properties of systems with many bodies. Sometimes, these properties can be tracked down to the properties of the elementary constituents, but in recent years there has been a flourishing of novel quantum systems which show behaviors of the whole system that have no explanation in terms of the constituting particles, but instead of their collective behavior and interaction. In this spirit, in our approach the fundamental theory has no geometric or gravitational microscopic degrees of freedom, no elements of the Lorentz group, etc. These are the obvious similarities and differences. Then there are several issues in which similar questions arise between our approach and other background-independent ones, where the connection is not clear. An example is diffeomorphism invariance and whether this can be an emergent symmetry, an issue also present, for example, in loop quantum gravity (see, e.g., [36]). Below we discuss some of these issues in a more general context.

This research program presents numerous new possibilities and new issues. We will now sketch out some of those, for the central issues of time and gravity, and quantum mechanics.

*Time and Lorentz invariance.* It is natural for a researcher in quantum gravity with a background in relativity to object to the kind of approach we are advocating here. The spin system has a fundamental Hamiltonian; that is, we are using non-relativistic quantum mechanics. How can we possibly reconcile this with time in relativity? In fact, there are a number of different possibilities, some of which we will sketch out here:

1. Effective finite light cones from non-relativistic quantum mechanics. The physics of the LR speed that we discussed in section 3.1 provides a way to have both non-relativistic quantum mechanics and finite light cones. These become reconciled simply because, even though there *is* a signal outside the light cone, it is exponentially suppressed. Note that the LR light cone is an effective light cone but cannot be called emergent as it is still at the level of the spin system. It needs to be investigated whether, combined with the string network mechanism and the emergent Maxwell equations, we not only have finite light cones but also have emergent Minkowski spacetime.
2. Microscopic versus emergent internal time. At least naively, condensed matter-type approaches to quantum gravity must break Lorentz invariance because of the presence of a lattice and the associated fundamental discreteness and preferred frame. Recently, substantial attention has been paid to Lorentz invariance violations and there are tight constraints on certain types of violations, with more data coming in in the near future.

Nonetheless, there are still numerous possibilities for Lorentz invariance violations that are not so constrained [40]. Perhaps a little less naively, the relevant question is: what is the physics seen by an observer that lives inside the system and who is at low energy and hence has no access to the Planckian lattice. For example, in analogue gravity [4], the external metric is flat Newtonian, while the internal one is Lorentzian.

An important issue to emphasize is that the models of [12] and [14], as well as condensed matter approaches to quantum gravity in general, assume the existence of a notion of time and of time evolution as given by a Hamiltonian, as opposed to the constrained evolution of canonical pure gravity. It is a general question for all condensed matter approaches to quantum gravity whether such evolution is consistent with the diffeomorphism invariance of general relativity. While it is not possible to settle this question without first finding out whether the condensed matter microscopic system has a low-energy phase which is general relativity, we can analyze the problem further, following the direction outlined in [15]. In general, there are two possible notions of time: the time related to the  $g_{00}$  component of the metric describing the geometry at low energy and the time parameter in the fundamental microscopic Hamiltonian. Let us call the first ‘geometric time’ and the second ‘fundamental time’. In our context, it is clear that the geometric time will only appear at low energy, when geometry appears. The problem of the emergence of geometric time is the same as the problem of the emergence of space, of geometry. The constrained evolution of general relativity refers to geometric time. By making the geometry not fundamental, we are able to make a distinction between the geometric and the fundamental time, which opens up the possibility that, while the geometric time is a symmetry, the fundamental time is real. It is important to note that the relation between geometric and fundamental time is nontrivial in the systems we are studying and that the existence of a fundamental time does not necessarily imply a preferred geometric time. We also note that, in the presence of matter in general relativity, a proper time can be identified. In addition, the system studied in [14] has matter and in that sense it is perhaps more natural that it also has a straightforward notion of time.

3. Lorentz invariance as a property of the matter. Another point, relevant to these kinds of approaches to quantum gravity, is that Lorentz invariance can be viewed as a symmetry of the geometry—the physics of Minkowski spacetime—or as a symmetry of the matter. The latter is how Lorentz viewed his transformations: when a charged particle moves, its field lines get deformed (see, e.g., [25, 39]). Viewing the Lorentz transformations as a property of the spacetime amounts to postulating that there is a maximum speed and that all massless particles travel at that speed. In Einstein’s theory, there is no reason why this is so, but introducing the postulate by introducing the Minkowski spacetime is a very elegant way to impose universality of the speed of light. Emergent gravity approaches do not have this option. Instead, it is a common feature, and perhaps the most important challenge to such approaches, that different particle species can travel at different speeds.

Obtaining relativistic spacetime out of a microscopic quantum mechanical system amounts to showing that diffeomorphisms is an emergent symmetry. It makes sense to start with Lorentz invariance as an emergent symmetry, as an intermediate step, especially given the recent attention paid to possible observational signatures of violations of Lorentz invariance [40] and the possibility that Lorentz invariance may be emergent [41]. Since research in quantum gravity has for long suffered from a lack of access to observations, it is a very attractive

feature of emergent gravity approaches that observations can prove them wrong, and in the near future.

*Quantum effects of dynamical geometry.* It is common to expect that a quantum theory of gravity will, at some level, be described by a quantum superposition of geometries. This superposition presents many mathematical and conceptual problems. A very important one is the emergence of our classical world: why do we not see superpositions of geometry in everyday life? This is in effect the measurement problem in a cosmological context, one of the topics of study in quantum cosmology. For us, the geometry is given by quantum states describing graph configurations, and we have indicated a mechanism for the emergence of metastable geometries from unitary evolution. As long as the metastable states do not present macroscopic entanglement, the coarse grained observables would behave as if there is no superposition. But why we do not have such macroscopic entanglement? This is an open problem but our approach makes it attackable. If the state we consider is the ground state of a gapped local Hamiltonian, all correlations, functions decay exponentially and the total amount of entanglement is very low. The behavior of the system away from equilibrium is more complicated, but lately there has been a flourishing of results about the time evolution of quantum many-body systems after a quantum quench [30], and our models can be studied in the same way.

At very short scales, however, it is possible to have an equal superposition of different geometries. Time evolution can, for instance, produce a state that locally looks like a superposition of the first and fourth states of equation (4). This superposition is tensored with the state of the square lattice everywhere else. Imagine that we perform an experiment on the time of arrival of particles to the upper right corner. When the system collapses in the first branch, we will have a certain time of arrival  $t_1$ . On the other hand, when the system collapses in the second branch, a particle can jump directly from the lower left corner to the upper right corner, yielding an arrival time  $t_2 < t_1$ . On a large scale, in every branch there are some of these shortcuts and they average out. But at a short scale, one can observe the superposition of the state with and without a shortcut. There is a second important quantum effect to consider. Since matter and geometry can be entangled, the evolution of the matter alone cannot be unitary. The evolution of the matter alone must be described in the setting of an open quantum system, and there is decoherence towards the spatial degrees of freedom. Therefore our model implies (at low couplings) a tiny deviation for unitarity. Moreover, the decoherence rate depends on the entanglement and thus on the average degree of the graph. Since high-connected graphs correspond to strong curvature, we expect to only observe decoherence in the presence of strong gravitational effects. The entanglement between matter and geometry has also consequences for the black hole information paradox. If we consider the matter escaping from our toy black hole, we will observe a mixed state, even though we have started with a pure state. Because of the open system scenario, the black hole behaves like a quantum channel that increases the entropy. In other words, the matter coming out is thermal because it is entangled with what is inside the black hole, both edges and particles. When all the matter has escaped, there is no matter to be entangled with anymore. This is what makes the problem in the usual treatments of the black hole information paradox: we have thermal radiation coming from the entanglement of matter with nothing! But in our scenario, the disappearance of the black hole only means that the graph rearranges in a way such that the region that hosted the black hole becomes homogeneous with the surrounding space. The radiation is still entangled with all the spatial degrees of freedom in that region. The black hole can disappear but the entanglement stays.

To conclude, we are only at the beginning. We need to give a detailed description of the open quantum system to the end of making predictions about the decoherence in the presence of strong fields. We also need to exploit entanglement theory in quantum many-body systems to give a precise description of the entanglement in these models, and proving that macroscopic entanglement is negligible.

Even when we have accomplished all this, we still do not have gravity! The way in which the graph curves in the presence of matter in our model has nothing to do with what gravity does. The Hamiltonian for the graph and its interaction with the matter is too simple. The gravity model will be very complicated, with many messy terms. We think that the fundamental theory is messy, and the emergent theory is simple and elegant. So we do not go to smaller scales to *find* the simple fundamental theory, but to *explain* the beauty of the emergent one.

### Acknowledgments

The authors are indebted to Abhay Ashtekar, Francesco Caravelli, Florian Conrady, Bei-Lok Hu, John Preskill and Lee Smolin for useful advice and comments. We are grateful to the Santa Fe Institute where this work was carried out and its members for their wonderful hospitality. This work was supported by the Alexander von Humboldt Foundation and NSERC grant no. RGPIN-312738-2007. Research at Perimeter Institute is supported by the Government of Canada through Industry Canada, and by the Province of Ontario through the Ministry of Research and Innovation.

### References

- [1] Hossenfelder S 2010 Experimental search for quantum gravity arXiv:[1010.3420](#)
- [2] Strominger A 1995 Les Houches lectures on black holes arXiv:[hep-th/9501071](#)
- [3] Banks T 1999 TASI lectures on matrix theory arXiv:[hep-th/9911068](#)  
Seiberg N 2006 Emergent spacetime arXiv:[hep-th/0601234](#)
- [4] Unruh W G 1995 Dumb holes and the effects of high frequencies on black hole evaporation *Phys. Rev. D* **51** 2827–38 (arXiv:[gr-qc/9409008](#))  
Visser M and Weinfurter S 2007 Analogue spacetimes: Toy models for ‘quantum gravity’ arXiv:[0712.0427](#)
- [5] Volovik G 2009 *The Universe in a Helium Droplet* (Oxford: Oxford University Press)
- [6] Horava P 2009 Quantum gravity at a Lifshitz point *Phys. Rev. D* **79** 084008
- [7] Jacobson T 1995 *Phys. Rev. Lett.* **75** 1260–3 (arXiv:[gr-qc/9504004](#))
- [8] Hu B L 2005 Can spacetime be a condensate? arXiv:[gr-qc/0503067](#)
- [9] Ambjorn J, Jurkiewicz J and Loll R 2009 The self-organized de Sitter universe *Int. J. Mod. Phys. D* **17** 2515–20 (arXiv:[0806.0397](#))
- [10] Isham C J 1993 Prima facie questions in quantum gravity arXiv:[gr-qc/9310031](#)  
Isham C J and Butterfield J 1999 On the emergence of time in quantum gravity *The Arguments of Time* ed J Butterfield (Oxford: Oxford University Press) (arXiv:[gr-qc/9901024](#))
- [11] Konopka T, Markopoulou F and Smolin L 2006 Quantum graphity arXiv:[hep-th/0611197](#)
- [12] Konopka T, Markopoulou F and Severini S 2008 Quantum graphity: a model of emergent locality *Phys. Rev. D* **77** 104029
- [13] Hama A, Markopoulou F, Premont-Schwarz I and Severini S 2008 Lieb–Robinson bounds and the speed of light from topological order *Phys. Rev. Lett.* **102** 017204 (arXiv:[0808.2495](#))
- [14] Hama A, Markopoulou F, Lloyd S, Caravelli F, Severini S and Markstrom K 2009 A quantum Bose–Hubbard model with evolving graph as toy model for emergent spacetime *Phys. Rev. D* **81** 104032 (arXiv:[0911.5075](#))



- [15] Markopoulou F 2008 Space does not exist, so time can *Third Prize Essay in Nature of Time Competition* <http://www.fqxi.org/community/essay/winners> (arXiv:0909.1861)
- [16] Levin M and Wen X G 2005 *Rev. Mod. Phys.* **77** 871  
Gu Z-C and Wen X-G 2006 arXiv:gr-qc/0606100v1
- [17] Lieb E H and Robinson D W 1972 *Commun. Math. Phys.* **28** 251
- [18] Caravelli F and Markopoulou F 2010 Properties of quantum graphity at low temperature arXiv:1008.1340
- [19] Rovelli C and Smolin L 1995 Discreteness of area and volume in quantum gravity *Nucl. Phys. B* **442** 593  
Ashtekar A and Lewandowski J 1997 *Class. Quantum Grav.* **14** A55  
Lewandowski J 1997 *Class. Quantum Grav.* **14** 71
- [20] Penrose R 1971 *Quantum Theory and Beyond* ed T Bastin (Cambridge: Cambridge University Press) p 875
- [21] Hasslacher B and Perry M J 1981 *Phys. Lett. B* **103** 21 (arXiv:gr-qc/0404018)
- [22] Oriti D 2007 *Quantum Gravity* ed B Fauser, J Tolksdorf and E Zeidler (Basel: Birkhauser) (arXiv:gr-qc/0512103)
- [23] Sorkin R 2003 Causal sets: discrete gravity (Notes for the Valdivia Summer School) *Proc. Valdivia Summer School* ed A Gombero and D Marolf (arXiv:gr-qc/0309009)
- [24] Markopoulou F 2000 *Class. Quantum Grav.* **17** 2059
- [25] Dreyer O 2009 Emergent general relativity *Approaches to Quantum Gravity—Toward a New Understanding of Space, Time and Matter* ed D Oriti (Cambridge: Cambridge University Press) (arXiv:gr-qc/0604075)
- [26] Konopka T 2008 Statistical mechanics of graphity models *Phys. Rev. D* **78** 044032 (arXiv:0805.2283)
- [27] Conrady F 2010 Space as a low-temperature regime of graphs arXiv:1009.3195
- [28] Tasaki H 1998 *Phys. Rev. Lett.* **80** 1373
- [29] Page D 1993 Average entropy of a subsystem *Phys. Rev. Lett.* **71** 1291 (arXiv:gr-qc/9305007)
- [30] Calabrese P and Cardy J 2006 *Phys. Rev. Lett.* **96** 136801  
Polkovnikov A, Sengupta K, Silva A and Vengalattore M 2009 arXiv:1007.5331  
Tsomokos D I et al 2009 *Phys. Rev. A* **80** 060302
- [31] Ashtekar A and Lewandowski J 2005 *Class. Quantum Grav.* **21** R53 (arXiv:gr-qc/0404018)  
Rovelli C 2005 *Quantum Gravity* (Cambridge: Cambridge University Press)
- [32] Ambjørn J and Loll R 1998 *Nucl. Phys. B* 536–407 (arXiv:hep-th/9805108)  
Ambjørn J, Jurkiewicz J and Loll R 2004 *Phys. Rev. Lett.* **93** 131301 (arXiv:hep-th/0404156)
- [33] Freidel L 2005 *Int. J. Theor. Phys.* **44** 1769 (arXiv:hep-th/0607032)
- [34] Rovelli C and Smolin L 1995 Discreteness of area and volume in quantum gravity *Nucl. Phys. B* **442** 593  
Ashtekar A and Lewandowski J 1997 *Class. Quantum Grav.* **14** 55  
Lewandowski J 1997 *Class. Quantum Grav.* **14** 71
- [35] Ambjorn J, Jurkiewicz J and Loll R 2009 The self-organized de sitter universe *Int. J. Mod. Phys. D* **17** 2515–20 (arXiv:0806.0397)
- [36] Dittrich B 2008 Diffeomorphism symmetry in quantum gravity models arXiv:0810.3594v2
- [37] Hama A, Markopoulou F and Riera A 2011 Trapped surfaces and emergent curved spacetime in the Bose–Hubbard model arXiv:1108.2013
- [38] Hama A, Markopoulou F, Prémont-Schwarz I and Severini S 2009 *Phys. Rev. Lett.* **102** 017204
- [39] Bell J S 1988 *Speakable and Unspeakable in Quantum Mechanics* (Cambridge: Cambridge University Press)
- [40] Mattingly D 2005 *Living Rev. Rel.* **8** 5 (arXiv:gr-qc/0502097)  
Hossenfelder S 2010 Experimental search for quantum gravity arXiv:1010.3420v1  
Amelino-Camelia G 2008 Quantum gravity phenomenology arXiv:0806.0339v1
- [41] Jacobson T and Wall A 2010 Black hole thermodynamics and Lorentz symmetry *Found. Phys.* **40** 1076–80 (arXiv:0804.2720)
- [42] D’Ariano G M and Tosini A 2010 Space-time and special relativity from causal networks arXiv:1008.4805
- [43] Lee S-S 2009 Emergence of gravity from interacting simplices *Int. J. Mod. Phys. A* **24** 4271–86
- [44] Caravelli F and Markopoulou F 2011 Properties of quantum graphity at low temperature *Phys. Rev. D* **84** 024002