

Inclusion of Matter in Inhomogeneous Loop Quantum Cosmology

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Based on: *Matter in inhomogeneous loop quantum cosmology: the Gowdy T^3 model* (Phys. Rev. D **83**, 084050 (2011)) [arXiv:1012.2324](https://arxiv.org/abs/1012.2324)

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Introduction

- Loop Quantum Cosmology (LQC)
 - Quantum approach for cosmological systems inspired by the Loop Quantum Gravity (LQG)
 - Satisfactory quantization of several homogeneous cosmological models
 - New quantum phenomenology \Rightarrow Resolution of initial singularity
- Hybrid quantization: Inhomogeneous models
 - Reduced model with only global constraints
 - It assumes that the most relevant effects of loop quantum geometry are in the homogeneous degrees of freedom
 - Combine LQC quantization for this homogeneous sector with a Fock quantization for inhomogeneities
- First model studied: vacuum Gowdy model with the three torus topology [Garay, Martín-Benito, Mena Marugán, '08]

Motivation

- Inclusion of a massless scalar field in the Gowdy \mathbb{T}^3 model
 - Minimally coupled
 - Same symmetries of the geometry
- Motivation
 - Inclusion of matter inhomogeneities in LQC.
 - Study of more realistic models, closer to the observed universe.
 - Scenario in which one can study some interesting features
 - Quantum effects of the inhomogeneities and the anisotropies on an FRW background.
 - Robustness of the Big Bounce scenario of LQC.
 - Changes in the evolution of the matter inhomogeneities due to quantum geometry effects.
 - *projection* to more symmetric quantum models.

Classical Settings

■ Reduced phase space

- Homogeneous sector: Bianchi I + homogeneous massless scalar field ϕ
- Inhomogeneous sector: Matter inhomogeneities and gravitational waves (propagating in $\theta \in S^1$)

■ Ashtekar-Barbero variables for Bianchi I

- $su(2)$ connection: \mathcal{A}^j ; densitized triad: p_j $i \in \{\theta, \sigma, \delta\}$

■ Satisfactory Fock quantization of the inhomogeneities:

- Unitary dynamics + Vacuum invariant under S^1 translations.
- Parametrization of the matter φ and gravitational ξ inhomogeneities
- Creation-annihilation variables (free m. s. f.): $\left(a_m^{(\alpha)*}, a_m^{(\alpha)}\right)$, $\alpha = \xi, \varphi$

■ Two global constraints remain:

- Diffeomorphism constraint: $C_\theta = C_\theta^\xi + C_\theta^\varphi$
- Densitized Hamiltonian constraint: $\mathcal{C} = \mathcal{C}_{\text{hom}} + \mathcal{C}_{\text{inh}}$

Hybrid Quantization: Kinematics

Kinemematical Hilbert space

$$\mathcal{H}_{\text{kin}} = \mathcal{H}_{\text{kin}}^{\text{hom}} \otimes \mathcal{H}_{\text{kin}}^{\text{inh}} = \mathcal{H}_{\text{kin}}^{\text{BI}} \otimes L^2(\mathbb{R}, d\phi) \otimes \mathcal{F}^{\xi} \otimes \mathcal{F}^{\varphi}$$

■ Fock Spaces \mathcal{F}^{α} :

- $a_m^{(\alpha)*}, a_m^{(\alpha)} \rightarrow \hat{a}_m^{(\alpha)\dagger}, \hat{a}_m^{(\alpha)}$: creation-annihilation operators.
- n-particle states: $|\mathbf{n}^{\alpha}\rangle = |\dots, n_{-m}^{\alpha}, \dots, n_m^{\alpha}, \dots\rangle$, $n_m^{\alpha} \in \mathbb{N}$, $\sum_m n_m^{\alpha} < \infty$

■ $\mathcal{H}_{\text{kin}}^{\text{hom-mat}} = L^2(\mathbb{R}, d\phi)$:

- Standard Schrödinger quantization: $\hat{\phi}, \hat{p}_{\phi} = -i\hbar\partial_{\phi}$

■ Bianchi I kinematical Hilbert space:

- LQC: Triads p_i and holonomies of connections $\mathcal{N}_{\mu_i}(c_i)$.
- Improved dynamics: minimum length, $\bar{\mu}_j$, in the holonomies.
- $\mathcal{H}_{\text{kin}}^{\text{BI}} = \overline{\text{span}\{|\lambda_{\theta}, \lambda_{\sigma}, v\rangle : \lambda_{\theta}, \lambda_{\sigma}, v \in \mathbb{R}\}}$ $v = 2\lambda_{\theta}\lambda_{\sigma}\lambda_{\delta}$
- Discrete inner product: $\langle\lambda_{\theta}, \lambda_{\sigma}, v|\lambda'_{\theta}, \lambda'_{\sigma}, v\rangle = \delta_{\lambda_{\theta}\lambda'_{\theta}}\delta_{\lambda_{\sigma}\lambda'_{\sigma}}\delta_{vv'}$.
- $\hat{\mathcal{N}}_{\pm\bar{\mu}_i}$: Scale λ_i such that shift v in ± 1 ; $\hat{p}_j : p_j \propto \text{sgn}(\lambda_j)\lambda_j^2$.

Operators on the inhomogeneous Hilbert space

■ Diffeomorphism constraint operator

$$\hat{\mathcal{C}}_\theta = \sum_{m=1}^{\infty} m \left(\hat{X}_m^\xi + \hat{X}_m^\varphi \right), \quad \hat{X}_m^\alpha = \hat{a}_m^{(\alpha)\dagger} \hat{a}_m^{(\alpha)} - \hat{a}_{-m}^{(\alpha)\dagger} \hat{a}_{-m}^{(\alpha)}.$$

$$\blacksquare \hat{\mathcal{C}}_\theta |\mathbf{n}^\xi\rangle \otimes |\mathbf{n}^\varphi\rangle \Rightarrow \sum_{m=1}^{\infty} m (X_m^\xi + X_m^\varphi) = 0, \quad X_m^\alpha = n_m^\alpha - n_{-m}^\alpha.$$

$$\blacksquare \mathcal{F}_p \equiv \text{proper subspace of } \mathcal{F}_\xi \otimes \mathcal{F}_\varphi.$$

■ Operators in $\hat{\mathcal{C}}_{\text{inh}}$

$$\blacksquare \hat{H}_0 = \sum_{\alpha \in \{\xi, \varphi\}} \sum_{m=1}^{\infty} m \hat{N}_m^\alpha, \quad \hat{N}_m^\alpha = \hat{a}_m^{(\alpha)\dagger} \hat{a}_m^{(\alpha)} + \hat{a}_{-m}^{(\alpha)\dagger} \hat{a}_{-m}^{(\alpha)}.$$

$$\blacksquare \hat{H}_{\text{int}} = \sum_{\alpha \in \{\xi, \varphi\}} \sum_{m=1}^{\infty} \frac{1}{m} \left(\hat{N}_m^\alpha + \hat{a}_m^{(\alpha)\dagger} \hat{a}_{-m}^{(\alpha)\dagger} + \hat{a}_m^{(\alpha)} \hat{a}_{-m}^{(\alpha)} \right).$$

Hamiltonian constraint operator $\hat{\mathcal{C}} = \hat{\mathcal{C}}_{\text{hom}} + \hat{\mathcal{C}}_{\text{inh}}$

- $\hat{\mathcal{C}}_{\text{hom}} = - \sum_{i \neq j} \sum_j \frac{\hat{\Theta}_i \hat{\Theta}_j}{16\pi G \gamma^2} - \frac{\hbar^2}{2} \left[\frac{\partial}{\partial \phi} \right]^2, \quad i, j \in \{\theta, \delta, \sigma\}.$
- $\hat{\mathcal{C}}_{\text{inh}} = 2\pi\hbar \widehat{|p_\theta|} \hat{H}_0 + \hbar \left[\frac{1}{|p_\theta|^{\frac{1}{4}}} \right]^2 \frac{(\hat{\Theta}_\delta + \hat{\Theta}_\sigma)^2}{16\pi\gamma^2} \left[\frac{1}{|p_\theta|^{\frac{1}{4}}} \right]^2 \hat{H}_{\text{int}}.$
- $\hat{\Theta}_i = \widehat{c_i p_i} = i\pi\gamma G \hbar \widehat{\sqrt{|v|}} \left[(\hat{\mathcal{N}}_{-2\bar{\mu}_i} - \hat{\mathcal{N}}_{2\bar{\mu}_i}) \widehat{\text{sgn}(p_i)} + \widehat{\text{sgn}(p_i)} (\hat{\mathcal{N}}_{-2\bar{\mu}_i} - \hat{\mathcal{N}}_{2\bar{\mu}_i}) \right] \widehat{\sqrt{|v|}}$
- Symmetric factor ordering:
 - Triad operators: $v = 0$ states decouple (kin. singularity resolution)
 - $\hat{\Theta}_j$ operators do not mix states with different sign of $\lambda_\theta, \lambda_\sigma, v$.
- $\tilde{\mathcal{H}}_{\text{kin}}^{\text{BI}}$: states such that $\lambda_\theta, \lambda_\sigma, v > 0 \Rightarrow \Lambda_\theta = \log \lambda_\theta, \Lambda_\sigma = \log \lambda_\sigma.$
- Superselection sectors:
 - in v : $v \in \mathcal{L}_\epsilon = \{\epsilon + 4k; k \in \mathbb{N}\}, \quad \epsilon \in (0, 4]$
 - in Λ_a : Given an initial $\Lambda_a^\star \Rightarrow \Lambda_a = \Lambda_a^\star + z_\epsilon, z_\epsilon \in \mathcal{Z}_\epsilon$

Physical Hilbert space

- Action of the Hamiltonian constraint
 - The coefficients do not depend on Λ_σ
 - It is a difference equation in $v \Rightarrow$ evolution equation in v
 - The solutions can be determined by a set of initial data on the section of minimum homogeneous volume
- Physical Hilbert space $\mathcal{H}_{\text{phy}} \Leftrightarrow$ Hilbert space of initial data

$$\mathcal{H}_{\text{p}} = \mathcal{H}_{\text{phys}}^{\text{BI}} \otimes L^2(\mathbb{R}, d\phi) \otimes \mathcal{F}_{\text{p}}$$

- $\mathcal{H}_{\text{phys}}^{\text{BI}} \equiv$ Physical Hilbert space of Bianchi I

Projection to LRS-Gowdy

- The model is symmetric under the interchange of σ and δ directions.
- Classical solutions with local rotational symmetry (LRS)
- General state: $|\Psi\rangle = \sum_{\Lambda_\theta, \Lambda_\sigma, v} |\Psi(\Lambda_\theta, \Lambda_\sigma, v)\rangle \otimes |\Lambda_\theta, \Lambda_\sigma, v\rangle$
- *Projection map:*

$$|\Psi(\Lambda_\theta, \Lambda_\sigma, v)\rangle \longrightarrow \sum_{\Lambda_\sigma} |\Psi(\Lambda_\theta, \Lambda_\sigma, v)\rangle \equiv |\psi(\Lambda_\theta, v)\rangle$$

Quantum Gowdy Model $\xrightarrow{\text{projection}}$ Quantum LRS-Gowdy Model

- *projection* over Λ_θ to get the isotropic Gowdy model fails
 - There is no classical inhomogeneous and isotropic solutions.

Conclusions

- Satisfactory quantization of the Gowdy T^3 model with linearly polarized gravitational waves and a massless scalar field.
- Hybrid quantization applied as in the vacuum model.
- Inclusion of the matter field:
 - Classical isotropic solutions of the homogeneous sector.
 - Two “copies” of inhomogeneities (mathematically speaking).
 - Matter inhomogeneities in LQC.
- Same results as in the vacuum model.
 - Standard Fock quantization of the inhomogeneities is recovered.
 - Classical singularity resolved at the kinematical level.
- Study of the *projection* to more symmetric systems.
- Possibility of analyzing the effects of the anisotropies and the inhomogeneities on a flat FRW model. (Work in progress)

Thanks for your attention!