# KLT relations from the Einstein-Hilbert Lagrangian 

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#### Abstract

The Kawai-Lewellen-Tye (KLT) relations derived from string theory tell us that perturbative gravity amplitudes are the "square" of the corresponding amplitudes in gauge theory. Starting from the light-cone Lagrangian for pure gravity we make these relations manifest off-shell, for three- and four-graviton vertices, at the level of the action. © 2007 Elsevier B.V. All rights reserved.


## 1. Introduction

The Kawai-Lewellen-Tye (KLT) relations relate tree-level amplitudes in closed and open string theories [1]. In the field theory limit the KLT relations, for three- and four-point amplitudes, reduce to ${ }^{1}$

$$
\begin{equation*}
M_{3}^{\mathrm{tree}}(1,2,3)=A_{3}^{\mathrm{tree}}(1,2,3) A_{3}^{\mathrm{tree}}(1,2,3), \quad M_{4}^{\mathrm{tree}}(1,2,3,4)=-i s_{12} A_{4}^{\mathrm{tree}}(1,2,3,4) A_{4}^{\mathrm{tree}}(1,2,4,3) \tag{1.1}
\end{equation*}
$$

where the $M_{n}$ represent gravity amplitudes and the $A_{n}$ are color-ordered [3,4] amplitudes in pure Yang-Mills theory ( $s_{i j} \equiv-\left(p_{i}+\right.$ $\left.p_{j}\right)^{2}$ ). Although the KLT relations apply only at the tree-level they have been used, with great success, in conjunction with unitarity based methods to derive loop amplitudes in gravity [2,5]. In particular, these relations have proven invaluable in studying the ultraviolet properties of $\mathcal{N}=8$ supergravity [6]. The question of whether the KLT relations are valid only for on-shell amplitudes or, more generally, at the level of the Lagrangian remains open [7]. This is the issue we focus on in this Letter.

The tree-level amplitudes take a very compact form in a helicity basis. Thus when attempting to derive the KLT relations starting from the gravity Lagrangian it seems natural to work in light-cone gauge where only the helicity states propagate. Tree-level amplitudes in which precisely two external legs carry negative helicity are called maximally helicity violating (MHV) amplitudes. A very simple expression for all the MHV amplitudes in Yang-Mills theory was given in [8]. An MHV-Lagrangian (also referred to as the CSW Lagrangian) where the fundamental vertices are off-shell versions of the MHV amplitudes was proposed in [9]. In [10] and [11] it was shown how this MHV-Lagrangian can be derived from the usual light-cone Yang-Mills Lagrangian by a suitable field redefinition.

In this Letter we perform a field redefinition, similar to that in [10,11], on the light-cone gravity Lagrangian. Although the shifted Lagrangian is not simply the sum of MHV-vertices, the off-shell KLT relations, to the order examined in this Letter, are manifest.

[^0]
## 2. Yang-Mills

We start by sketching schematically, the proposal of [10,11] for Yang-Mills. The light-cone Yang-Mills Lagrangian is of the form

$$
\begin{equation*}
L \sim L_{+-}+L_{++-}+L_{+--}+L_{++--} \tag{2.1}
\end{equation*}
$$

where the indices, in no particular order, refer to helicity. The field redefinition maps the first two terms (the kinetic and one cubic term) into a purely kinetic term. This transformation also generates an infinite series of higher order terms producing exactly the MHV-Lagrangian

$$
\begin{equation*}
L_{Y M} \sim L_{+-}+L_{+--}+L_{++--}+L_{+++--}+L_{++++--}+\cdots+L_{(+)^{n}--}+\cdots \tag{2.2}
\end{equation*}
$$

Again, this is merely a formal way of writing the Lagrangian. For example, $L_{++--}$receives contributions from the two inequivalent orderings $\operatorname{tr}(A \bar{A} A \bar{A})$ and $\operatorname{tr}(A A \bar{A} \bar{A})$ where $A$ and $\bar{A}$ are gluons of helicity ${ }^{2}+1$ and -1 , respectively. Each trace is multiplied by an off-shell continuation (cf. Appendix A) of the appropriate Parke-Taylor amplitude $[4,8]$

$$
\begin{equation*}
\frac{\langle k l\rangle^{4}}{\prod_{i=1}^{n}\langle i(i+1)\rangle}, \quad n+1 \equiv 1 \tag{2.3}
\end{equation*}
$$

We will not go into details regarding the derivation of these results which can be found in [10-13]. The analysis in the gravity case is completely analogous and is presented in detail in Section 3. The hope is that a similar field redefinition in pure gravity will generate interaction terms which make KLT factorization manifest. The purpose of this Letter is to examine this issue.

## 3. Gravity in light-cone gauge

We follow closely, in this section, the light-cone formulation of gravity in [14]. Here, we only review the key features of this formulation and refer the reader to Appendix C in [14] for a detailed derivation of the results presented below. The Einstein-Hilbert action reads

$$
\begin{equation*}
S_{\mathrm{EH}}=\int d^{4} x \mathcal{L}=\frac{1}{2 \kappa^{2}} \int d^{4} x \sqrt{-g} R \tag{3.1}
\end{equation*}
$$

where $g=\operatorname{det} g_{\mu \nu}$ and $R$ is the curvature scalar. Light-cone gauge is chosen by setting

$$
\begin{equation*}
g_{--}=g_{-i}=0, \quad i=1,2 \tag{3.2}
\end{equation*}
$$

Our conventions and notation are explained in Appendix A. The metric is parameterized as follows

$$
\begin{equation*}
g_{+-}=-e^{\frac{\psi}{2}}, \quad g_{i j}=e^{\psi} \gamma_{i j} \tag{3.3}
\end{equation*}
$$

The field $\psi$ is real while $\gamma_{i j}$ is a $2 \times 2$ real, symmetric, unimodular matrix. The $R_{-i}=0$ constraint allows us to eliminate $g^{-i}$. From the $R_{--}=0$ constraint we find

$$
\begin{equation*}
\psi=\frac{1}{4} \frac{1}{\partial_{-}^{2}}\left(\partial_{-} \gamma^{i j} \partial_{-} \gamma_{i j}\right) \tag{3.4}
\end{equation*}
$$

The Lagrangian density now reads

$$
\begin{equation*}
\mathcal{L}=\frac{1}{2 \kappa^{2}} \sqrt{-g}\left(2 g^{+-} R_{+-}+g^{i j} R_{i j}\right) \tag{3.5}
\end{equation*}
$$

We expand this to find [15]

$$
\begin{align*}
\mathcal{L}= & \frac{1}{2 \kappa^{2}}\left\{\mathrm{e}^{\psi}\left(\frac{3}{2} \partial_{+} \partial_{-} \psi-\frac{1}{2} \partial_{+} \gamma^{i j} \partial_{-} \gamma_{i j}\right)-\mathrm{e}^{\frac{\psi}{2}} \gamma^{i j}\left(\frac{1}{2} \partial_{i} \partial_{j} \psi-\frac{3}{8} \partial_{i} \psi \partial_{j} \psi-\frac{1}{4} \partial_{i} \gamma^{k l} \partial_{j} \gamma_{k l}+\frac{1}{2} \partial_{i} \gamma^{k l} \partial_{k} \gamma_{j l}\right)\right. \\
& \left.-\frac{1}{2} \mathrm{e}^{-\frac{3}{2} \psi} \gamma^{i j} \frac{1}{\partial_{-}} R_{i} \frac{1}{\partial_{-}} R_{j}\right\} \tag{3.6}
\end{align*}
$$

where

$$
\begin{equation*}
R_{i}=\mathrm{e}^{\psi}\left(-\frac{1}{2} \partial_{-} \gamma^{j k} \partial_{i} \gamma_{j k}+\frac{3}{2} \partial_{-} \partial_{i} \psi-\frac{1}{2} \partial_{i} \psi \partial_{-} \psi\right)-\partial_{k}\left(\mathrm{e}^{\psi} \gamma^{j k} \partial_{-} \gamma_{i j}\right) \tag{3.7}
\end{equation*}
$$

This is the closed form of the Lagrangian.

[^1]
### 3.1. The perturbative expansion

In order to obtain a perturbative expansion of the metric we choose

$$
\gamma_{i j}=\left(\mathrm{e}^{\kappa H}\right)_{i j}, \quad H=\frac{1}{\sqrt{2}}\left(\begin{array}{cc}
h+\bar{h} & -i(h-\bar{h})  \tag{3.8}\\
-i(h-\bar{h}) & -h-\bar{h}
\end{array}\right),
$$

where $h$ and $\bar{h}$ represent gravitons of helicity +2 and -2 , respectively. The light-cone Lagrangian density for pure gravity, to order $\kappa^{2}$ [14], reads $^{3}$

$$
\begin{align*}
\mathcal{L}= & \bar{h} \square h+2 \kappa \bar{h} \partial_{-}^{2}\left(\frac{\bar{\partial}}{\partial_{-}} h \frac{\bar{\partial}}{\partial_{-}} h-h \frac{\bar{\partial}^{2}}{\partial_{-}^{2}} h\right)+2 \kappa h \partial_{-}^{2}\left(\frac{\partial}{\partial_{-}} \bar{h} \frac{\partial}{\partial_{-}} \bar{h}-\bar{h} \frac{\partial^{2}}{\partial_{-}^{2}} \bar{h}\right) \\
& +2 \kappa^{2}\left\{\frac{1}{\partial_{-}^{2}}\left(\partial_{-} h \partial_{-} \bar{h}\right) \frac{\partial \bar{\partial}}{\partial_{-}^{2}}\left(\partial_{-} h \partial_{-} \bar{h}\right)+\frac{1}{\partial_{-}^{3}}\left(\partial_{-} h \partial_{-} \bar{h}\right)\left(\partial \bar{\partial} h \partial_{-} \bar{h}+\partial_{-} h \partial \bar{\partial} \bar{h}\right)\right. \\
& -\frac{1}{\partial_{-}^{2}}\left(\partial_{-} h \partial_{-} \bar{h}\right)\left(2 \partial \bar{\partial} h \bar{h}+2 h \partial \bar{\partial} \bar{h}+9 \bar{\partial} h \partial \bar{h}+\partial h \bar{\partial} \bar{h}-\frac{\partial \bar{\partial}}{\partial_{-}} h \partial_{-} \bar{h}-\partial_{-} h \frac{\partial \bar{\partial}}{\partial_{-}} \bar{h}\right) \\
& -2 \frac{1}{\partial_{-}}\left(2 \bar{\partial} h \partial_{-} \bar{h}+h \partial_{-} \bar{\partial} \bar{h}-\partial_{-} \bar{\partial} h \bar{h}\right) h \partial \bar{h}-2 \frac{1}{\partial_{-}}\left(2 \partial_{-} h \partial \bar{h}+\partial_{-} \partial h \bar{h}-h \partial_{-} \partial \bar{h}\right) \bar{\partial} h \bar{h} \\
& -\frac{1}{\partial_{-}}\left(2 \bar{\partial} h \partial_{-} \bar{h}+h \partial_{-} \bar{\partial} \bar{h}-\partial_{-} \bar{\partial} h \bar{h}\right) \frac{1}{\partial_{-}}\left(2 \partial_{-} h \partial \bar{h}+\partial_{-} \partial h \bar{h}-h \partial_{-} \partial \bar{h}\right) \\
& \left.-h \bar{h}\left(\partial \bar{\partial} h \bar{h}+h \partial \bar{\partial} \bar{h}+2 \bar{\partial} h \partial \bar{h}+3 \frac{\partial \bar{\partial}}{\partial_{-}} h \partial_{-} \bar{h}+3 \partial_{-} h \frac{\partial \bar{\partial}}{\partial_{-}} \bar{h}\right)\right\} . \tag{3.9}
\end{align*}
$$

As in (2.1) the three-vertex terms are of the form $(-,+,+)$ and $(+,-,-)$. In analogy to Yang-Mills, a solution to the self-duality condition

$$
\begin{equation*}
R_{\mu \nu \rho \sigma}=\frac{i}{2} \epsilon_{\mu \nu}^{\alpha \beta} R_{\alpha \beta \rho \sigma} \tag{3.10}
\end{equation*}
$$

is

$$
\begin{equation*}
\bar{h}=0, \quad \square h+2 \kappa \partial_{-}^{2}\left(\frac{\bar{\partial}}{\partial_{-}} h \frac{\bar{\partial}}{\partial_{-}} h-h \frac{\bar{\partial}^{2}}{\partial_{-}^{2}} h\right)=0 \tag{3.11}
\end{equation*}
$$

where the second relation is the $\bar{h}$ equation of motion (at $\bar{h}=0$ ). Thus, as in Yang-Mills, we will map the first two terms in (3.9) to a free theory. Further discussions regarding this point may be found in [12].

### 3.2. The field redefinition

We seek a transformation $(h, \bar{h}) \rightarrow(C, \bar{C})$ such that ${ }^{4}$

$$
\begin{equation*}
K=-\bar{h} \partial_{+} \partial_{-} h+\bar{h} V(h)=-\bar{C} \partial_{+} \partial_{-} C+\bar{C} \partial \bar{\partial} C \tag{3.12}
\end{equation*}
$$

where

$$
\begin{equation*}
V(h)=\partial \bar{\partial} h+\kappa \partial_{-}^{2}\left(\frac{\bar{\partial}}{\partial_{-}} h \frac{\bar{\partial}}{\partial_{-}} h-h \frac{\bar{\partial}^{2}}{\partial_{-}^{2}} h\right) . \tag{3.13}
\end{equation*}
$$

The remaining three- and four-point vertices in (3.9) all involve exactly two negative helicity gravitons. Since MHV amplitudes also involve exactly two negative helicity legs, we aim to preserve this structure. ${ }^{5}$ In analogy with Yang-Mills, we choose $h$ to be a function of $C$ alone while $\bar{h}$ is chosen to be a function of both $C$ and $\bar{C}$. This field redefinition is not unique and we will comment on this below.

To find the explicit transformation, which is in fact a canonical transformation on the phase space with coordinates $\left(C\right.$, $\left.\pi_{C}\right)$, we start with a generating function of the form $G\left(C, \pi_{h}\right)=\int g(C) \pi_{h}$. Then

$$
\begin{equation*}
\pi_{C} \equiv \frac{\partial L}{\partial\left(\partial_{+} C\right)}=\partial_{-} \bar{C}=\frac{\delta G}{\delta C}=\int \frac{\delta g}{\delta C} \pi_{h}, \quad h=\frac{\delta G}{\delta \pi_{h}}=g(C) \tag{3.14}
\end{equation*}
$$

[^2]Since $\pi_{h}=\partial_{-} \bar{h}$ we have

$$
\begin{equation*}
\partial_{-} \bar{C}(y)=\int d^{3} x \partial_{-} \bar{h}(x) \frac{\delta h(x)}{\delta C(y)}, \tag{3.15}
\end{equation*}
$$

where the integral is performed on a surface of constant $x^{+}$. The Lagrangian density then reads (here and below we drop surface terms)

$$
\begin{equation*}
\mathcal{L}=-\bar{C} \partial_{+} \partial_{-} C+\bar{C} \partial \bar{\partial} C=\partial_{-} \bar{C} \partial_{+} C-\partial_{-} \bar{C} \frac{\partial \bar{\partial}}{\partial_{-}} C . \tag{3.16}
\end{equation*}
$$

Using (3.15) the Lagrangian becomes

$$
\begin{equation*}
L=\int d^{3} x \partial_{-} \bar{h}(x) \partial_{+} h(x)-\int d^{3} x \int d^{3} y \partial_{-} \bar{h}(y) \frac{\partial \bar{\partial}}{\partial_{-}} C(x) \frac{\delta h(y)}{\delta C(x)} . \tag{3.17}
\end{equation*}
$$

We want this to be equal to

$$
\begin{equation*}
L=\int d^{3} x\left(\partial_{-} \bar{h}(x) \partial_{+} h(x)-\partial_{-} \bar{h}(x) \frac{1}{\partial_{-}} V(h(x))\right), \tag{3.18}
\end{equation*}
$$

implying that

$$
\begin{equation*}
\frac{\partial \bar{\partial}}{\partial_{-}} h(x)+\kappa \partial_{-}\left(\frac{\bar{\partial}}{\partial_{-}} h \frac{\bar{\partial}}{\partial_{-}} h-h \frac{\bar{\partial}^{2}}{\partial_{-}^{2}} h\right)(x)=\int d^{3} y \frac{\partial \bar{\partial}}{\partial_{-}} C(y) \frac{\delta h(x)}{\delta C(y)} . \tag{3.19}
\end{equation*}
$$

In momentum space, this becomes

$$
\begin{equation*}
\frac{p \bar{p}}{p_{-}} h\left(p_{-}\right)-\int d^{3} m \frac{m \bar{m}}{m_{-}} C(m) \frac{\delta h(p)}{\delta C(m)}=-\kappa \int d^{3} k d^{3} l \delta^{(3)}(p-k-l)\left(k_{-}+l_{-}\right)\left(\frac{\bar{k} \bar{l}}{k_{-} l_{-}}-\frac{\bar{l}^{2}}{l_{-}^{2}}\right) h(k) h(l) . \tag{3.20}
\end{equation*}
$$

For $h$, we choose the ansatz

$$
\begin{equation*}
h(p)=\sum_{n=1}^{\infty} \int \prod_{i=1}^{n} d^{3} k_{i} Z^{(n)}\left(p_{1}, k_{1}, \ldots, k_{n}\right) C\left(k_{1}\right) \cdots C\left(k_{n}\right), \tag{3.21}
\end{equation*}
$$

so (3.20) implies

$$
\begin{align*}
& \int d^{3} k d^{3} l\left(\frac{p \bar{p}}{p_{-}}-\frac{k \bar{k}}{k_{-}}-\frac{l \bar{l}}{l_{-}}\right) Z^{(2)}(p, k, l) C(k) C(l) \\
& \quad=-\kappa \int d^{3} k d^{3} l\left(k_{-}+l_{-}\right)\left(\frac{\bar{k} \bar{l}}{k_{-} l_{-}}-\frac{\bar{l}^{2}}{l_{-}^{2}}\right) C(k) C(l) \delta^{(3)}(p-k-l) . \tag{3.22}
\end{align*}
$$

Thus

$$
\begin{align*}
& Z^{(1)}(p, k)=\delta^{(3)}(p-k), \\
& Z^{(2)}(p, k, l)=\frac{\kappa}{2}\left(k_{-}+l_{-}\right) \frac{\frac{\bar{l}^{2}}{l^{2}}+\frac{\bar{k}^{2}}{k_{-}^{2}}-2 \frac{\bar{k} \bar{l}}{k_{-} l_{-}}}{\frac{p \overline{\bar{L}}}{p-}-\frac{k \overline{\bar{c}}}{k_{-}}-\frac{l \bar{l}}{l_{-}}} \delta^{(3)}(p-k-l)=-\frac{\kappa}{2} \frac{p_{-}^{2}}{k_{-} l_{-}} \frac{[k l]}{\langle k l\rangle} \delta^{(3)}(p-k-l) . \tag{3.23}
\end{align*}
$$

From (3.15) we also find

$$
\begin{equation*}
p_{-} \bar{h}(p)=p_{-} \bar{C}(p)-\int d^{3} k d^{3} l k_{-}\left(Z^{(2)}(-k,-p, l)+Z^{(2)}(-k, l,-p)\right) \bar{C}(k) C(l)+\cdots, \tag{3.24}
\end{equation*}
$$

which can be rewritten as

$$
\begin{equation*}
\bar{h}(p)=\bar{C}(p)+\kappa \int d^{3} k d^{3} l \frac{k_{-}^{3}}{p_{-}^{2} l_{-}} \frac{[k l]}{\langle k l\rangle} \bar{C}(k) C(l)+\cdots \tag{3.25}
\end{equation*}
$$

It is straightforward to work out a recursion relation for the coefficients $Z^{(n)}$ which can then be solved to any desired order. We will not present the details here.

### 3.3. The shifted gravity action

After performing the field redefinition described in the previous section we find that the gravity action, to order $\kappa^{2}$, is

$$
\begin{align*}
& \int d^{4} p \bar{C}(-p) p^{2} C(p)+\kappa \int d^{4} p d^{4} k d^{4} l \frac{\langle k l\rangle^{6}}{\langle l p\rangle^{2}\langle p k\rangle^{2}} C(p) \bar{C}(k) \bar{C}(l) \delta^{(4)}(p+k+l) \\
& +\kappa^{2} \int d^{4} p d^{4} q d^{4} k d^{4} l \frac{\langle k l\rangle^{8}[k l]}{\langle k l\rangle\langle k p\rangle\langle k q\rangle\langle l p\rangle\langle l q\rangle\langle p q\rangle^{2}} C(p) C(q) \bar{C}(k) \bar{C}(l) \delta^{(4)}(p+q+k+l) \\
& \quad+\kappa^{2} \int d^{4} p d^{4} q d^{4} k d^{4} l\left(J(p, q, k, l) p^{2}+K(p, q, k, l) k^{2}\right) C(p) C(q) \bar{C}(k) \bar{C}(l) \delta^{(4)}(p+q+k+l) \tag{3.26}
\end{align*}
$$

We stress that the coefficients in the action above are off-shell. Note that the four-graviton amplitude does not receive exchange contributions due to the structure of the action at the cubic level after the field redefinitions (3.21) and (3.25). The functions $J$ and $K$ turn out to be fairly complicated but are irrelevant for on-shell four-point scattering since the third line vanishes on-shell. In particular, when interaction vertices are proportional to the free equations of motion they can be eliminated by a suitable field redefinition [16]. The required field redefinitions are ${ }^{6}$

$$
\begin{align*}
& C(p) \rightarrow C(p)-\kappa^{2} \int d^{4} q d^{4} k d^{4} l K(k, q,-p, l) C(k) C(q) \bar{C}(l) \delta^{(4)}(-p+q+k+l), \\
& \bar{C}(p) \rightarrow \bar{C}(p)-\kappa^{2} \int d^{4} q d^{4} k d^{4} l J(-p, q, k, l) \bar{C}(k) \bar{C}(l) C(q) \delta^{(4)}(-p+q+k+l) \tag{3.27}
\end{align*}
$$

and these eliminate the third line in (3.26). The light-cone action for gravity to order $\kappa^{2}$ thus reads

$$
\begin{align*}
& \int d^{4} p \bar{C}(-p) p^{2} C(p)+\kappa \int d^{4} p d^{4} k d^{4} l \frac{\langle k l\rangle^{6}}{\langle l p\rangle^{2}\langle p k\rangle^{2}} C(p) \bar{C}(k) \bar{C}(l) \delta^{(4)}(p+k+l) \\
& \quad+\kappa^{2} \int d^{4} p d^{4} q d^{4} k d^{4} l \frac{\langle k l\rangle^{8}[k l]}{\langle k l\rangle\langle k p\rangle\langle k q\rangle\langle l p\rangle\langle l q\rangle\langle p q\rangle^{2}} C(p) C(q) \bar{C}(k) \bar{C}(l) \delta^{(4)}(p+q+k+l) . \tag{3.28}
\end{align*}
$$

These off-shell vertices clearly factorize into products of off-shell MHV vertices in Yang-Mills. In particular this confirms, offshell, the relations (1.1) for three- and four-point vertices. It will be interesting to see if this KLT factorization extends to higher orders in the action where non-MHV vertices appear.

In contrast to the Yang-Mills case, the MHV vertices in gravity appear only after a further field redefinition (3.27) that removes interaction vertices proportional to the free equations of motion. This was to be expected given that the gravity Lagrangian, unlike Yang-Mills, does not stop at quartic order and that the MHV gravity amplitudes are non-holomorphic [17]. Furthermore MHV vertices in the gravity Lagrangian are not sufficient to compute all the non-MHV diagrams, at least for our choice of field variables. For example the 5-point amplitude $M^{\text {tree }}(+,+,-,-,-)$ has contributions from the MHV vertices but also from a direct contact vertex present in the original Lagrangian. ${ }^{7}$ The five-point MHV amplitude $M^{\text {tree }}(+,+,+,-,-)$ is special in that it has three contributions: one term from the original Lagrangian and two from the field redefinition acting on the three- and four-point vertices. Otherwise, as in Yang-Mills, all $n$-point $(n>5)$ MHV amplitudes are generated by the field redefinitions alone.

The discussion in the main body of this Letter dealt with light-cone gravity at tree-level. At the loop level, field redefinitions have to be considered with much greater care. If the Jacobian of the field redefinition is not unity it will lead to additional interaction terms [16]. Even if the Jacobian is classically one there may be anomalies which lead to additional interaction terms as proposed in the context of the MHV Lagrangian for Yang-Mills in [10]; see also the discussion in [12,13].

An interesting question is whether the Lagrangians of $\mathcal{N}=8$ supergravity and $\mathcal{N}=4$ super Yang-Mills share a similar relationship. Since there exist superfield formulations, in light-cone gauge, for both these theories [19] a similar analysis is certainly worth performing.

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[^3]
## Appendix A. Conventions and notation

We work with the metric $(-,+,+,+)$ and define

$$
\begin{equation*}
x^{ \pm}=\frac{1}{\sqrt{2}}\left(x^{0} \pm x^{3}\right), \quad \partial_{ \pm}=\frac{1}{\sqrt{2}}\left(\partial_{0} \pm \partial_{3}\right) \tag{A.1}
\end{equation*}
$$

$x^{+}$plays the role of light-cone time and $\partial_{+}$the light-cone Hamiltonian. $\partial_{-}$is now a spatial derivative and its inverse, $\frac{1}{\partial_{-}}$, is defined using the prescription in [20]. We define

$$
\begin{array}{ll}
x=\frac{1}{\sqrt{2}}\left(x^{1}+i x^{2}\right), & \bar{\partial} \equiv \frac{\partial}{\partial x}=\frac{1}{\sqrt{2}}\left(\partial_{1}-i \partial_{2}\right) \\
\bar{x}=\frac{1}{\sqrt{2}}\left(x^{1}-i x^{2}\right), & \partial \equiv \frac{\partial}{\partial \bar{x}}=\frac{1}{\sqrt{2}}\left(\partial_{1}+i \partial_{2}\right) \tag{A.2}
\end{array}
$$

A four-vector $p_{\mu}$ may be expressed as a bispinor $p_{a \dot{a}}$ using the $\sigma^{\mu}=(-\mathbf{1}, \boldsymbol{\sigma})$ matrices

$$
p_{a \dot{a}} \equiv p_{\mu}\left(\sigma^{\mu}\right)_{a \dot{a}}=\left(\begin{array}{cc}
-p_{0}+p_{3} & p_{1}-i p_{2}  \tag{A.3}\\
p_{1}+i p_{2} & -p_{0}-p_{3}
\end{array}\right)=\sqrt{2}\left(\begin{array}{cc}
-p_{-} & \bar{p} \\
p & -p_{+}
\end{array}\right)
$$

The determinant of this matrix is

$$
\begin{equation*}
\operatorname{det}\left(p_{a \dot{a}}\right)=-2\left(p \bar{p}-p_{+} p_{-}\right)=-p^{\mu} p_{\mu} \tag{A.4}
\end{equation*}
$$

When the vector $p_{\mu}$ is light-like we have $p_{+}=\frac{p \bar{p}}{p_{-}}$which is the on-shell condition. We then define holomorphic and antiholomorphic spinors ${ }^{8}$

$$
\begin{equation*}
\lambda_{a}=\frac{2^{\frac{1}{4}}}{\sqrt{p}_{-}}\binom{p_{-}}{-p}, \quad \tilde{\lambda}_{\dot{a}}=-\left(\lambda_{a}\right)^{*}=-\frac{2^{\frac{1}{4}}}{\sqrt{p}_{-}}\binom{p_{-}}{-\bar{p}} \tag{A.5}
\end{equation*}
$$

such that $\lambda_{a} \tilde{\lambda}_{\dot{a}}$ agrees with (A.3) on-shell. We define the off-shell holomorphic and anti-holomorphic spinor products [13]

$$
\begin{equation*}
\langle i j\rangle=\sqrt{2} \frac{p^{i} p_{-}^{j}-p^{j} p_{-}^{i}}{\sqrt{p_{-}^{i} p_{-}^{j}}}, \quad[i j]=\sqrt{2} \frac{\bar{p}^{i} p_{-}^{j}-\bar{p}^{j} p_{-}^{i}}{\sqrt{p_{-}^{i} p_{-}^{j}}} \tag{A.6}
\end{equation*}
$$

Their product is

$$
\begin{equation*}
\langle i j\rangle[i j]=s_{i j} \equiv-\left(p_{i}+p_{j}\right)^{2} \tag{A.7}
\end{equation*}
$$

## References

[1] H. Kawai, D.C. Lewellen, S.H.H. Tye, Nucl. Phys. B 269 (1986) 001.
[2] Z. Bern, L.J. Dixon, M. Perelstein, J.S. Rozowsky, Nucl. Phys. B 546 (1999) 423, hep-th/9811140.
[3] F.A. Berends, W. Giele, Nucl. Phys. B 294 (1987) 700.
[4] L. Dixon, TASI (1995) 539, hep-ph/9601359.
[5] Z. Bern, L.J. Dixon, M. Perelstein, J.S. Rozowsky, Phys. Lett. B 444 (1998) 273, hep-th/9809160.
[6] Z. Bern, J. Carrasco, L. Dixon, H. Johansson, D. Kosower, R. Roiban, hep-th/0702112.
[7] Z. Bern, A.K. Grant, Phys. Lett. B 457 (1999) 23, hep-th/9904026;
Z. Bern, Living Rev. Rel. 5 (2002) 5, gr-qc/0206071.
[8] S.J. Parke, T.R. Taylor, Phys. Rev. Lett. 56 (1986) 2459;
F.A. Berends, W.T. Giele, Nucl. Phys. B 306 (1988) 759.
[9] F. Cachazo, P. Svrcek, E. Witten, JHEP 0409 (2004) 006, hep-th/0403047.
[10] A. Gorsky, A. Rosly, JHEP 0601 (2006) 101, hep-th/0510111.
[11] P. Mansfield, JHEP 0603 (2006) 037, hep-th/0511264.
[12] J. Ettle, T. Morris, JHEP 0608 (2006) 003, hep-th/0605121;
J. Ettle, C. Fu, J. Fudger, P. Mansfield, T. Morris, hep-th/0703286.
[13] A. Brandhuber, B. Spence, G. Travaglini, K. Zoubos, arXiv: 0704.0245 [hep-th].
[14] S. Ananth, L. Brink, R. Heise, H.G. Svendsen, Nucl. Phys. B 753 (2006) 195, hep-th/0607019.
[15] J. Scherk, J.H. Schwarz, Gen. Relativ. Gravit. 6 (1975) 537;
I. Bengtsson, M. Cederwall, O. Lindgren, GOTEBORG-83-55, 1983.

[^4][16] A. Salam, J.A. Strathdee, Phys. Rev. D 2 (1970) 2869;
G.'t Hooft, M. Veltman, Diagrammar, CERN Report No. 73-9, 1973;
H. Georgi, Nucl. Phys. B 361 (1991) 339;
C. Arzt, Phys. Lett. B 342 (1995) 189.
[17] F.A. Berends, W.T. Giele, H. Kuijf, Phys. Lett. B 211 (1988) 91.
[18] N.E.J. Bjerrum-Bohr, D.C. Dunbar, H. Ita, W.B. Perkins, K. Risager, JHEP 0601 (2006) 009, hep-th/0509016.
[19] L. Brink, O. Lindgren, B.E.W. Nilsson, Nucl. Phys. B 212 (1983) 401;
S. Ananth, L. Brink, P. Ramond, JHEP 0407 (2004) 082, hep-th/0405150;
S. Ananth, L. Brink, P. Ramond, JHEP 0505 (2005) 003, hep-th/0501079.
[20] S. Mandelstam, Nucl. Phys. B 213 (1983) 149.


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    ${ }^{1}$ For higher-point generalizations see [2].

[^1]:    2 The helicity label assumes that the particle is outgoing.

[^2]:    $\overline{3}$ As seen in Appendix C of [14], a field redefinition which removes occurrences of $\partial_{+}$from the interaction terms has been performed.
    4 Note that the d'Alembertian is $\square=2\left(\partial \bar{\partial}-\partial_{+} \partial_{-}\right)$. See Appendix A for further details.
    5 We point out that higher order terms in (3.6) do not possess this structure.

[^3]:    ${ }^{6}$ This field redefinition changes the structure $h=h(C)$ to $h=h(C, \bar{C})$ but affects only four- and higher-point vertices. This demonstrates the non-uniqueness of the field redefinition in Section 3.2.
    7 In [18], the five-point non-MHV graph is simply a sum of MHV-exchange diagrams. In our case there is also a direct contribution: this is not surprising since, in our Lagrangian, we have eliminated the three-vertex $M(+,+,-)$ and so do not have a contribution equivalent to $D_{2}$ in Eq. (3.14) of that reference.

[^4]:    8 When working with a Lorentzian signature, choosing $\tilde{\lambda}_{\dot{a}}= \pm\left(\lambda_{a}\right)^{*}$ ensures that $p_{a \dot{a}}$ is real.

