

# Proposed cavity Josephson plasmonics with complex-oxide heterostructures

*Y. Laplace<sup>1</sup>, S. Fernandez-Pena<sup>2</sup>, S. Gariglio<sup>2</sup>, J.M. Triscone<sup>2</sup>, A. Cavalleri<sup>1,3</sup>*

<sup>1</sup> Max Planck Institute for the Structure and Dynamics of Matter, Hamburg, Germany

<sup>2</sup> Département de Physique de la Matière Condensée, University of Geneva, 24 Quai Ernest-Ansermet, 1211 Genève 4, Switzerland

<sup>3</sup>Department of Physics, Oxford University, Clarendon Laboratory, Oxford, United Kingdom

## Abstract

We discuss how complex-oxide heterostructures that include high- $T_c$  superconducting cuprates can be used to realize an array of sub-millimeter cavities that support Josephson plasmon polaritons. These cavities have several attractive features for new types of light matter interaction studies and we show that they promote “ultrastrong” coupling between THz frequency radiation and Josephson plasmons. Cavity electrodynamics of Josephson plasmons allows to manipulate the superconducting order-parameter phase coherence. As an example, we discuss how it can be used to cool superconducting phase fluctuations with light.

Light in optical cavities has been extensively used to dress quantum states of matter, opening up fields as diverse as optomechanics, exciton-polariton condensation and Cavity Quantum Electrodynamics at the single atom-photon level. However, this extraordinary degree of control has not been applicable to many complex condensed matter systems of current interest, such as High  $T_c$  superconducting cuprates. These materials are in fact often grown in bulk crystals difficult to process and to incorporate in photonic devices.

Recently, complex oxide heterostructuring has shown to be a formidable platform for engineering new electronic states of matter, as exemplified by the emergent phenomena appearing at oxide interfaces<sup>i</sup>. These advances in materials synthesis, based either on Pulsed Laser Deposition or on Molecular Beam Epitaxy, have improved the quality of materials to a level comparable to semiconductor heterostructures. In this paper we address how these new capabilities can be applied to a new class of cavity electrodynamic experiments that control unconventional quantum orders. As an example, we propose the design of a cavity aimed at optically dressing, manipulating and even cooling the order parameter phase of a high- $T_c$  superconductor. This could be important because cuprate superconductors are strongly affected by the loss of phase coherence near the transition temperature  $T_c$ , with a finite Cooper pair density thought to persist up to temperature scales in excess of  $T_c$ <sup>ii iii iv</sup>. Hence, an important challenge involves the suppression of such phase fluctuations.

High- $T_c$  cuprate superconductors are layered materials, with superconducting planes coupled by Cooper pair tunneling across non-superconducting regions. The excitation of the relative phase between the layers ( $\varphi_n = \chi_{n+1} - \chi_n$ ,  $\varphi_{n-1} = \chi_n - \chi_{n-1}$ , etc.. see figure 1a) corresponds to collective plasma oscillations of the superconducting

condensate. These modes are known as Josephson plasma oscillations, or Josephson plasmons<sup>v vi</sup>. Josephson plasmons have a characteristic resonance frequency  $\nu_p$ , the Josephson Plasma resonance (JPR), in the GHz to THz frequency range depending on the cuprate family.

In the present case, the control of these low frequency modes is not to be realized with light fields at optical frequencies<sup>vii,viii</sup>, which would destroy the condensate by exciting quasiparticles across the superconducting gap  $2\Delta$ . Rather, radiation at THz frequencies will be used to affect the strength and spatial distribution of the condensate phase without changing the density of Cooper pairs<sup>ix x xi xii xiii</sup>. As the cavity mode must couple to the interlayer Josephson plasma oscillations, the radiation should propagate with the electric field polarized perpendicularly to the superconducting planes. Also, in order to apply cavity cooling techniques such as sideband cooling<sup>viii</sup>, the free spectral range of the cavity must be larger than the typical fluctuation frequency, that is in the THz-frequency range, hence the cavity should have sub-millimeter dimensions. The cavity design based on oxide heterostructures and shown schematically in figure 1 would fulfill all these requirements.

Let us consider a complex oxide heterostructure involving a high- $T_c$  superconducting thin film (e.g.  $\text{La}_{2-x}\text{Sr}_x\text{CuO}_4$ ), cladded between lattice-matched insulating films ( $\text{La}_2\text{CuO}_4$ ), with a conducting oxide at the bottom (e.g.  $\text{SrRuO}_3$ ) and capped with a metal patch (e.g. gold) (figure 1b). This design can be realized with standard c-axis grown thin film heterostructuring, in combination with etching techniques<sup>i</sup>.

Due to the strong sub-wavelength confinement of the electromagnetic field between the top and bottom metals ( $D \ll \lambda$ ), resonant modes that propagate along the film can be excited. These are highly localized within the volume defined by this mesa, and have

their electric field polarized along the  $z(c)$ -direction<sup>xiv</sup> (see figure 1c). This heterostructure corresponds then to a periodically structured array of cavities with sub-millimeter spacing, which greatly enhances the coupling efficiency of the incoming radiation as compared to a single cavity<sup>xv</sup>.

Let us first consider a “bare” cavity without the superconducting film and based on Au/La<sub>2</sub>CuO<sub>4</sub>/SrRuO<sub>3</sub>. The reflection spectrum of a periodic array of such cavities was computed here using the Fourier modal method<sup>xvi</sup>. We consider radiation impinging at normal incidence ( $k_x=k_y=0$ ,  $k_z=2\pi\nu/c$ ) with the electric field polarized along the  $y$ -direction (TM polarization). We modelled the optical properties of the different materials from experimental results found in the literature, including anisotropy and dissipation (La<sub>2-x</sub>Sr<sub>x</sub>CuO<sub>4</sub><sup>xvii xviii xix xx xxi</sup>, SrRuO<sub>3</sub><sup>xxii</sup> and Au<sup>xxiii</sup>). The computations are performed at fixed thickness  $D=1.2\mu\text{m}$  between the Au patch and SrRuO<sub>3</sub> and fixed periodicity  $p=33\mu\text{m}$ . The thicknesses of the Au patch and SrRuO<sub>3</sub> are 200nm and 500 nm respectively. The results for these bare cavities are shown in figure 2a.

For a given patch width  $w$ , the reflectivity as a function of frequency displays a lorentzian dip centered at the cavity resonance frequency  $\nu_c(w)$  (fig 2a left panel). Because of the strong impedance mismatch at the patch edges, each patch acts as a cavity, with resonance frequency  $\nu_{n,m} = c/(2\sqrt{\epsilon_\infty}w) \sqrt{n^2 + m^2}$  ( $n$  and  $m$  are integers denoting the number of nodes of the electric field of the cavity mode in the  $x$  and  $y$  direction respectively,  $\epsilon_\infty \approx 27$  is the dielectric constant of the insulating La<sub>2</sub>CuO<sub>4</sub><sup>xvii</sup> and  $w$  is the width of the top metallic square patch)<sup>xiv</sup>. Hence, this resonance corresponds to the  $(n, m) = (0,1)$  fundamental mode of the cavity, whose electric field distribution is shown on figure 1c. In fig 2a, we show the reflectivity of the heterostructure as a function of frequency and  $w$  (note that on the abscissa, the patch width  $w$  has been converted to the “inverse patch width” frequency defined by the

relation  $\nu(w) = c/(2\sqrt{\epsilon_\infty}w)$ . We see that the cavity resonance frequency scales linearly with the inverse patch width, as expected from the formula above for  $\nu_{n,m}$ .

In figure 2b, we show a similar calculation performed on a cavity that includes a film of the optimally doped high- $T_c$  superconductor  $\text{La}_{2-x}\text{Sr}_x\text{CuO}_4$  ( $x=17\%$ ,  $T_c=36\text{K}$ ,  $\nu_p = 2.3\text{THz}$ ). The anisotropic dielectric permittivity of  $\text{La}_{2-x}\text{Sr}_x\text{CuO}_4$  is expressed as a diagonal dielectric tensor. Although the model realistically includes anisotropies and dissipation, the results of the computation are dominated by the electrostatics of the  $c$ -axis superconducting component, whose dielectric permittivity in this frequency range can be described by  $\epsilon_c^{SC} = \epsilon_\infty \left(1 - \frac{\nu_p^2}{\nu^2}\right)$ , with  $\epsilon_\infty \approx 27$  and  $\nu_p = 2.3\text{THz}$ . As shown in fig2b, by tuning the fundamental mode of the cavity  $\nu_c(w)$  across the JPR frequency  $\nu_p=2.3\text{THz}$  we observe a clear avoided crossing of the two bare dispersions  $\nu = \nu_p$  and  $\nu = \nu_c(w)$ , which indicates strong coupling between the cavity field and the JPR.

At the frequencies for which the JPR and the cavity are resonant ( $\nu_c = \nu_p$ ), the reflectivity displays two absorptions of similar intensities whose frequencies are separated by an amount  $2\Omega$ . This frequency splitting can be thought of as a Rabi splitting <sup>xxiv</sup> and quantifies the rate at which energy is exchanged between the electromagnetic field and the Josephson plasmon.

The behavior of such heterostructures can be quite generally captured by a simple effective medium model <sup>xxv</sup>. In our case, the effective dielectric permittivity  $\epsilon_{eff}(\omega)$  of the insulator/high- $T_c$ /insulator heterostructure can be written as  $\frac{1}{\epsilon_{eff}(\omega)} = \frac{1-f}{\epsilon_\infty} + \frac{f}{\epsilon_c^{SC}(\omega)}$  ( $f$  is the filling fraction, defined as the ratio between the thickness  $d$  of the superconducting film and the total thickness  $D$  of the cavity :  $f = \frac{d}{D}$  ( $0 \leq f \leq 1$ )). For a cavity tuned at  $\nu_c(w)$ , the frequency of the hybridized resonances are given by  $\epsilon_{eff}(\nu)\nu^2 = \epsilon_\infty\nu_c(w)^2$ , leading to the two branches  $\{\nu_+(w), \nu_-(w)\}$  plotted on fig 2b for

the corresponding filling fraction  $f \approx 17\%$  (black solid lines). We see that the effective medium model matches well the results of the full numerical computation.

We next analyze the strength of the coupling, evidenced by the size of the Rabi splitting. In figure 3a, we compute the reflectivity as a function of filling fraction  $f$  for a cavity tuned at  $\nu_c = \nu_p$ . Already for relatively small filling fractions (as small as  $\sim 3\%$ ), the Rabi splitting  $2\Omega$  of the two hybridized modes is observable, and exceeds the linewidth broadening originating from the cavity dissipation and JPR damping. In fig3b, we show the frequency splitting  $2\Omega$  normalized to the bare cavity frequency  $\nu_c = \nu_p$  yielding the normalized Rabi splitting. We observe that, even at small filling fractions, coupling in this heterostructure can be characterized as “ultrastrong” according to the usual prescription of a normalized Rabi splitting greater than  $\sim 20\%$ .

Note that the optical properties of a dissipationless and isotropic optical excitation can be described as:  $\epsilon = \epsilon_\infty(1 - \nu_p^2/(\nu^2 - \nu_0^2))$ , where  $\nu_0$  is the frequency of the transition and  $\nu_p$  is its oscillator strength. When a material carrying this optical excitation is placed inside the cavity with a filling fraction  $f$ , a Rabi splitting of  $2\Omega$  is observed when the

cavity is tuned at  $\nu_c = \sqrt{\nu_0^2 + \nu_p^2}$ , yielding a normalized Rabi splitting given by

$$\frac{2\Omega}{\nu_c} = \sqrt{1 + \sqrt{f} \frac{\nu_p}{\sqrt{\nu_p^2 + \nu_0^2}}} - \sqrt{1 - \sqrt{f} \frac{\nu_p}{\sqrt{\nu_p^2 + \nu_0^2}}}$$

obtained from the effective medium model described previously.

Light-matter interaction studies have to date been realized almost exclusively for cavities tuned to optical frequencies and coupling to low oscillator strength excitations, that is for  $\nu_0 \gg \nu_p$ , such as the case of interband transitions in quantum wells <sup>xxiv</sup>. Hence,

typical normalized Rabi splittings of  $\sim\sqrt{f}\frac{\nu_p}{\nu_0}\ll 1$  fall in the weak to moderate coupling regime. Recently, high-density 2-dimensional electron gases with electronic transitions of  $\nu_0$  and  $\nu_p$  both in the THz frequency range have reached the ultra-strong coupling regime <sup>xxv xxvi</sup>.

Here, for the superconducting material,  $\nu_0 = 0$ , as the condensate response peaks at zero frequency and is characterized by the zero-frequency delta function in the conductivity. Hence, the normalized Rabi splitting reduces to  $\sqrt{1+\sqrt{f}}-\sqrt{1-\sqrt{f}}$ , and unlike the more conventional cases is independent of either  $\nu_0$  or  $\nu_p$ . For realistic filling fractions, the ultra-strong coupling regime is easily reached.

So far, we analyzed the linear electro-dynamical properties of these oxide heterostructures and described how the cavities can be used to optically dress the Josephson plasmons. In the following, we discuss how such cavities can as well cool thermally excited Josephson plasmons in high-T<sub>c</sub> superconductors with cavity sideband cooling techniques, as done in the case of a single Josephson junction <sup>xxvii</sup>.

The electro-dynamics of this cavity/high-T<sub>c</sub> heterostructure is intrinsically nonlinear due to the Josephson relation between the current density and the phase difference between the planes:  $j_n = j_c \sin(\varphi_n) = j_c \sin\left(\frac{2ed^*}{\hbar} \int_{-\infty}^t E_n(t') dt'\right)$  with  $j_c$  denoting the critical current density,  $d^*$  the distance between two adjacent planes and  $E_n$  the electric field between the n<sup>th</sup> and the (n+1)<sup>th</sup> superconducting planes. Due of this nonlinearity, we find by following [xxvii,xxviii] that in the limit of small optical field intensities the coupled dynamics of the superconducting phase difference  $\varphi_n$  ( $\varphi_n = \chi_{n+1} - \chi_n$ ) and of the optical field is described by the following Hamiltonian:

$$H_{\varphi_n} = E_J \left( \frac{1}{8\pi^2 v_p^2} \left( \frac{\partial \varphi_n}{\partial t} \right)^2 + 1 - \cos(\varphi_n) \right)$$

$$H_A = E_J \frac{D}{d^*} \left( \frac{1}{16\pi^2 v_p^2} \left( \frac{\partial A}{\partial t} \right)^2 + \frac{v_c^2}{4v_p^2} A^2 \right)$$

$$H_n^{int} = E_J \cos(\varphi_n) \frac{A^2}{4}$$

Here,  $A(t)$  is the amplitude of the vector potential of the optical field inside the cavity expressed as  $\vec{A} = A(t) \cos(\frac{\pi}{w} y) \vec{z}$  ( $0 \leq y \leq w$ ) for the fundamental mode shown in figure 1c and  $E_J = \frac{\hbar j_c w^2}{2e}$  is the Josephson energy. This Hamiltonian describes the dynamics of two oscillators, namely the optical cavity field and the superconducting phase ( $H_A, H_{\varphi_n}$ ), coupled via the nonlinear interaction term  $H_n^{int}$ .

Provided that a static current bias  $j$  is applied along the c-axis, the shift of the equilibrium phase difference  $\varphi_n \rightarrow \varphi_n + \varphi_0$  allows one to tune the coupling term between the phase and the cavity as  $H_{int}/E_J = -\sin(\varphi_0) \varphi_n A^2/4 - \cos(\varphi_0) \varphi_n^2 A^2/8$ . The first of these coupling terms is formally equivalent to the coupling between the radiation pressure ( $A^2$ ) and the position of an end-mirror ( $\varphi_n$ ) in a Fabry-Perot cavity<sup>Error! Bookmark not defined.</sup>, typically used to cool macroscopic objects in optomechanics<sup>viii</sup>. Following the analogy with optomechanics, the second term would correspond to the coupling of a membrane-in-the-middle cavity, which has recently attracted much attention<sup>xxix xxx</sup>. In this situation, cooling and even squeezing have been predicted<sup>xxxi</sup>.

The above Hamiltonian describes the dynamics of a single phase difference  $\varphi_n$  between a given set of neighboring planes that is coupled to the electromagnetic field  $\vec{A}$  of the cavity. In the situation considered here, we do have multiple superconducting planes and therefore a set of phases  $\{\varphi_n\}_n$ . However, the coupling among them originates from capacitive charge screening and is known to be weak compared to the coupling to the

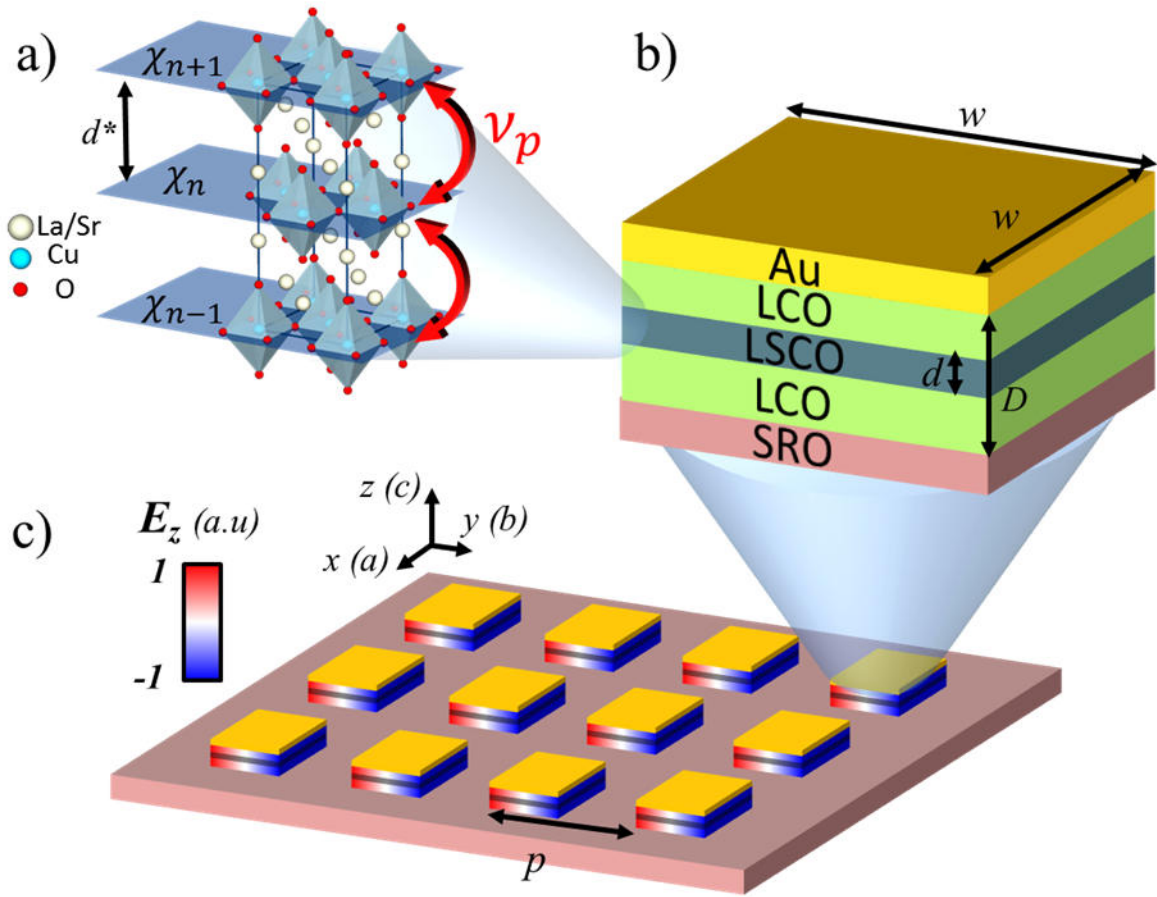


cavity mode rooted in the inductive screening of its magnetic field, therefore we can consider them to be independent<sup>xxxii</sup>. Hence, the physical situation described here is that of a set of independent phases, each of them being coupled to a single optical field  $\vec{A}$ . The corresponding Hamiltonian in our case is given by :  $H = H_A + \sum_n (H_{\phi_n} + H_n^{int})$ . This is reminiscent of the physics of *superradiance* and we may expect enhanced cooling efficiency due to collective phase synchronization<sup>xxxiii xxxiv xxxv</sup>. As for the choice of the superconductor, phase cooling could be achieved at any doping. Nevertheless, in the underdoped regime of the high-Tc superconductors, one could expect unconventional electronic properties resulting in macroscopic changes like the increase of  $T_c$ .

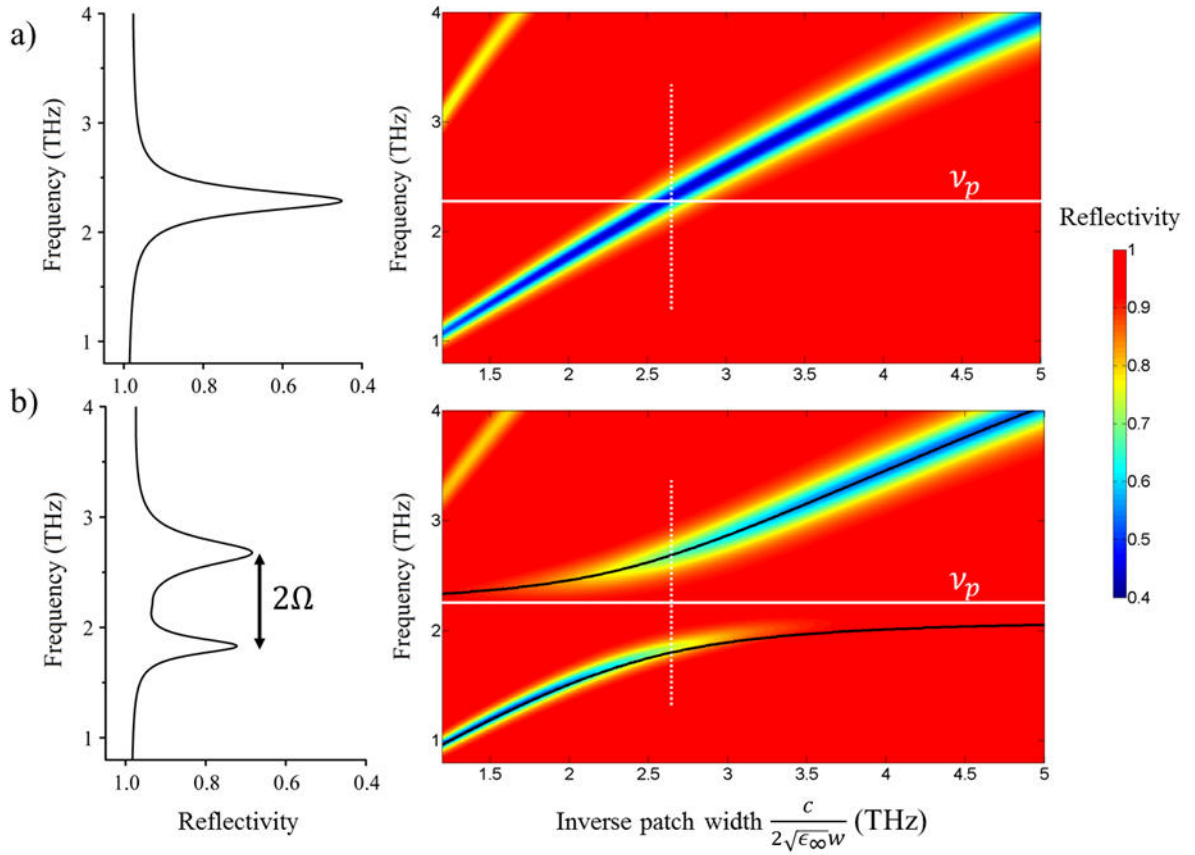
In summary, we have discussed how oxide heterostructuring can be used to generate arrays of optical cavities to dress and manipulate Josephson plasmons in high-Tc superconductors. The conditions discussed here may allow for experiments in which the interlayer phase fluctuations of high-Tc superconductors are cooled with light. More generally, within this framework, other sorts of unconventional electronic order could be dressed in such complex-oxide based optical cavities, leading to a new platform for light/complex-matter interaction experiments.

## Acknowledgments

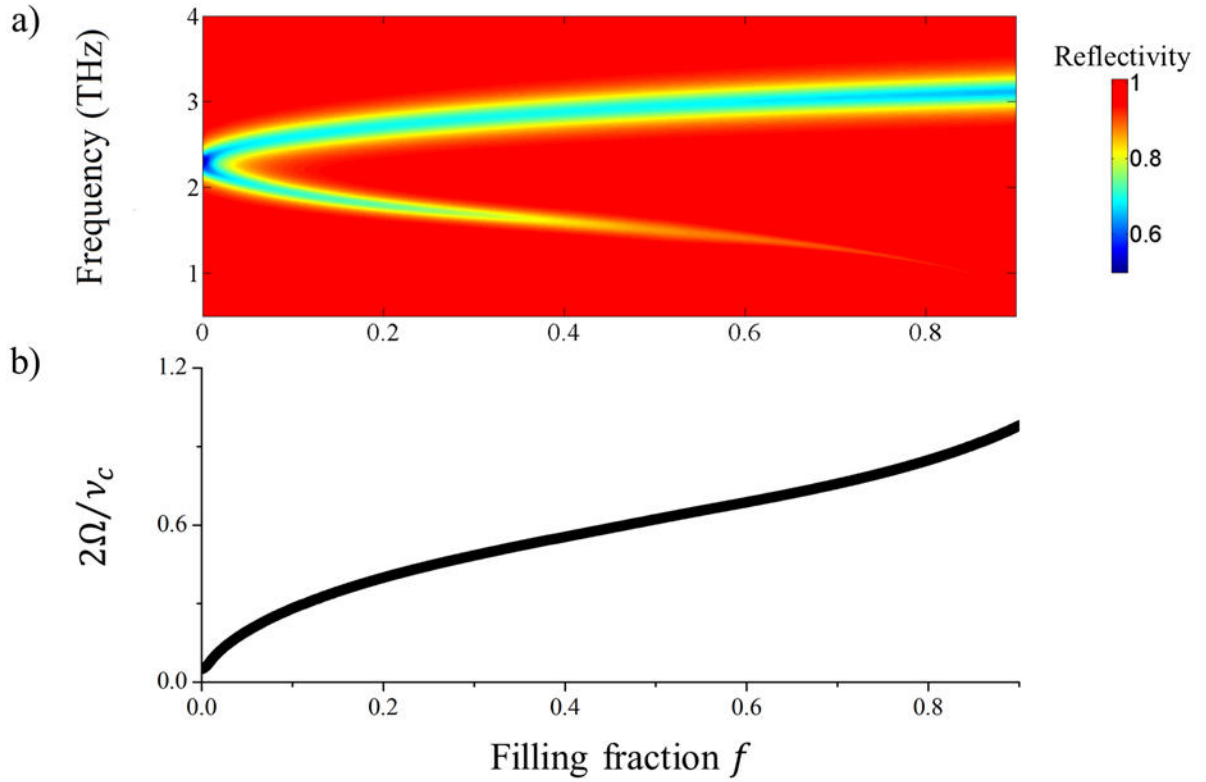
We are thankful to D. Jaksch, S. Clark, S. Denny, M. Eckstein and L. Zhang for discussions and suggestions. This work was supported by a grant (Grant No. ERC-319286 QMAC) from the European Research Council.



**Figure 1:** **a)** the layered high- $T_c$  superconductor  $\text{La}_{2-x}\text{Sr}_x\text{CuO}_4$ :  $\chi_n$  is the phase of the superconducting order parameter of the  $n^{\text{th}}$  superconducting plane. Red arrows indicate the Josephson plasma oscillations between the planes at the characteristic frequency  $\nu_p$  **b)** a single cavity/high- $T_c$  mesa : Au (yellow)/ $\text{La}_2\text{CuO}_4$  (“LCO”, green)/ $\text{La}_{2-x}\text{Sr}_x\text{CuO}_4$  (“LSCO”, dark blue)/ $\text{La}_2\text{CuO}_4$  (“LCO”, green)/ $\text{SrRuO}_3$  (“SRO”, red) **c)** the proposed heterostructure : array of cavity/high- $T_c$  mesas. The electric field distribution of the  $(n,m)=(0,1)$  fundamental mode of the “bare” cavities is shown (that is without the superconductor, see text for details). The corresponding notations are used throughout the text.



**Figure 2 :** **a) Right:** Reflectivity as a function of frequency and inverse patch width for the bare cavity ( $D=1.2\mu\text{m}$ ,  $p=33\mu\text{m}$ ). The white solid line shows the JPR frequency  $\nu_p$ . **Left:** Reflectivity cut of the bare cavity tuned at  $\nu_p$  ( $w=11\mu\text{m}$  : corresponding to the white dashed line on the right panel) **b) Right:** Reflectivity as a function of frequency and inverse patch width for the cavity/high- $T_c$  heterostructure ( $d=200\text{nm}$ ,  $D=1.2\mu\text{m}$ ,  $p=33\mu\text{m}$ , corresponding to a filling fraction  $f = \frac{d}{D} \approx 17\%$ ). The white solid line shows the JPR frequency  $\nu_p$ . The black solid lines are the hybridized resonances  $\{\nu_+(w), \nu_-(w)\}$  (see text for details) **Left:** Reflectivity cut of the cavity/ high- $T_c$  heterostructure when the bare cavity is tuned at  $\nu_p$  ( $w=11\mu\text{m}$  : corresponding to the white dashed line on the right panel)



**Figure 3:** **a)** Reflectivity as a function of frequency and filling fraction  $f$  for a cavity tuned at  $\nu_c = \nu_p$  (computed at fixed  $D=1.2\mu\text{m}$ ,  $p=33\mu\text{m}$ ,  $w=11\mu\text{m}$ ) **b)** Normalized Rabi splitting  $2\Omega/\nu_c$  as a function of the filling fraction  $f$  computed from Figure 3a)

- 
- <sup>i</sup> J. Mannhart, and D. G. Schlom, *Science* 327, 1607 (2010)
- P. Zubko, S. Gariglio, M. Gabay, P. Ghosez, and J.-M. Triscone, *Annu. Rev. Condens. Matter Phys.* 2, 141 (2011)
- H. Y. Hwang, Y. Iwasa, M. Kawasaki, B. Keimer, N. Nagaosa, and Y. Tokura, *Nature Materials* 11, 103 (2012)
- <sup>ii</sup> V. J. Emery, S. A. Kivelson, *Nature* 374, 434 - 437 (1994)
- <sup>iii</sup> J. Corson, R. Mallozzi, J. Orenstein, J. N. Eckstein & I. Bozovic, *Nature* 398, 221-223 (1999)
- <sup>iv</sup> L. S. Bilbro, R. Valdés Aguilar, G. Logvenov, O. Pelleg, I. Bozović, N. P. Armitage, *Nature Physics* 7, 298–302 (2011)
- <sup>v</sup> Lawrence, W. E. & Doniach, S. in *Proc. 12th Int. Conf. Low Temp. Phys.* (ed. Kanda, E.) 361–362 (Tokyo Keigaku Publishing, 1971);
- Kleiner, R., Steinmeyer, F., Kunkel, G. & Muller, P., *Phys. Rev. Lett.* 68, 2394–2397 (1992);
- Matsuda, Y., Gaifullin, M. B., Kumagai, K., Kadowaki, K. & Mochiku, T., *Phys. Rev. Lett.* 75, 4512–4515 (1995);
- Tsui, O. K. C., Ong, N. P. & Peterson, J. B., *Phys. Rev. Lett.* 76, 819–822 (1996).
- <sup>vi</sup> Tamasaku, K., Nakamura, Y. & Uchida, S., *Phys. Rev. Lett.* 69, 1455–1458 (1992)
- <sup>vii</sup> V. E. Shapiro, *JETP*, Vol. 43, No. 4, p. 763, April 1976
- <sup>viii</sup> F. Diedrich, J. C. Bergquist, Wayne M. Itano, and D. J. Wineland, *Phys. Rev. Lett.* 62, 403-406 (1989);
- J. Kippenberg, K. J. Vahala, *Science* 321, 1172 (2008);
- M. Aspelmeyer & al., *J. Opt.Soc. Am.* B27, A189 (2010);
- F.Marquardt, S. M. Girvin, *Physics* 2, 40 (2009)
- <sup>ix</sup> A. Dienst, M. C. Hoffmann, D. Fausti, J. C. Petersen, S. Pyon, T. Takayama, H. Takagi & A. Cavalleri, *Nature Photonics* 5, 485–488 (2011)
- <sup>x</sup> A. Dienst, E. Casandruc, D. Fausti, L. Zhang, M. Eckstein, M. Hoffmann, V. Khanna, N. Dean, M. Gensch, S. Winnerl, W. Seidel, S. Pyon, T. Takayama, H. Takagi and A. Cavalleri, *Nature Materials* 12, 535–541 (2013)
- <sup>xi</sup> D. Fausti R. I. Tobey, N. Dean, S. Kaiser, A. Dienst, M. C. Hoffmann, S. Pyon, T. Takayama, H. Takagi, A. Cavalleri, *Science* 331, 189 (2011)
- <sup>xii</sup> W. Hu, S. Kaiser, D. Nicoletti, C. R. Hunt, I. Gierz, M. C. Hoffmann, M. Le Tacon, T. Loew, B. Keimer & A. Cavalleri, *Nature Materials* 13, 705–711 (2014)

- 
- xiii S. Kaiser, C. R. Hunt, D. Nicoletti, W. Hu, I. Gierz, H. Y. Liu, M. Le Tacon, T. Loew, D. Haug, B. Keimer, and A. Cavalleri, *Phys. Rev. B* 89, 184516 (2014)
- xiv Y. Todorov, L. Tosetto, J. Teissier, A. M. Andrews, P. Klang, R. Colombelli, I. Sagnes, G. Strasser, and C. Sirtori *Optics Express* 18, 13886 (2010)
- xv Cheryl Feuillet-Palma, Yanko Todorov, Angela Vasanelli & Carlo Sirtori, *Scientific Reports* 3, 1361, (2013)
- xvi V. Liu & S. Fan, *Computer Physics Communications* **183**, 2233-2244 (2012)
- xvii S. V. Dordevic, Seiki Komiya, Yoichi Ando, and D. N. Basov, *Phys. Rev. Lett.* 91, 167401 (2003)
- xviii S. V. Dordevic, Seiki Komiya, Yoichi Ando, Y. J. Wang, and D. N. Basov, *Phys. Rev. B* 71, 054503 (2005)
- xix S. V. Dordevic, E. J. Singley, D. N. Basov, Seiki Komiya, Yoichi Ando, E. Bucher, C. C. Homes, and M. Strongin, *Phys. Rev. B* 65, 134511 (2002)
- xx S. Uchida, T. Ido, H. Takagi, T. Arima, Y. Tokura, and S. Tajima, *Phys. Rev. B* 43, 7942 – Published 1 April 1991
- xxi W. J. Padilla, M. Dumm, Seiki Komiya, Yoichi Ando, and D. N. Basov, *Phys. Rev. B* 72, 205101 (2005)
- xxii J. S. Dodge, C. P. Weber, J. Corson, J. Orenstein, Z. Schlesinger, J. W. Reiner, and M. R. Beasley *Phys. Rev. Lett.* 85, 4932 (2000)
- xxiii M. A. Ordal, L. L. Long, R. J. Bell, S. E. Bell, R. R. Bell, R. W. Alexander, Jr., and C. A. Ward, *Applied Optics*, Vol. 22, Issue 7, pp. 1099-1119 (1983)
- xxiv G. Khitrova, H. M. Gibbs, F. Jahnke, M. Kira, and S. W. Koch, *Rev. Mod. Phys.* 71, 1591 (1999)
- xxv Y. Todorov, A. M. Andrews, R. Colombelli, S. De Liberato, C. Ciuti, P. Klang, G. Strasser, and C. Sirtori, *Phys. Rev. Lett.* 105, 196402 (2010)
- xxvi G. Scalari, C. Maissen, D. Turčinková, D. Hagenmüller, S. De Liberato, C. Ciuti, C. Reichl, D. Schuh, W. Wegscheider, M. Beck, J. Faist, *Science*, 335, 1323 (2012)
- xxvii J. Hammer, M. Aprili, and I. Petkovic, *Phys. Rev. Lett.* 107, 017001 (2011)
- xxviii M. V. Fistul and A. V. Ustinov, *Phys. Rev. B* 75, 214506 (2007)
- xxix A M Jayich, J C Sankey, B M Zwickl, C Yang, J D Thompson, S M Girvin, A A Clerk, F Marquardt and J G E Harris, *New Journal of Physics* 10, 095008 (2008)

---

xxx J. D. Thompson, B. M. Zwickl, A. M. Jayich, Florian Marquardt, S. M. Girvin & J. G. E. Harris, *Nature* 452, 72-75 (2008)

xxxi A. Nunnenkamp, K. Børkje, J. G. E. Harris, and S. M. Girvin, *Phys. Rev. A* 82, 021806(R) (2010)

xxxii D. van der Marel, A. A. Tsvetkov, *Phys. Rev. B* 64, 024530 (2001);

Xiao Hu and Shi-Zeng Lin, *Supercond. Sci. Technol.* 23 053001 (2010)

Sergey Savel'ev, V A Yampol'skii, A L Rakhmanov and Franco Nori, *Rep. Prog. Phys.* 73 026501 (2010)

xxxiii Kohjiro Kobayashi, David Stroud, *Physica C: Superconductivity* 469, 5–6, 216-224 (2009)

xxxiv L. Ozyuzer, A. E. Koshelev, C. Kurter, N. Gopalsami, Q. Li, M. Tachiki, K. Kadowaki, T. Yamamoto, H. Minami, H. Yamaguchi, T. Tachiki, K. E. Gray, W.-K. Kwok, U. Welp, *Science* 318, 1291 (2007)

xxxv Christopher J. Wood, Troy W. Borneman, and David G. Cory *Phys. Rev. Lett.* 112, 050501 (2014)