Symmetries of post-Newtonian expansions

Joaquim Gomis

Departament de Física Quàntica i Astrofísica and Institut de Ciències del Cosmos (ICCUB), Universitat de Barcelona, Martí Franquès, ES-08028 Barcelona, Spain

Axel Kleinschmidt*

Max-Planck-Institut für Gravitationsphysik (Albert-Einstein-Institut), Am Mühlenberg 1, DE-14476 Potsdam, Germany

Jakob Palmkvist

Department of Mathematical Sciences, Chalmers University of Technology and the University of Gothenburg, SE-412 96 Göteborg, Sweden

Patricio Salgado-Rebolledo

Instituto de Física, Pontificia Universidad Católica de Valparaíso Casilla 4059, Valparaiso, Chile (Dated: October 31, 2019)

In this note we introduce an infinite-dimensional space on which an infinite-dimensional generalization of the Galilei group acts. Standard Minkowski space can be modelled in this space and its symmetries yield an embedding of the Poincaré group in the infinite extension. The extension has an interpretation in terms of post-Newtonian corrections to Galilei symmetries. We also construct particle and string actions that are invariant under these transformations.

Non-relativistic physics can be obtained from relativistic physics in an expansion in 1/c where c is the velocity of light [1]. At the level of symmetries, it is well-known that the strict $c \to \infty$ limit of the relativistic Poincaré group yields the Galilei group of non-relativistic symmetries via Wigner-Inönü contraction [2]. The non-relativistic limit of general relativity is Newtonian gravity although this limit is hard to perform at the level of the action. Corrections to the non-relativistic Newtonian theory at higher order in 1/c are famously important for the original experimental evidence for general relativity, see for example [3, 4] for discussions. Higher-order (parametrized) post-Newtonian corrections to the Keplerian two-body motion are of central importance in current investigations of binary gravitational wave sources [5–9] and post-Newtonian corrections have also been investigated in the context of cosmology and structure formation, see for instance [10].

One common feature of truncated post-Newtonian expansions is that they inherently do not preserve the full relativistic symmetries. As an early example, the quantum-mechanical Breit equation describes corrections of order v/c to the two-body problem for electrons but it is not invariant under Lorentz transformations [11]. More recently, due to the possible applications of Newtonian gravities to non-relativistic holography for strongly coupled systems in condensed matter physics, non-relativistic gravities that are invariant under various classes of non-relativistic symmetry Lie algebras have been investigated [12, 13]. Subsequently, extensions of the Bargmann algebra by a finite number of generators were first presented in [14–16] and these describe the next-to-leading orders in the non-relativistic expansion

of gravity. The construction can also be generalized to infinite-dimensional algebras [17–19] using different approaches.

In this note, we shall exhibit a universal scheme for obtaining post-Newtonian expansions of non-relativistic systems by means of their symmetry algebras, starting from the Galilei algebra \mathfrak{G} . Our scheme can be viewed as either stemming from (affine) Kac-Moody algebras [19] or from a Lie algebra expansion [20–22], and it is based on an infinite-dimensional algebra, which we will refer to as \mathfrak{G}_{∞} . This algebra can be viewed as including the full formal power series in 1/c of relativistic systems. The analysis in [18] is in the context of the large c expansion of general relativity [23–26], while our approach is purely kinematical and provides an extension of special relativity that describes post-Newtonian physics. We introduce an infinite-dimensional space on which this symmetry acts and identify subspaces of finite co-dimension that are mapped to one another by the usual action of the Poincaré algebra as a specific combination of transformations in \mathfrak{G}_{∞} . The infinite-dimensional algebra \mathfrak{G}_{∞} admits finite-dimensional quotients corresponding to working up to a finite order in 1/c.

The present paper is mainly concerned with outlining the kinematic underpinnings of this symmetry algebra. Truncations of it have been used in a slightly different guise to construct non-relativistic gravity theories [14–16] that should properly be identified as post-Newtonian gravities. In a follow-up paper [27], we shall study field-theory implementations of the symmetry structure in the framework of three-dimensional Chern–Simons theory and also the post-Newtonian corrections of the non-vibrating non-relativistic string theory [28].

Generalized Minkowski space-time and its infinite-dimensional symmetry

Our starting point is the (d+1)-dimensional Poincaré algebra $\mathfrak{iso}(d,1)$ written in a non-covariant form with generators $\{p_0,j_{ab},j_{a0},p_a\}$ and non-trivial commutation relations

$$[j_{ab}, j_{cd}] = 4\delta_{[c[b}j_{a]d]}, \quad [j_{ab}, p_c] = 2\delta_{c[b}p_{a]},$$

$$[j_{a0}, j_{b0}] = j_{ab}, \quad [p_0, j_{a0}] = p_a,$$

$$[j_{a0}, p_b] = -\delta_{ab}p_0, \quad [j_{ab}, j_{c0}] = 2\delta_{c[b}j_{a]0}.$$
(1)

Roman indices $a=1,\ldots,d$ are vector indices of the spatial rotation algebra $\mathfrak{so}(d)$ that is generated by j_{ab} . The other generators will be referred to as time translation (p_0) , boost (j_{a0}) and translation (p_a) .

We can construct an associated infinite-dimensional algebra by applying the method of infinite Lie algebra expansions [27, 29] to the previous algebra. The generators are $H^{(m)} = c^{-2m+1} \otimes p_0$, $P_a^{(m)} = c^{-2m} \otimes p_a$, $P_a^{(m)} = c^{-2m-1} \otimes j_{a0}$ and $P_a^{(m)} = c^{2m} \otimes j_{ab}$, which satisfy the algebra \mathfrak{G}_{∞} :

$$\begin{split} [J_{ab}^{(m)},J_{cd}^{(n)}] &= 4\delta_{[c[b}J_{a]d]}^{(m+n)}\,,\\ [J_{ab}^{(m)},B_{c}^{(n)}] &= 2\delta_{c[b}B_{a]}^{(m+n)}\,,\,\, [J_{ab}^{(m)},P_{c}^{(n)}] = 2\delta_{c[b}P_{a]}^{(m+n)}\,,\\ [H^{(m)},B_{a}^{(n)}] &= P_{a}^{(m+n)}\,,\,\, [B_{a}^{(m)},P_{b}^{(n)}] = -\delta_{ab}H^{(m+n+1)}\,,\\ [B_{a}^{(m)},B_{b}^{(n)}] &= J_{ab}^{(m+n+1)}\,. \end{split} \tag{2}$$

This algebra can also be obtained from the positive modes of an affine Kac–Moody [30] associated to the Galilei algebra [19].

Thinking of the collection of all boost generators $B_a^{(m)}$ and rotation generators $J_{ab}^{(m)}$ for $m \geq 0$ as generating a generalized Lorentz algebra \mathfrak{L}_{∞} , we introduce generalized Minkowski space as the formal coset space $\exp \mathfrak{G}_{\infty}/\exp \mathfrak{L}_{\infty}$. We put local coordinates on this infinite-dimensional space by introducing coordinates $t_{(m)}$ and $x_{(m)}^a$ dual to the generalized translation generators $H^{(m)}$ and $P_a^{(m)}$.

We now consider the action of an infinitesimal transformation of the coordinates on generalized Minkowski space under a general transformation in \mathfrak{G}_{∞} of (2) with parameters $\alpha_{(m)}^{ab}$ (for $J_{ab}^{(m)}$), $\epsilon_{(m)}$ (for $H^{(m)}$), $\epsilon_{(m)}^{a}$ (for $P_{a}^{(m)}$) and $v_{(m)}^{a}$ (for $P_{a}^{(m)}$). The infinitely many coordinates transform as

$$\delta x_{(m)}^{a} = \epsilon_{(m)}^{a} + \sum_{n=0}^{m} \left(v_{(n)}^{a} t_{(m-n)} - \delta_{bc} \alpha_{(n)}^{ab} x_{(m-n)}^{c} \right) ,$$

$$\delta t_{(m)} = \epsilon_{(m)} + \sum_{n=0}^{m-1} \delta_{ab} v_{(n)}^{a} x_{(m-1-n)}^{b} . \tag{3}$$

Restricting only to the 'zero modes' these transformations take the form of the usual Galilei transformations on $(t_{(0)}, x_{(0)}^a)$. However, they differ when including higher modes.

Recovering standard Minkowski space

We now introduce the coordinates

$$X^{a} = \sum_{m=0}^{\infty} c^{-2m} x_{(m)}^{a}, \qquad X^{0} = \sum_{m=0}^{\infty} c^{-2m+1} t_{(m)}, \quad (4)$$

that are formal power series in 1/c. Note that the coordinates $t_{(m)}$ have dimension of $[L^{2m}/T^{2m-1}]$, while the coordinates $x^a_{(m)}$ have dimension of $[L^{2m+1}/T^{2m}]$. This ensures that the coordinates X^0 and X^a have dimensions of time and length and it is compatible with the fact that these coordinates are associated to the Lie algebra generators of (2) that emerge from a particular expansion of the translation algebra in powers of 1/c. Expansions of coordinates in 1/c have already appeared in the literature in [18]. In the following we shall sometimes write \vec{X} for the spatial vector with components X^a and use the dot product to denote scalar products with respect to δ_{ab} .

In order to understand the action of the symmetry algebra (2) on these coordinates we focus on the boosts [31] and introduce a similar collective series-expanded parameter

$$\theta^a = \sum_{m=0}^{\infty} c^{-2m-1} v_{(m)}^a \,. \tag{5}$$

The action of this parameter on the coordinates (4) becomes from (3)

$$\delta \vec{X} = \vec{\theta} X^0 \,, \qquad \delta X^0 = \vec{\theta} \cdot \vec{X} \,. \tag{6}$$

These transformations resemble formally the standard Lorentz boosts except for the fact that θ^a contains an infinity of independent components $v^a_{(m)}$. In order to remedy this we let

$$v_{(m)}^{a} = \frac{1}{2m+1} v^{2m+1} n^{a}, \qquad (7)$$

with the same unit vector n^a for all m and a single additional parameter v. By replacing (7) in the definition (6), we obtain

$$\delta \vec{X} = \vec{\theta} X^0 = \sum_{m=0}^{\infty} \frac{1}{2m+1} \left(\frac{v}{c}\right)^{2m+1} \vec{n} X^0,$$

$$\delta X^0 = \vec{\theta} \cdot \vec{X} = \sum_{m=0}^{\infty} \frac{1}{2m+1} \left(\frac{v}{c}\right)^{2m+1} \vec{n} \cdot \vec{X}$$
(8)

and we recover the usual expression of the rapidity parameter in terms of the boost velocity, i.e., $\theta^a = \theta n^a$ with $\theta = \arctan(v/c)$ upon using the series expansion of arctanh [32]. The linear transformation in θ in (8) agrees with the standard infinitesimal Lorentz boost on Minkowski space.

While we have thus arrived at a formal agreement, the definition (4) requires additional discussion. For

fixed (X^0, X^a) this defines a subspace of generalized Minkowski space of co-dimension d+1. The transformations (6) we have just analysed move between different embedded such subspaces. For general parameters $v_{(m)}^a$ they also move the individual $(t_{(m)}, x_{(m)}^a)$ for $m \geq 0$ around in a non-uniform manner, while (7) results in a more uniform transformation of all $(t_{(m)}, x_{(m)}^a)$ as it only uses d independent parameters. We thus arrive at the conclusion that we should identify standard Minkowski space as the space of subspaces of codimension d+1 labelled by the coordinates (4). We note that identifying Minkowski space by picking local coordinates $(ct_{(0)}, x_{(0)}^a) = (X^0, X^a)$ while setting the 'higher' coordinates to zero is not suitable as the transformations of (2) either do not preserve this choice or lead to just Galilean boosts instead of relativistic Lorentz boosts.

Post-Newtonian expansion and finite expansions

We now compare the structure of the transformations (3) to those of an expansion in 1/c of standard Lorentz boosts. For coordinates (T, \vec{X}) on Minkowski space, a finite boost with rapidity parameter $\vec{\theta} = \theta \vec{n}$ and $\theta = \operatorname{arctanh}(v/c)$ acts to order $1/c^2$ by

$$T' = T + \frac{1}{c}\vec{\theta} \cdot \vec{X} + \frac{1}{2}\vec{\theta}^{2}T + \dots$$

$$= T + c^{-2}\vec{v}_{(0)} \cdot \vec{X} + \frac{1}{2}c^{-2}\vec{v}_{(0)}^{2}T + O(c^{-3}), \quad (9)$$

$$\vec{X}' = \vec{X} + c\vec{\theta}T + \frac{1}{2}(\vec{\theta} \cdot \vec{X})\vec{\theta} + \frac{c}{6}\vec{\theta}^{2}\vec{\theta}T + \dots$$

$$= \vec{X} + \vec{v}_{(0)}T + c^{-2}\vec{v}_{(1)}T + \frac{c^{-2}}{2}(\vec{v}_{(0)} \cdot \vec{X})\vec{v}_{(0)}$$

$$+ \frac{c^{-2}}{6}\vec{v}_{(0)}^{2}\vec{v}_{(0)}T + O(c^{-3}), \quad (10)$$

where the expansion of

$$\vec{\theta} = \operatorname{arctanh}(v/c)\vec{n} = \frac{1}{c}\vec{v}_{(0)} + \frac{1}{c^3}\vec{v}_{(1)} + \dots$$
 (11)

was introduced. When dropping consistently all terms of higher order, these transformations close.

The step of dropping all higher order terms can be circumvented by formally expanding also the Minkowski coordinates according to

$$T = t_{(0)} + c^{-2}t_{(1)} + \dots, \quad \vec{X} = \vec{x}_{(0)} + c^{-2}\vec{x}_{(1)} + \dots$$
 (12)

and reading off the coefficients at fixed order in 1/c. This leads to

$$\begin{split} t'_{(0)} &= t_{(0)} \,, \qquad \vec{x}'_{(0)} &= \vec{x}_{(0)} + \vec{v}_{(0)} t_{(0)} \,, \\ t'_{(1)} &= t_{(1)} + \vec{v}_{(0)} \cdot \vec{x}_{(0)} + \frac{1}{2} \vec{v}_{(0)}^2 t_{(0)} \,, \\ \vec{x}'_{(1)} &= \vec{x}_{(1)} + \vec{v}_{(1)} t_{(0)} + \vec{v}_{(0)} t_{(1)} + \frac{1}{2} \vec{v}_{(0)}^2 \vec{x}_{(0)} + \frac{1}{6} \vec{v}_{(0)}^2 \vec{v}_{(0)} t_{(0)} \,. \end{split} \tag{13}$$

These finite transformations close without dropping any higher order terms. Moreover, the transformations (13) coincide exactly with the finite transformations obtained from (3) when consistently quotienting the algebra \mathfrak{G}_{∞} by the ideal spanned by all generators with superscript larger than N=1. More generally, the algebra (2) can be truncated consistently by setting to zero all generators with superscript larger than some fixed integer $N \geq 0$. The case N=0 corresponds to the Galilei algebra and the first line in (13) indeed corresponds to the standard Galilean boost on the zero modes while the transformation of $(t_{(1)}, \vec{x}_{(1)})$ build up higher polynomials. This pattern persists when increasing N to N+1: the lower transformations up to $(t_{(N)}, \vec{x}_{(N)})$ are unchanged and the new $(t_{(N+1)}, \vec{x}_{(N+1)})$ transform with higher degree polynomials. The transformations (13) generate a group as the truncation of the algebra (2) to any fixed N is a consistent Lie algebra quotient. We can recover the post-Newtonian expansion of the Lorentz boost to any given order c^{-2N} from the transformation of the highest coordinates $(t_{(N)}, \vec{x}_{(N)})$.

The quotient algebra obtained with N=1 is exactly the one that has appeared in [14] in the context of nonrelativistic gravity. Here, we have obtained it from a purely kinematical analysis of post-Newtonian expansions.

Invariant metric and particle dynamics

The subspace of the generalized Minkowski space defined for fixed (X^0, X^a) has an invariant metric under the transformations (3) given by

$$ds^2 = -dX^{0^2} + d\vec{X}^2. (14)$$

Using (4) we have

$$ds^{2} = \sum_{m,n=0}^{\infty} c^{-2(m+n)} \left(-c^{2} dt_{(m)} dt_{(n)} + d\vec{x}_{(m)} \cdot d\vec{x}_{(n)} \right).$$
(15)

Note that the term of order c^0 is

$$-2dt_{(0)}dt_{(1)} + d\vec{x}_{(0)} \cdot d\vec{x}_{(0)} \tag{16}$$

and is an invariant metric for the Bargmann algebra. We also recognize it as the metric of (d+2)-dimensional Minkowski space written in light-cone coordinates. This is in agreement with the fact that the Bargmann algebra in d+1 dimensions is a subalgebra of the Poincaré algebra in d+2 dimensions.

From the invariant metric we can construct a particle action for a massive particle by considering

$$S_{\text{part.}} = -mc \int ds = -mc \int d\tau \sqrt{-\dot{X}^{\mu} \dot{X}_{\mu}},$$

= $S_{(0)} + S_{(1)} + S_{(2)} + \dots$ (17)

with

$$\begin{split} S_{(0)} &= -m \, c^2 \int d\tau \, \dot{t}_{(0)} \,, \\ S_{(1)} &= m \int d\tau \left[-\dot{t}_{(1)} + \frac{\dot{\vec{x}}_{(0)}^2}{2 \, \dot{t}_{(0)}} \right], \\ S_{(2)} &= \frac{m}{c^2} \int d\tau \left[-\dot{t}_{(2)} + \frac{\dot{\vec{x}}_{(0)} \cdot \dot{\vec{x}}_{(1)}}{\dot{t}_{(0)}} - \frac{\dot{t}_{(1)} \dot{\vec{x}}_{(0)}^2}{2 \, \dot{t}_{(0)}^2} + \frac{\dot{\vec{x}}_{(0)}^4}{8 \, \dot{t}_{(0)}^3} \right], \end{split}$$
(18)

where τ is the affine embedding parameter and the dot denotes a τ -derivative. The first action leading to non-trivial dynamics is $S_{(1)}$, which is invariant under the Bargmann algebra. The variable $t_{(1)}$ can be eliminated, leading to a quasi-invariance under the Galilei algebra and to the standard action of the massive non-relativistic particle invariant under worldline diffeomorphisms.

The subsequent actions $S_{(n)}$ with n>1 describe non-relativistic particles plus post-Newtonian corrections; they are individually invariant under (3). We will illustrate this by considering the action $S_{(2)}$, which includes the coordinates $(t_{(0)}, \vec{x}_{(0)}, t_{(1)}, \vec{x}_{(1)}, t_{(2)})$. Due to the extra coordinate $t_{(2)}$, this action is invariant under a central extension of the symmetry algebra realized by the transformations (13) (this symmetry has been considered in the (2+1)-dimensional case in [15, 16]), which is obtained by adding the transformation law

$$t'_{(2)} = t_{(2)} + \vec{v}_{(0)} \cdot \vec{x}_{(1)} + \vec{v}_{(1)} \cdot \vec{x}_{(0)} + \vec{v}_{(1)} \cdot \vec{v}_{(0)} t_{(0)}$$

$$+ \frac{1}{2} \vec{v}_{(0)}^2 t_{(1)} + \frac{1}{6} \vec{v}_{(0)}^2 \vec{v}_{(0)} \cdot \vec{x}_{(0)} + \frac{1}{24} \vec{v}_{(0)}^4 t_{(0)}.$$
 (19)

Neglecting the total derivative $\dot{t}_{(2)}$ leads to an action that is only quasi-invariant under the transformations (13). In fact, its variation gives total derivatives of the form $d(\vec{v}_{(0)} \cdot \vec{x}_{(1)})/d\tau$ and $d(\vec{v}_{(1)} \cdot \vec{x}_{(0)})/d\tau$, which are in correspondence with central extensions

$$[B_a^{(0)}, P_b^{(1)}] = -\delta_{ab}H^{(2)}, \quad [B_a^{(1)}, P_b^{(0)}] = -\delta_{ab}H^{(2)}$$
(20)

of the algebra truncated at N=1. This is in complete analogy with the analysis of the non-relativistic particle, where the invariance under the Bargmann algebra can be deduced from the quasi-invariance of the action under Galilean transformations, which leads to the central extension $[B_a^{(0)}, P_b^{(0)}] = -\delta_{ab}H^{(1)}$ [33].

Defining the canonical momenta $\vec{p}^{(m)} = \partial L/\partial \dot{\vec{x}}_{(m)}$ and $E^{(m)} = -\partial L/\partial \dot{t}_{(m)}$ for $(t_{(0)}, \vec{x}_{(0)}, t_{(1)}, \vec{x}_{(1)}, t_{(2)})$, satisfying the Poisson brackets

$$\{t_{(m)}, E^{(n)}\} = -\delta_m^n, \quad \{x_{(m)}^a, p_b^{(n)}\} = \delta_m^n \delta_b^a,$$
 (21)

leads to the primary constraints

$$\vec{p}^{(1)} \cdot \vec{p}^{(0)} - \frac{1}{2}E^{(1)2} - \frac{m}{c^2}E^{(0)} = 0,$$

$$\frac{1}{2}\vec{p}^{(1)2} - \frac{m}{c^2}E^{(1)} = 0, \quad \frac{m}{c^2} - E(2) = 0.$$
(22)

These 'generalized mass-shell conditions' define first-class constraints associated to the reparametrization invariance of the action. They can be made second-class by introducing a suitable gauge-fixing [34], which in this case can be chosen as

$$\frac{1}{c^{2m}}t_{(m)} = t = \tau \quad \text{(for all } m),$$
 (23)

including fixing the affine parameter. This allows in particular to express the generalized momenta E^m in terms p(m). Finally by projecting the action on the hyperplane

$$\frac{1}{c^{2m}}x^a_{(m)} = x^a \quad \text{(for all } m),$$
 (24)

the action $S_{(2)}$ becomes

$$S_{(2)} = \int dt \left[-mc^2 + \frac{m}{2}\dot{\vec{x}}^2 + \frac{m}{8c^2}\dot{\vec{x}}^4 \right].$$
 (25)

Note that the conditions (23) and (24) break the symmetry (3) of the action. The energy and momentum for this action are given by

$$E = mc^2 + \frac{m}{2}\dot{\vec{x}}^2 + \frac{3m}{8c^2}\dot{\vec{x}}^4, \quad \vec{P} = m\dot{\vec{x}} + \frac{m}{2c^2}\dot{\vec{x}}^2\vec{x}, \quad (26)$$

which correspond to the usual non-relativistic relations of energy and momentum plus their first post-Newtonian corrections. Also note that each term in the expansion (26) in 1/c can be recovered from the canonical momenta defined in (21). Indeed, for any m, $c^{2m}E^{(m)}$ has dimension of energy, while $c^{2m}\vec{p}^{(m)}$ has dimension of linear momentum. One can verify that after gauge fixing and imposing (24), these quantities reduce to the terms of E and \vec{p} in (26) order by order. More general actions $S_{(n)}$ incorporate more generalized momenta and first-class constraints, leading to post-Newtonian corrections up to order $1/c^{2n-2}$ for a non-relativistic particle [27].

One can also use the coordinate expansion (4) in the Nambu–Goto action for a bosonic relativistic string. In the same way one can consider the Nambu–Goto action

$$S_{\text{string}} = -\frac{T}{c} \int d\tau d\sigma \left[\left(\dot{X}^{\mu} X'_{\mu} \right)^{2} - \dot{X}^{\mu} \dot{X}_{\mu} X'^{\nu} X'_{\nu} \right]^{1/2}$$

$$= -T \int d\tau d\sigma \left[\sqrt{\left| \dot{t}_{(0)} \vec{x}'_{(0)} - t'_{(0)} \dot{\vec{x}}'_{(0)} \right|^{2}} + \dots \right], \tag{27}$$

where the next-to-leading term, proportional to $1/c^2$, is straightforward to compute but rather cumbersome and therefore omitted here. Unlike the case of the relativistic particle, the first term in this expansion is exactly invariant under the Galilei algebra [28, 35], while the term proportional to $1/c^2$ is also invariant under (13) without the need of adding central extensions.

Summary

In this note, we have outlined a procedure for systematically describing symmetries of post-Newtonian expansions through the embedding in an enlarged space. The necessity of working up to a certain order in the 1/c-expansion is replaced by exactly closing symmetry transformations on an extended space. Invariant non-relativistic particles and strings can be easily considered in this language. The same type of symmetry algebras as considered have appeared in the context of non-relativistic gravity theories [14–16]. In this context, the results shown in [18] naturally arise when gauging this infinite-dimensional symmetry or, in other words, from a differentiable manifold whose tangent space is the generalized Minkowski space presented here.

We acknowledge discussions with Eric Bergshoeff, Roberto Casalbuoni, Jaume Gomis, Niels Obers and Diederik Roest. JG acknowledges the hospitality and support of the Van Swinderen Institute where this work was finished. JG also has been supported in part by MINECO FPA2016-76005-C2-1-P and Consolider CPAN, and by the Spanish government (MINECO/FEDER) under project MDM-2014-0369 of ICCUB (Unidad de Excelencia Marà de Maeztu). PS-R acknowledges DI-VRIEA for financial support through Proyecto Postdoctorado 2019 VRIEA-PUCV.

- * International Solvay Institutes, ULB-Campus Plaine CP231, BE-1050 Brussels, Belgium
- [1] In any physical problem, the expansion involves a dimensionless ratio v/c where the velocity v depends on the system under study, e.g. v could be the velocity of a particle or a relative velocity. In theories with extended objects the non-relativistic expansion is not unique [28, 36].
- [2] E. Inönü and E. P. Wigner, Proc. Nat. Acad. Sci. 39, 510 (1953).
- [3] G. Dautcourt, Acta Phys. Polon. B 65, 637 (1964).
- [4] G. Dautcourt, Acta Physica Polonica. Series B **21**, 755 (1990).
- [5] L. Blanchet, T. Damour, B. R. Iyer, C. M. Will, and A. Wiseman, Phys. Rev. Lett. 74, 3515 (1995), arXiv:gr-qc/9501027 [gr-qc].
- [6] A. Buonanno and T. Damour, Phys. Rev. **D59**, 084006 (1999), arXiv:gr-qc/9811091 [gr-qc].
- [7] L. Blanchet, Living Rev. Rel. 17, 2 (2014), arXiv:1310.1528 [gr-qc].
- [8] C. M. Will, Living Rev. Rel. 17, 4 (2014), arXiv:1403.7377 [gr-qc].
- [9] T. Damour, P. Jaranowski, and G. Schäfer, Phys. Rev. D91, 084024 (2015), arXiv:1502.07245 [gr-qc].

- [10] J. Ehlers and T. Buchert, Gen. Rel. Grav. 29, 733 (1997), arXiv:astro-ph/9609036 [astro-ph].
- [11] G. Breit, Phys. Rev. 39, 616 (1932).
- [12] R. Andringa, E. Bergshoeff, S. Panda, and M. de Roo, Class. Quant. Grav. 28, 105011 (2011), arXiv:1011.1145 [hep-th].
- [13] J. Hartong, E. Kiritsis, and N. A. Obers, Phys. Rev. D92, 066003 (2015), arXiv:1409.1522 [hep-th].
- [14] D. Hansen, J. Hartong, and N. A. Obers, Phys. Rev. Lett. 122, 061106 (2019), arXiv:1807.04765 [hep-th].
- [15] N. Ozdemir, M. Ozkan, O. Tunca, and U. Zorba, JHEP 05, 130, arXiv:1903.09377 [hep-th].
- [16] E. Bergshoeff, J. M. Izquierdo, T. Ortín, and L. Romano, JHEP 08, 048, arXiv:1904.08304 [hep-th].
- [17] O. Khasanov and S. Kuperstein, J. Phys. **A44**, 475202 (2011), arXiv:1103.3447 [hep-th].
- [18] D. Hansen, J. Hartong, and N. A. Obers, (2019), arXiv:1904.05706 [gr-qc].
- [19] J. Gomis, A. Kleinschmidt, and J. Palmkvist, JHEP 09, 109, arXiv:1907.00410 [hep-th].
- [20] J. A. de Azcarraga, J. M. Izquierdo, M. Picon, and O. Varela, Nucl. Phys. B662, 185 (2003), arXiv:hep-th/0212347 [hep-th].
- [21] M. Hatsuda and M. Sakaguchi, Prog. Theor. Phys. 109, 853 (2003), arXiv:hep-th/0106114 [hep-th].
- [22] F. Izaurieta, E. Rodriguez, and P. Salgado, J. Math. Phys. 47, 123512 (2006), arXiv:hep-th/0606215 [hep-th].
- [23] R. De Pietri, L. Lusanna, and M. Pauri, Class. Quant. Grav. 12, 219 (1995), arXiv:gr-qc/9405046 [gr-qc].
- [24] G. Dautcourt, Class. Quant. Grav. 14, A109 (1997), arXiv:gr-qc/9610036 [gr-qc].
- [25] W. Tichy and E. E. Flanagan, Phys. Rev. D84, 044038 (2011), arXiv:1101.0588 [gr-qc].
- [26] D. Van den Bleeken, Class. Quant. Grav. 34, 185004 (2017), arXiv:1703.03459 [gr-qc].
- [27] J. Gomis, A. Kleinschmidt, J. Palmkvist, and P. Salgado-Rebolledo, in preparation.
- [28] C. Batlle, J. Gomis, and D. Not, JHEP 02, 049, arXiv:1611.00026 [hep-th].
- [29] D. M. Peafiel and L. Ravera, J. Math. Phys. 58, 081701 (2017), arXiv:1611.05812 [hep-th].
- [30] This affine Kac–Moody algebra possesses an integer grading different from the superscripts of the generators.
- [31] We restrict all discussions to boosts for simplicity; the same analysis applies to rotations and translations.
- [32] We also note that at this point we have evaluated the formal power series in 1/c and convergence requires v < c.
- [33] The relation among quasi-invariance of a Lagrangian and extension of a symmetry algebra was studied in [37, 38].
- [34] M. Henneaux and C. Teitelboim, Quantization of gauge systems (Princeton university press, 1994).
- [35] J. Gomis and P. K. Townsend, JHEP **02**, 105, arXiv:1612.02759 [hep-th].
- [36] A. Barducci, R. Casalbuoni, and J. Gomis, Phys. Rev. D98, 085018 (2018), arXiv:1804.10495 [hep-th].
- [37] J.-M. Lévy-Leblond, Communications in Mathematical Physics **12**, 64 (1969).
- [38] G. Marmo, G. Morandi, A. Simoni, and E. C. G. Sudarshan, Phys. Rev. D37, 2196 (1988).