# Accelerated Gravitational Point Set Alignment with Altered Physical Laws* 

Vladislav Golyanik ${ }^{1}$<br>${ }^{1}$ MPI for Informatics

Christian Theobalt ${ }^{1}$<br>${ }^{2}$ University of Kaiserslautern

Didier Stricker ${ }^{2,3}$<br>${ }^{3}$ DFKI


#### Abstract

This work describes Barnes-Hut Rigid Gravitational Approach (BH-RGA) - a new rigid point set registration method relying on principles of particle dynamics. Interpreting the inputs as two interacting particle swarms, we directly minimise the gravitational potential energy of the system using non-linear least squares. Compared to solutions obtained by solving systems of second-order ordinary differential equations, our approach is more robust and less dependent on the parameter choice. We accelerate otherwise exhaustive particle interactions with a Barnes-Hut tree and efficiently handle massive point sets in quasilinear time while preserving the globally multiply-linked character of interactions. Among the advantages of $B H-R G A$ is the possibility to define boundary conditions or additional alignment cues through varying point masses. Systematic experiments demonstrate that BH-RGA surpasses performances of baseline methods in terms of the convergence basin and accuracy when handling incomplete, noisy and perturbed data. The proposed approach also positively compares to the competing method for the alignment with prior matches.


## 1. Introduction

Alignment of point sets provided in different reference frames is an essential algorithmic component in multiple fields including but not limited to visual odometry [33], 3D reconstruction and augmented reality [40], [39], robot navigation and localisation [46], computer graphics [48], CAD modelling [52], cultural heritage [8] and medical technology [11, 25]. Even though rigid point set registration is a well-studied area $[6,11,43,20,12,18,25,50,41,38$, 14, 21, 15], multiple challenges are remaining - on the top of the list are the broader basin of convergence, the robustness to disturbing effects in the samples (noise, missing data, clustered outliers) and handling of massive point sets. We are considering rigid alignment (RA) relative to a fixed reference frame $\mathscr{R}$ (perhaps an observer's reference frame), and coordinates of one of the point sets $\mathcal{X}$ are known in $\mathscr{R}$.

[^0]

Figure 1: An overview of our BH-RGA approach. In every external iteration, we compute a Barnes-Hut tree for the reference and template jointly and specify an optimisation problem. Next, an NLLS solver runs for a few internal iterations while requiring $\mathcal{O}(M \log N)$ time to approximate gravitational potential energy of globally multiply-linked particle interactions.

We refer to $\mathcal{X}$ as a reference. The objective of RA is the recovery of a non-injective, non-surjective correspondence function $f: \mathcal{Y} \rightarrow \mathcal{X}$ together with the rigid transformation parameters, i.e., rotation $\mathbf{R}$ and translation $\mathbf{t}(6 \mathrm{DoF})$ registering the template point set $\mathcal{Y}$ to $\mathcal{X}$.

All RA techniques can be divided into local and multiply-linked policies. An RA method is called globally multiply-linked if and only if every point of $\mathcal{X}$ interacts with every point of $\mathcal{Y}$, with an example of the recently proposed particle dynamics based gravitational approach (GA) [21]. In GA, every point of $\mathcal{Y}$ interacts with every point of $\mathcal{X}$ through a force field. The method was shown to be more robust to large amounts of noise. This robustness comes at the cost of quadratic complexity and parameter sensitivity which impedes its applicability to real-world large and dense point sets. Moreover, the solution based on physical simulation is slow and unstable. We believe that considerable improvements are possible in the field of physics-based RA. In practical applications, we are interested in fast convergence and well-posedness with respect to the parameter settings, and less in the physically accurate modelling. On the other hand, feasible acceleration techniques for particle dynamics based RA remained unexplored so far and are worth to investigate.

Contributions. In the proposed exposition, we present a stable, robust and accelerated globally multiply-linked particle dynamics based RA approach and overcome the limitations of previous works, see Fig. 1 for an overview. Recall that behind the scenes, GA finds transformations corre-
sponding to the locally minimal gravitational potential energy (GPE). Along with that, the physically accurate GPE is not the optimal form to minimise by general-purpose iterative optimisation techniques (e.g., Gauss-Newton method).

Thus, we first apply a negative elementwise reciprocal transform (Sec. 4) and enable efficient solution in the framework of non-linear least squares (NLLS). The proposed transform changes the form of the energy functional to a weighted sum of squared residuals and inverts physics while preserving the advantages of multiply-linked particle dynamics (e.g., notion of the particle mass). Thanks to the new form of GPE, optimisation requires fewer iterations and can be performed with a general-purpose GaussNewton solver, thus eliminating the need for the explicit updates of forces, acceleration and displacements. Moreover, the number of parameters is reduced by the factor of three while showing a wider convergence basin and outperforming the baseline GA [21] in the accuracy.

Second, we adopt an acceleration technique based on the idea of Barnes and Hut [5] initially developed for $N$ body simulations [1] (Sec. 4.2). Applied to RA, a BarnesHut (BH) tree allows to efficiently handle large point sets with hundreds of thousands of points while preserving the globally multiply-linked nature of interactions. In other words, this happens not at the expense of constraining interactions to local vicinities but rather by accumulating contributions of reference points into clusters. The BH tree encompasses multiple extractable representations of $\mathcal{X}$ with spatially varying cluster configurations.

We call the new method Barnes-Hut Rigid Gravitational Approach (BH-RGA). Finally, we design a set of systematic tests for RA with heightened complexity and variety. We compare $\mathrm{BH}-\mathrm{RGA}$ to multiple RA methods [ $6,18,38,32,21,23$ ] on our extensive evaluation benchmark and apply it to real data originating from cultural heritage and a lidar sensor (Sec. 5).

## 2. Related Work

Early point set registration algorithms were motivated by emerging 3D scanners producing partial point clouds that need to be aligned. The seminal iterative closest point (ICP) algorithm for aligning two point clouds [6, 11] alternates between transformation estimation [28,29] and local correspondence inference [17]. Even though ICP is easy to implement, its fundamentally heuristic local correspondence search makes it prone to erroneous local convergence if badly initialised, and sensitive to outliers. Different improvements were subsequently proposed for ICP, ranging from accelerating policies for nearest neighbour search [26, 41] and relaxation of one-to-one correspondences [20] to more efficient optimisation schemes [45, 18].

Another class of methods models source and target point clouds as probability density functions [12, 38, 32], such
as Gaussian Mixture Models (GMMs). Coherent Point Drift (CPD) [38] explicitly incorporates a model of uniform noise. A wider convergence basin for alignment was obtained by Gaussian mixture decoupling [16] or combining GMM representation and continuous domain mapping with a support vector machine [10]. The multiply-linked Kernel Correlation (KC) approach minimises the Renyi's quadratic entropy of the joint system of the reference and the transformed template [50]. In contrast to BH-RGA, only local neighbourhoods are involved in one-to-many interactions.

More recently, approaches relying on analogies to physical processes gained attention [14, 21, 3]. Deng et al. cast point sets into the Schrödinger distance transform representation, and align them by minimising a geodesic distance on a unit Hilbert sphere [14]. Golyanik et al. [21] interpret point sets as rigid swarms of particles with masses moving in a gravitational force field. An optimal alignment corresponds to the state of minimal GPE of the system. Quadratic complexity and parameter sensitivity impede applicability of GA in practice. We develop the idea of global multiplylinking further and accelerate all-to-all interactions with a BH tree resulting in quasilinear computational complexity. We are altering the law of particle dynamics and simulate an inverse world. In the inverse world, a state with locally optimal GPE can be found iteratively by NLLS more easily. Thus, we remove viscosity (an energy dissipator in GA), and the gravitational constant becomes redundant as a multiplicative factor in the energy.

Large point sets in RA are often treated by subsampling [25], and rarely raw, let alone in the globally multiplylinked fashion. Some methods naturally allow embedding of prior matches [6,23]. Similarly to GA, BH-RGA uses varying masses to define different boundary conditions. Particle masses can be set proportional to the reliabilities of superimposed prior correspondences. Some methods additionally use colours as an alignment cue [13]. In BH-RGA, point intensities can be likewise mapped to masses.

## 3. Newtonian Gravitational Approach [21]

Consider globally multiply-linked gravitational alignment of two point sets, with the reference $\left[\mathbf{x}_{j}\right]=\mathbf{X} \in$ $\mathbb{R}^{D \times N}, j \in\{1, \ldots, N\}$ and the template $\left[\mathbf{y}_{i}\right]=\mathbf{Y} \in$ $\mathbb{R}^{D \times M}, i \in\{1, \ldots, M\} . N$ and $M$ denote the number of points in the reference and template, respectively, and $D$ is point set dimensionality. The objective is to recover the parameter set $\vartheta=\{\mathbf{R}, \mathbf{t}\}$, i.e., rotation $\mathbf{R}\left(\mathbf{R}^{-1}=\mathbf{R}^{\top}\right.$, $\operatorname{det}(\mathbf{R})=1$ ) and translation $\mathbf{t}$ aligning $\mathbf{Y}$ to $\mathbf{X}$. In [21], point sets are aligned by minimising the mutual gravitational potential energy (GPE) $\mathbf{U}$ of the corresponding system of particles in the force field induced by $\mathbf{X}$ :

$$
\begin{equation*}
\mathbf{U}(\mathbf{R}, \mathbf{t})=-G \sum_{i, j} \frac{m_{\mathbf{y}_{i}} m_{\mathbf{x}_{j}}}{\left\|\mathbf{R} \mathbf{y}_{i}+\mathbf{t}-\mathbf{x}_{j}\right\|_{2}+\epsilon} \tag{1}
\end{equation*}
$$

where $m_{\mathbf{y}_{i}}$ and $m_{\mathbf{x}_{j}}$ denote masses, $G$ is the gravitational constant and $\epsilon$ is a softening parameter for preventing local near-field singularities. For the reasons stressed in [24], we do not include scaling in (1) and our energy functionals (5) and (9). Golyanik et al. [21] minimise (1) implicitly by updating the forces $\overrightarrow{\mathbf{f}}_{i}$ acting on particles $\mathbf{y}_{i}$, accelerations (using Newton's second law of motion $\overrightarrow{\mathbf{f}}_{i}=m_{\mathbf{y}_{i}} \vec{a}_{i}$ ), velocities $v_{i}^{t+1}$ and individual point displacements $d_{i}^{t+1}$ [1]:

$$
\begin{gather*}
\overrightarrow{\mathbf{f}}_{i}=-G m_{\mathbf{y}_{i}} \sum_{j} m_{\mathbf{x}_{j}}\left(\left\|\mathbf{y}_{i}-\mathbf{x}_{j}\right\|^{2}+\epsilon^{2}\right)^{-3 / 2} \hat{\mathbf{n}}_{i, j}-\eta v_{i}^{t} \\
v_{i}^{t+1}=v_{i}+\Delta t \frac{\overrightarrow{\mathbf{f}}_{i}}{m_{\mathbf{y}_{i}}} \text { and } d_{i}^{t+1}=\Delta t v_{i}^{t+1} \tag{2}
\end{gather*}
$$

In (2), $\eta$ denotes a dissipation constant which determines the portion of the kinetic energy which is dragged from the system per template particle. In (3), $\Delta t$ is the the forward integration step (time interval of the simulation). In every iteration, once updated, the unconstrained displacements are added to the current positions. A consensus rigid transformation is found using Procrustes analysis [29], which relates the previous and current poses. The algorithm converges when the difference in the GPE of two last successive system states is below some small threshold $\rho_{E}$ or terminates when the maximum number of iterations is reached.

## 4. The Proposed Approach

### 4.1. Our Gravitational Potential Energy Functional

GPE (1) includes a reciprocal relation and is minimised implicitly by ordinary differential equations (ODE) of second order by updating forces acting on the template particles, point accelerations and displacements. The forces are inversely proportional to distances, and at the optimal alignment, the GPE is limited by $-\infty$. For the transformation parameters $\mathbf{R}$ and $\mathbf{t}$ to converge, additional regularisation constants are introduced in (2), i.e., softening distance $\epsilon$ to avoid near-field singularities, and an energy dissipation constant $\eta$ to, de facto, prevent infinite oscillations. The combination of all these properties (reciprocal relation, regularising energy dissipation term, $\epsilon$ ) makes GA ill-posed with respect to the parameter values. Thus, every scenario demands different parameters - otherwise, the method will not converge. Another disadvantage of the implicit solution with second-order ODEs is that a large number of iterations is required until convergence ( $300-1000$ iterations are not rare). Moreover, $\mathbf{R}$ is updated from unconstrained point displacements regularised by Procrustes fitting. Ideally, we would like to optimise directly in the $S O(D)$ group.

We propose to directly optimise a GPE and avoid implicit solution by the second-order ODEs. Moreover, we eliminate all disadvantages of (1) and reduce the number of parameters while preserving the advantages of gravitational
alignment. Since we are solving a computer vision problem and are not interested in accurate physical simulations, we are free to alter laws of simulated physics, as long as it brings us closer to the desired properties.

Consider a negative elementwise reciprocal transform operator $\eta^{-}(\cdot)$ which acts on every element $c$ of an arbitrary sum or a vector as $\eta^{-}\left(-\frac{1}{c}\right)=c . \eta^{-}(\cdot)$ preserves function monotonicity and is reversible ${ }^{i}$. We apply $\eta^{-}(\cdot)$ to (1) and obtain:

$$
\begin{equation*}
\xi^{-}(\mathbf{U}(\mathbf{R}, \mathbf{t}))=\sum_{i, j} \frac{1}{G m_{\mathbf{y}_{i}} m_{\mathbf{x}_{j}}}\left\|\mathbf{R} \mathbf{y}_{i}+\mathbf{t}-\mathbf{x}_{j}\right\|_{2}+\epsilon \tag{4}
\end{equation*}
$$

Eq. (4) changes the model of the simulated world. In the inverse world, the potential between two particles is inversely proportional to the product of their masses and directly proportional to the distance between the points. The meaning of a mass has changed - now, the mass is comprehended as a property of the matter so that the less its value, the more significant is the interaction. Nevertheless, we can preserve the notion of mass and replace $\frac{1}{m_{\mathbf{y}_{i}}^{\text {old }}}=m_{\mathbf{y}_{i}}^{\text {new }}$ and $\frac{1}{m_{\mathbf{x}_{j}}^{\text {old }}}=m_{\mathbf{x}_{j}}^{\text {new }}$. Note that the softening parameter $\epsilon$ can be omitted, due to the vanished near-field singularities. Moreover, we get rid of $G$ because it is a global multiplicative constant in the energy. Thus, our final form of GPE $\mathbf{E}$ is ${ }^{\text {ii }}$

$$
\begin{equation*}
\mathbf{E}(\mathbf{R}, \mathbf{t})=\sum_{i} \sum_{j} m_{\mathbf{y}_{i}} m_{\mathbf{x}_{j}}\left\|\mathbf{R} \mathbf{y}_{i}+\mathbf{t}-\mathbf{x}_{j}\right\|_{2} \tag{5}
\end{equation*}
$$

With the new GPE (5), the optimal alignment is still achieved when GPE is locally minimal, while there is no reciprocity. It is noteworthy that in both cases, i.e, $\mathbf{U}(\mathbf{R}, \mathbf{t})$ and $\xi^{-}(\mathbf{U}(\mathbf{R}, \mathbf{t}))$, the further are two particles apart from each other, the higher is the GPE between them:

$$
\begin{align*}
& \lim _{\|\mathbf{R Y}+\mathbf{T}\|_{\mathcal{F}} \rightarrow\|\mathbf{X}\|_{\mathcal{F}}} \quad \mathbf{U}(\mathbf{R}, \mathbf{t})=-\infty, \text { and }  \tag{6}\\
& \lim _{\|\mathbf{R Y}+\mathbf{T}\|_{\mathcal{F}} \rightarrow\|\mathbf{X}\|_{\mathcal{F}}} \xi^{-}(\mathbf{U}(\mathbf{R}, \mathbf{t}))=0, \tag{7}
\end{align*}
$$

where $\|\cdot\|_{\mathcal{F}}$ denotes Frobenius norm, and we use a shorthand $\mathbf{T}_{D \times M}=\left[\begin{array}{llll}\mathbf{t} & \mathbf{t} & \ldots & \mathbf{t}\end{array}\right]$. Note that (5) still represents a globally multiply-linked GPE, with no explicit encoding of correspondences. These two factors reveal the deceptive visual similarity of the transformed GPE with the classic ICP formulation [6]. The further difference is that ICP requires alternating between correspondence and transformation updates, whereas (5) does not change throughout the optimisation (when not accelerated).

[^1]

Figure 2: Building a BH tree exemplified. Each point is of different shape and colour for visualisation purposes. A new element either resides in a non-occupied leaf or initiates cell partitioning (while updating cluster masses and centres of mass) until it is placed in its own external leaf.

### 4.2. Acceleration with a Barnes-Hut $2^{D}$-Tree

To surpass the quadratic computational complexity while computing potentials between points, we introduce a new variant of a BH tree or a $2^{D}$-tree which is applied in the acceleration of $N$-body problems in astrodynamics [5]. We approximate exhaustive interactions of every template point $\mathbf{y}_{i}$ with $N$ reference points by interaction with $\mathcal{O}(\log N)$ many clusters which are fetched from the BH tree which we build. Whereas previous approaches de facto restrict interactions to local neighbourhoods [50, 38], we preserve global multiply-linking in our approximation.

Suppose $\mathbf{y}_{i}$ interacts with $K$ sufficiently distant particles $\mathbf{x}_{k}, k \in\{1, \ldots, K\}$ and $\mathbf{x}_{k}$ are sufficiently close to each other. Let us denote $\hat{\mathbf{y}}_{i}=\mathbf{R} \mathbf{y}_{i}+\mathbf{t}$ for briefness. Consider that variance of the distances from $\mathbf{y}_{i}$ to every $\mathbf{x}_{k}$ is below some small $\zeta$. If located sufficiently far away, the total impact of all elements $\mathbf{x}_{k}$ in a volume of space $V$ to $\mathbf{y}_{i}$ can be well approximated by the impact of the combined single particle $\tilde{\mathbf{x}}$. The mass of $\tilde{\mathbf{x}}$ equals to the mass integral of $\mathbf{x}_{k}$ over $V$, and the position of $\tilde{\mathbf{x}}$ equals to the centre of mass of the elements $\mathbf{x}_{k}$ in $V$ [5]:

$$
\begin{equation*}
\sum_{k=1}^{K} m_{\mathbf{y}_{i}} m_{\mathbf{x}_{k}}\left\|\hat{\mathbf{y}}_{i}-\mathbf{x}_{k}\right\|_{2} \approx m_{\mathbf{y}_{i}}\left[\sum_{k=1}^{K} m_{\mathbf{x}_{k}}\right]\left\|\hat{\mathbf{y}}_{i}-\tilde{\mathbf{x}}\right\|_{2} . \tag{8}
\end{equation*}
$$

Building BH Tree. We build a joint BH tree on the point set union $\mathcal{X} \cup \mathcal{Y}$. Since template points are not influenced by each other, we set all $m_{\mathbf{y}_{i}}$ to zeroes. This allows us to include all points of the template into the tree but exclude their effect (mass) to other template points while calculating potentials. If $\mathcal{X}$ and $\mathcal{Y}$ contain duplicated points, the subdivision in a BH tree will continue infinitely long. In practice, the depth of a BH tree is restricted to avoid infinite splits.

A BH tree is initialised as a root node with $2^{D}$ empty external nodes. An insertion of every new $\mathbf{y}_{i}$ always results in a new leaf (an occupied external node) and can initiate clusters (internal nodes). The size of daughter cells is always the half size of the parent cell along all dimensions. Every $\mathbf{y}_{i}$ affects the mass and the centre of mass of every cluster
which includes $\mathbf{y}_{i}$. Fig. 2 schematically visualises BH tree building. We strictly follow the insertion rules proposed in [5]. Additionally, see our supplement for more details.
Using BH Tree. The potential at position $\mathbf{y}_{i}$ is calculated by traversing the BH tree starting from the root towards the leaves and requires a single parameter - the distance threshold $\gamma$. Suppose $l$ is the size of the current examined cell, and $\mu$ is the distance from $\mathbf{y}_{i}$ to the centre of the cell. If $l / \mu<\gamma^{-1}$, then the cell's influence on $\mathbf{y}_{i}$ is accumulative. Otherwise, subcells on the next level are visited, and the examination is performed recursively until the condition $l / \mu<\gamma^{-1}$ is satisfied or the tree is entirely traversed [5].

To obtain a non-zero potential for the current $\mathbf{y}_{i}$, we restore its original mass. After the potential is computed, we set $m_{\mathbf{y}_{i}}$ again to zero in order to avoid interaction between template points. The earlier the node is used, the higher is the amount of individual particle-particle interactions which are bypassed. The time and memory requirements of BH tree building for a given $\mathcal{X} \cup \mathcal{Y}$ depend on how homogeneous point distribution is and have an upper bound, and the speed of GPE calculation depends on $\gamma$ (see Sec. 4.4).

Fig. 3 visualises cluster configurations fetched during registration for a single template point and one instance of the clean-500 experiment (see Sec. 5.1) with $\gamma=5.0$. The input point clouds contain 817 points each, and the cardinality of the configurations fetched from the BH octree lies in the range $[49 ; 62]$ points. Some of the clusters are massless and can be removed from the respective configuration as they comprise only massless template points. Distinct is the trajectory of the exemplary point over all iterations. Some distant particles are joined into clusters which are persistent over all iterations, whereas compositions of other clusters change dynamically from iteration to iteration while the template undergoes a rigid transformation.
GPE with a BH tree. The final GPE of BH-RGA with $2^{D}$ tree acceleration for particle interactions reads

$$
\begin{equation*}
\mathbf{E}(\mathbf{R}, \mathbf{t})=\sum_{\forall \mathbf{y}_{i}} m_{\mathbf{y}_{i}} \sum_{\mathbf{k}_{j} \in \mathcal{K}\left(\mathbf{y}_{i}\right)} m_{\mathbf{k}_{j}}\left\|\mathbf{R} \mathbf{y}_{i}+\mathbf{t}-\mathbf{k}_{j}\right\|_{2}, \tag{9}
\end{equation*}
$$

where $\mathbf{k}_{j} \in \mathcal{K}\left(\mathbf{y}_{i}\right)$ denotes fetchable representations of the reference for every $\mathbf{y}_{i}$ with clusters $\mathbf{k}_{j}$ and masses $m_{\mathbf{k}_{j}}$.

### 4.3. Gravitational Potential Energy Optimisation

The proposed approach is summarised in Alg. 1. BHRGA minimises a sum of squared residuals $f_{r}(\vartheta)=$ $m_{\mathbf{y}_{i}} m_{\mathbf{k}_{j}}\left\|\mathbf{R} \mathbf{y}_{i}+\mathbf{t}-\mathbf{k}_{j}\right\|_{2}^{2}$ weighted by the product of masses. The number and composition of residuals depend on the BH tree configuration and $\gamma$. Since minimising GPE (9) is an overconstrained non-linear optimisation problem, we solve it in the least-squares sense with the LevenbergMarquardt (LM) algorithm [34, 36] by iteratively linearising the GPE in the vicinity of the current solution. For the


Figure 3: Clusters fetched during alignment (clean-500 experiment, see Sec. 5.1) from a BH octree: a/ cluster configuration at the beginning; b/ overlayed cluster configurations for all seven iterations. The colours in $\mathbf{a} / \mathbf{a n d} \mathbf{b} /$ encode the distance to the exemplary template particle (the darker, the further away), and massless clusters are shown in black; c/ cluster configuration in the first iteration, for a single template point (shown as a triangle) and three different $\gamma$ values. The colour scheme is given on the right. Each colour encodes a mass range of the cluster. Initially, all points were assigned unit masses.

```
Algorithm 1 Barnes-Hut Rigid Gravitational Approach
Input: reference \(\mathbf{X} \in \mathbb{R}^{D \times N}\) and template \(\mathbf{Y} \in \mathbb{R}^{D \times M}\)
Output: rigid transformation \(\vartheta=\{\mathbf{R}, \mathbf{t}\}\) aligning \(\mathbf{Y}\) to \(\mathbf{X}\)
    Initialisation: \(\quad \mathbf{R}=\mathbf{I}, \mathbf{t}=\mathbf{0}\); masses \(m_{\mathbf{y}_{i}}\) and \(m_{\mathbf{x}_{j}}\), the
    distance threshold \(\gamma\) for cluster fetching from a \(2^{D}\)-tree
    while not converged do
        build a \(2^{D}\)-tree on the union of point sets \(\{\mathbf{X}, \mathbf{R Y}+\mathbf{T}\}\),
        \(\mathbf{T}_{D \times M}=\left[\begin{array}{llll}\mathbf{t} & \mathbf{t} & \ldots & \mathbf{t}\end{array}\right]\)
        for every template point \(\mathbf{y}_{i}\) do
            fetch a set of clusters \(\mathcal{K}\left(\mathbf{y}_{i}\right)\) from the \(2^{D}\)-tree
        end for
        minimise (9), i.e., the \(\operatorname{GPE} \mathbf{E}(\mathbf{R}, \mathbf{t})=\)
        \(\sum_{\forall \mathbf{y}_{i}} m_{\mathbf{y}_{i}} \sum_{\mathbf{k}_{j} \in \mathcal{K}\left(\mathbf{y}_{i}\right)} m_{\mathbf{k}_{j}}\left\|\mathbf{R} \mathbf{y}_{i}+\mathbf{t}-\mathbf{k}_{j}\right\|_{\epsilon}\)
        with non-linear least squares
    : end while
```

higher resistance to outliers, we apply Huber loss in the evaluation of every $f_{r}(\vartheta)$ defined as:

$$
\|a\|_{\varepsilon}= \begin{cases}\frac{1}{2} a^{2}, & \text { if }|a| \leq \varepsilon  \tag{10}\\ \varepsilon\left(|a|-\frac{1}{2} \varepsilon\right), & \text { otherwise }\end{cases}
$$

The BH tree is renewed after every successful $\Delta \vartheta$ update, and the objective is formulated for the new set of $f_{r}(\vartheta)$ in every external iteration. The number of internal LM iterations is kept in the range $[3 ; 10]$, see Fig. 1. BH-RGA converges when the difference in two consecutive GPE values is below some threshold $\rho_{E}$.

### 4.4. Computational Complexity

Building a BH tree jointly for the union $\{\mathcal{X}, \mathcal{Y}\}$ in every iteration is performed in $\mathcal{O}((M+N) \log (M+N))$ time. The complexity of fetching a set of clusters from the reference for one template point is $\mathcal{O}(\log N)$. In total, this operation is performed for all $M$ template points, and we arrive at $\mathcal{O}(M \log N)$. Thus, the computational complexity of BH-RGA is $\mathcal{O}((M+N) \log (M+N)+M \log N) \approx$ $\mathcal{O}(M \log N)$ (assuming $M \sim N)$. This is a significant improvement compared to the quadratic complexity of GA [21], especially for large point sets.

### 4.5. Defining Boundary Conditions With Masses

Setting varying masses of template and reference points in BH-RGA allows for different effects such as embedding of prior correspondences, additional cues and counterbalancing inhomogeneous point densities.
Sparse Prior Matches. Different masses in BH-RGA can encode prior correspondences as the optimal alignment is more likely when points with larger masses are close to each other. Suppose $(i, j) \in N_{c} \subset \mathbb{N}^{2}$ is the set of points for which correspondences are known in advance. The GPE of BH-RGA which considers prior matches is then specified as

$$
\mathbf{E}^{p}(\mathbf{R}, \mathbf{t})=\left\{\begin{array}{l}
\mathbf{E}(\mathbf{R}, \mathbf{t}) \text { as in (9), } \forall \mathbf{y}_{i}: i \notin N_{c}  \tag{11}\\
m_{\mathbf{y}_{i}}^{p} m_{\mathbf{x}_{j}}^{p}\left\|\mathbf{R} \mathbf{y}_{i}+\mathbf{t}-\mathbf{x}_{j}\right\|_{2}, \text { else }
\end{array}\right.
$$

with $m^{p}=m_{\mathbf{y}_{i}}^{p}=m_{\mathbf{x}_{j}}^{p}$ being, as a rule, several orders of magnitude larger than default masses (depending on the total mass distribution and the reliability of prior matches). Note that for points in $N_{c}$, multiply-linking and cluster fetching are avoided. We refer to anchor points as two subsets of $\mathcal{X}$ and $\mathcal{Y}$ for which it is known that they correspond among each other. Anchor points can be embedded into BH-RGA by setting their masses to $m_{p}$, with no changes in (9). Thus, the effect of anchor points in BH-RGA is closer to alignment in alignment or a hierarchical registration.
Point Colours. Additional alignment cues such as point colour can be converted to masses in BH-RGA. For the GPE to be locally minimal in this setting, points with similar masses have to coincide or be close to each other.
Handling Varying Point Densities. Handling of point sets with inhomogeneous densities has been indicated as challenging for many approaches, with specialised techniques and pre-processing steps (such as re-sampling) required to handle the setting efficiently [30]. For BH-RGA, we propose to normalise total point mass per unit volume in linear time. In volumetric mass normalisation (VMN), the bounding box of a point set is split into multiple non-overlapping voxels, and in every voxel, the pre-defined mass is uniformly distributed among all points belonging to it.

## 5. Experiments

In this section, we describe the evaluation methodology and results. BH-RGA is implemented in $\mathrm{C}++$ with eigen3 [27] and Galois [42] libraries. For NLLS, we use ceres solver [2]. The experiments are performed on a system with 32 Gb RAM and quad-core Intel $i 7-6700 \mathrm{~K}$ processor running at 4 GHz . We compare our BH-RGA to Iterative Closest Point (ICP) [6], LM-ICP [18], Coherent Point Drift (CPD) [38], GMM Registration (GMR) [32] and GA [21]. CPD and GMR are evaluated both as $6 \mathrm{DoF}(\mathbf{R}, \mathbf{t})$ and 7 DoF ( $\mathbf{R}, \mathbf{t}$ and scale $s$ ) variants. We also compare BHRGA against rigid Extended CPD (E-CPD) guided by prior matches [23]. ICP is taken from the Matlab repository [51], and implementations of several methods are publicly available $[37,31,35]$. In all experiments, we report the mean root-mean-square error (RMSE) and the standard deviation of the RMSE denoted by $\sigma$.

### 5.1. Quantitative Evaluation

Main Tests. In the tests with synthetic data, we use a subsampled Stanford bunny [47]. In the first clean-500 test, we sample the 3 D rotation space by the angular displacement of $\frac{2 \pi}{10}$ radians and synthesise 500 different initial configurations (the duplicated states are avoided). In this experiment, we test the methods for their ability to resolve rotation in scenarios with noiseless data. In contrast, the second test is designed to evaluate the ability of RA to converge under severe amounts of noise. We add $100 \%$ of uniform noise to the clean-500 and obtain N500-U100 dataset. The range of the noise generating function is a bounding sphere of the template. Several methods demonstrate poor performance on N500-U100, and we also generate N500-U50 dataset with $50 \%$ of uniform noise according to the same rule.

Next, we choose a single initial configuration which can be successfully resolved by all methods and add three different noise patterns to the template - $100 \%$ of uniform noise, $100 \%$ of Gaussian noise (the range volume is the same as in N500-U100) and $100 \%$ of the Gaussian noise with the mean values coinciding with the template point locations. The resulting datasets are abbreviated as $100 \mathrm{U}, 100 \mathrm{G}$ and 100 GS , respectively. Table 1 summarises the outcomes of the experiments with the new datasets. If RMSE is below 0.1 , the registration is assumed to succeed. This value is not set arbitrarily - in the case of successful alignments, RMSE was, in most cases, much lower than 0.1. Otherwise, it was in most cases considerably larger than 0.1 .

Upon our expectation, BH-RGA is not the most accurate method on the clean-500 dataset (it is the second-best regarding the success rate). Perfectly matching data is, however, not common in practical applications. With an increasing noise, the relative accuracy and performance of BH-RGA compared to all other methods increases. While CPD ( 7 DoF ) achieves a lower RMSE than BH-RGA on
clean-500, the situation inverts on N500-U50 and N500U100, with the success rates of BH-RGA close to those of CPD (and the lowest RMSE compared to CPD). At the same time, ICP, LM-ICP, CPD ( 6 DoF ) and GMR considerably lower their success rates on N500-U50 and N500U100. BH-RGA was the only successful method on U100 - it resolves all cases with the lowest RMSE. The secondbest method is CPD ( 7 DoF , 48 cases out of 50 ), while demonstrating a $50 \%$ to $66 \%$ higher RMSE on N500-U100 and $U 100$, respectively. GMR is able to hold on par with ICP/LM-ICP only in the 6 DoF mode. On G100 and GS100, most methods successfully recover transformations for all cases, except GMR ( 6 DoF ) and GA which recover $52 \%$, $25 \%, 2 \%$ and $20 \%$ of transformations, respectively.

The performance of baseline GA [21] in our tests is, in most cases, inferior to the results of other methods. Our experimental setting is different from the one proposed in [21] where the angle of initial misalignment is randomly selected in the range $\left[0 ; \frac{\pi}{2}\right]$. Golyanik el al. [21] report that GA was starting to fail in the angular misalignment range $\left[\frac{\pi}{4}, \frac{\pi}{2}\right]$. In our clean-500 and N500- experiments, the angles are sampled with the $36^{\circ}$ step implying that starting from the second value (out of ten), the angle is always $\geq 72^{\circ}$. Regarding $U 100, G 100$ and $G S 100$, the excessive amount of noise hinders GA to converge, though the angle of initial misalignment is in the convergence basin of GA.
Perturbed Data. Next, we propose a test for evaluating how robust are the methods to the corrupted data - the U256 and G256 experiments. We perturb template points with uniform and Gaussian noise. In total, there are 256 magnitude indexes generating states with the increasing degree of perturbation. The maximum displacement in every state is equal to the scaled magnitude index times the length of the bounding box of the object in the $x$-direction. In both experiments, no point of the template physically coincides with any reference point at the optimum. RMSE between the reference and the template transformed by the recovered $\mathbf{R}, \mathbf{t}$ is reported in Table 2. CPD and GMR have higher RMSE than ICP and BH-RGA, i.e., both probabilistic methods are sensitive to point perturbations, which agrees with our expectations. ICP shows the lowest RMSE. Especially for small perturbations, conditions for ICP are optimal. BHRGA slightly concedes to ICP as its RMSE is a little higher for small perturbation magnitudes. On the other side, BHRGA has the smallest $\sigma$.
Prior Matches and Anchor Points. We repeat the N500U50 test with up to three prior matches and anchor points, and compare BH-RGA to Extended CPD (E-CPD) [23], see Table 3 for the summary. The latter approach is a correspondence-conditioned CPD variant. Especially with only one or two reliable prior matches, E-CPD has advantages [23]. Thus, CPD with no prior matches resolves $17 \%$ of the cases and its success rate increases to $18 \%$ and $75 \%$

Table 1: Summary of the quantitative evaluation of the compared methods on clean-500, N500-U50, N500-U100, U100, G100 and GS100 datasets.

| methods and metrics |  | clean-500 | N500-U50 | N500-U100 | U100 | G100 | GS100 |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| ICP [6] | success (in \%) | 62 (12.4\%) | 36 (7.2\%) | 19 (3.8\%) | 33 (66\%) | 50 (100\%) | 50 (100\%) |
|  | RMSE ( $\sigma$ ) | 0.005 (0.016) | 0.022 (0.3) | 0.042 (0.031) | 0.091 (0.081) | 0.007 (0.002) | 0.002 (5E-4) |
| LM-ICP [18] | success (in \%) | 123 (24.6\%) | 82 (16.4\%) | 72 (14.4\%) | 49 (98\%) | 50 (100\%) | 50 (100\%) |
|  | RMSE ( $\sigma$ ) | $0.002(6.3 E-4)$ | 0.015 (0.009) | 0.023 (0.017) | 0.025 (0.021) | 0.006 (0.003) | 0.003 (0.001) |
| CPD (7 DoF) [38] | success (in \%) | 130 (26\%) | 128 (25.6\%) | 109 (21.8\%) | 48 (96\%) | 50 (100\%) | 50 (100\%) |
|  | RMSE ( $\sigma$ ) | 0.04 (6E-5) | 0.064 (0.003) | 0.088 (0.003) | 0.098 (0.066) | 0.027 (1.7E-3) | 0.046 (1.7E-3) |
| CPD (6 DoF) [38] | success (in \%) | 143 (28.6\%) | 98 (19.6\%) | 62 (12.4\%) | 48 (96\%) | 50 (100\%) | 50 (100\%) |
|  | RMSE ( $\sigma$ ) | 0.006 (0.009) | 0.025 (0.014) | 0.034 (0.017) | 0.061 (0.148) | 0.01 (0.003) | $7.5 E-3$ (2.3E-3) |
| GMR (7 DoF) [32] | success (in \%) | 126 (25.2\%) | 113 (22.6\%) | 0 (0\%) | 0 (0\%) | 50 (100\%) | 50 (100\%) |
|  | RMSE ( $\sigma$ ) | $7 E-5$ (8E-5) | 0.084 (0.005) | $\mathrm{n} / \mathrm{a}$ | $\mathrm{n} / \mathrm{a}$ | 0.04 (7.5E-3) | $1.5 E-3$ (4.7E-4) |
| GMR (6 DoF) [32] | success (in \%) | 131 (26.2\%) | 79 (15.8\%) | 87 (17.4\%) | 36 (72\%) | 26 (52\%) | 25 (25\%) |
|  | RMSE ( $\sigma$ ) | 8E-5 (2E-4) | $5 E-4(2 E-4)$ | $6 E-4(2 E-4)$ | 0.16 (0.077) | 0.37 (0.4) | 0.37 (0.38) |
| GA [21] | success (in \%) | 21 (4.2\%) | 6 (1.2\%) | 3 (0.6\%) | 19 (38\%) | 1 (2\%) | 10 (20\%) |
|  | RMSE ( $\sigma$ ) | 0.029 (0.021) | 0.049 (0.025) | 0.03 (0.012) | 0.164 (0.082) | 0.289 (0.087) | 0.163 (0.072) |
| BH-RGA (ours) | success (in \%) | 132 (26.4\%) | 132 (26.4\%) | 100 (20\%) | 50 (100\%) | 50 (100\%) | 50 (100\%) |
|  | RMSE ( $\sigma$ ) | 0.009 (0.004) | 0.032 (0.013) | 0.059 (0.021) | 0.056 (0.017) | 0.04 (0.01) | 0.022 (0.006) |

Table 2: RMSE and $\sigma$ (in parentheses) in $U 256$ and G256 experiments.

| method | $U 256$ | $G 256$ |
| :---: | :---: | :---: |
| ICP [6] | $\mathbf{9 E - 3}(\mathbf{7 E - 3})$ | $\mathbf{0 . 0 1 5}(\mathbf{0 . 0 1 2 )}$ |
| LM-ICP [18] | $0.077(0.041)$ | $0.113(0.063)$ |
| CPD (7 DoF) [38] | $0.051(9 \mathrm{E}-3)$ | $0.064(0.021)$ |
| CPD (6 DoF) [38] | $0.016(0.014)$ | $0.045(0.058)$ |
| GMR (7 DoF) [32] | $0.019(0.028)$ | $0.065(0.051)$ |
| GMR (6 DoF) [32] | $0.853(1.16)$ | $1.027(1.226)$ |
| GA [21] | $0.149(0.143)$ | $0.207(0.158)$ |
| BH-RGA (ours) | $\mathbf{0 . 0 1 5 ( \mathbf { E E } - 3 )}$ | $\mathbf{0 . 0 1 9 ( \mathbf { 9 . 5 E - 3 } )}$ |

with one and two prior matches, respectively. BH-RGA resolves $25 \%$ with no priors and $\sim 39 \%$ with one or two available anchor points. Recall that anchor points have higher masses and interact with all other points, which is the reason that the success with three anchor points drops to $33 \%$. In contrast, BH-RGA resolves all cases with two or three prior correspondences and outperforms E-CPD. Prior matches are not involved in the interaction with other points, which is the strongest boundary condition.

Runtime and Convergence. We evaluate the runtime of BH-RGA for the inputs of different sizes with different thresholds $\gamma$. For each approximation level of the BH tree, we report the achieved accuracy and runtime, among others.

We take a frame from the SINTEL sleeping2 sequence [9] as a reference and its translated and rotated clone as a template. The translation roughly amounts to one-third of the point cloud extent in the $x$-direction, and the rotation is either $5^{\circ}$ or $24^{\circ}$ around all axes simultaneously. In total, there are seven versions of the point cloud obtained by subsampling, with the cardinalities of $c a .5 k, 12 k, 25 k, 50 k$, $105 k, 205 k$ and $446 k$ (the original resolution). For each combination of the resolution and template transformation, we test $\gamma \in[0.0625 ; 64.0]$. The runtime evaluation metrics

Table 3: Comparison of E-CPD [23] and BH-RGA.

| evaluated configuration |  |  | RMSE | $\sigma$ |
| :---: | :---: | :---: | :---: | :---: |
|  | [no priors | 0.08 | 0.013 | cases (success rate) |
| E-CPD [23], | 1 prior | 0.084 | 0.012 | $89(17 \%)$ ] |
|  | 2 priors | 0.076 | 0.014 | $376(75 \%)$ |
|  | 3 priors | 0.056 | 0.018 | $488(97 \%)$ |
|  | [no priors | $\mathbf{0 . 0 5}$ | $\mathbf{0 . 0 1 8}$ | $\mathbf{1 2 4}(\mathbf{2 5 \%})$ ] |
| BH-RGA (ours), | 1 prior | $\mathbf{0 . 0 4 3}$ | $\mathbf{0 . 0 1 3}$ | $\mathbf{1 9 9 ( 4 0 \% )}$ |
| anchor points | 2 priors | $\mathbf{0 . 0 1 1}$ | $\mathbf{0 . 0 1 7}$ | $191(38 \%)$ |
|  | 3 priors | $\mathbf{0 . 0 0 8 6}$ | $\mathbf{0 . 0 0 5}$ | $165(33 \%)$ |
| BH-RGA (ours), | 1 prior | $\mathbf{0 . 0 1 9 1}$ | $\mathbf{1 . 1 E - 4}$ | $\mathbf{4 3 5}(\mathbf{8 7 \%})$ |
| prior matches | 2 priors | $\mathbf{0 . 0 0 5 5}$ | $\mathbf{1 . 7 2 E - 5}$ | $\mathbf{5 0 0}(\mathbf{1 0 0 \%})$ |
|  | 3 priors | $\mathbf{0 . 0 0 0 1}$ | $\mathbf{4 . 5 E - 9}$ | $\mathbf{5 0 0}(\mathbf{1 0 0 \%})$ |

for $5^{\circ}$ rotations are summarised in Fig. 4, i.e., runtimes (full and per iteration), runtime ratio of the BH tree generation, RMSE and the number of residuals.

Several patterns can be noticed in the runtime statistics. When $\gamma$ is fixed, RMSE is similar for a different number of points. Conversely, with the increasing $\gamma$, the alignment accuracy increases for all resolutions, and smaller and smaller runtime fraction is spent on building BH tree while not being influenced much by the number of points. For $\gamma>2.0$, BH tree runtime ratio is $\leq 0.5 \%$. In comparison, the solver runtime ratio is always $\sim 80 \%$ suggesting that an up to fivefold acceleration with a Gauss-Newton solver on a GPU is possible. Next, as $\gamma$ increases, the number of iterations decreases and stabilises. For $\gamma \geq 6$, the number of iterations is $\leq 4$. The average number of residuals (number of point-to-cluster interactions) reaches $\approx 1.8 E 7$ for $\gamma=6.0$ and $446 k$ points. The graph metrics for the case of $24^{\circ}$ rotations coarsely follow the shapes of the graph metrics for the visualised case of $5^{\circ}$ rotations. Larger misalignments require larger $\gamma$ for successful registration and the overall runtime increases by the factor of two (note that in this experiment, we do not perform a translational pre-alignment). The periteration runtime increases by $\sim 40 \%$, and the number of iterations for $\gamma \geq 6$ increases by one-two.



Figure 5: BH-RGA is applicable to partially overlapping real-world data. Vestalin is aligned with the colour cue and three anchor points, and for pedestal, we use two anchor points. For the lidar data, we apply VMN helping to counterbalance varying point densities. For the Vestalin and the universe, the initialisation is shown on the left, and the alignments are shown on the right. For the pedestal and driving scenario, the initialisations are on the top, and the alignments are on the bottom.

Sec. 4.5) which initialises masses uniformly per unit volume. It can be well noticed on buildings and cars that the scans are accurately registered. In this experiment, no anchor points or further prior knowledge are used. Both scans contain $c a .120 k$ points, and the registration is performed in $\approx 90$ seconds on a CPU without subsampling or parallelisation, with $\gamma=1.5 \cdot 10^{4}$ (at the original point set scale).

## 6. Concluding Remarks

The main conclusion of this paper is that particle dynamics based RA can be accelerated using a BH tree and stabilised by altering laws of simulated physics. Our formulation significantly reduces the number of parameters in the gravitational alignment - by the factor of three - and eliminates the need for parameter tuning. Even for large point sets of dimensionality going beyond the extents feasible for other general-purpose RA methods, the proposed BH-RGA performs global multiply-linked updates. This contrasts with other known techniques which de facto only follow locally multiply-linked policies for the approximation of point interactions. As a result, BH-RGA is notably robust to noise and achieves state-of-the-art performance whenever existing techniques do. Furthermore, BH-RGA accepts such additional cues as prior matches, anchor points and even point intensities, and efficiently handles different point densities with VMN.

There are multiple avenues for further investigation of particle dynamics based RA. The most promising ones are multibody registration, parallelisation and adaption of BHRGA for CAD and automotive applications.

Our supplementary material contains more details on the $2^{D}$-tree generation and projective scene flow visualisation used in Fig. 4. Besides, we visualise statistics of resolved misalignments for clean-500 and N500-U100 experiments.

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[^1]:    ${ }^{\mathrm{i}}$ Reciprocal transform is one of the widely-used transforms in signal processing and applied mathematics [7]. It has several modifications such as reciprocal-wedge transform in image processing [49].
    ${ }^{\text {ii }}$ Next, we simplify the notation for masses and omit the superscript ${ }^{\text {new }}$

