# Electron-positron annihilation into two photons in an intense plane-wave field 

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#### Abstract

The process of electron-positron annihilation into two photons in the presence of an intense classical plane wave of an arbitrary shape is investigated analytically by employing light-cone quantization and by taking into account the effects of the plane wave exactly. We introduce a general description of second-order 2 -to- 2 scattering processes in a plane-wave background field, indicating the necessity of considering the localization of the colliding particles, and that is achieved by means of wave packets. We define a local cross section in the background field, which generalizes the vacuum cross section and which, though not being directly an observable, allows for a comparison between the results in the plane wave and in vacuum without relying on the shape of the incoming wave packets. Two possible cascade or two-step channels have been identified in the annihilation process and an alternative way of representing the two-step and one-step contributions via a "virtuality" integral has been found. Finally, we compute the total local cross section to leading order in the coupling between the electron-positron field and the quantized photon field, excluding the interference between the two leading-order diagrams arising from the exchange of the two final photons, and express it in a relatively compact form, which contains the dependence on the plane-wave field only via the dressed fermion momenta. In contrast to processes in a background field initiated by a single particle, the pair annihilation into two photons, in fact, also occurs in vacuum. Our result naturally embeds the vacuum part and reduces to the vacuum expression, known in the literature, in the case of a vanishing laser field.


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## I. INTRODUCTION

With the development of high-power laser technology the verification of the nonlinear-QED predictions for various phenomena in an intense background field of a macroscopic extent is becoming attainable in laboratory experiments [1-5]. Among QED processes in an intense laser field, two first-order ones, nonlinear Compton scattering $\left(e^{-} \Rightarrow\right.$ $e^{-} \gamma$ ) [6-18] and nonlinear Breit-Wheeler pair production $\left(\gamma \Rightarrow e^{-} e^{+}\right)$[8, 19-26] have been extensively investigated theoretically (see also the reviews $[2-4,27]$ ), where by a double-line arrow we highlight the fact that a process happens in a background field, which in general has to be taken into account nonperturbatively. Recently, nonlinear Compton scattering was also probed experimentally and signatures of quantum effects were observed [28, 29] (see [30] for a related experiment carried out in crystals). Moreover, these reactions are the only QED effects included in common implementations of the QED Particle-In-Cell (PIC) scheme [31-33], which is a standard tool for the numerical investigation of the interaction between a laser field of extreme intensity ( $\left.\gtrsim 10^{23} \mathrm{~W} / \mathrm{cm}^{2}\right)$ and matter, in particular, of the dynamics of the electron-positron plasma, produced in this interaction [34-43] (an electron-positron plasma interacting with a background field can also arise in a collision of a high-density electron beam with a target [44] and in some astrophysical scenarios [45-49]).

Other channels of the first-order processes, i.e., electron-positron annihilation into one photon $\left(e^{-} e^{+} \Rightarrow \gamma\right)$ and photon absorption $\left(e^{-} \gamma \Rightarrow e^{-}\right)$are sizable only in a small part of the phase space of the incoming particles $[8,50-52]$. Therefore, if electron-positron annihilation and photon absorption are to be also included into the consideration of the evolution of a many-particle system in an intense laser field, which may involve different geometries of particle collisions, it is necessary to assess the next-order processes, i.e., $e^{-} e^{+} \Rightarrow \gamma \gamma$ and $e^{-} \gamma \Rightarrow e^{-} \gamma$, respectively.

However, a complete evaluation of a tree-level second-order process in an external laser field is not straightforward. For instance, first theoretical calculations for trident process, i.e., seeded electron-positron pair production ( $e^{-} \Rightarrow$ $e^{-} e^{-} e^{+}$, were performed long ago [53, 54]. It was demonstrated that the total probability can be decomposed into a two-step contribution, which is related to the physical situation of the intermediate electron being real and which can be reconstructed as a combination of the corresponding nonlinear Compton and Breit-Wheeler probabilities, and a one-step contribution, for which the intermediate electron is virtual and which was computed in part. Later, first experiments on trident were also carried out $[55,56]$. But only recently, via a series of works, a full evaluation of

[^0]

Figure 1. The leading-order Feynman diagrams for electron-positron annihilation into two photons in a plane-wave field: (a) the direct diagram, and (b) the exchange diagram. Double lines represent dressed (Volkov) wave functions and propagators (see the main text for details).
trident process was presented for the constant-crossed and general plane-wave background field cases [57-62] (for an estimation, the one-step part of trident is sometimes taken into account with the use of the Weizsäcker-Williams approximation [63, 64], see also [59]). A result for double Compton scattering ( $e^{-} \Rightarrow e^{-} \gamma \gamma$ ) has been obtained in a similar fashion [65-69]. As to $2 \Rightarrow 2$ reactions, considerations existing in the literature are limited to very specific cases, like a monochromatic or an almost monochromatic laser field, the weak-field limit, a circular laser polarization, and/or so-called resonance processes (see, e.g., [70-74]).

Here, we consider electron-positron annihilation into two photons, with the two leading-order Feynman diagrams shown in Fig. 1. We present the first analytical results for a total cross section (in a sense explained below) of $e^{-} e^{+} \Rightarrow \gamma \gamma$ in a laser pulse represented as a classical plane-wave (or null) field of a general shape. We provide an exact expression for the contribution of the individual diagrams in Fig. 1, without taking into account the interference between them. Keeping possible applications of our result to many-body-evolution numerical codes in mind, we define the cross section in such a way that it has the meaning of a local quantity, and we also write it in terms of the dressed momenta of the colliding particles in the plane wave. Furthermore, we use the example of $e^{-} e^{+} \Rightarrow \gamma \gamma$ for establishing general features of the description of second-order 2-to-2 collision processes in a plane-wave background field.

In contrast to nonlinear trident pair production and nonlinear double Compton scattering, the reaction $e^{-} e^{+} \Rightarrow \gamma \gamma$ does occur already in vacuum. This may pose a technical problem, since the two parts (vacuum and field-dependent one) have different numbers of momentum conservation delta functions. Therefore, one might encounter a difficulty of dealing with the different number of volume factors and of comparing and combining the two parts. We show that it is possible to incorporate both into a single expression for the total (local) cross section, which, in the limit of a vanishing external field, reduces to the result, known in the literature for the vacuum case. Moreover, unlike the mentioned second-order processes initiated by a single particle, for $e^{-} e^{+} \Rightarrow \gamma \gamma$ the intermediate fermion becomes real not in one but in two different cases corresponding to the physical situations in which either the electron or the positron first emits a final photon before annihilating with the other particle into the second final photon. Using the Schwinger proper time representation for the electron propagator, we express the two-step and one-step contributions in a form, which has an advantage that it is concise and involves only integrals with limits independent of any variable. Another additional feature of 2-to- 2 processes in a plane wave is the importance of taking into account the fact of the localization of the incoming particles, which we carry out by introducing normalized wave packets. The underlying reason is that the collision of two particles in a plane wave is effectively a three-body collision and it is important at which moment each participant arrives at the collision region and if a collision region, as a microscopic region where all participants are for a certain time and significantly interact, does exist at all.

This paper is organized as follows. In Sec. II we introduce the formalism. In Sec. III we consider the annihilation into two photons of an electron and a positron, which are described by wave packets. We find out the approximations, that one needs to make in order to introduce a cross section, and provide a general expression for the cross section of the reaction $e^{-} e^{+} \Rightarrow \gamma \gamma$. In Sec. IV the one- and two-step contributions to the cross section are investigated. In Sec. V we elaborate on the evaluation of the integrals for the process under consideration. The final result is presented in Sec. VI and the limit of a vanishing background field is considered in Sec. VII. The conclusions are presented in Sec. VIII. Five appendices contain explanations of the notation and technical details.

Throughout the paper, Heaviside and natural units are used ( $\epsilon_{0}=\hbar=c=1$ ), m and $e<0$ denote the electron mass and charge, respectively, $\alpha=e^{2} /(4 \pi) \approx 1 / 137$ is the fine-structure constant.

## II. FORMALISM

The formalism, that we employ, combines light-cone quantization [75-78] and Furry picture [79, 80] (a detailed discussion of the formalism is provided in [81]). With the quantization on the light-cone, a plane-wave background and
particularly momentum conservation laws are naturally included into the calculations (see [60, 69] for an application of light-cone quantization to trident and double Compton scattering). Also, the light-cone representation of the electromagnetic interaction via three types of vertices (see Appendix A), or, equivalently, the representation of the electron propagator (and also of the photon one) as combination of noninstantaneous and instantaneous terms (this can be done within the instant-form quantization as well [82, 83], see also [66, 67, 84]) allows one to write the spinor prefactors via fermion dressed momenta (see below), and, as a consequence, the final expressions formally have no explicit dependence on the background field and asymptotic fermion momenta. In this respect, the obtained result is similar to the ones usually derived in vacuum, where the final expressions depend on the particle four-momenta in the form of Mandelstam variables [85].

The laser field is described classically by the field tensor $F^{\mu \nu}(\phi)$, which is a function of the scalar product $\phi=k_{0} x$, with $k_{0}^{\mu}$ being the characteristic wave four-vector of the field or, in the quantum language, the characteristic fourmomentum of a laser photon $\left(k_{0}^{2}=k_{0}^{\mu} k_{0 \mu}=0\right)$ and $x^{\mu}$ being a position four-vector. We assume that $F^{\mu \nu}(\phi)$ does not contain a constant-term (zero-frequency) contribution, but only a $\phi$-dependent part, which vanishes asymptotically (as $\phi \rightarrow \pm \infty)$. Then the most general form of $F^{\mu \nu}(\phi)$ is given by

$$
\begin{equation*}
F^{\mu \nu}(\phi)=\sum_{i=1,2} f_{i}^{\mu \nu} \psi_{i}^{\prime}(\phi) \tag{1}
\end{equation*}
$$

where $f_{i}^{\mu \nu}=k_{0}^{\mu} a_{i}^{\nu}-k_{0}^{\nu} a_{i}^{\mu}$, the four-vectors $a_{i}^{\mu}$ define the amplitude of the field in two polarization directions $\left(k_{0} a_{i}=0\right.$, $a_{1} a_{2}=0$ ), and the functions $\psi_{i}^{\prime}(\phi)=\mathrm{d} \psi_{i}(\phi) /(\mathrm{d} \phi)$ characterize its shape $\left[\left|\psi_{i}^{\prime}(\phi)\right| \lesssim 1, \psi_{i}^{\prime}( \pm \infty)=0\right]$. In the following, the indices $i, j, k, \ldots$ always take the values 1,2 .

The light-cone coordinates are defined via specifying the light-cone basis $\left\{\eta^{\mu}, \bar{\eta}^{\mu}, e_{1}^{\mu}, e_{2}^{\mu}\right\}$ (see Appendix A for details). Below, we employ the canonical light-cone basis, which is defined as [86]

$$
\begin{equation*}
\eta^{\mu}=\frac{k_{0}^{\mu}}{m}, \quad \bar{\eta}^{\mu}=\frac{q^{\mu}}{q^{+}}-\frac{q^{2} \eta^{\mu}}{2 q^{+2}}, \quad e_{1}^{\mu}=\frac{q_{\lambda} f_{1}^{\lambda \mu}}{m q^{+} \sqrt{-a_{1}^{2}}}, \quad e_{2}^{\mu}=\frac{q_{\lambda} f_{2}^{\lambda \mu}}{m q^{+} \sqrt{-a_{2}^{2}}}, \tag{2}
\end{equation*}
$$

where the four-vector $q^{\mu}$ is such that $q^{+} \neq 0$. The calculations are greatly simplified if one chooses

$$
\begin{equation*}
q^{\mu}=p_{1}^{\mu}+p_{2}^{\mu} \tag{3}
\end{equation*}
$$

which implies $\boldsymbol{p}_{2}^{\perp}+\boldsymbol{p}_{1}^{\perp}=\boldsymbol{k}_{2}^{\perp}+\boldsymbol{k}_{1}^{\perp}=\mathbf{0}$. Here, $p_{1}^{\mu}$ and $p_{2}^{\mu}$ are the asymptotic four-momenta of the incoming electron and positron outside the plane wave, respectively, whereas $k_{1}^{\mu}$ and $k_{2}^{\mu}$ are the four-momenta of the final photons (see Fig. 1 and note that in the following we employ wave packets for the electron and positron and therefore $p_{1}^{\mu}$ and $p_{2}^{\mu}$ will be ultimately identified with the central four-momenta).

Since $\eta^{\mu}=k_{0}^{\mu} / m$, the laser phase is $\phi=m x^{+}$and the field $F^{\mu \nu}(\phi)$ depends only on the light-cone time. With the adoption of the light-cone gauge $A^{+}(x)=0$, the four-vector potential for $F^{\mu \nu}(\phi)$ reads

$$
\begin{equation*}
A^{\mu}(\phi)=\sum_{i} a_{i}^{\mu} \psi_{i}(\phi) \tag{4}
\end{equation*}
$$

In the following, we assume $A^{\mu}(-\infty)=0$, which implies $\psi_{i}(-\infty)=0$ [together with the fact of the absence of a constant-term contribution in $F^{\mu \nu}(\phi)$ this implies that also $A^{\mu}(\infty)=0$ and therefore $\psi_{i}(\infty)=0$.

The solution of the Dirac equation with the classical field (4) is the Volkov solution [87]. We write the positive-energy one in the form [84]

$$
\begin{equation*}
\psi_{p \sigma}(x)=K_{p}(\phi) \frac{u_{p \sigma}}{\sqrt{2 p^{+}}} \mathrm{e}^{i S_{p}(x)} \tag{5}
\end{equation*}
$$

with

$$
\begin{equation*}
K_{p}(\phi)=\left[\gamma \pi_{p}(\phi)+m\right] \frac{\gamma^{+}}{2 p^{+}}, \quad S_{p}(x)=-p x-\mathcal{S}_{p}(\phi), \quad \mathcal{S}_{p}(\phi)=\int_{-\infty}^{\phi} \mathrm{d} \beta\left(\frac{e p A(\beta)}{m p^{+}}-\frac{e^{2} A^{2}(\beta)}{2 m p^{+}}\right) \tag{6}
\end{equation*}
$$

and the negative-energy one in an analogous way (see Appendix A). Note that the phase $S_{p}(x)$ is the classical action of an electron in the plane wave and that the dressed momentum $\pi_{p}^{\mu}(\phi)=-\partial^{\mu} S_{p}(x)-e A^{\mu}(\phi)$ is the corresponding solution of the Lorentz equation. It is given by

$$
\begin{equation*}
\pi_{p}^{\mu}(\phi)=p^{\mu}-e A^{\mu}(\phi)+\eta^{\mu}\left(\frac{e p A(\phi)}{p^{+}}-\frac{e^{2} A^{2}(\phi)}{2 p^{+}}\right) \tag{7}
\end{equation*}
$$

such that $\pi_{p}^{2}(\phi)=p^{2}, \pi_{p}^{+}(\phi)=p^{+}$. The free Dirac bispinor $u_{p \sigma}$ is normalized such that $\bar{u}_{p \sigma} u_{p \sigma^{\prime}}=2 m \delta_{\sigma \sigma^{\prime}}, \bar{u}_{p \sigma} \gamma^{\mu} u_{p \sigma^{\prime}}=$ $2 p^{\mu} \delta_{\sigma \sigma^{\prime}}, \sum_{\sigma} u_{p \sigma} \bar{u}_{p \sigma}=\gamma p+m$ [85].

The fermion field $\psi(x)$ is expanded in the basis set of the Volkov wave functions (5) (and analogous ones for negative-energy states) and, as a consequence, in all diagrams free fermion lines are replaced with the corresponding Volkov ones [79, 80] (details on the quantization are given in Appendix A).

Though in electrodynamics, quantized on the light cone, there are three types of vertices, for our purposes it is convenient to combine them in the form of propagators. Then we have only the usual three-point QED vertex, but each electron and photon Feynman propagator consists of two terms [82, 83] (see also [66, 67, 84]), in particular, for the electron propagator $G\left(x_{2}, x_{1}\right)$ we have: $G\left(x_{2}, x_{1}\right)=G^{(\mathrm{ni})}\left(x_{2}, x_{1}\right)+G^{(\mathrm{in})}\left(x_{2}, x_{1}\right)$, with $G^{(\mathrm{ni})}\left(x_{2}, x_{1}\right)$ being a noninstantaneous (propagating) part

$$
\begin{equation*}
G^{(\mathrm{ni})}\left(x_{2}, x_{1}\right)=\int \frac{\mathrm{d}^{4} p}{(2 \pi)^{4}} \mathrm{e}^{-i p\left(x_{2}-x_{1}\right)-i \mathcal{S}_{p}\left(\phi_{2}, \phi_{1}\right)} K_{p}\left(\phi_{2}\right) \frac{\gamma \tilde{p}+m}{p^{2}-m^{2}+i \epsilon} \bar{K}_{p}\left(\phi_{1}\right), \tag{8}
\end{equation*}
$$

and with $G^{(\mathrm{in})}\left(x_{2}, x_{1}\right)$ being an instantaneous part

$$
\begin{equation*}
G^{(\mathrm{in})}\left(x_{2}, x_{1}\right)=\int \frac{\mathrm{d}^{4} p}{(2 \pi)^{4}} \mathrm{e}^{-i p\left(x_{2}-x_{1}\right)-i \mathcal{S}_{p}\left(\phi_{2}, \phi_{1}\right)} \frac{\gamma^{+}}{2 p^{+}} . \tag{9}
\end{equation*}
$$

Here, $\mathcal{S}_{p}\left(\phi_{2}, \phi_{1}\right)=\mathcal{S}_{p}\left(\phi_{2}\right)-\mathcal{S}_{p}\left(\phi_{1}\right)$ and

$$
\begin{equation*}
\tilde{p}^{\mu}=\left(p^{+}, \frac{\boldsymbol{p}^{\perp 2}+m^{2}}{2 p^{+}}, \boldsymbol{p}^{\perp}\right) \tag{10}
\end{equation*}
$$

such that $\tilde{p}^{2}=m^{2}$.
Below, we will employ the classical intensity parameters [3, 8]

$$
\begin{equation*}
\xi_{i}=\frac{|e| \sqrt{-a_{i}^{2}}}{m} \tag{11}
\end{equation*}
$$

We also introduce $\xi=\sqrt{\xi_{1}^{2}+\xi_{2}^{2}}$. Other parameters characterizing the scattering process are the quantum nonlinearity parameters, which are defined as $\chi_{i j}=p_{i}^{+} \xi_{j} / m$ for the fermions, and analogously for the photons [3, 8]. Note that by considering the interaction with the quantized photon field to leading order, we implicitly assume that the quantum nonlinearity parameters are much smaller than $1 / \alpha^{3 / 2} \approx 1600$, such that this interaction can be treated perturbatively. This assumption is reasonable for current and near-future laser-based setups (for discussions of the fully nonperturbative regime, see, e.g., [88-93]).

For a process with two incoming particles, the classical intensity parameters and the quantum nonlinearity parameters do not exhaust the list of quantities, that are necessary for describing the scattering (even when considering an observable obtained by averaging/summing over the discrete quantum numbers and by integrating over the final momenta). We introduce the additional parameters $t_{i}(\phi)$, which are given by [81]

$$
\begin{equation*}
t_{i}(\phi)=\frac{|e| \pi_{1 \mu}(\phi) f_{i}^{\mu \nu} \pi_{-2 \nu}(\phi)}{\xi_{i} m^{3}\left(p_{1}^{+}+p_{2}^{+}\right)} \tag{12}
\end{equation*}
$$

where $\pi_{1}^{\mu}(\phi)=\pi_{e, p_{1}}^{\mu}(\phi)$ and $-\pi_{-2}^{\mu}(\phi)=\pi_{-e, p_{2}}^{\mu}(\phi)$ are the dressed four-momenta of the electron and the positron, respectively. The asymptotic values of $t_{i}(\phi)$ are denoted as $t_{i}$, they have been employed in the literature before [8].

The parameters $t_{i}(\phi)$ have a particularly clear physical interpretation if we use the canonical light-cone basis (2) with $q^{\mu}$ from Eq. (3). With this choice, we have $\boldsymbol{\pi}_{e, p_{1}}^{\perp}(\phi)+\boldsymbol{\pi}_{-e, p_{2}}^{\perp}(\phi)=\boldsymbol{p}_{1}^{\perp}+\boldsymbol{p}_{2}^{\perp}=\mathbf{0}$ and $t_{i}(\phi)=\pi_{1}^{i}(\phi) / m$, i.e., $t_{1}(\phi)$ and $t_{2}(\phi)$ correspond to the transverse dressed momentum components of the incoming particles (with respect to the laser-pulse propagation direction).

## III. CROSS SECTION

Strictly speaking, a collision of an electron and a positron in the presence of a finite-duration laser pulse is a three-body process and the result of the collision depends on the time of arrival of each participant at the collision region and on whether a collision region, as a microscopic region where all participants are for a certain time and significantly interact there, can be defined at all. Thus, in the most general setup, one cannot rely on the description of the incoming particles via monochromatic plane waves, since they have an infinite temporal and spatial extent.

Therefore, in order to consistently describe the reaction $e^{-} e^{+} \Rightarrow \gamma \gamma$, we represent the electron and the positron as normalized wave packets with central on-shell four-momenta $p_{1}^{\mu}$ and $p_{2}^{\mu}$, respectively. A positive-energy wave packet $\Psi_{p}(x)$ with the central four-momentum $p^{\mu}$ is constructed according to

$$
\begin{equation*}
\Psi_{p}(x)=\int \frac{\tilde{\mathrm{d}}^{3} q}{(2 \pi)^{3}} f_{p}(q) \psi_{q}(x) \tag{13}
\end{equation*}
$$

where $f_{p}(q)$ is the momentum distribution density and $\psi_{q}(x)$ is the positive-energy Volkov state (5) with fourmomentum $q^{\mu}$ (for the definition of $\tilde{\mathrm{d}}^{3} q$ see Appendix A). Note that Volkov states are on-shell such that $q^{-}=$ $\left(\boldsymbol{q}^{\perp 2}+m^{2}\right) /\left(2 q^{+}\right)$, i.e., $f_{p}(q)$ depends on $\boldsymbol{q}^{\perp}$ and $q^{+}$only, but for simplicity, we write $f_{p}(q)$ as a function of $q^{\mu}$. The fact that the function $f_{p}(q)$ is centered around the on-shell four-momentum $p^{\mu}$ has to be intended analogously. Correspondingly, one can also define negative-energy wave packets. We refer to Appendix B for further details about the general properties of the wave packets $\Psi_{p}(x)$.

The polarization degrees of freedom of both incoming (outgoing) particles are averaged (summed) in the final expressions, with the assumption of the initial states being unpolarized, and therefore, for notational brevity, we suppress the subscripts for these degrees of freedom.

As already mentioned, the final photon four-momenta are $k_{1}^{\mu}$ and $k_{2}^{\mu}\left(k_{1}^{2}=k_{2}^{2}=0\right)$. The $S$-matrix element corresponding to the diagrams in Fig. 1 can be written as

$$
\begin{equation*}
S_{f i}=i \int \frac{\tilde{\mathrm{~d}}^{3} q_{2}}{(2 \pi)^{3}} \frac{\tilde{\mathrm{~d}}^{3} q_{1}}{(2 \pi)^{3}} f_{2}^{*}\left(q_{2}\right) f_{1}\left(q_{1}\right) \int \mathrm{d}^{4} x_{2} \mathrm{~d}^{4} x_{1} \tilde{T}\left(x_{2}, x_{1}, q_{2}, q_{1}\right) \tag{14}
\end{equation*}
$$

where $f_{1}(q)$ and $f_{2}^{*}(q)$ are the electron and positron wave-packet momentum distributions, respectively, and

$$
\begin{align*}
& \tilde{T}\left(x_{2}, x_{1}, q_{2}, q_{1}\right)=\int \frac{\mathrm{d}^{4} p_{3}}{(2 \pi)^{4}} \frac{M^{\text {direct }}\left(\phi_{2}, \phi_{1}, q_{2}, q_{1}\right)}{4 \sqrt{k_{2}^{+} k_{1}^{+} q_{2}^{+} q_{1}^{+}}} \exp \left[i\left(k_{2}-p_{3}-q_{2}\right) x_{2}+i\left(k_{1}+p_{3}-q_{1}\right) x_{1}\right. \\
&\left.-i \mathcal{S}_{p_{3}}\left(\phi_{2}, \phi_{1}\right)-i \mathcal{S}_{q_{1}}\left(\phi_{1}\right)+i \mathcal{S}_{-q_{2}}\left(\phi_{2}\right)\right] \\
&+\left\{\gamma_{1} \leftrightarrow \gamma_{2}\right\} \tag{15}
\end{align*}
$$

with

$$
\begin{equation*}
M^{\text {direct }}\left(\phi_{2}, \phi_{1}, q_{2}, q_{1}\right)=-e^{2} \bar{v}_{q_{2}}\left[K_{-q_{2} p_{3}}^{\mu}\left(\phi_{2}\right) \frac{\gamma \tilde{p}_{3}+m}{p_{3}^{2}-m^{2}+i \epsilon} K_{p_{3} q_{1}}^{\nu}\left(\phi_{1}\right)+\frac{K_{-q_{2} q_{1}}^{\mu \nu}\left(\phi_{2}, \phi_{1}\right)}{2 p_{3}^{+}}\right] u_{q_{1}} \epsilon_{2 \mu}^{*} \epsilon_{1 \nu}^{*} . \tag{16}
\end{equation*}
$$

Here and below, $\phi_{i}=k_{0} x_{i}=m x_{i}^{+}$and the term $\left\{\gamma_{1} \leftrightarrow \gamma_{2}\right\}$ corresponds to the exchange diagram with the photon quantum numbers swapped (see Fig. 1b). Also, the functions $K_{p p^{\prime}}^{\mu}(\phi)$ and $K_{p p^{\prime}}^{\mu \nu}\left(\phi, \phi^{\prime}\right)$ are given by

$$
\begin{equation*}
K_{p p^{\prime}}^{\mu}(\phi)=\bar{K}_{p}(\phi) \gamma^{\mu} K_{p^{\prime}}(\phi), \quad K_{p p^{\prime}}^{\mu \nu}\left(\phi, \phi^{\prime}\right)=\bar{K}_{p}(\phi) \gamma^{\mu} \gamma^{+} \gamma^{\nu} K_{p^{\prime}}\left(\phi^{\prime}\right) \tag{17}
\end{equation*}
$$

where $\bar{K}_{p}(\phi)=\gamma^{0} K_{p}^{\dagger}(\phi) \gamma^{0}$ and a dagger denotes the Hermitian conjugate.
In the following and analogously to the vacuum case (see, e.g., $[94,95]$ ), we assume the momentum distributions of the electron and the positron being sufficiently narrowly peaked around the central four-momenta and the detectors not being sensitive enough to resolve the final momenta within the widths of such distributions, such that we can in particular replace the four-momenta $q_{i}^{\mu}$ with the central ones in relatively slowly varying functions, i.e.,

$$
\begin{equation*}
M^{\text {direct }}\left(\phi_{2}, \phi_{1}, q_{2}, q_{1}\right) / \sqrt{q_{2}^{+} q_{1}^{+}} \approx M^{\text {direct }}\left(\phi_{2}, \phi_{1}\right) / \sqrt{p_{2}^{+} p_{1}^{+}} \tag{18}
\end{equation*}
$$

where $M^{\text {direct }}\left(\phi_{2}, \phi_{1}\right)=M^{\text {direct }}\left(\phi_{2}, \phi_{1}, p_{2}, p_{1}\right)$, and we do the same for the exchange term as well.
The total probability, obtained as the modulus squared of Eq. (14), averaged over the initial polarization states and summed over all final polarization and momentum states, can be written as

$$
\begin{align*}
W & \approx \frac{1}{4} \sum_{\mathrm{qn}}\left|\int \mathrm{~d}^{4} x_{2} \mathrm{~d}^{4} x_{1} F_{2}^{*}\left(x_{2}\right) F_{1}\left(x_{1}\right) \tilde{T}\left(x_{2}, x_{1}, p_{2}, p_{1}\right)\right|^{2} \\
& =\frac{1}{4} \sum_{\text {qn }} \int \mathrm{d}^{4} x_{2} \mathrm{~d}^{4} x_{1} \mathrm{~d}^{4} x_{2}^{\prime} \mathrm{d}^{4} x_{1}^{\prime} F_{2}^{*}\left(x_{2}\right) F_{2}\left(x_{2}^{\prime}\right) F_{1}\left(x_{1}\right) F_{1}^{*}\left(x_{1}^{\prime}\right) \tilde{T}\left(x_{2}, x_{1}, p_{2}, p_{1}\right) \tilde{T}^{*}\left(x_{2}^{\prime}, x_{1}^{\prime}, p_{2}, p_{1}\right), \tag{19}
\end{align*}
$$

where the abbreviation "qn" indicates that the sum/integral is taken over the discrete quantum numbers of the initial and final particles and the momenta of the final photons. Also, in Eq. (19) we have introduced the electron and positron wave packet amplitudes $F_{1}\left(x_{1}\right)$ and $F_{2}^{*}\left(x_{2}\right)$ in configuration space, which are defined analogously to the vacuum case [94], e.g., for an electron we have

$$
\begin{equation*}
F_{p}(x)=\int \frac{\tilde{\mathrm{d}}^{3} q}{(2 \pi)^{3}} f_{p}(q) \exp \left[-i(q-p) x-i \mathcal{S}_{q}(\phi)+i \mathcal{S}_{p}(\phi)\right] \tag{20}
\end{equation*}
$$

for a given four-momentum distribution $f_{p}(q)$. The wave packet in configuration space is given by

$$
\begin{equation*}
f_{p}(x)=\int \frac{\tilde{\mathrm{d}}^{3} q}{(2 \pi)^{3}} f_{p}(q) \exp \left[-i q x-i \mathcal{S}_{q}(\phi)\right]=F_{p}(x) \exp \left[-i p x-i \mathcal{S}_{p}(\phi)\right] \tag{21}
\end{equation*}
$$

(for a positron, the expressions are analogous). Note that $\left|f_{p}(x)\right|^{2}=\left|F_{p}(x)\right|^{2}$ is the (time-dependent) particle density. The properties of the particle density $\left|F_{p}(x)\right|^{2}$ are discussed in Appendix B and we only recall here that for a narrow wave packet, on the condition that also $\left|f_{p}(x)\right|^{2}$ is sufficiently peaked in configuration space, the center of the distribution $\left|f_{p}(x)\right|^{2}$ follows the classical trajectory of an electron in a given plane wave (see Appendix B for further details).

In principle, Eq. (19) is the expression one needs to employ in order to evaluate the total probability of the process under consideration. However, depending on the widths of the wave packets and on the formation lengths of the integrals in the space-time variables, one can achieve further simplifications.

The first step is to assume that the wave packets are sufficiently narrow (in momentum space), that on the formation length of a single-vertex process (essentially, a process obtained by cutting the propagator line, see Fig. 1) one can neglect the interference among the wave packets, i.e.,

$$
\begin{equation*}
F_{1}\left(x_{1}\right) F_{1}^{*}\left(x_{1}^{\prime}\right)=F_{1}\left(X_{1}-\delta_{1} / 2\right) F_{1}^{*}\left(X_{1}+\delta_{1} / 2\right) \approx\left|f_{1}\left(X_{1}\right)\right|^{2} \tag{22}
\end{equation*}
$$

and analogously for the positron wave packet amplitudes, where

$$
\begin{equation*}
X_{i}^{\mu}=\left(x_{i}^{\mu}+x_{i}^{\prime \mu}\right) / 2, \quad \delta_{1}^{\mu}=x_{1}^{\prime \mu}-x_{1}^{\mu}, \quad \delta_{2}^{\mu}=x_{2}^{\mu}-x_{2}^{\prime \mu} \tag{23}
\end{equation*}
$$

We point out that the assertion in Eq. (22) [and the corresponding one for $F_{2}\left(x_{2}^{\prime}\right) F_{2}^{*}\left(x_{2}\right)$ ] is a more complicated statement than in vacuum, in the sense that the typical scale of $\delta_{1}^{\mu}$ (and of $\delta_{2}^{\mu}$ for the positron) depends in general on the form and on the intensity of a considered background field, and Eq. (22) results from an interplay between the scale introduced by the field and the scale of the wave packets (details are given in Appendix C).

Under the approximation (22) and an analogous for the positron, the total probability (19) reads

$$
\begin{equation*}
W \approx \int \mathrm{~d}^{4} X_{2} \mathrm{~d}^{4} X_{1}\left|f_{2}\left(X_{2}\right)\right|^{2}\left|f_{1}\left(X_{1}\right)\right|^{2} W\left(X_{2}, X_{1}\right) \tag{24}
\end{equation*}
$$

with the two-point probability distribution

$$
\begin{equation*}
W\left(X_{2}, X_{1}\right)=\frac{1}{4} \sum_{\text {qn }} \int \mathrm{d}^{4} \delta_{2} \mathrm{~d}^{4} \delta_{1} \tilde{T}\left(x_{2}, x_{1}, p_{2}, p_{1}\right) \tilde{T}^{*}\left(x_{2}^{\prime}, x_{1}^{\prime}, p_{2}, p_{1}\right) . \tag{25}
\end{equation*}
$$

An additional simplification is attained under the assumption, that on a typical distance between $X_{1}^{\mu}$ and $X_{2}^{\mu}$ (in essence, on the typical distance between the two single-vertex processes, see Fig. 1) the wave packets do not change significantly, i.e.,

$$
\begin{equation*}
\left|f_{2}\left(X_{2}\right)\right|^{2}\left|f_{1}\left(X_{1}\right)\right|^{2}=\left|f_{2}(x+\delta / 2)\right|^{2}\left|f_{1}(x-\delta / 2)\right|^{2} \approx\left|f_{2}(x)\right|^{2}\left|f_{1}(x)\right|^{2} \tag{26}
\end{equation*}
$$

where

$$
\begin{equation*}
x^{\mu}=\left(X_{2}^{\mu}+X_{1}^{\mu}\right) / 2, \quad \delta^{\mu}=X_{2}^{\mu}-X_{1}^{\mu} \tag{27}
\end{equation*}
$$

Then Eq. (24) transforms into

$$
\begin{equation*}
W \approx \int \mathrm{~d}^{4} x\left|f_{2}(x)\right|^{2}\left|f_{1}(x)\right|^{2} W(x) \tag{28}
\end{equation*}
$$

where

$$
\begin{equation*}
W(x)=W(\phi)=\frac{1}{4} \sum_{\mathrm{qn}} \int \mathrm{~d}^{4} \delta \mathrm{~d}^{4} \delta_{2} \mathrm{~d}^{4} \delta_{1} \tilde{T}\left(x_{2}, x_{1}, p_{2}, p_{1}\right) \tilde{T}^{*}\left(x_{2}^{\prime}, x_{1}^{\prime}, p_{2}, p_{1}\right) . \tag{29}
\end{equation*}
$$

Eq. (28) is the approximation that is commonly used for the description of scattering in vacuum and that allows to define a cross section, a quantity, which characterizes the process itself without relying on the precise shape of the wave packets [94, 95]. We stress that in a background field the assumption (26) can be restrictive as the intermediate particle may become real and hence $\delta^{\mu}$ can have a macroscopic scale, i.e., of the order of the extension of the background field (see Appendix C for details).

Now, it is worth pointing out an additional difference with the vacuum case. In the latter case, in fact, the quantity $W(x)$ is independent on the coordinates and therefore non-negative [94, 95]. In contrast to this, the quantity $W(\phi)$ here explicitly depends on the light-cone time (via $\phi$ ) and it can be negative for some values of $\phi$. Thus, generally speaking, the quantity

$$
\begin{equation*}
w(x)=\left|f_{2}(x)\right|^{2}\left|f_{1}(x)\right|^{2} W(\phi) \tag{30}
\end{equation*}
$$

cannot be interpreted as a probability per unit time and unit volume. However, it can be seen as a quantity, which generalizes this probability and which contains interference effects among contributions from different points of the particles trajectory in the plane wave, and therefore may become negative. This is somewhat similar to the relation between a classical phase-space distribution and the Wigner distribution, with the latter generalizing the former and, indeed, being also potentially negative [96].

Furthermore, we can define a generalized (local) cross section, which, though not being directly an observable quantity, since it can become negative, is a useful theoretical tool for investigating the influence of the external field on the scattering process. We follow the approach in the instant-form quantization in vacuum, where the cross section is obtained from the probability per unit time and unit volume by dividing it by the factor $\left|g_{2}(x)\right|^{2}\left|g_{1}(x)\right|^{2} I /\left(p_{2}^{0} p_{1}^{0}\right)$, where $I=\sqrt{\left(p_{2} p_{1}\right)^{2}-m^{4}}$ and $g_{i}(x)$ are wave packets in the instant form [85, 95]. Then in our case we can analogously introduce the local cross section as

$$
\begin{equation*}
\sigma(\phi)=\frac{p_{2}^{+} p_{1}^{+}}{\left|f_{2}(x)\right|^{2}\left|f_{1}(x)\right|^{2} I(\phi)} w(x)=\frac{p_{2}^{+} p_{1}^{+}}{I(\phi)} W(\phi), \tag{31}
\end{equation*}
$$

where the invariant $I(\phi)$ reads

$$
\begin{equation*}
I(\phi)=\sqrt{\left[\pi_{-e, p_{2}}(\phi) \pi_{e, p_{1}}(\phi)\right]^{2}-m^{4}} \tag{32}
\end{equation*}
$$

Below, we explicitly verify (except for the interference term, as has been pointed out in the introduction) that in the absence of the background field the cross section (31) reduces to the one, known for the vacuum case in the instant-form quantization. One should also keep in mind that the choice of the invariant $I(\phi)$ implies that the cross section is normalized to the flux coming into the point $x^{\mu}$ inside the laser field and in this sense is a local quantity. This can be useful, for instance, in the analysis of the importance of the studied process in the development of QED cascades where the colliding particles are produced inside the field. However, if one would like to consider a beambeam collision experiment in the presence of a laser field, then the use of the vacuum counterpart $I$ in place of $I(\phi)$ could be more convenient. The total probability $W$ is of course independent of this choice. Indeed, we emphasize that in order to obtain a physically observable quantity, one has ultimately to rely on the probability $W$ in Eq. (28) or, of course, on the corresponding differential probabilities.

With the use of Eqs. (15) and (16), we obtain for the cross section:

$$
\begin{equation*}
\sigma(\phi)=\int_{0}^{p_{2}^{+}+p_{1}^{+}} \frac{\mathrm{d} k_{1}^{+}}{2 \pi} \int \frac{\mathrm{~d}^{2} k_{1}^{\perp}}{(2 \pi)^{2}} \frac{1}{32 k_{2}^{+} k_{1}^{+} I(\phi)} \int \mathrm{d} \delta^{+} \mathrm{d} \delta_{2}^{+} \mathrm{d} \delta_{1}^{+} \frac{1}{4} \sum_{\sigma_{i}, \lambda_{i}} \tilde{M}\left(\phi_{2}, \phi_{1}, p_{2}, p_{1}\right) \tilde{M}^{*}\left(\phi_{2}^{\prime}, \phi_{1}^{\prime}, p_{2}, p_{1}\right), \tag{33}
\end{equation*}
$$

where $k_{2}^{(+, \perp)}=\left(p_{2}+p_{1}-k_{1}\right)^{(+, \perp)}$, $\sigma_{i}$ and $\lambda_{i}$ denote the polarization states of the incoming and outgoing particles, respectively, and we have divided the result by 2 , in order to adjust it for the double counting of the final states of the two identical particles. The reduced matrix element $\tilde{M}\left(\phi_{2}, \phi_{1}, p_{2}, p_{1}\right)$ is given by

$$
\begin{align*}
\tilde{M}\left(\phi_{2}, \phi_{1}, p_{2}, p_{1}\right)=\int \frac{\mathrm{d} p_{3}^{-}}{2 \pi} M^{\text {direct }}\left(\phi_{2}, \phi_{1}\right) \exp [ & i\left(k_{2}^{-}-p_{3}^{-}-p_{2}^{-}\right) x_{2}^{+}+i\left(k_{1}^{-}+p_{3}^{-}-p_{1}^{-}\right) x_{1}^{+} \\
& \left.-i \mathcal{S}_{p_{3}}\left(\phi_{2}, \phi_{1}\right)-i \mathcal{S}_{p_{1}}\left(\phi_{1}\right)+i \mathcal{S}_{-p_{2}}\left(\phi_{2}\right)\right] \\
& +\left\{\gamma_{1} \leftrightarrow \gamma_{2}\right\} \tag{34}
\end{align*}
$$

where $p_{3}^{(+, \perp)}=\left(p_{1}-k_{1}\right)^{(+, \perp)}=\left(k_{2}-p_{1}\right)^{(+, \perp)}$.
The quantity $\tilde{M}\left(\phi_{2}, \phi_{1}, p_{2}, p_{1}\right)$ in Eq. (34) contains four distinct terms because $M^{\text {direct }}\left(\phi_{2}, \phi_{1}\right)$ alone consists of a noninstantaneous and an instantaneous contributions. Taking the modulus squared yields 16 terms. However, only 8 of them are different after we sum over the states of the final photons, i.e.,

$$
\begin{equation*}
\sigma(\phi)=\sigma^{\mathrm{dd}}(\phi)+\sigma^{\mathrm{ee}}(\phi)+\sigma^{\mathrm{de}}(\phi)+\sigma^{\mathrm{ed}}(\phi)=2 \sigma^{\mathrm{dd}}(\phi)+2 \sigma^{\mathrm{de}}(\phi) \tag{35}
\end{equation*}
$$

where $\sigma^{\mathrm{dd}}(\phi)$ is the contribution, arising from squaring the amplitude for the direct diagram (see Fig. 1a), and can be written as

$$
\begin{equation*}
\sigma^{\mathrm{dd}}(\phi)=\sigma^{\mathrm{nndd}}(\phi)+\sigma^{\mathrm{nidd}}(\phi)+\sigma^{\mathrm{indd}}(\phi)+\sigma^{\mathrm{iidd}}(\phi) \tag{36}
\end{equation*}
$$

with $\sigma^{\text {nndd }}(\phi)$ originating from squaring the noninstantaneous direct term, $\sigma^{\text {nidd }}(\phi)$ from the product of the noninstantaneous and complex-conjugated instantaneous direct terms, etc. The other contributions in Eq. (35) can be written down analogously. In the following, we only consider $\sigma^{\mathrm{dd}}(\phi)$. We note that for the differential quantities the interference terms 'de' and 'ed' lead to an enhancement of the cross section by a factor of two in the case of the final photons being in the same state. On the other hand, at least in an ultrarelativistic setup, the available phase space is typically so large that one might expect that the integrated interference term $\sigma^{\text {de }}(\phi)$ should give a negligible contribution. Moreover, in the vacuum case, the interference contribution for the total cross section is relatively large only for mildly relativistic collisions [85]. If we assume a similar behavior in our case, then we should expect that the term $\sigma^{\text {de }}(\phi)$ can be nonnegligible only for some $\sqrt{s(\phi)} \sim m$, where the invariant mass squared $s(\phi)$ in the field is defined as

$$
\begin{equation*}
s(\phi)=\left[\pi_{-e, p_{2}}(\phi)+\pi_{e, p_{1}}(\phi)\right]^{2}=\frac{m^{2}\left(p_{2}^{+}+p_{1}^{+}\right)^{2}}{p_{2}^{+} p_{1}^{+}}\left[1+\boldsymbol{t}^{\perp 2}(\phi)\right] \tag{37}
\end{equation*}
$$

with $\boldsymbol{t}^{\perp 2}(\phi)=t_{1}^{2}(\phi)+t_{2}^{2}(\phi)$. It follows that if $p_{1}^{+} \sim p_{2}^{+}$and $\boldsymbol{t}^{\perp 2}(\phi) \sim 1$, the interference term might provide a somewhat sizable contribution. However, in the highly nonlinear regime, i.e., in the regime of $\xi \gg 1$, which we are interested in here, the function $\boldsymbol{t}^{\perp 2}(\phi)$ is rapidly oscillating and the condition $\boldsymbol{t}^{\perp 2}(\phi) \sim 1$ is fulfilled only in some narrow ranges $\sim 1 / \xi \ll 1$ of $\phi$ 's. Therefore, if one considers dynamics over one/several laser period(s), one might expect that on average the term $\sigma^{\mathrm{de}}(\phi)$ can be neglected.

Summing over the final photon polarizations results in the replacement

$$
\begin{equation*}
\epsilon_{i}^{\mu} \epsilon_{i}^{* \nu} \rightarrow-g^{\mu \nu} \tag{38}
\end{equation*}
$$

(we discard the terms proportional to $k_{i}^{\mu}$ and $k_{i}^{\nu}$ due to the Ward identity).
Averaging over the polarization states of the initial particles results in the replacements [85]

$$
\begin{equation*}
u_{1} \bar{u}_{1} \rightarrow \rho_{1}, \quad v_{2} \bar{v}_{2} \rightarrow \rho_{2}^{(-)}=-\rho_{-2} \tag{39}
\end{equation*}
$$

and taking the trace over the bispinor part of $\tilde{M}\left(\phi_{2}, \phi_{1}, p_{2}, p_{1}\right) \tilde{M}^{*}\left(\phi_{2}^{\prime}, \phi_{1}^{\prime}, p_{2}, p_{1}\right)$. The quantities $\rho_{1}$ and $\rho_{2}^{(-)}$denote the electron and positron density matrices, respectively. In the case of the initial particles being unpolarized, we have

$$
\begin{equation*}
\rho_{1}=\frac{1}{2}\left(\gamma p_{1}+m\right), \quad \rho_{2}^{(-)}=-\rho_{-2}=\frac{1}{2}\left(\gamma p_{2}-m\right) . \tag{40}
\end{equation*}
$$

Upon squaring the noninstantaneous part of the direct diagram, we obtain:

$$
\begin{equation*}
\frac{1}{4} \sum_{\sigma_{i}, \lambda_{i}} \tilde{M}^{\mathrm{nd}} \tilde{M}^{\mathrm{nd} *}=-8 e^{4} m^{4} \int \frac{\mathrm{~d} p_{3}^{-}}{2 \pi} \frac{\mathrm{~d} p_{3}^{\prime-}}{2 \pi} \exp \left(i \Phi^{\mathrm{dd}}\right) \frac{\tilde{\mathcal{M}}^{\mathrm{nndd}}}{\left(p_{3}^{2}-m^{2}+i \epsilon\right)\left(p_{3}^{\prime 2}-m^{2}-i \epsilon\right)}, \tag{41}
\end{equation*}
$$

with $p_{3}^{(+, \perp)}=p_{3}^{(+, \perp)}$. The phase $\Phi^{\mathrm{dd}}$ reads

$$
\begin{equation*}
\Phi^{\mathrm{dd}}=\left(k_{2}^{-}-p_{2}^{-}\right) \delta_{2}^{+}-\left(k_{1}^{-}-p_{1}^{-}\right) \delta_{1}^{+}-p_{3}^{-}\left(x_{2}^{+}-x_{1}^{+}\right)+p_{3}^{\prime-}\left(x_{2}^{\prime+}-x_{1}^{\prime+}\right)+\Phi_{F}^{\mathrm{dd}} \tag{42}
\end{equation*}
$$

with the field-dependent part $\Phi_{F}^{\mathrm{dd}}$ given by [we use the canonical light-cone basis (2)]

$$
\begin{align*}
\Phi_{F}^{\mathrm{dd}}= & \frac{m}{p_{3}^{+}} \sum_{i} \xi_{i} k_{1}^{i}\left(\delta_{2}^{+} I_{2 i}+\delta_{1}^{+} I_{1 i}\right) \\
& -\frac{m^{2}}{p_{3}^{+}} \sum_{i} t_{i} \xi_{i}\left(\frac{k_{2}^{+}}{p_{2}^{+}} \delta_{2}^{+} I_{2 i}+\frac{k_{1}^{+}}{p_{1}^{+}} \delta_{1}^{+} I_{1 i}\right)-\frac{m^{2}}{2 p_{3}^{+}} \sum_{i} \xi_{i}^{2}\left(\frac{k_{2}^{+}}{p_{2}^{+}} \delta_{2}^{+} J_{2 i}+\frac{k_{1}^{+}}{p_{1}^{+}} \delta_{1}^{+} J_{1 i}\right), \tag{43}
\end{align*}
$$

where

$$
\begin{equation*}
I_{j i}=\frac{1}{2} \int_{-1}^{1} \mathrm{~d} \lambda \psi_{i}\left(m X_{j}^{+}+\frac{1}{2} m \delta_{j}^{+} \lambda\right), \quad J_{j i}=\frac{1}{2} \int_{-1}^{1} \mathrm{~d} \lambda \psi_{i}^{2}\left(m X_{j}^{+}+\frac{1}{2} m \delta_{j}^{+} \lambda\right) \tag{44}
\end{equation*}
$$

For the products of the noninstantaneous and instantaneous direct terms and vice versa, we obtain correspondingly

$$
\begin{equation*}
\frac{1}{4} \sum_{\sigma_{i}, \lambda_{i}} \tilde{M}^{\mathrm{nd}} \tilde{M}^{\mathrm{id} *}=-2 e^{4}\left[2 m^{2}+s(\phi)\right] \delta\left(\delta_{2}^{+}+\delta_{1}^{+}-2 \delta^{+}\right) \int \frac{\mathrm{d} p_{3}^{-}}{2 \pi} \exp \left(i \Phi^{\mathrm{dd}}\right) \frac{\tilde{\mathcal{M}}^{\text {nidd }}}{p_{3}^{2}-m^{2}+i \epsilon} \tag{45}
\end{equation*}
$$

and

$$
\begin{equation*}
\frac{1}{4} \sum_{\sigma_{i}, \lambda_{i}} \tilde{M}^{\mathrm{id}} \tilde{M}^{\mathrm{nd} *}=-2 e^{4}\left[2 m^{2}+s(\phi)\right] \delta\left(\delta_{2}^{+}+\delta_{1}^{+}+2 \delta^{+}\right) \int \frac{\mathrm{d} p_{3}^{\prime-}}{2 \pi} \exp \left(i \Phi^{\mathrm{dd}}\right) \frac{\tilde{\mathcal{M}}^{\mathrm{indd}}}{{p_{3}^{\prime 2}-m^{2}-i \epsilon}^{2} . .} \tag{46}
\end{equation*}
$$

Finally, the product of the two instantaneous direct terms is given by

$$
\begin{equation*}
\frac{1}{4} \sum_{\sigma_{i}, \lambda_{i}} \tilde{M}^{\mathrm{id}} \tilde{M}^{\mathrm{id} *}=e^{4} \delta\left(\delta_{2}^{+}+\delta_{1}^{+}\right) \delta\left(\delta^{+}\right) \exp \left(i \Phi^{\mathrm{dd}}\right) \tilde{\mathcal{M}}^{\mathrm{iidd}} \tag{47}
\end{equation*}
$$

The quantities $\tilde{\mathcal{M}}^{\text {nndd }}, \tilde{\mathcal{M}}^{\text {nidd }}, \tilde{\mathcal{M}}^{\text {indd }}$, and $\tilde{\mathcal{M}}^{\text {iidd }}$ are the traces of the corresponding bispinor parts, which are subsequently rearranged with the use of momentum relations in the background field. Details and explicit expressions are provided in Appendix D. The prefactors in Eqs. (41), (45), and (46) are chosen in such a way, that $\tilde{\mathcal{M}}^{\text {nndd }}=$ $\tilde{\mathcal{M}}^{\text {indd }}=\tilde{\mathcal{M}}^{\text {nidd }}=1$ in the limit of a vanishing laser field.

## IV. ONE-STEP AND TWO-STEP CONTRIBUTIONS

As it has been pointed out in the introduction, in contrast to the vacuum case, the probability of a tree-level secondorder process in an external field [and hence the cross section (33)] contains contributions with the intermediate particle being virtual, as well as real, and it can be written as a sum of so called one-step and two-step or cascade terms [53, $54,58-62,67,69]$. If the intermediate particle is real, generally speaking, the propagation distance may be arbitrarily large inside the field. This causes at least two problems: for sufficiently large distances, the approximation (26) may break down and also radiative corrections to the electron/photon propagator may become sizable. On the other hand, in principle, one can recover the two-step contribution as a combination of the two corresponding first-order processes, therefore, it is the one-step contribution that is the most nontrivial.

Let us single out the one-step contribution from the cross section (33). In our approach, we employ the Schwinger proper time representation for the denominators of the electron propagators. This allows us to avoid the use of the Heaviside step functions and to write the two-step and one-step contributions as integrals with limits independent of any variable. But let us first highlight the main ideas of the common approach employed in the literature.

Note that the two-step contribution is contained in the 'nndd' term [60, 69]. For the 'nndd' term (41), let us consider the integrals in $p_{3}^{-}$and ${p_{3}^{\prime}}^{-}$:

$$
\begin{equation*}
I_{p_{3}, p_{3}^{\prime}}=\int \frac{\mathrm{d} p_{3}^{-}}{2 \pi} \frac{\mathrm{~d} p_{3}^{\prime-}}{2 \pi} \frac{\mathrm{e}^{-i p_{3}^{-}\left(x_{2}^{+}-x_{1}^{+}\right)} \mathrm{e}^{i p_{3}^{\prime-}\left(x_{2}^{\prime+}-x_{1}^{\prime+}\right)}}{\left(p_{3}^{2}-m^{2}+i \epsilon\right)\left({p_{3}^{\prime}}^{2}-m^{2}-i \epsilon\right)} \tag{48}
\end{equation*}
$$

Evaluating each of the integrals separately and then combining the results, one obtains:

$$
\begin{equation*}
I_{p_{3}, p_{3}^{\prime}}=\frac{1}{\left(2 p_{3}^{+}\right)^{2}} \exp \left[-i \tilde{p}_{3}^{-}\left(\delta_{2}^{+}+\delta_{1}^{+}\right)\right]\left[\theta\left(p_{3}^{+}\right) \theta\left(x_{2}^{+}-x_{1}^{+}\right) \theta\left(x_{2}^{\prime+}-x_{1}^{\prime+}\right)+\theta\left(-p_{3}^{+}\right) \theta\left(x_{1}^{+}-x_{2}^{+}\right) \theta\left(x_{1}^{\prime+}-x_{2}^{\prime+}\right)\right] \tag{49}
\end{equation*}
$$

The product $\theta\left(x_{2}^{+}-x_{1}^{+}\right) \theta\left(x_{2}^{\prime+}-x_{1}^{\prime+}\right)$ can be written as $[60,62]$

$$
\begin{equation*}
\theta\left(x_{2}^{+}-x_{1}^{+}\right) \theta\left(x_{2}^{\prime+}-x_{1}^{\prime+}\right)=\theta\left(\delta^{+}\right)\left[1-\theta\left(\frac{\left|\delta_{2}^{+}+\delta_{1}^{+}\right|}{2}-\delta^{+}\right)\right] \tag{50}
\end{equation*}
$$

In Eq. (50), a two-step contribution is usually associated with the first term, and the second term is referred to as a one-step contribution. Recalling the definition of $\delta^{+}$[see Eq. (27)], we conclude that the function $\theta\left(\delta^{+}\right)$identifies the
two-step contribution corresponding to the electron emitting a photon first and then annihilating with the positron into the second photon. Using an analogous transformation for the product $\theta\left(x_{1}^{+}-x_{2}^{+}\right) \theta\left(x_{1}^{\prime+}-x_{2}^{\prime+}\right)$ in Eq. (49), one obtains a two-step contribution $\propto \theta\left(-\delta^{+}\right)$, which corresponds to the positron emitting a photon first and then annihilating with the electron into the second photon. The total two-step contribution can be written as

$$
\begin{equation*}
I_{p_{3}, p_{3}^{\prime}}^{\mathrm{two} \text { step }}=\frac{1}{\left(2 p_{3}^{+}\right)^{2}} \exp \left[-i \tilde{p}_{3}^{-}\left(\delta_{2}^{+}+\delta_{1}^{+}\right)\right] \theta\left(p_{3}^{+} \delta^{+}\right) \tag{51}
\end{equation*}
$$

and the one-step contribution, originating from the 'nndd' term, as

$$
\begin{align*}
I_{p_{3}, p_{3}^{\prime}}^{\text {one-step }}=-\frac{1}{\left(2 p_{3}^{+}\right)^{2}} \exp \left[-i \tilde{p}_{3}^{-}\left(\delta_{2}^{+}+\delta_{1}^{+}\right)\right] & {\left[\theta\left(p_{3}^{+}\right) \theta\left(\delta^{+}\right) \theta\left(\frac{\left|\delta_{2}^{+}+\delta_{1}^{+}\right|}{2}-\delta^{+}\right)\right.} \\
& \left.+\theta\left(-p_{3}^{+}\right) \theta\left(-\delta^{+}\right) \theta\left(\frac{\left|\delta_{2}^{+}+\delta_{1}^{+}\right|}{2}+\delta^{+}\right)\right] \tag{52}
\end{align*}
$$

Now, let us show an alternative way of representing the two-step and one-step contributions in Eqs. (51) and (52), respectively. We employ the following proper-time representation for the denominators:

$$
\begin{equation*}
\frac{1}{p_{3}^{2}-m^{2}+i \epsilon}=-i \int_{0}^{\infty} \mathrm{d} s \mathrm{e}^{i\left(p_{3}^{2}-m^{2}+i \epsilon\right) s}, \quad \frac{1}{{p_{3}^{\prime}}^{2}-m^{2}-i \epsilon}=i \int_{0}^{\infty} \mathrm{d} t \mathrm{e}^{-i\left(p_{3}^{\prime 2}-m^{2}-i \epsilon\right) t} \tag{53}
\end{equation*}
$$

Below, we do not write the terms with $i \epsilon$ for brevity. The integrals in $p_{3}^{-}$and $p_{3}^{\prime-}$ yield [see Eq. (48)]

$$
\begin{equation*}
\int \frac{\mathrm{d} p_{3}^{-}}{2 \pi} \frac{\mathrm{~d} p_{3}^{\prime-}}{2 \pi} \rightarrow \delta\left[2 p_{3}^{+} s-\left(x_{2}^{+}-x_{1}^{+}\right)\right] \delta\left[2 p_{3}^{+} t-\left(x_{2}^{\prime+}-x_{1}^{\prime+}\right)\right] . \tag{54}
\end{equation*}
$$

In place of $s$ and $t$, we introduce the variables $\tau$ and $v[97,98]$ :

$$
\begin{equation*}
\tau=s+t, \quad v=\frac{s-t}{s+t}, \quad \int_{0}^{\infty} \mathrm{d} s \mathrm{~d} t \rightarrow \int_{-1}^{1} \mathrm{~d} v \int_{0}^{\infty} \mathrm{d} \tau \frac{\tau}{2} \tag{55}
\end{equation*}
$$

In terms of the new variables the delta functions in Eq. (54) can be written as

$$
\begin{equation*}
\delta\left[2 p_{3}^{+} s-\left(x_{2}^{+}-x_{1}^{+}\right)\right] \delta\left[2 p_{3}^{+} t-\left(x_{2}^{\prime+}-x_{1}^{\prime+}\right)\right]=\delta\left(\delta^{+}-p_{3}^{+} \tau\right) \delta\left(\delta_{2}^{+}+\delta_{1}^{+}-2 p_{3}^{+} v \tau\right) \tag{56}
\end{equation*}
$$

and the initial quantity $I_{p_{3}, p_{3}^{\prime}}$ in Eq. (48) reads

$$
\begin{equation*}
I_{p_{3}, p_{3}^{\prime}}=\int_{-1}^{1} \mathrm{~d} v \int_{0}^{\infty} \mathrm{d} \tau \frac{\tau}{2} \delta\left(\delta^{+}-p_{3}^{+} \tau\right) \delta\left(\delta_{2}^{+}+\delta_{1}^{+}-2 p_{3}^{+} v \tau\right) \exp \left[-i \tilde{p}_{3}^{-}\left(\delta_{2}^{+}+\delta_{1}^{+}\right)\right] \tag{57}
\end{equation*}
$$

Evaluating the integrals in $\tau$ and $v$, one obtains that

$$
\begin{equation*}
I_{p_{3}, p_{3}^{\prime}}=\frac{1}{\left(2 p_{3}^{+}\right)^{2}} \exp \left[-i \tilde{p}_{3}^{-}\left(\delta_{2}^{+}+\delta_{1}^{+}\right)\right] \theta\left(p_{3}^{+} \delta^{+}\right) \theta\left(1-\left|\frac{\delta_{2}^{+}+\delta_{1}^{+}}{2 \delta^{+}}\right|\right) \tag{58}
\end{equation*}
$$

where the first $\theta$-function comes from the integral in $\tau$ and the second one comes from the integral in $v$. We notice that Eq. (58) is the same as Eq. (51), apart from the presence of the second $\theta$-function. Then, the two-step contribution can be written as

$$
\begin{equation*}
I_{p_{3}, p_{3}^{\prime}}^{\mathrm{two-step}}=\int \mathrm{d} v \int_{0}^{\infty} \mathrm{d} \tau \frac{\tau}{2} \delta\left(\delta^{+}-p_{3}^{+} \tau\right) \delta\left(\delta_{2}^{+}+\delta_{1}^{+}-2 p_{3}^{+} v \tau\right) \exp \left[-i \tilde{p}_{3}^{-}\left(\delta_{2}^{+}+\delta_{1}^{+}\right)\right] \tag{59}
\end{equation*}
$$

which agrees with Eq. (51) upon the evaluation of the integrals in $\tau$ and $v$ [note that the limits of the integration in $v$ are extended to be $(-\infty, \infty)]$. The difference between Eqs. (57) and (59) is the one-step contribution:

$$
\begin{equation*}
I_{p_{3}, p_{3}^{\prime}}^{\text {one-step }}=-\int_{\Gamma_{v}} \mathrm{~d} v \int_{0}^{\infty} \mathrm{d} \tau \frac{\tau}{2} \delta\left(\delta^{+}-p_{3}^{+} \tau\right) \delta\left(\delta_{2}^{+}+\delta_{1}^{+}-2 p_{3}^{+} v \tau\right) \exp \left[-i \tilde{p}_{3}^{-}\left(\delta_{2}^{+}+\delta_{1}^{+}\right)\right] \tag{60}
\end{equation*}
$$

with $\Gamma_{v}=(-\infty,-1) \cup(1, \infty)$. In the following, we consider the one-step contribution and therefore employ Eq. (60). The final expression can be easily transformed into the result for the two-step contribution [Eq. (59)] or for the sum of both contributions [Eq. (57)].

## V. EVALUATION OF THE INTEGRALS

For the 'nidd' and 'indd' terms in Eqs. (45) and (46), respectively, we also employ the proper-time representation, e.g., we have

$$
\begin{equation*}
\int \frac{\mathrm{d} p_{3}^{-}}{2 \pi} \frac{\mathrm{e}^{-i p_{3}^{-}\left(x_{2}^{+}-x_{1}^{+}\right)}}{p_{3}^{2}-m^{2}+i \epsilon}=-i \int_{0}^{\infty} \mathrm{d} \tau \frac{1}{2} \delta\left(\delta^{+}-p_{3}^{+} \tau\right) \exp \left[-i \tilde{p}_{3}^{-}\left(\delta_{2}^{+}+\delta_{1}^{+}\right)\right] \tag{61}
\end{equation*}
$$

for the 'nidd' term and an analogous expression for the 'indd' term [note that for the 'iidd' term no proper-time representation is required, since there are no noninstantaneous parts of the propagators and integrals in the ' - , momentum components, see Eq. (47)]. After that, we notice that each of the four terms, which we need to compute, contains two delta functions [see Eqs. (41), (45), (46), (47), (60), and (61)], and they allow us to evaluate the integrals in $\delta^{+}$and $\delta_{2}^{+}$in Eq. (33). In place of $\delta_{1}^{+}$we introduce

$$
\begin{equation*}
\rho=\frac{m^{2} p_{2}^{+}}{k_{2}^{+} p_{3}^{+}} \delta_{2}^{+}+\frac{m^{2} p_{1}^{+}}{k_{1}^{+} p_{3}^{+}} \delta_{1}^{+} \tag{62}
\end{equation*}
$$

and we also rescale $\tau$ as

$$
\begin{equation*}
m^{2} \tau \rightarrow \tau \tag{63}
\end{equation*}
$$

such that the rescaled variable is dimensionless. Then the direct-direct parts of the total cross section are given by

$$
\begin{align*}
\sigma^{\mathrm{nndd}}(\phi) & =\frac{2 \pi^{2} r_{e}^{2}}{I(\phi)\left(p_{2}^{+}+p_{1}^{+}\right)} \int_{0}^{p_{2}^{+}+p_{1}^{+}} \frac{\mathrm{d} k_{1}^{+}}{2 \pi} \int \frac{\mathrm{~d}^{2} k_{1}^{\perp}}{(2 \pi)^{2}} \int_{\Gamma_{v}} \mathrm{~d} v \int_{0}^{\infty} \mathrm{d} \tau \int \mathrm{~d} \rho \tau \exp \left(i \Phi_{v}^{\mathrm{dd}}\right) \tilde{\mathcal{M}}^{\mathrm{nndd}}  \tag{64}\\
\sigma^{\{\mathrm{ni}\} \mathrm{dd}}(\phi) & =\frac{i \pi^{2} r_{e}^{2}\left[2 m^{2}+s(\phi)\right]}{2 m^{2} I(\phi)\left(p_{2}^{+}+p_{1}^{+}\right)} \int_{0}^{p_{2}^{+}+p_{1}^{+}} \frac{\mathrm{d} k_{1}^{+}}{2 \pi} \int \frac{\mathrm{~d}^{2} k_{1}^{\perp}}{(2 \pi)^{2}} \int_{0}^{\infty} \mathrm{d} \tau \int \mathrm{~d} \rho\left[\exp \left(i \Phi_{1}^{\mathrm{dd}}\right) \tilde{\mathcal{M}}^{\text {nidd }}-\exp \left(i \Phi_{-1}^{\mathrm{dd}}\right) \tilde{\mathcal{M}}^{\text {indd }}\right]  \tag{65}\\
\sigma^{\mathrm{iidd}}(\phi) & =-\frac{\pi^{2} r_{e}^{2}}{I(\phi)\left(p_{2}^{+}+p_{1}^{+}\right)} \int_{0}^{p_{2}^{+}+p_{1}^{+}} \frac{\mathrm{d} k_{1}^{+}}{2 \pi} \int \frac{\mathrm{~d}^{2} k_{1}^{\perp}}{(2 \pi)^{2}} \int \mathrm{~d} \rho \exp \left(i \Phi^{\mathrm{iidd}}\right) \tag{66}
\end{align*}
$$

where $r_{e}=\alpha / m$ is the classical electron radius, and the 'nidd' and 'indd' terms have been combined as

$$
\begin{equation*}
\sigma^{\{\mathrm{ni}\} \mathrm{dd}}(\phi)=\sigma^{\mathrm{indd}}(\phi)+\sigma^{\mathrm{nidd}}(\phi) \tag{67}
\end{equation*}
$$

The phase $\Phi_{v}^{\mathrm{dd}}$ is given by

$$
\begin{align*}
\Phi_{v}^{\mathrm{dd}}= & -\frac{\rho}{2 m^{2}}\left(\boldsymbol{k}_{1}^{\perp 2}+2 \boldsymbol{k}_{1}^{\perp} \boldsymbol{P}^{\perp}\right)+\frac{\rho\left(1+t_{1}^{2}+t_{2}^{2}\right)}{4}\left[\frac{k_{2}^{+2}}{p_{2}^{+2}}(u-1)-\frac{k_{1}^{+2}}{p_{1}^{+2}}(u+1)\right] \\
& +\frac{\rho}{2} \sum_{i} t_{i} \xi_{i}\left[\frac{k_{2}^{+2}}{p_{2}^{+2}}(u-1) I_{2 i}-\frac{k_{1}^{+2}}{p_{1}^{+2}}(u+1) I_{1 i}\right]+\frac{\rho}{4} \sum_{i} \xi_{i}^{2}\left[\frac{k_{2}^{+2}}{p_{2}^{+2}}(u-1) J_{2 i}-\frac{k_{1}^{+2}}{p_{1}^{+2}}(u+1) J_{1 i}\right] \tag{68}
\end{align*}
$$

where

$$
\begin{equation*}
u=\left[\left(\frac{k_{2}^{+}}{p_{2}^{+}}+\frac{k_{1}^{+}}{p_{1}^{+}}\right)-\frac{4 v \tau}{\rho}\right] /\left(\frac{k_{2}^{+}}{p_{2}^{+}}-\frac{k_{1}^{+}}{p_{1}^{+}}\right) \tag{69}
\end{equation*}
$$

and

$$
\begin{equation*}
P^{i}=\frac{1}{2} m \xi_{i}\left[\frac{k_{2}^{+}}{p_{2}^{+}}(u-1) I_{2 i}-\frac{k_{1}^{+}}{p_{1}^{+}}(u+1) I_{1 i}\right]-\frac{2 m t_{i} v \tau}{\rho} \tag{70}
\end{equation*}
$$

the phases in Eq. (65) are the same as $\Phi_{v}^{\mathrm{dd}}$, but with $v=1$ and $v=-1$, respectively, and the phase $\Phi^{\mathrm{iidd}}$ is given by

$$
\begin{equation*}
\Phi^{\mathrm{iidd}}=-\frac{\rho \boldsymbol{k}_{1}^{\perp 2}}{2 m^{2}}+\frac{\rho k_{2}^{+} k_{1}^{+}}{2 p_{2}^{+} p_{1}^{+}}\left[1+\sum_{i}\left(t_{i}+\xi_{i} I_{i}\right)^{2}+\sum_{i} \xi_{i}^{2}\left(J_{i}-I_{i}^{2}\right)\right] \tag{71}
\end{equation*}
$$



Figure 2. The employed deformation of the integration contour for the 'iidd' term. As $\epsilon \rightarrow 0$, the integrals over the two half-lines combine together into a principle value integral, and the integral over the semicircle results in a term proportional to the residue at $\rho=0$ (see the main text for details). The choice of the deformation into the lower half-plane is dictated by the fact that after the deformation the integral in $\boldsymbol{k}_{1}^{\perp}$ has to be convergent for a finite value of $\epsilon$.
where

$$
\begin{equation*}
I_{i}=\frac{1}{2} \int_{-1}^{1} \mathrm{~d} \lambda \psi_{i}\left(\phi+\frac{k_{2}^{+} k_{1}^{+}}{2 m\left(p_{2}^{+}+p_{1}^{+}\right)} \rho \lambda\right), \quad J_{i}=\frac{1}{2} \int_{-1}^{1} \mathrm{~d} \lambda \psi_{i}^{2}\left(\phi+\frac{k_{2}^{+} k_{1}^{+}}{2 m\left(p_{2}^{+}+p_{1}^{+}\right)} \rho \lambda\right) \tag{72}
\end{equation*}
$$

The old variables $\delta_{1}^{+}$and $\delta_{2}^{+}$are expressed via the new ones as

$$
\begin{equation*}
\delta_{1}^{+}=\frac{p_{3}^{+} k_{1}^{+}}{2 m^{2} p_{1}^{+}}(1+u) \rho, \quad \delta_{2}^{+}=\frac{p_{3}^{+} k_{2}^{+}}{2 m^{2} p_{2}^{+}}(1-u) \rho \tag{73}
\end{equation*}
$$

The integrals in $\boldsymbol{k}_{1}^{\perp}$ are Gauss-type (Fresnel) integrals and can be evaluated analytically [note that the exponential prefactors in Eqs. (64), (65), and (66) do not depend on $\boldsymbol{k}_{1}^{\perp}$; see Appendix D for details]. However, before being able to perform an integral in $\boldsymbol{k}_{1}^{\perp}$, we need to change the order of the integrations and, strictly speaking, we have to ensure that upon those changes the integrals remain convergent. It can be seen from Eqs. (68) and (71) that $\rho=0$ is a possible problematic point. Then, assuming that, if necessary, the integration contour for $\rho$ is deformed from $(-\infty, \infty)$ into a new appropriately chosen contour $\Gamma_{\rho}$ we obtain that

$$
\begin{equation*}
\int \frac{\mathrm{d}^{2} k_{1}^{\perp}}{(2 \pi)^{2}} \int_{\Gamma_{\rho}} \mathrm{d} \rho \exp \left[-i \frac{\rho}{2 m^{2}}\left(\boldsymbol{k}_{1}^{\perp 2}+2 \boldsymbol{k}_{1}^{\perp} \boldsymbol{P}^{\perp}\right)\right]=-\frac{i m^{2}}{2 \pi} \int_{\Gamma_{\rho}} \frac{\mathrm{d} \rho}{\rho} \exp \left(i \frac{\rho \boldsymbol{P}^{\perp 2}}{2 m^{2}}\right), \tag{74}
\end{equation*}
$$

where one should put $\boldsymbol{P}^{\perp}=\mathbf{0}$ for the 'iidd' term. In order to specify $\Gamma_{\rho}$, let us consider the 'iidd' term and the other two separately. We start with the 'iidd' term [Eq. (66)].

If follows from Eq. (74), that upon the exchange of the integrations the integral in $\boldsymbol{k}_{1}^{\perp}$ yields an infinite volume factor, if $\boldsymbol{P}^{\perp}=\mathbf{0}$ and $\rho=0$. Therefore, we indeed need to deform the contour, such that the new contour $\Gamma_{\rho}$ does not go through the point $\rho=0$. One of the possibilities is to shift the integration line by $i \epsilon$ off the real axis. This results in an $i \epsilon$ prescription for $\rho[99,100]$. However, since the singularity is only at $\rho=0$, it is enough to deform the contour locally by introducing a semicircle of radius $\epsilon$, as shown in Fig. 2. Then, as $\epsilon \rightarrow 0$, the integral over the two half-lines results in the principal value integral, and the integral over the semicircle yields $i \pi C_{-1}$, with $C_{-1}$ being the residue at $\rho=0$ [101].

For the other terms [Eqs. (64) and (65)], the vector $\boldsymbol{P}^{\perp}$ is given by Eq. (70). As a result, upon setting $\rho=0$, the integral in $\boldsymbol{k}_{1}^{\perp}$ is evaluated not to an infinite volume factor, but to a delta function. Therefore, we argue that the deformation of the contour for $\rho$ is not required for these terms and $\Gamma_{\rho}=(-\infty, \infty)$. We justify this by reproducing the vacuum results, known from the literature, if the external field is set to zero (see below).

We also point out that if one makes the replacement $\rho \rightarrow-\rho$, then

$$
\begin{equation*}
\Phi_{v}^{\mathrm{dd}} \rightarrow-\Phi_{-v}^{\mathrm{dd}},\left.\quad \tilde{\mathcal{M}}^{\text {nndd }} \rightarrow \tilde{\mathcal{M}}^{\text {nndd }}\right|_{v \rightarrow-v}, \quad \tilde{\mathcal{M}}^{\text {indd }} \rightarrow \tilde{\mathcal{M}}^{\text {nidd }}, \quad \tilde{\mathcal{M}}^{\text {nidd }} \rightarrow \tilde{\mathcal{M}}^{\text {indd }} \tag{75}
\end{equation*}
$$

Therefore, the integral in $v$ can be reduced to an integral over the interval $(1, \infty)$ [alternatively, the integral in $\rho$ can be reduced to an integral over $(0, \infty)$; we use the first option below].

As the last steps, we notice that after the integration in $\boldsymbol{k}_{1}^{\perp}$, upon rescaling $\rho$ as $\rho k_{2}^{+} k_{1}^{+} /\left(p_{2}^{+} p_{1}^{+}\right) \rightarrow \rho$ for the 'iidd' term, the integral in $k_{1}^{+}$can be also evaluated analytically and only a single integral in $\rho$ remains in this term, which can be also written as an integral over $(0, \infty)$.

## VI. FINAL RESULT

After all steps described above are carried out, one obtains the final expressions for the direct-direct contributions to the total cross section:

$$
\begin{align*}
\sigma^{\mathrm{nndd}}(\phi) & =\frac{r_{e}^{2} m^{2}}{I(\phi)\left(p_{2}^{+}+p_{1}^{+}\right)} \operatorname{Im} \int_{0}^{p_{2}^{+}+p_{1}^{+}} \mathrm{d} k_{1}^{+} \int_{1}^{\infty} \mathrm{d} v \int_{0}^{\infty} \mathrm{d} \tau \int \frac{\mathrm{~d} \rho}{\rho} \tau \exp \left(i \Phi_{v}^{\mathrm{dd}}\right) \tilde{\mathcal{M}}^{\mathrm{nndd}}  \tag{76}\\
\sigma^{\{\mathrm{ni}\} \mathrm{dd}}(\phi) & =\frac{r_{e}^{2}\left[2 m^{2}+s(\phi)\right]}{4 I(\phi)\left(p_{2}^{+}+p_{1}^{+}\right)} \operatorname{Re} \int_{0}^{p_{2}^{+}+p_{1}^{+}} \mathrm{d} k_{1}^{+} \int_{0}^{\infty} \mathrm{d} \tau \int \frac{\mathrm{~d} \rho}{\rho} \exp \left(i \Phi_{1}^{\mathrm{dd}}\right) \tilde{\mathcal{M}}^{\mathrm{nidd}}  \tag{77}\\
\sigma^{\mathrm{iidd}}(\phi) & =\frac{r_{e}^{2} m^{2}}{2 I(\phi)}\left(\int_{0}^{\infty} \frac{\mathrm{d} \rho}{\rho} \sin \Phi^{\mathrm{iidd}}+\frac{\pi}{2}\right) \tag{78}
\end{align*}
$$

where $\operatorname{Im}$ and Re denote an imaginary and a real part, respectively, expressions for the quantities $\tilde{\mathcal{M}}^{\text {nndd }}$ and $\tilde{\mathcal{M}}^{\text {nidd }}$ are provided in Appendix D , and the phase $\Phi_{v}^{\mathrm{dd}}$ is given by

$$
\begin{align*}
\Phi_{v}^{\mathrm{dd}}= & \frac{\rho}{4}\left[\frac{k_{2}^{+2}}{p_{2}^{+2}}(u-1)-\frac{k_{1}^{+2}}{p_{1}^{+2}}(u+1)\right]+\frac{\rho}{8}\left(u^{2}-1\right) \sum_{i}\left(\frac{k_{2}^{+}}{p_{2}^{+}} \zeta_{2 i}-\frac{k_{1}^{+}}{p_{1}^{+}} \zeta_{1 i}\right)^{2} \\
& +\frac{\rho}{4} \sum_{i}\left[\frac{k_{2}^{+2}}{p_{2}^{+2}}(u-1)\left(\zeta_{2 i}^{(2)}-\zeta_{2 i}^{2}\right)-\frac{k_{1}^{+2}}{p_{1}^{+2}}(u+1)\left(\zeta_{1 i}^{(2)}-\zeta_{1 i}^{2}\right)\right] \tag{79}
\end{align*}
$$

with

$$
\begin{equation*}
\zeta_{j i}=\frac{1}{2} \int_{-1}^{1} \mathrm{~d} \lambda t_{i}\left(m X_{j}^{+}+\frac{1}{2} m \delta_{j}^{+} \lambda\right), \quad \zeta_{j i}^{(2)}=\frac{1}{2} \int_{-1}^{1} \mathrm{~d} \lambda t_{i}^{2}\left(m X_{j}^{+}+\frac{1}{2} m \delta_{j}^{+} \lambda\right) \tag{80}
\end{equation*}
$$

and $X_{2}^{+}=x^{+}+p_{3}^{+} \tau / m^{2}, X_{1}^{+}=x^{+}-p_{3}^{+} \tau / m^{2}$. For the 'iidd' term, the phase $\Phi^{\text {iidd }}$ is given by

$$
\begin{equation*}
\Phi^{\mathrm{iidd}}=\rho\left[1+\sum_{i} \zeta_{i}^{2}+\sum_{i}\left(\zeta_{i}^{(2)}-\zeta_{i}^{2}\right)\right] \tag{81}
\end{equation*}
$$

where

$$
\begin{equation*}
\zeta_{i}=\frac{1}{2} \int_{-1}^{1} \mathrm{~d} \lambda t_{i}\left(\phi+\frac{p_{2}^{+} p_{1}^{+}}{m\left(p_{2}^{+}+p_{1}^{+}\right)} \rho \lambda\right), \quad \zeta_{i}^{(2)}=\frac{1}{2} \int_{-1}^{1} \mathrm{~d} \lambda t_{i}^{2}\left(\phi+\frac{p_{2}^{+} p_{1}^{+}}{m\left(p_{2}^{+}+p_{1}^{+}\right)} \rho \lambda\right) \tag{82}
\end{equation*}
$$

## VII. ZERO-FIELD LIMIT

In general, the integrals in Eqs. (76), (77), and (78) have to be computed numerically. However, in the case of a vanishing plane-wave field, one should be able to evaluate them analytically and thus recover the result, known from the literature [85]. Since this derivation is different from and also somewhat less trivial than the one usually presented, we show explicitly how the vacuum expressions are obtained.

Let us start with the 'iidd' term in Eq. (78), which is the simplest out of three. If the external field is set to zero, then $\Phi^{\text {iidd }}=\left(1+\boldsymbol{t}^{\perp 2}\right) \rho$, where $\boldsymbol{t}^{\perp 2}=t_{1}^{2}+t_{2}^{2}$. The integral in $\rho$ reduces to the Dirichlet integral and we obtain that

$$
\begin{equation*}
\sigma^{\mathrm{iidd}}=-\frac{\pi r_{e}^{2}}{4 \sqrt{\mu(\mu-1)}} \tag{83}
\end{equation*}
$$

where $\mu$ is the scaled invariant mass squared: $\mu=s /\left(4 m^{2}\right)$, with $s=\left(p_{2}+p_{1}\right)^{2}$.

The other two contributions require some more manipulations. Upon setting the laser field to zero, the quantities $\tilde{\mathcal{M}}^{\text {nndd }}$, and $\tilde{\mathcal{M}}^{\text {nidd }}$ are equal to unity, and the phase $\Phi_{v}^{\text {dd }}$ reduces to

$$
\begin{equation*}
\Phi_{v}^{\mathrm{dd}}=\frac{1}{\rho}+\frac{1}{4} a^{2} v^{2} \tau^{2} \rho-b v \tau \tag{84}
\end{equation*}
$$

where

$$
\begin{equation*}
a=2 \sqrt{\frac{\boldsymbol{t}^{\perp 2}\left(1+\boldsymbol{t}^{\perp 2}\right) k_{2}^{+} k_{1}^{+}}{p_{2}^{+} p_{1}^{+}}}, \quad b=\left(\frac{k_{2}^{+}}{p_{2}^{+}}+\frac{k_{1}^{+}}{p_{1}^{+}}\right)\left(1+\boldsymbol{t}^{\perp 2}\right) \tag{85}
\end{equation*}
$$

and we have rescaled $\rho$ as $\rho /\left(2 \boldsymbol{t}^{\perp 2} v^{2} \tau^{2}\right) \rightarrow \rho$. After that, the integrals are evaluated in the order shown in Eqs. (76) and (77). Details are presented in Appendix E. The results are given by

$$
\begin{equation*}
\sigma^{\mathrm{nndd}}=-\frac{\pi r_{e}^{2}}{4 \mu \sqrt{\mu(\mu-1)}} \tag{86}
\end{equation*}
$$

and

$$
\begin{equation*}
\sigma^{\{\mathrm{ni}\} \mathrm{dd}}=\frac{\pi r_{e}^{2}}{4 \mu(\mu-1)}\left(\mu+\frac{1}{2}\right) \ln \left(\frac{\sqrt{\mu}+\sqrt{\mu-1}}{\sqrt{\mu}-\sqrt{\mu-1}}\right) \tag{87}
\end{equation*}
$$

where $\ln$ indicates the natural logarithm. Combining all three terms together, we obtain that

$$
\begin{equation*}
\sigma^{\mathrm{dd}}=\frac{\pi r_{e}^{2}}{4 \mu^{2}(\mu-1)}\left[\mu\left(\mu+\frac{1}{2}\right) \ln \left(\frac{\sqrt{\mu}+\sqrt{\mu-1}}{\sqrt{\mu}-\sqrt{\mu-1}}\right)-(\mu+1) \sqrt{\mu(\mu-1)}\right] \tag{88}
\end{equation*}
$$

which is the same as the corresponding cross section in Ref. [85].
We point out, that initially the cross section $\sigma^{\text {dd }}$ has been defined within the light-cone quantization formalism. However, the obtained expression (88) is the same as the one derived within the instant-form quantization, which supports the way of defining the cross section on the light cone, that we have suggested.

Another important remark is the fact that the 'nndd' term in Eq. (76) does not contain the two-step contribution. Nevertheless, the complete result has been recovered, which means that the two-step contribution vanishes in vacuum, as it has to be, if the two-step contribution indeed corresponds to the physical situation of the intermediate fermion becoming real. In fact, one can verify this directly by setting the integration interval for the virtuality $v$ to $(-\infty, \infty)$ and confirming that the integral vanishes (one should be aware that in this case it is necessary to recover the $i \epsilon$ prescription for $\tau$ in order to shift the pole $v=0$ off the real axis).

## VIII. CONCLUSIONS AND OUTLOOK

We have investigated analytically the process of annihilation of an electron-positron pair into two photons in the presence of an intense plane-wave field, as a characteristic example of $2 \Rightarrow 2$ reactions. The external field has been taken into account exactly in the calculations by working in the Furry picture, and light-cone quantization has been employed, in order to have a formalism automatically incorporating the properties of the plane-wave background.

Though the presented description of the scattering based on the use of wave packets is tailored to the reaction $e^{-} e^{+} \Rightarrow \gamma \gamma$ in a laser pulse, it applies to a general second-order 2-to-2 reaction in an intense background field. We have seen that it is convenient to introduce the concept of a local cross section, which although not being a measurable quantity, is a useful tool especially for comparison of the results in a laser field and the corresponding ones in vacuum. Indeed, the local cross section in a plane-wave field is a qualitatively different entity with respect to its vacuum limit, since it bears the dependence on the light-cone moment of the collision and may also become negative in some regions of the parameter space. Therefore, the cross section in the external field cannot be seen as an observable, but instead could be interpreted as a quantity, which extends the concept of the classical cross section, similar to the relation between the Wigner distribution and the classical phase-space distribution.

In contrast to processes in a plane wave initiated by a single particle, in fact, the pair annihilation into two photons does also occur in vacuum. The vacuum part has an additional momentum-conserving delta function at each vertex, which is hidden, if one works in the Furry picture (see [102] for a discussion of splitting the amplitude of a secondorder tree-level process in a laser field into different parts). Our definition of the cross section and also the analytical evaluation of Gauss-type integrals in the transverse momentum components of the final particles effectively remove
those delta functions and allow one to write the total local cross section without a formal split into a vacuum and a field-dependent parts. We have also ensured that by setting the external field to zero, the vacuum cross section is recovered.

A distinct feature of second-order tree-level processes in an intense background is a nonvanishing contribution from the cascade or two-step channels, which correspond to the intermediate particle becoming real. In contrast to 1-to-3 reactions, 2 -to- 2 reactions have not one, but two cascade channels, which in the case of $e^{-} e^{+} \Rightarrow \gamma \gamma$ correspond to either the electron or the positron emitting first a photon and then annihilating with the other particle into the second photon. Though the different contributions can be treated in a standard fashion, which involves the use of Heaviside step functions, we have demonstrated a concise way of representing them via virtuality integrals with different integration limits being independent of any variable.

We have explicitly evaluated the total cross section (without taking into account the interference term between the direct and the exchange amplitudes) and presented the result in a form, which should be particularly suitable for numerical computations. In fact, as many as possible integrals have been performed analytically, and the result itself is expressed in terms of the dressed fermion momenta. Even though the interference term, which has not been calculated here, might have a sizable effect for some values of $\phi$, each range of these values within a laser pulse shrinks as $\xi$ grows. Hence, in a highly nonlinear (ultrarelativistic) regime one might expect that on average the interference term is negligible, if dynamics over the duration of $\Delta \phi \gtrsim 1$ in terms of the laser phase is considered (even for realistic ultrashort pulses, the condition on the pulse duration $\Delta \phi_{\text {pulse }} \gtrsim 1$ is still fulfilled).

The final result for the total cross section contains integrals which, generally speaking, need to be evaluated numerically. Commonly, a simplification is achieved in the constant-crossed and locally-constant field cases [8]. For the reaction $e^{-} e^{+} \Rightarrow \gamma \gamma$, however, the usual symmetry of processes in a plane-wave background is not preserved owing to the transverse momentum components of the incoming particles, which are encoded in the new parameters $t_{i}(\phi)$, appearing in the final result. Therefore, it might not be possible to reduce the number of the integrals even in a constant-crossed field. A detailed study of special cases and the numerical analysis are left for a future investigation.

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## Appendix A: Light-cone quantization

We define the light-cone coordinates in a covariant way using light-cone basis $\left\{\eta^{\mu}, \bar{\eta}^{\mu}, e_{1}^{\mu}, e_{2}^{\mu}\right\}$, with the four-vectors of this basis satisfying the following properties [98]:

$$
\begin{equation*}
\eta^{2}=\bar{\eta}^{2}=0, \quad \eta \bar{\eta}=1, \quad \eta e_{i}=\bar{\eta} e_{i}=0, \quad e_{i} e_{j}=-\delta_{i j} \tag{A1}
\end{equation*}
$$

Then an arbitrary four-vector $a^{\mu}$ can be written as

$$
\begin{equation*}
a^{\mu}=a^{+} \bar{\eta}^{\mu}+a^{-} \eta^{\mu}+a^{1} e_{1}^{\mu}+a^{2} e_{2}^{\mu} \tag{A2}
\end{equation*}
$$

where

$$
\begin{equation*}
a^{+}=a \eta, \quad a^{-}=a \bar{\eta}, \quad a^{1}=-a e_{1}, \quad a^{2}=-a e_{2} . \tag{A3}
\end{equation*}
$$

The metric tensor is given by

$$
\begin{equation*}
g^{\mu \nu}=\eta^{\mu} \bar{\eta}^{\nu}+\bar{\eta}^{\mu} \eta^{\nu}-e_{1}^{\mu} e_{1}^{\nu}-e_{2}^{\mu} e_{2}^{\nu} \tag{A4}
\end{equation*}
$$

which can be written in the matrix form as

$$
g^{\mu \nu}=\left(\begin{array}{rrrr}
0 & 1 & 0 & 0  \tag{A5}\\
1 & 0 & 0 & 0 \\
0 & 0 & -1 & 0 \\
0 & 0 & 0 & -1
\end{array}\right)
$$

(note that the order of the components is $+,-, 1,2$ ). The scalar product of two four-vectors $a^{\mu}$ and $b^{\mu}$ is

$$
\begin{equation*}
a b=a^{+} b^{-}+a^{-} b^{+}+a^{i} b_{i}=a^{+} b^{-}+a^{-} b^{+}-\boldsymbol{a}^{\perp} \boldsymbol{b}^{\perp} \tag{A6}
\end{equation*}
$$

where $\boldsymbol{a}^{\perp}=\left(a^{1}, a^{2}\right)$ and $\boldsymbol{b}^{\perp}=\left(b^{1}, b^{2}\right)$. For the quantization in the presence of a plane-wave field $A^{\mu}(x)=A^{\mu}\left(k_{0} x\right)$, we choose $\eta^{\mu}=k_{0}^{\mu} / m$. We also need to specify the signs of scalar products. In order to do that, we assume the signature $(+,-,-,-)$ for the metric tensor in the instant form. Then, we have $p^{+}>0, p^{2}=m^{2}$ for an on-shell fermion with four-momentum $p^{\mu}$.

The derivation of the light-front Hamiltonian is analogous to the one in the vacuum case (see [76-78]), however, the background field $A^{\mu}\left(k_{0} x\right)$ is included in the zeroth-order Hamiltonian $H_{0}[79,80]$. The result is [76-78, 81]

$$
\begin{equation*}
H=H_{0}+V_{1}+V_{2}+V_{3} \tag{A7}
\end{equation*}
$$

with

$$
\begin{align*}
& H_{0}=\int \mathrm{d}^{2} x^{\perp} \mathrm{d} x^{-}\left[\bar{\psi} \gamma^{-} i \partial_{-} \psi+e \bar{\psi} \gamma^{+} \psi \mathcal{A}^{-}+\frac{1}{2}\left(\partial_{-} \mathcal{A}^{-}\right)^{2}+\frac{1}{2}\left(\partial_{1} \mathcal{A}_{2}-\partial_{2} \mathcal{A}_{1}\right)^{2}\right] \\
& V_{1}=e \int \mathrm{~d}^{2} x^{\perp} \mathrm{d} x^{-} \bar{\psi} \gamma^{\mu} \psi \mathcal{A}_{\mu} \\
& V_{2}=\frac{e^{2}}{2} \int \mathrm{~d}^{2} x^{\perp} \mathrm{d} x^{-} \mathcal{A}_{\mu} \bar{\psi} \gamma^{\mu} \frac{\gamma^{+}}{i \partial_{-}} \gamma^{\nu} \psi \mathcal{A}_{\nu}  \tag{A8}\\
& V_{3}=\frac{e^{2}}{2} \int \mathrm{~d}^{2} x^{\perp} \mathrm{d} x^{-} \bar{\psi} \gamma^{+} \psi \frac{1}{\left(i \partial_{-}\right)^{2}} \bar{\psi} \gamma^{+} \psi
\end{align*}
$$

where $\psi$ and $\mathcal{A}^{\mu}$ are the electron and photon fields, respectively, to be quantized (in fact, only the projection $\psi_{+}=\Lambda_{+} \psi$ is an independent degree of freedom, where $\Lambda_{+}=\gamma^{-} \gamma^{+} / 2$, and $\mathcal{A}^{\mu}$ has only two independent components $[76,77]$ ).

The Dirac equation for the electron field $\psi$ is $[\gamma(i \partial-e A)-m] \psi=0$, as a result, in the interaction picture we obtain the following expansion of $\psi(x)$ via the Volkov wave functions (see [81, 103] for discussions of the completeness of the Volkov solutions on the light cone)

$$
\begin{equation*}
\psi(x)=\sum_{\sigma} \int \frac{\tilde{\mathrm{d}}^{3} p}{(2 \pi)^{3}}\left[a_{p \sigma} \psi_{p \sigma}(x)+b_{p \sigma}^{\dagger} \psi_{p \sigma}^{(-)}(x)\right] \tag{A9}
\end{equation*}
$$

where

$$
\begin{equation*}
\frac{\tilde{\mathrm{d}}^{3} p}{(2 \pi)^{3}}=\frac{\mathrm{d}^{2} p^{\perp}}{(2 \pi)^{2}} \frac{\mathrm{~d} p^{+}}{2 \pi} \theta\left(p^{+}\right) \tag{A10}
\end{equation*}
$$

$a_{p \sigma}, b_{p \sigma}\left(a_{p \sigma}^{\dagger}, b_{p \sigma}^{\dagger}\right)$ are the annihilation (creation) operators, with the anticommutation relations

$$
\begin{equation*}
\left\{a_{p \sigma}, a_{p^{\prime} \sigma^{\prime}}^{\dagger}\right\}=\left\{b_{p \sigma}, b_{p^{\prime} \sigma^{\prime}}^{\dagger}\right\}=(2 \pi)^{3} \delta^{(+, \perp)}\left(p-p^{\prime}\right) \delta_{\sigma \sigma^{\prime}} \tag{A11}
\end{equation*}
$$

$\psi_{p \sigma}(x)$ are the positive-energy Volkov wave functions (5), and $\psi_{p \sigma}^{(-)}(x)$ are the negative energy ones:

$$
\begin{equation*}
\psi_{p \sigma}^{(-)}(x)=\frac{K_{-p}(\phi) v_{p \sigma}}{\sqrt{2 p^{+}}} \mathrm{e}^{i S_{-p}(x)}, \tag{A12}
\end{equation*}
$$

with the free Dirac bispinor $v_{p \sigma}$ defined such that $\bar{v}_{p \sigma} v_{p \sigma^{\prime}}=-2 m \delta_{\sigma \sigma^{\prime}}, \bar{v}_{p \sigma} \gamma^{\mu} v_{p \sigma^{\prime}}=2 p^{\mu} \delta_{\sigma \sigma^{\prime}}, \sum_{\sigma} v_{p \sigma} \bar{v}_{p \sigma}=\gamma p-m$ [85].
The quantized part $\mathcal{A}^{\mu}(x)$ of the photon field is represented in the same way, as in the vacuum case [75-77]:

$$
\begin{equation*}
\mathcal{A}^{\mu}(x)=\sum_{\lambda} \int \frac{\tilde{\mathrm{d}}^{3} k}{(2 \pi)^{3}}\left[c_{k \lambda} \phi_{k \lambda}^{\mu}(x)+c_{k \lambda}^{\dagger} \phi_{k \lambda}^{* \mu}(x)\right] \tag{A13}
\end{equation*}
$$

where the creation and annihilation operators obey the relation

$$
\begin{equation*}
\left[c_{k \lambda}, c_{k^{\prime} \lambda^{\prime}}^{\dagger}\right]=(2 \pi)^{3} \delta^{(+, \perp)}\left(k-k^{\prime}\right) \delta_{\lambda \lambda^{\prime}} \tag{A14}
\end{equation*}
$$

and $\phi_{k \lambda}^{\mu}(x)$ is given by

$$
\begin{equation*}
\phi_{k \lambda}^{\mu}(x)=\frac{\epsilon_{k \lambda}^{\mu}}{\sqrt{2 k^{+}}} \mathrm{e}^{-i k x} \tag{A15}
\end{equation*}
$$

with the polarization four-vectors $\epsilon_{k \lambda}^{\mu}$ satisfying the conditions

$$
\begin{equation*}
\epsilon_{k \lambda}^{\mu} \epsilon_{k \lambda^{\prime} \mu}^{*}=-\delta_{\lambda \lambda^{\prime}}, \quad k_{\mu} \epsilon_{k \lambda}^{\mu}=0, \quad \sum_{\lambda} \epsilon_{k \lambda}^{\mu} \epsilon_{k \lambda}^{* \nu}=-g^{\mu \nu}+\frac{\eta^{\mu} k^{\nu}+\eta^{\nu} k^{\mu}}{k^{+}} \tag{A16}
\end{equation*}
$$

## Appendix B: Wave packets

A positive-energy wave packet $\Psi_{p}(x)$ with the central momentum $p^{\mu}$ (the polarization degree of freedom is suppressed) is constructed according to Eq. (13). The density $f_{p}(q)$ is defined such that $\Psi_{p}(x)$ is normalized to one particle:

$$
\begin{equation*}
\int \mathrm{d}^{2} x^{\perp} \mathrm{d} x^{-}\left|\Psi_{p}(x)\right|^{2}=\int \frac{\tilde{\mathrm{d}}^{3} q}{(2 \pi)^{3}}\left|f_{p}(q)\right|^{2}=1 \tag{B1}
\end{equation*}
$$

The four-current density is defined as $J_{e}^{\mu}(x)=\bar{\Psi}_{p}(x) \gamma^{\mu} \Psi_{p}(x)$ [85]. By assuming that $f_{p}(q)$ is peaked around the four-momentum $p^{\mu}$ and by taking into account that the bispinor part of the wave packet is slowly varying with $q^{\mu}$, we obtain that

$$
\begin{equation*}
J_{e}^{\mu}(x) \approx\left|f_{p}(x)\right|^{2} \frac{\pi_{e, p}^{\mu}(\phi)}{p^{+}} \tag{B2}
\end{equation*}
$$

The subscript $e$ denotes the electron current density, and $\pi_{e, p}^{\mu}(\phi)=\pi_{p}^{\mu}(\phi)$ [see Eq. (7)].
For positrons, one obtains that

$$
\begin{equation*}
J_{-e}^{\mu}(x) \approx\left|f_{p}(x)\right|^{2} \frac{\pi_{-e, p}^{\mu}(\phi)}{p^{+}} \tag{B3}
\end{equation*}
$$

where

$$
\begin{equation*}
\pi_{-e, p}^{\mu}(\phi)=-\pi_{-p}^{\mu}(\phi) \tag{B4}
\end{equation*}
$$

Physically, the quantity $\left|f_{p}(x)\right|^{2}$ has a particularly transparent form in the considered case of a narrow wave packet in momentum space. We indicate as (we focus on the positive-energy states, for the negative-energy ones all considerations are analogous)

$$
\begin{equation*}
h_{p}(x)=\int \frac{\tilde{\mathrm{d}}^{3} q}{(2 \pi)^{3}} f_{p}(q) \exp (-i q x) \tag{B5}
\end{equation*}
$$

the asymptotic form of the particle density for $\phi \rightarrow-\infty$, where the field-dependent part of the phase vanishes. By expanding the phase $q x$ up to leading order in $\boldsymbol{q}^{\perp}-\boldsymbol{p}^{\perp}$ and $q^{+}-p^{+}$, one neglects the spreading of the wave packet and it is easy to see that if the function $h_{p}(x)$ is peaked at $x^{+}=0$ around the point $\boldsymbol{x}^{\perp}=\mathbf{0}$ and $x^{-}=0$, then for a generic $x^{+}$it will be peaked at $\boldsymbol{x}^{\perp}=\boldsymbol{p}^{\perp} x^{+} / p^{+}$and $x^{-}=\left(m^{2}+\boldsymbol{p}^{\perp 2}\right) x^{+} / 2 p^{+2}$, as expected for a freely-propagating wave packet. By carrying out the same calculation with the full wave packet $f_{p}(x)$ [see Eq. (21)], one obtains

$$
\begin{equation*}
\left|f_{p}(x)\right|^{2}=\left|F_{p}(x)\right|^{2} \approx\left|\int \frac{\tilde{\mathrm{~d}}^{3} q}{(2 \pi)^{3}} f_{p}(q) \exp \left[-i(q-p) x-i \nabla_{p}^{\perp} \mathcal{S}_{p}(\phi)\left(\boldsymbol{q}^{\perp}-\boldsymbol{p}^{\perp}\right)-i \partial_{p}^{+} \mathcal{S}_{p}(\phi)\left(q^{+}-p^{+}\right)\right]\right|^{2} \tag{B6}
\end{equation*}
$$

Now, by recalling that the phase of a positive-energy Volkov state corresponds to the classical action of an electron in the corresponding plane wave, one obtains that $\left|f_{p}(x)\right|^{2} \approx\left|h_{p}\left(x_{p}\right)\right|^{2}$, where $x_{p}^{\mu}=\left[0, x_{p}^{-}\left(x^{+}\right), \boldsymbol{x}_{p}^{\perp}\left(x^{+}\right)\right]$, with

$$
\begin{align*}
& x_{p}^{-}\left(x^{+}\right)=x^{-}-\frac{m^{2}+\boldsymbol{p}^{\perp 2}}{2 p^{+2}} x^{+}+\int_{-\infty}^{\phi} d \beta\left(\frac{e \boldsymbol{p}^{\perp} \boldsymbol{A}^{\perp}(\beta)}{p^{+2}}-\frac{e^{2} \boldsymbol{A}^{\perp 2}(\beta)}{2 p^{+2}}\right)  \tag{B7}\\
& \boldsymbol{x}_{p}^{\perp}\left(x^{+}\right)=\boldsymbol{x}^{\perp}-\frac{\boldsymbol{p}^{\perp}}{p^{+}} x^{+}+\frac{e}{p^{+}} \int_{-\infty}^{\phi} d \beta \boldsymbol{A}^{\perp}(\beta) \tag{B8}
\end{align*}
$$

which indicates that the function $\left|f_{p}(x)\right|^{2}$ is centered around the classical trajectory of the electron in the plane wave under consideration.

## Appendix C: Conditions for the approximations for the wave packets

Here we provide a discussion about the approximations given in Eqs. (22) and (26).

For the case of a plane-wave field $A^{\mu}\left(k_{0} x\right)$, since the dependence of the field on $x^{\mu}$ is only via the light-cone time $x^{+}$, the conditions for the approximations for the components $x^{-}, \boldsymbol{x}^{\perp}$, that one needs to make in order to obtain the final expression (28), are ultimately the same as in vacuum, i.e., related only to the resolution of the detector and the widths of the wave packets.

In order to get an idea about the conditions for the light-cone time in Eq. (22), i.e., for the variable $\delta_{1}^{+}$, let us first consider the approximation $F_{1}\left(X_{1}-\delta_{1} / 2\right) \approx F_{1}\left(X_{1}\right)$ formally (for the positron wave-packet the consideration below proceeds analogously). Let us assume to work in the highly nonlinear regime, i.e., $\xi \gg 1$. Then the integrand in Eq. (20) is highly oscillating, and the first-order process is expected to form on the scale $m\left|\delta_{1}^{+}\right| \sim 1 / \xi \ll 1[8]$. Requiring the correction to the phase in Eq. (20) due to $\delta_{1}^{+}$to be small, and keeping only linear terms in $\delta_{1}^{+}$and in the widths $\Delta p_{1}^{+}$and $\Delta \boldsymbol{p}_{1}^{\perp}$ of the wave packet, one arrives at the following condition:

$$
\begin{equation*}
\left|\frac{\boldsymbol{\pi}_{1}^{\perp}\left(\Phi_{1}\right) \Delta \boldsymbol{p}_{1}^{\perp}}{2 p_{1}^{+}} \delta_{1}^{+}-\frac{m^{2}+\boldsymbol{\pi}_{1}^{\perp 2}\left(\Phi_{1}\right)}{4 p_{1}^{+2}} \Delta p_{1}^{+} \delta_{1}^{+}\right| \ll 1 \tag{C1}
\end{equation*}
$$

where $\Phi_{1}=m X_{1}^{+}$. Estimating $\left|\boldsymbol{\pi}_{1}^{\perp}\left(\Phi_{1}\right)\right|$ to be $\sim m \xi$ and assessing each term separately, we obtain that the conditions are

$$
\begin{equation*}
\left|\Delta \boldsymbol{p}_{1}^{\perp}\right| \ll \frac{m \chi_{1}}{\xi}, \quad \Delta p_{1}^{+} \ll \frac{p_{1}^{+} \chi_{1}}{\xi^{2}} \tag{C2}
\end{equation*}
$$

where $\chi_{1}=p_{1}^{+} \xi / m$. One notices that for typical values $\chi_{1} \lesssim 1$ the derived conditions can be quite restrictive.
We point out, however, that the above relations are generic, in the sense that one requires that the neglected phase is always much smaller than unity independently of the relevant integration regions. It implies that they rather apply to differential probabilities, i.e., not integrated over the final photon momenta. The actual conditions for the total probability are expected to be less restrictive. In order to see this, let us consider a subprocess of the reaction $e^{-} e^{+} \Rightarrow \gamma \gamma$, in particular, nonlinear Compton scattering $e^{-} \Rightarrow e^{-} \gamma$. Assuming the initial electron to be described by the wave packet distribution $f_{1}\left(q_{1}\right)$, and replacing $q_{1}^{\mu}$ with the central momentum $p_{1}^{\mu}$ in slowly varying functions, one again encounters the necessity of making the approximation (22). However, if we proceed with the evaluation without making this approximation, we obtain that the total probability is given by

$$
\begin{equation*}
W_{e^{-} \Rightarrow e^{-} \gamma} \approx \int \frac{\mathrm{d} k_{1}^{+}}{2 \pi} \frac{\tilde{\mathrm{~d}}^{3} q_{1}^{\prime}}{(2 \pi)^{3}} \frac{\tilde{\mathrm{~d}}^{3} q_{1}}{(2 \pi)^{3}} \int \mathrm{~d}^{4} X_{1} \mathrm{~d} \delta_{1}^{+} f_{1}\left(q_{1}\right) \exp \left[i S_{q_{1}}\left(X_{1}\right)\right] f_{1}^{*}\left(q_{1}^{\prime}\right) \exp \left[-i S_{q_{1}^{\prime}}\left(X_{1}\right)\right]\{\ldots\} \exp \left[i \Phi_{q_{1}}\left(X_{1}^{+}, \delta_{1}^{+}\right)\right] \tag{C3}
\end{equation*}
$$

where the combination $\{\ldots\}$ indicates a prefactor, which is nonessential for the considerations below, and where

$$
\begin{equation*}
\Phi_{q_{1}}\left(X_{1}^{+}, \delta_{1}^{+}\right)=\frac{m k_{1}^{+}}{q_{1}^{+}\left(q_{1}^{+}-k_{1}^{+}\right)} \varphi\left(X_{1}^{+}, \delta_{1}^{+}\right) \tag{C4}
\end{equation*}
$$

with $k_{1}^{+}$being the ' + ' momentum component of the final photon and $\varphi\left(X_{1}^{+}, \delta_{1}^{+}\right)$being a function independent of the momenta. One can approximate $\Phi_{q_{1}}\left(X_{1}^{+}, \delta_{1}^{+}\right)$with $\Phi_{p_{1}}\left(X_{1}^{+}, \delta_{1}^{+}\right)$, if $\varphi\left(X_{1}^{+}, \delta_{1}^{+}\right)$is fastly oscillating and $\Delta p_{1}^{+} \ll$ $p_{1}^{+}, p_{1}^{+}-k_{1}^{+}$(note that $k_{1}^{+}$is an integration variable, therefore, strictly speaking, one needs to limit the integration interval for $k_{1}^{+}$, or consider the differential probability, or ensure that the interval with $\Delta p_{1}^{+} \sim p_{1}^{+}-k_{1}^{+}$does not give a significant contribution, which in general has to be done numerically). Then one arrives at

$$
\begin{equation*}
W_{e^{-} \Rightarrow e^{-} \gamma} \approx \int \mathrm{d}^{4} X_{1}\left|f_{1}\left(X_{1}\right)\right|^{2} W_{e^{-} \Rightarrow e^{-} \gamma}\left(X_{1}^{+}\right) \tag{C5}
\end{equation*}
$$

i.e., the result obtained if the approximation (22) is made in the first place. Note, however, that the conditions of the validity of Eq. (C5) are different from the ones given in Eq. (C2) as the only required conditions (after integrating over the transverse momenta of the final electron and photon) are on $\Delta p_{1}^{+}$, they do not depend on $\xi$, and, indeed, are less restrictive than those in Eq. (C2). It is also worth pointing out that for the reaction $e^{-} \Rightarrow e^{-} \gamma$ different momentum components of the wave packet in fact do not interfere with each other (due to the momentum conservation relations) and the total probability (C3) can be simplified to the average of the probability for a definite-momentum initial state over the modulus squared $\left|f_{1}\left(q_{1}\right)\right|^{2}[17]$.

By considering the full process $e^{-} e^{+} \Rightarrow \gamma \gamma$, one should expect to be able to relax the formal conditions (C2) as well. We conclude that the approximation (22) (and the analogous one for the positron wave-packet amplitudes) should be understood as an effective one, i.e., arising from the consideration of the total probability (19). In this sense, in order to obtain the true conditions of the validity of the approximation (22) one needs to perform the (numerical) evaluation of the whole expression with the wave packets.

The approximation in Eq. (26) is qualitatively different than that in Eq. (22), since it is an approximation for the particle densities (which are classical concepts), rather than for the wave packets themselves. However, it can be related to the approximation (22). For assessing the conditions for the approximation in Eq. (26), we refer to Eq. (58). The second $\theta$-function in this equation establishes a connection between $\delta^{+}$and $\delta_{1}^{+}, \delta_{2}^{+}$. It can be written as

$$
\begin{equation*}
\left[\theta\left(1-\left|\frac{\delta_{2}^{+}+\delta_{1}^{+}}{2 \delta^{+}}\right|\right)-1\right]+1 \tag{C6}
\end{equation*}
$$

where the first term corresponds to the one-step contribution, and the second one to the two-step contribution (for details see Sec. IV). Then, one sees that the one-step contribution is nonzero only if $\left|\delta^{+}\right|<\left|\delta_{2}^{+}+\delta_{1}^{+}\right| / 2$, therefore, the approximation (26) should be acceptable if the electron and the positron densities do not change significantly over the phase $m\left|\delta^{+}\right|<m\left|\delta_{2}^{+}+\delta_{1}^{+}\right| / 2 \ll 1 / \xi$. In the two-step contribution, on the other hand, no limit on $\delta^{+}$is imposed, and consequently $m \delta^{+}$can in principle be of the order of the total laser phase. Hence, if we are to employ the cross section (33), in general, we need to restrict ourselves to the evaluation of the one-step contribution alone, unless we consider incoming wave packets, which are broader in configuration space than the laser pulse.

We emphasize the semi-quantitative nature of the above considerations restating the importance of performing numerical calculations in order to ascertain precisely the conditions under which the approach based on the local cross section in Eq. (33) is applicable.

## Appendix D: Traces

The initial traces for the four terms, constituting the direct-direct part of the cross section, are given by

$$
\begin{gather*}
\mathcal{M}^{\mathrm{nndd}}=\frac{1}{8 m^{4}} \operatorname{Tr}\left\{\rho_{-2} K_{-p_{2} p_{3}}^{\varkappa}\left(\phi_{2}\right)\left(\gamma \tilde{p}_{3}+m\right) K_{p_{3} p_{1}}^{\lambda}\left(\phi_{1}\right) \rho_{1} K_{p_{1} p_{3}}^{\mu}\left(\phi_{1}^{\prime}\right)\left(\gamma \tilde{p}_{3}+m\right) K_{p_{3},-p_{2}}^{\nu}\left(\phi_{2}^{\prime}\right)\right\} g_{\varkappa \nu} g_{\lambda \mu},  \tag{D1}\\
\mathcal{M}^{\text {nidd }}=\frac{1}{2 m^{2}+s(\phi)} \operatorname{Tr}\left\{\rho_{-2} K_{-p_{2} p_{3}}^{\varkappa}\left(\phi_{2}\right)\left(\gamma \tilde{p}_{3}+m\right) K_{p_{3} p_{1}}^{\lambda}\left(\phi_{1}\right) \rho_{1} \frac{K_{p_{1},-p_{2}}^{\mu \nu}\left(\phi_{1}^{\prime}, \phi_{2}^{\prime}\right)}{2 p_{3}^{+}}\right\} g_{\varkappa \nu} g_{\lambda \mu}  \tag{D2}\\
\mathcal{M}^{\text {indd }}=\frac{1}{2 m^{2}+s(\phi)} \operatorname{Tr}\left\{\rho_{-2} \frac{K_{-p_{2} p_{1}}^{\varkappa \lambda}\left(\phi_{2}, \phi_{1}\right)}{2 p_{3}^{+}} \rho_{1} K_{p_{1} p_{3}}^{\mu}\left(\phi_{1}^{\prime}\right)\left(\gamma \tilde{p}_{3}+m\right) K_{p_{3},-p_{2}}^{\nu}\left(\phi_{2}^{\prime}\right)\right\} g_{\varkappa \nu} g_{\lambda \mu}  \tag{D3}\\
\mathcal{M}^{\text {iidd }}=-\operatorname{Tr}\left\{\rho_{-2} \frac{K_{-p_{2} p_{1}}^{\varkappa \lambda}\left(\phi_{2}, \phi_{1}\right)}{2 p_{3}^{+}} \rho_{1} \frac{K_{p_{1},-p_{2}}^{\mu \nu}\left(\phi_{1}^{\prime}, \phi_{2}^{\prime}\right)}{2 p_{3}^{+}}\right\} g_{\varkappa \nu} g_{\lambda \mu} \tag{D4}
\end{gather*}
$$

[note that $\phi_{1}^{\prime}=\phi_{2}^{\prime}$ and $\phi_{1}=\phi_{2}$ for the 'nidd' and 'indd' terms, respectively, see Eqs. (45) and (46), and both relations are valid for the 'iidd' term, see Eq. (47)]. In principle, the traces can be evaluated with the use of the standard techniques [85]. Alternative approaches have also been suggested [81, 84]. The results can be written in a manifestly Lorentz-invariant form [81]:

$$
\begin{align*}
\mathcal{M}^{\mathrm{nndd}}= & \frac{-1}{4 p_{2}^{+} p_{1}^{+} p_{3}^{+2} m^{4}}\left[\frac{1}{2}\left(p_{1}^{+2}+p_{3}^{+2}\right) \Delta_{1}^{2}+2 k_{1}^{+} p_{1}^{+} k_{1} Z_{1}-2 p_{1}^{+} p_{3}^{+} m^{2}\right]\left[\frac{1}{2}\left(p_{2}^{+2}+p_{3}^{+2}\right) \Delta_{2}^{2}-2 k_{2}^{+} p_{2}^{+} k_{2} Z_{2}+2 p_{2}^{+} p_{3}^{+} m^{2}\right] \\
& +\frac{2 m^{2} k_{2}^{+} k_{1}^{+}}{p_{3}^{+2}} \Delta_{1} \Delta_{2}+\frac{2\left(p_{1}^{+}+p_{3}^{+}\right)\left(p_{3}^{+}-p_{2}^{+}\right)}{p_{2}^{+} p_{1}^{+} p_{3}^{+2}}\left[\Delta_{1} \Delta_{2}\left(k_{2}^{+} k_{1}^{+} Z_{1} Z_{2}-k_{2}^{+} p_{1}^{+} k_{1} Z_{2}+k_{1}^{+} p_{2}^{+} k_{2} Z_{1}-p_{2}^{+} p_{1}^{+} k_{2} k_{1}\right)\right. \\
& \left.+k_{2}^{+} p_{1}^{+} k_{1} \Delta_{2} \Delta_{1} Z_{2}-k_{1}^{+} p_{2}^{+} k_{2} \Delta_{1} Z_{1} \Delta_{2}-k_{2}^{+} k_{1}^{+} \Delta_{1} Z_{2} Z_{1} \Delta_{2}+p_{2}^{+} p_{1}^{+} k_{1} \Delta_{2} k_{2} \Delta_{1}\right], \tag{D5}
\end{align*}
$$

$$
\begin{align*}
\mathcal{M}^{\text {nidd }}=\frac{2}{p_{3}^{+2}\left[2 m^{2}+s(\phi)\right]}[ & m^{2}\left(p_{3}^{+}-k_{1}^{+}\right)\left(p_{3}^{+}+k_{2}^{+}\right)+\frac{1}{4}\left(p_{1}^{+}+p_{3}^{+}\right)\left(p_{3}^{+}-p_{2}^{+}\right) \Delta_{1} \Delta_{2} \\
& -\frac{1}{2} k_{2}^{+}\left(p_{1}^{+}+p_{3}^{+}\right) \Delta_{1} Z_{2}-\frac{1}{2} k_{1}^{+}\left(p_{3}^{+}-p_{2}^{+}\right) Z_{1} \Delta_{2} \\
& -\frac{1}{2} p_{2}^{+}\left(p_{1}^{+}+p_{3}^{+}\right) k_{2} \Delta_{1}+\frac{1}{2} p_{1}^{+}\left(p_{3}^{+}-p_{2}^{+}\right) k_{1} \Delta_{2} \\
& \left.+k_{2}^{+} k_{1}^{+} Z_{1} Z_{2}-k_{2}^{+} p_{1}^{+} k_{1} Z_{2}+k_{1}^{+} p_{2}^{+} k_{2} Z_{1}-p_{2}^{+} p_{1}^{+} k_{2} k_{1}\right] \tag{D6}
\end{align*}
$$

where

$$
\begin{equation*}
\Delta_{1}^{\mu}=\Delta_{p_{1}}^{\mu}\left(\phi_{1}^{\prime}, \phi_{1}\right), \quad Z_{1}^{\mu}=Z_{p_{1}}^{\mu}\left(\phi_{1}^{\prime}, \phi_{1}\right), \quad \Delta_{2}^{\mu}=\Delta_{-p_{2}}^{\mu}\left(\phi_{2}, \phi_{2}^{\prime}\right), \quad Z_{2}^{\mu}=Z_{-p_{2}}^{\mu}\left(\phi_{2}, \phi_{2}^{\prime}\right) \tag{D8}
\end{equation*}
$$

with

$$
\begin{equation*}
\Delta_{p}^{\mu}\left(\phi, \phi^{\prime}\right)=\pi_{p}^{\mu}(\phi)-\pi_{p}^{\mu}\left(\phi^{\prime}\right), \quad Z_{p}^{\mu}\left(\phi, \phi^{\prime}\right)=\left[\pi_{p}^{\mu}(\phi)+\pi_{p}^{\mu}\left(\phi^{\prime}\right)\right] / 2 \tag{D9}
\end{equation*}
$$

(note that, in order to be consistent, we should have used $\Delta_{-2}^{\mu}$ and $Z_{-2}^{\mu}$, but for clarity the minus signs are suppressed). In Eq. (D5) a combination of four four-vectors stands for the product of two scalar products, e.g., $k_{1} \Delta_{2} k_{2} \Delta_{1}=$ $\left(k_{1} \cdot \Delta_{2}\right)\left(k_{2} \cdot \Delta_{1}\right)$ and analogously for the other combinations.

The results in Eqs. (D5), (D6), and (D7) can be cast into a more convenient form with the use of momentum relations for the dressed momenta. First, we notice that, since ' + ' and ' $\perp$ ' momentum components are conserved in the plane wave, the relations

$$
\begin{equation*}
\left[\pi_{p^{\prime}}(\phi)+k-\pi_{p}(\phi)\right]^{(+, \perp)}=0 \tag{D10}
\end{equation*}
$$

hold, where $k^{\mu}$ and $\pi_{p}(\phi)^{\mu}$ are the photon and fermion four-momenta, respectively, which come into the point $x^{\mu}$ and $\pi_{p^{\prime}}(\phi)$ is the outgoing fermion four-momentum. For an analogous combination of the '-' components, in each vertex we have the integral

$$
\begin{equation*}
\int \mathrm{d} x^{+}\left[\pi_{p^{\prime}}^{-}(\phi)+k^{-}-\pi_{p}^{-}(\phi)\right] \mathrm{e}^{i \Phi\left(x^{+}\right)}=-i \int \mathrm{~d} x^{+} \partial_{+}\left[\mathrm{e}^{i \Phi\left(x^{+}\right)}\right] \tag{D11}
\end{equation*}
$$

where $\Phi\left(x^{+}\right)=\left(p^{\prime-}+k^{-}-p^{-}\right) x^{+}+\mathcal{S}_{p^{\prime}}(\phi)-\mathcal{S}_{p}(\phi)$. Assuming that the boundary terms must not affect observables, we obtain the full four-momentum conservation law (see [14, 27, 58, 66, 84] for similar considerations)

$$
\begin{equation*}
\pi_{p^{\prime}}^{\mu}(\phi)+k^{\mu}-\pi_{p}^{\mu}(\phi)=0 \tag{D12}
\end{equation*}
$$

which, strictly speaking, holds only inside the integral in $x^{+}$. With the use of Eq. (D12), one can derive the following momentum relations [81]:

$$
\begin{equation*}
k \pi_{p^{\prime}}(\phi)=\frac{1}{2}\left(p^{2}-{p^{\prime}}^{2}-k^{2}\right), \quad k \pi_{p}(\phi)=-\frac{1}{2}\left({p^{\prime}}^{2}-k^{2}-p^{2}\right), \quad \pi_{p^{\prime}}(\phi) \pi_{p}(\phi)=-\frac{1}{2}\left(k^{2}-{p^{\prime}}^{2}-p^{2}\right) \tag{D13}
\end{equation*}
$$

The relations (D13) allow one to extract instantaneous parts, i.e., terms $\propto\left(p_{3}^{2}-m^{2}\right)$ and $\left(p_{3}^{\prime 2}-m^{2}\right)$ from the 'nndd' contribution (D5) and include them into the 'indd' and 'nidd' contributions, respectively. Subsequently, the instantaneous parts can be extracted from the 'nidd' and 'indd' contributions and combined with the 'iidd' contribution. These rearrangements are significantly simplified if one employs the coordinate system, defined by Eqs. (2) and (3). The result is (see [81] for details)

$$
\begin{gather*}
\tilde{\mathcal{M}}^{\text {ndd }}=-\frac{1}{4 p_{2}^{+} p_{1}^{+} p_{3}^{+2} m^{4}}\left[\frac{1}{2}\left(p_{1}^{+2}+p_{3}^{+2}\right) \boldsymbol{\Delta}_{1}^{\perp 2}+2 p_{1}^{+} p_{3}^{+} m^{2}\right]\left[\frac{1}{2}\left(p_{2}^{+2}+p_{3}^{+2}\right) \boldsymbol{\Delta}_{2}^{\perp 2}-2 p_{2}^{+} p_{3}^{+} m^{2}\right]+\frac{k_{2}^{+} k_{1}^{+}}{4 p_{3}^{+2} m^{2}} \boldsymbol{\Delta}_{1}^{\perp} \boldsymbol{\Delta}_{2}^{\perp} \\
+\frac{\left(p_{1}^{+}+p_{3}^{+}\right)\left(p_{3}^{+}-p_{2}^{+}\right)}{8 p_{2}^{+2} p_{1}^{++} p_{3}^{+2} m^{4}}\left\{\boldsymbol{\Delta}_{1}^{\perp} \boldsymbol{\Delta}_{2}^{\perp}\left[\mathcal{Z}^{\perp 2}+k_{2}^{+2} p_{1}^{+2}\left(m^{2}+\frac{1}{4} \boldsymbol{\Delta}_{2}^{\perp 2}\right)+k_{1}^{+2} p_{2}^{+2}\left(m^{2}+\frac{1}{4} \boldsymbol{\Delta}_{1}^{\perp 2}\right)\right]\right. \\
 \tag{D14}\\
\left.\quad-2 \mathcal{Z}^{\perp} \boldsymbol{\Delta}_{1}^{\perp} \mathcal{Z}^{\perp} \boldsymbol{\Delta}_{2}^{\perp}\right\},
\end{gather*}
$$

$$
\begin{align*}
& \tilde{\mathcal{M}}^{\text {nidd }}=\frac{1}{p_{2}^{+} p_{1}^{+} p_{3}^{+2}\left[2 m^{2}+s(\phi)\right]}\left\{p_{3}^{+2}\left[2 p_{2}^{+} p_{1}^{+}+\left(p_{2}^{+}+p_{1}^{+}\right)^{2}\right] m^{2}+p_{3}^{+2} p_{2}^{+} p_{1}^{+}\left(\boldsymbol{\Delta}_{2}^{\perp 2}+\boldsymbol{\Delta}_{1}^{\perp 2}\right)\right. \\
&+\left[\mathcal{Z}^{\perp}+\frac{1}{2} p_{3}^{+}\left(p_{1}^{+}+p_{3}^{+}\right) \boldsymbol{\Delta}_{1}^{\perp}+\frac{1}{2} p_{3}^{+}\left(p_{3}^{+}-p_{2}^{+}\right) \boldsymbol{\Delta}_{2}^{\perp}\right]^{2} \\
&\left.-\left[\frac{1}{2} k_{2}^{+}\left(p_{1}^{+}+p_{3}^{+}\right) \boldsymbol{\Delta}_{1}^{\perp}-\frac{1}{2} k_{1}^{+}\left(p_{3}^{+}-p_{2}^{+}\right) \boldsymbol{\Delta}_{2}^{\perp}\right]^{2}\right\},  \tag{D15}\\
& \tilde{\mathcal{M}}^{\text {indd }}=\left.\tilde{\mathcal{M}}^{\text {nidd }}\right|_{\boldsymbol{\Delta}_{1}^{\perp} \rightarrow-\boldsymbol{\Delta}_{1}^{\perp}, \boldsymbol{\Delta}_{2}^{\perp} \rightarrow-\boldsymbol{\Delta}_{2}^{\perp}}, \quad \tilde{\mathcal{M}}^{\text {iidd }}=-2, \tag{D16}
\end{align*}
$$

where

$$
\begin{equation*}
\mathcal{Z}^{\perp}=k_{2}^{+} p_{1}^{+} \boldsymbol{Z}_{2}^{\perp}-k_{1}^{+} p_{2}^{+} \boldsymbol{Z}_{1}^{\perp} \tag{D17}
\end{equation*}
$$

Note that the final expressions do not depend on the vector $\boldsymbol{k}_{1}^{\perp}$. This facilitates the analytical evaluation of the integral in this variable, as we have mentioned in the main text.

## Appendix E: Integrals for the zero-field limit

Here, we present the evaluation of the integrals, given in Eqs. (76) and (77), for the case of a vanishing laser field. The phase $\Phi_{v}^{\mathrm{dd}}$ is given in Eq. (84).

The integral in $\rho$ evaluates to a Bessel function of first kind, in particular [104],

$$
\begin{equation*}
\int \frac{\mathrm{d} \rho}{\rho} \exp \left(\frac{i}{\rho}+\frac{i}{4} a^{2} v^{2} \tau^{2} \rho\right)=2 i \pi J_{0}(a v \tau) \tag{E1}
\end{equation*}
$$

For the integrals in $\tau$, formally, one needs to recover the $i \epsilon$ prescription, in order to make them convergent at infinity. On the other hand, we can rotate the integration contour clockwise by $\pi / 2$ and then make replacement $\tau \rightarrow-i \tau$, after that the $i \epsilon$ prescription is not necessary (note that $b>a$ ). One obtains [104]

$$
\begin{equation*}
\int_{0}^{\infty} \mathrm{d} \tau \tau J_{0}(a v \tau) \mathrm{e}^{-i b v \tau}=-\frac{b}{\left(b^{2}-a^{2}\right)^{3 / 2} v^{2}}, \quad \int_{0}^{\infty} \mathrm{d} \tau J_{0}(a \tau) \mathrm{e}^{-i b \tau}=-\frac{i}{\sqrt{b^{2}-a^{2}}} \tag{E2}
\end{equation*}
$$

The integral in $v$ is elementary in the case of a vanishing external field. The evaluation of the integrals in $k_{1}^{+}$is also straightforward. Afterward, one needs to express the result in terms of $\mu$, which can be written as

$$
\begin{equation*}
\mu=\frac{\left(p_{2}^{+}+p_{1}^{+}\right)^{2}}{4 p_{2}^{+} p_{1}^{+}}\left(1+\boldsymbol{t}^{\perp 2}\right) \tag{E3}
\end{equation*}
$$

We obtain:

$$
\begin{equation*}
\int_{0}^{p_{2}^{+}+p_{1}^{+}} \mathrm{d} k_{1}^{+} \frac{b}{\left(b^{2}-a^{2}\right)^{3 / 2}}=\frac{p_{2}^{+}+p_{1}^{+}}{4 \mu}, \quad \int_{0}^{p_{2}^{+}+p_{1}^{+}} \mathrm{d} k_{1}^{+} \frac{1}{\sqrt{b^{2}-a^{2}}}=\frac{p_{2}^{+}+p_{1}^{+}}{4 \sqrt{\mu(\mu-1)}} \ln \left(\frac{\sqrt{\mu}+\sqrt{\mu-1}}{\sqrt{\mu}-\sqrt{\mu-1}}\right) . \tag{E4}
\end{equation*}
$$

Combining everything together, one recovers the final expressions, presented in Sec. VII.
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