


## Algorithms for Quantum Simulation at Finite Energies

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We introduce two kinds of quantum algorithm to explore microcanonical and canonical properties of many-body systems. The first is a hybrid quantum algorithm that, given an efficiently preparable state, computes expectation values in a finite energy interval around its mean energy. This algorithm is based on a filtering operator, similar to quantum phase estimation, which filters out energies outside the desired energy interval. However, instead of performing this operation on a physical state, it recovers the physical values by performing interferometric measurements without the need to prepare the filtered state. We show that the computational time scales polynomially with the number of qubits, the inverse of the prescribed variance, and the inverse error. In practice, the algorithm does not require the evolution for long times, but instead a significant number of measurements in order to obtain sensible results. Our second algorithm is a quantum assisted Monte Carlo sampling method to compute other quantities that approach the expectation values for the microcanonical and canonical ensembles. Using classical Monte Carlo techniques and the quantum computer as a resource, this method circumvents the sign problem that plagues classical quantum Monte Carlo simulations, as long as one can prepare states with suitable energies. All algorithms can be used with small quantum computers and analog quantum simulators, as long as they can perform the interferometric measurements. We also show that this last task can be greatly simplified at the expense of performing more measurements.

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### I. INTRODUCTION

The advent of quantum simulators [1,2] opens many exciting opportunities to probe and understand fundamental problems in physics, ranging from condensed matter to high-energy physics and quantum chemistry [3–5]. Feynman’s original proposal in 1982 was to build a universal digital quantum computer that can imitate any physical system. Although tremendous progress has been made, building a universal quantum computer that will fulfill Feynman’s vision is still a long-term task. However, both near-term (noisy) quantum computers and analog quantum simulators can already help us to address some of those problems. The latter, where the interaction is engineered directly according to the physical Hamiltonian under investigation, are particularly advanced in different platforms, like cold atoms in optical lattices [6,7], trapped ions [8], Rydberg atoms [9], quantum dots [10],

superconductors [11], photons [12], etc. In particular, very controlled experiments can be carried out with around 50 qubits [13–26] and it is expected that this number will be significantly increased in the coming years.

There are many questions that seek answers from quantum simulators, especially physical properties of ground, nonequilibrium, and finite temperature states. Most of the theoretical work on quantum simulations has focused on the dynamics of many-body quantum systems, as well as on their properties at zero temperature. Since the first algorithm [2] that showed how the dynamics could be efficiently simulated, large improvements have been achieved, leading to a very economic algorithm [27]. In practice, in analog quantum simulators the dynamics are naturally implemented by letting the system evolve according to the engineered Hamiltonian [3]. For ground state problems, the situation is quite different since determining its properties is very demanding and, in general, it requires exponential time in  $N$ , the number of qubits to be simulated [28]. A quantum simulator can still be of significant help since the corresponding classical simulator requires exponential resources both in time and memory, whereas the quantum one achieves a moderate speed-up albeit with polynomial memory. The first algorithms [29–31] used quantum phase estimation to project onto an eigenstate

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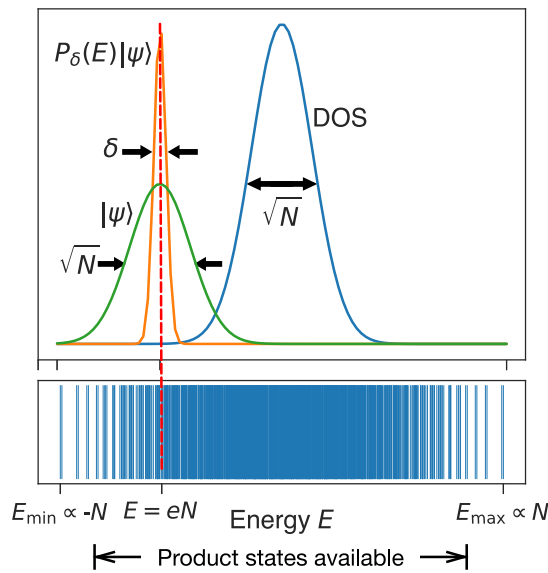
of the Hamiltonian, and have been successively improved [32–35]. In particular, in Ref. [34] a cosine-filtering operator is used to prepare a state close to the ground state with a very small variance. It is similar in its conception to quantum phase estimation, but has a better scaling for that purpose. This idea has also been used in the context of tensor networks [36] to estimate the amount of entanglement required to achieve small energy variances along the whole spectrum of a Hamiltonian. Although all these quantum algorithms were originally designed for scalable quantum computers, proposals to use them with analog quantum simulators have recently been put forward [37]. This can be very convenient for small simulators, as the exponential scaling of the resources still limits their applicability to large systems. Other heuristic algorithms, like adiabatic [38–40] and variational [41,42] strategies, can be very useful and overcome the exponential scaling in certain cases [4].

Quantum algorithms for excited states or finite temperature are more scarce. Chowdhury and Somma [43] showed how to realize the imaginary time evolution operator to produce a Gibbs state, whereas other algorithms

propose sampling techniques [44–46]. Phase estimation can also be directly used to prepare states at different energies, and thus address quantum statistical questions in the microcanonical ensemble. All those algorithms may work well in practice. However, as for classical ones, they require an exponential time in  $N$ , although only polynomial memory resources. Additionally, it remains challenging to implement most of them with the existing small quantum computers or analog quantum simulators.

In this paper we introduce and test two different types of quantum algorithm to determine physical properties in an energy interval or at finite temperature (see Fig. 1 for a graphical summary). The idea underlying our proposal relies on the cosine filter of Refs. [34,36] to target states with small energy variance at selected energies, in which the observables could be measured. Preparing those states may nevertheless be challenging in practice. We overcome this obstacle by showing that observations obtained after running quantum simulators for different stroboscopic evolution times are sufficient to determine the values of interesting quantities, without the need to prepare the filtered state at all. This is in the spirit of the time series approach

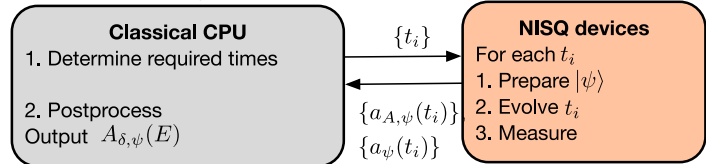
### Filtering $|\psi\rangle$ at finite energy density $e$



### Main results

#### ① Efficient hybrid quantum-classical algorithms for microcanonical observables (Secs. III–IV)

Given an easy to prepare state  $|\psi\rangle$ , an observable  $A$ , and  $\delta$



#### ② Heuristic quantum-assisted Monte Carlo algorithms for microcanonical and canonical observables (Sec. V)

For inverse temperature  $\beta$ , an observable  $A$ , and  $\delta$

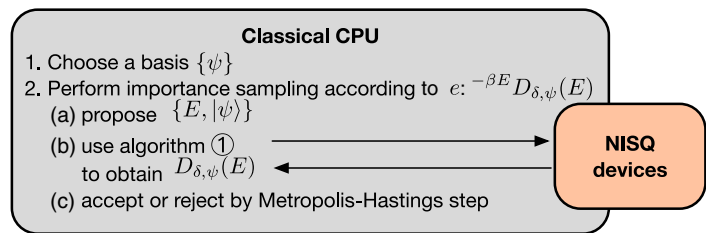


FIG. 1. Graphical summary of the concept and main results of the paper. Left: our algorithms compute properties of a state with narrow energy variance (a schematic spectral distribution is shown in orange), which would result from applying a filter onto an easy to prepare state, e.g., a product state (green curve). For local Hamiltonians, for which the density of states in the thermodynamic limit approaches a Gaussian distribution of width proportional to  $\sqrt{N}$  (blue curve), the accessible energy densities can lie on the tails of the spectrum. The box under the graph shows the actual energy spectrum for a  $N = 10$  Ising chain. Right: the main contribution of our paper is the proposal of two sets of algorithms, schematically summarized in this figure. The first (top) is a provably efficient hybrid quantum-classical algorithm for computing expectation values of observables in filtered states, as illustrated on the left. It is based on repeated preparation, evolution, and measurements on an easy-to-prepare state, run by a quantum device, and postprocessing by classical computation. The second algorithm is a quantum assisted sampling algorithm that can be used to compute microcanonical and canonical properties (summarized at the bottom for the canonical case). This algorithm performs classical Monte Carlo importance sampling where the quantum device is used to determine the sampling probabilities efficiently. It requires shorter coherence time than the first algorithm, at the expense of more measurements.

introduced in Ref. [47] and later used for the related question of estimating the (binned averages of) Hamiltonian eigenvalues [48] (see also extensions based on Hamiltonian querying and Chebyshev expansions [49,50]). Our algorithms explicitly target finite energy properties, such as microcanonical expectation values, for which we can demonstrate efficiency. The quantities required in our method can be obtained with interferometric measurements, which involve the conditional evolution depending on the state of a single qubit, and are especially suited for quantum simulators. More concretely, we consider the following algorithms.

1. Our first algorithm is an efficient hybrid quantum-classical algorithm that targets the physical properties of states in an energy interval around that of any state that can be efficiently prepared (see Fig. 1 for an illustration). We prove that this algorithm can be carried out in time that is polynomial in  $N$ , the inverse error, and the inverse width of the filtering operations. The later is related to the energy variance of the targeted state. To our knowledge, no classical algorithm achieves this polynomial scaling.
2. Our second algorithm combines the quantum simulation with classical Monte Carlo methods, and provides practical methods to obtain both microcanonical and canonical expectation values of observables (see Fig. 1 for an illustration). These quantum assisted Monte Carlo methods use quantum simulators to compute the sampling probabilities that are required in quantum Monte Carlo methods. Remarkably, they circumvent the sign problem [51], the main obstacle of applying such methods to many physics problems, so long as one can prepare (a product or other kind of) states of suitable energies.

Let us briefly mention that our proposal is different from other emerging families of quantum methods that make use of hybrid schemes, such as variational quantum eigensolvers [41,42] and, in particular, hybrid ansatzes [52] or quantum subspace diagonalization [53] (see also Refs. [54–57]). Different to our algorithms, such methods typically target ground state problems and, while some of them [53] also use a superposition of a state evolved to different times, the coefficients of the superposition are variational parameters that need to be optimized, unlike in our proposal, where the coefficients are fixed by the filter. An even more significant difference is that these methods are not proven to be efficient.

Finally, the interferometric methods required for the algorithms presented here have been used in the context of the Loschmidt echo [58] in NMR [59]. They have also been proposed for ions [60], atoms [61], and,

more recently, to perform phase estimation with such systems [37]. We give several alternative procedures to simplify that task that can be applied in different situations. First, we show that one can replace phase estimation by the ability of both preparing catlike states [62] and having access to two additional internal states. This capability already exists in different platforms [26,63–71]. We then show that one can perform a similar procedure but without the requirement of the two additional states. Then, we give a procedure that does not require the preparation of catlike states, but proceeds with a sequence of measurements. Finally, we give an even simpler method that works for Hamiltonians possessing certain symmetries, like  $XY$  or Hubbard models. The last two methods have an advantage with respect to the previous methods since no catlike state needs to be evolved, and thus the method is more resilient against decoherence. However, they require a larger number of measurements.

The structure of this paper is as follows. In Sec. II we introduce the models and the basic idea of the cosine filter. In Sec. III, we present the first algorithm and show how one can use a quantum simulator to efficiently compute certain expectation values around fixed energies. In Sec. IV we give more practical methods for the same purpose and we test them numerically. In Sec. V, we present a family of quantum assisted Monte Carlo algorithms for microcanonical and canonical observables that combine classical Monte Carlo and quantum simulators. We numerically explore their performance and demonstrate that they are robust against certain noises. In the appendices we present the methods to replace interferometric measurements, describe the specific model we used in our numerical study, give details of the proof regarding the polynomial scaling of our algorithm, and investigate how much one has to decrease the variance in order to converge to the microcanonical and canonical results for a nonexactly solvable model.

## II. SETUP AND COSINE FILTERING

### A. Setup

We consider  $N$  spins on a lattice and a Hamiltonian

$$H = \sum_{n=1}^N h_n, \quad (1)$$

and denote by  $E_{\min}$  and  $E_{\max}$  the minimum and maximum eigenvalues of  $H$ . We assume that  $|E_{\min}|, |E_{\max}| < N/2$ , so that the spectrum of  $H$  lies within the interval  $[-N/2, N/2]$ . The Hamiltonian  $H$  could be local in any spatial dimension, i.e., the term  $h_n$  acts only on the  $n$ th lattice site and its neighbors. However, we emphasize that the algorithms proposed here can also be applied to more general setups, where  $h_n$  has long-range interactions or even

with more complicated Hamiltonians, although some of our estimations will rely on the original form (1) or in its locality. The main requirement is that the evolution generated by  $H$  can be efficiently implemented with the quantum simulator. More specifically, given an initial state,  $\psi$ , and an observable  $A$ , one can efficiently determine

$$a_{A,\psi}(t) = \langle \psi | A e^{-iHt} | \psi \rangle, \quad (2a)$$

$$a_\psi(t) = \langle \psi | e^{-iHt} | \psi \rangle, \quad (2b)$$

with a sufficiently small error. Furthermore, if one can measure those quantities, one also has access to

$$a_{A,\psi}(t_1, t_2) = \langle \psi | e^{iHt_1} A e^{-iHt_2} | \psi \rangle, \quad (3)$$

since  $a_{A,\psi}(t_1, t_2) = a_{A,\psi(t_1)}(t_2 - t_1)$ . The value of Eq. (3) can be readily measured using a quantum computer (see, e.g., Ref. [72]). For analog quantum simulators, this may not be possible. In Appendix A we give a series of alternatives to obtain such quantities. In general, by repeating the experiment  $L$  times, the error will be additive and scale as  $L^{-1/2}$ .

### B. Initial states

In order for the simulation algorithm to be efficient, we choose as  $\psi$  a state that can be prepared by the simulator. The simplest is a product state

$$|p\rangle = |p_1, p_2, \dots, p_N\rangle, \quad (4)$$

where the  $p_n$  are normalized states, e.g., for qubits,

$$|p_n\rangle = \cos(\theta_n)|0\rangle + e^{i\varphi_n} \sin(\theta_n)|1\rangle. \quad (5)$$

In general, we denote the mean energy and variance of  $H$  in the state  $\psi$  by

$$E_\psi = \langle \psi | H | \psi \rangle, \quad (6a)$$

$$\sigma_\psi^2 = \langle \psi | (H - E_\psi)^2 | \psi \rangle. \quad (6b)$$

In case  $H$  is local and the state  $\psi$  has finite correlation length, both  $E_\psi$  and  $\sigma_\psi^2$  will scale as  $N$ .

The energies of states that can be efficiently prepared will determine the range of energies that our algorithm can efficiently explore. If we restrict ourselves to product states, the mean energy  $E_p$  does not cover the whole spectrum of  $H$ ; there are energies  $E_{p,\min}$  and  $E_{p,\max}$  such that we can always choose a state  $p$  with  $E_p$  in the interval  $[E_{p,\min}, E_{p,\max}]$  but never outside (see Fig. 1 for an illustration). It has been shown that the range of the interval can be extensive in  $N$  for a local Hamiltonian (see Appendix C for more details). Finding product states within this interval amounts to solving a mean-field problem, and thus can be done efficiently on classical computers. In order to access

energies outside this interval, one can consider other states,  $\psi$ , that are still easy to prepare but can cover a wider range of energies. In particular, we could consider products of spin blocks, matrix product states [73–75], or states obtained through adiabatic evolution or variational methods. Alternatively, the state  $\psi$  could be prepared by starting from a product state and running the quantum simulator with a different Hamiltonian for some time.

### C. Cosine filter

Following Ref. [34], we define the cosine-filtering operator

$$P_\delta(E) = \left[ \cos\left(\frac{H - E}{N}\right) \right]^{\lfloor N^2/\delta^2 \rfloor_2}, \quad (7)$$

where we use  $\lfloor \cdot \rfloor_2$  to indicate the nearest even integer. Here  $\delta$  can take arbitrary values, including decreasing (e.g.,  $\delta \sim 1/N$ ) or constant ( $\delta \sim 1$ ) with  $N$ , which will be considered in the following sections. In order to interpret the action of this operator, it is useful to approximate [34]

$$P_\delta(E) \simeq e^{-(H-E)^2/2\delta^2}, \quad (8)$$

as long as the spectrum of the operator that appears in the argument of the cosine lies in the interval  $[-\pi/2, \pi/2]$  (in fact, this is also true in a larger interval; see Appendix E). Thus, it basically projects out the eigenstates of  $H$  that have an energy  $E'$  with  $|E' - E| \gg \delta$ , and thus acts as a filter around  $E$  [76]. By definition,  $0 < P_\delta(E) \leq 1$ .

As in Ref. [34], we approximate

$$\cos^M(X) \approx \sum_{m=-\lfloor x\sqrt{M} \rfloor}^{\lfloor x\sqrt{M} \rfloor} c_m e^{-i2mX} \quad (9)$$

up to an error (in the operator norm) bounded by  $2e^{-x^2/2}$  for  $\|X\|_\infty \leq 1$ , and where

$$c_m = \frac{1}{2^M} \binom{M}{M/2 - m}. \quad (10)$$

We can use this expansion to express cosine filter (7) in terms of the evolution operator  $e^{-iHt}$  for certain times  $t$ . For  $|E| \leq N/2$ , we take  $X = (H - E)/N$ , so that

$$P_\delta(E) \simeq \sum_{m=-R}^R c_m e^{-i(H-E)t_m}, \quad (11)$$

where

$$R = xN/\delta, \quad t_m = 2m/N. \quad (12)$$

The idea will be, as in Ref. [36], to apply Eq. (9) to certain states in order to filter them around some energy  $E$ , and

then to obtain expectation values of observables with the resulting states. However, instead of preparing the state, we first express the desired values in terms of Eqs. (2), and then use the quantum simulator to measure those values separately. The number of measurements will be  $2R$  times the number of repetitions required to obtain a prescribed accuracy. Each of the runs of the simulator will be for a time  $t \leq 2x/\delta$ . In the end, we perform the multiplications and sum classically.

#### D. Ising model

In order to benchmark our algorithms, we use a rather trivial model for which we can obtain numerical results for values of  $N \sim 100$  qubits, which should be attainable in present or planned quantum simulators. Let us take an even number  $N$  of fermionic modes and a Hamiltonian

$$H = \frac{g}{2} \sum_{n=1}^N (a_n + a_n^\dagger)(a_{n+1} - a_{n+1}^\dagger) + h \sum_n (a_n^\dagger a_n - 1/2), \quad (13)$$

where the  $a_n$  are annihilation operators of the vacuum  $|\text{vac}\rangle$  and we have chosen periodic boundary conditions for the fermions:  $a_{N+1} = a_1$ . Through the Jordan-Wigner transformation, this corresponds to the Ising Hamiltonian

$$H = \frac{g}{2} \sum_{n=1}^N \sigma_{n,x} \sigma_{n+1,x} + \frac{h}{2} \sum_{n=1}^N \sigma_{n,z}, \quad (14)$$

where the  $\sigma_{x,z}$  are Pauli operators, with appropriate boundary conditions.

Defining the operators in the momentum representation

$$b_k = \frac{1}{\sqrt{N}} \sum_{n=1}^N e^{i2\pi kn/N} a_n, \quad (15)$$

where  $k = -N/2 + 1, \dots, N/2$ , we perform some computations with (Fock) states of the form

$$|k\rangle = b_{k_1}^\dagger \cdots b_{k_\ell}^\dagger |\text{vac}\rangle, \quad (16)$$

where  $k_1 < k_2 < \cdots < k_\ell$ . They form an orthonormal basis and, even though they are not eigenstates of  $H$ , they are easy to deal with for large values of  $N$ . In Appendix B we provide the analytical formulas that allow us to obtain exact numerical results for this model. The minimal eigenvalue of  $H$ ,  $E_{\min}$ , and the minimal energy attained by a state of form (16),  $E_{p,\min}$ , can be easily computed with those formulas.

Even though we are interested in energies above  $E_{\min}$  and finite temperatures, we note that at zero temperature Hamiltonian (13) features a phase transition at  $g = h$ . For  $h \gg g$ , the ground state is the vacuum, whereas for  $h \ll g$ , it is a superposition of the vacuum with states where excitations occur in pairs of momenta  $\pm k$ .

### III. EFFICIENT QUANTUM ALGORITHM FOR OBSERVABLES AT FINITE ENERGY

Given a state  $\psi$  and an observable  $A$ , we define

$$A_{\delta,\psi}(E) = \frac{\langle \psi | [AP_\delta(E) + P_\delta(E)A] | \psi \rangle}{2\langle \psi | P_\delta(E) | \psi \rangle}, \quad (17a)$$

$$A'_{\delta,\psi}(E) = \frac{\langle \psi | P_\delta(E)AP_\delta(E) | \psi \rangle}{\langle \psi | P_\delta(E)^2 | \psi \rangle}. \quad (17b)$$

Both quantities are related to the microcanonical expectation value of  $A$ . In particular, if  $\langle E | \psi \rangle \neq 0$ , where  $|E\rangle$  is the eigenstate of  $H$  corresponding to the energy  $E$ , Eq. (17b) converges to that value in the limit  $\delta \rightarrow 0$ .

The numerators and denominators of Eqs. (17) can be expressed in terms of  $a_{A,\psi}(t_n, t_m)$ . Thus, the quantum algorithm uses the quantum simulator to determine those quantities up to the required precision, classically computes  $c_m$  and  $\exp(iEt_m)$ , and then (classically) performs the required sums and multiplications. We will show that both Eqs. (17a) and (17b) can be efficiently computed using a quantum simulator that has access to Eqs. (2). But for that, we have to first explain what we mean by ‘‘efficiently’’ and also formulate the problem more precisely.

We say that a state  $\psi$  can be efficiently prepared if, for any prescribed error  $\epsilon > 0$ , we can obtain a state  $\varphi$  with  $\|\varphi - \psi\|^2 < \epsilon$  in time

$$T = \text{poly}(N, 1/\epsilon). \quad (18)$$

Furthermore, we say that the quantum simulator can efficiently measure  $A$  if it can perform measurements to obtain  $a_{A,\psi}(t', t)$  (with  $t' \leq t$ ) with an error smaller than  $\epsilon$  in time

$$T = \text{poly}(N, t, 1/\epsilon). \quad (19)$$

Note that this basically requires an efficient procedure to evolve according to the Hamiltonian and the possibility of performing interferometric measurements.

*Result.* If a quantum simulator can efficiently prepare  $\psi$  and measure  $A$ , with  $\|A\|_\infty \leq 1$ , then, for any  $\epsilon, \delta > 0$ , one can always find

$$E \in [E_\psi - r\sigma_\psi, E_\psi + r\sigma_\psi] \quad (20)$$

with  $r = \{3 \log[2(1 + 2\sigma_\psi^2/\delta^2)]\}^{1/2}$ , so that one can obtain Eqs. (17) up to an additive error  $\epsilon$  in time

$$T = \text{poly}(N, 1/\delta, 1/\epsilon), \quad (21)$$

including the cost of finding the value  $E$ .

Note that, for  $\delta = \text{poly}(1/N)$ , the result can still be obtained in polynomial time. Furthermore, since  $\sigma_\psi \leq N/2$ ,  $E$  will differ from  $E_\psi$  by at most a constant times  $\sigma_\psi \log^{1/2}(N)$ . If  $\sigma_\psi \propto O(\sqrt{N})$ , as it occurs in states with

finite correlation length and local Hamiltonians, this difference will only scale as  $\sqrt{N \log N}$ . The reader may wonder why we need to introduce an interval, instead of simply fixing  $E$  to some value, for instance  $E = E_\psi$ . The reason is that the spectrum of  $H$  is discrete and it may well be that  $\langle \psi | P_\delta(E) | \psi \rangle$  is exponentially small in  $N$ . As we show in Appendix C, this issue can be avoided if we are allowed to vary  $E$  in a small interval. We are not aware of any classical algorithm that can achieve this scaling.

The result can be proven by expressing the numerators and denominators of Eqs. (17a) and (17b) as a function of Eqs. (2) by means of Eq. (9), and then showing that both, as well as their quotient, can be computed with the required accuracy in time (21). Here, we show it explicitly only for Eq. (17a), but it can be done in the same way for Eq. (17b). We define

$$p = \langle \psi | [AP_\delta(E) + P_\delta(E)A] | \psi \rangle, \quad (22a)$$

$$q = 2\langle \psi | P_\delta(E) | \psi \rangle, \quad (22b)$$

and denote by  $\Delta p$ ,  $\Delta q$  the bounds on their error; that is, if the measured values are  $\tilde{p}$ ,  $\tilde{q}$ , they fulfill  $|\tilde{p} - p| < \Delta p$  (and analogously for  $q$ ). Let us first argue that if we require  $\Delta p$  to scale polynomially with  $N^{-1}$ ,  $\delta$ , and  $\epsilon$ , we can reach it with the quantum simulator in time (21). The reason is that, using Eqs. (9) and (12), we only need to determine the  $2xN/\delta$  values  $a_{A,\psi}(t_m)$ , and each will require running the quantum simulator for a time not larger than  $2x/\delta$ . Furthermore, we have to repeat the procedure a number of times in order to reduce the error. The time to perform those tasks scales as Eq. (21), given that, by assumption, the quantum simulator can efficiently measure  $A$ . An analogous argument applies to  $\Delta q$ .

For a given  $E$ , the total error  $|\tilde{p}/\tilde{q} - p/q|$  will be upper bounded by

$$\frac{\Delta p + p\Delta q/q}{q - \Delta q} \leq \frac{\Delta x + 2\Delta q/q}{q - \Delta q}, \quad (23)$$

as long as  $\Delta q < q$ , where we have used the fact that  $p \leq 2$ . One can readily check that if  $\Delta p = \epsilon q/3$  and  $\Delta q = \epsilon q^2/6$  then the error will be bounded by  $\epsilon$ . Thus, the problem is reduced to proving that

$$q = \text{poly}(N, 1/\delta), \quad (24)$$

since in this case,  $\Delta p$  and  $\Delta q$  will scale polynomially with  $N^{-1}$ ,  $\delta$ , and  $\epsilon$ , which, as argued above, can be accomplished with Eq. (21). As for the cost of finding the value  $E$ , we show in Appendix C that there always exists an interval of energies of size  $\Delta E \geq \delta^2/6N$  within interval (20), where  $q \geq (1/4)[\delta^2/(\delta^2 + 2\sigma_\psi^2)]^{3/2}$ . Since  $\sigma_\psi \leq N/2$ , then in that interval  $q$  fulfills Eq. (24). Thus, the procedure consists of dividing interval (20) into  $24Nr\sigma_\psi/\delta^2$  equal slices and picking an energy  $E$  in each of them. At

least one is then guaranteed to fulfill Eq. (24) and thus we will be able to determine Eqs. (17) with an error smaller than  $\epsilon$  in time (21).

## IV. PRACTICAL COMPUTATIONS

In practice, one can use the quantum simulator much more efficiently than what has been presented in the previous section, and also employ it to access other physical properties. In this section we propose and analyze several algorithms to compute different quantities related to the microcanonical and canonical quantum statistical ensembles. We also illustrate them with some examples for the model of Sec. IID.

### A. Local density of states

The simplest quantity is a broadened version of the local density of states,

$$D_{\delta,\psi}(E) = \langle \psi | P_\delta(E) | \psi \rangle, \quad (25)$$

which (up to a factor) converges to that quantity in the limit  $\delta \rightarrow 0$ . Using Eq. (9), we can express

$$D_{\delta,\psi}(E) \approx \sum_{m=-R}^R c_m a_\psi(t_m) e^{i(E-E_\psi)t_m}, \quad (26)$$

where  $R$  and  $t_m$  are defined as in Eq. (12).

As before, our algorithm uses the quantum simulator to determine  $a_\psi(t_m)$ . The method can be made more efficient by noting that norm of  $X_0 = (H - E_\psi)/(\tilde{r}\sigma_\psi)$  will be bounded by one in the subspace where the state  $\psi$  has most of its weight if we choose  $\tilde{r} \sim 1$ . Thus, we can use expansion (9) with  $X = X_0$  and  $M = \tilde{r}^2\sigma_\psi^2/\delta^2$  to obtain Eq. (25), but now with  $R = x\tilde{r}\sigma_\psi/\delta$  and  $t_m = 2m/(\tilde{r}\sigma_\psi)$ . For  $\sigma_\psi \leq \sqrt{N}$ , the number of required measurements will significantly decrease with respect to Eq. (12). In that case, we let  $\tilde{r}\sigma_\psi = r\sqrt{N}$ , so that

$$R = xr\sqrt{N}/\delta, \quad t_m = 2m/(r\sqrt{N}). \quad (27)$$

In Fig. 2 we plot  $D_{\delta,k}(E_k)$  on the logarithmic scale for  $N = 100$  spins,  $\delta = 0.1$ , and Hamiltonian (13) with  $g = 1$ ,  $h = 2$ , for 50 randomly generated states  $|k\rangle$  (16). We take  $x = 3$  and compare the results of the original (12) (circles) and the optimized alternatives (27) (crosses for  $r = 0.4$ , plus symbols for  $r = 1$ ). For this value of  $\delta$ , the improved method yielding Eq. (27) with  $r = 0.4$  corresponds to 120 measurements of  $a_\psi(t)$  and a maximum value of  $t = 60$ . For  $\delta = 1$ , it requires 12 measurements with a maximum value of  $t = 6$ , which is very reasonable for present experiments. We observe that, for  $r = 1$ , one already obtains an error of the order of  $10^{-3}$ , whereas for  $r = 0.4$ , it is about  $10^{-2}$ , which is what one could expect with imperfect devices.

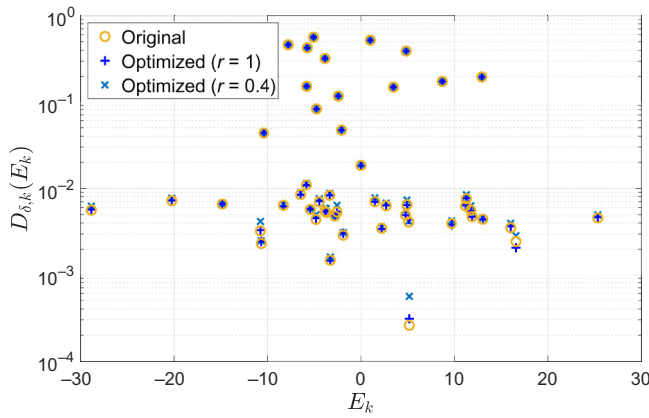


FIG. 2. Local density of states as a function of the mean energy for Hamiltonian (13) for  $N = 100$  spins,  $g = 1$ ,  $h = 2$ ,  $\delta = 0.1$ , and 50 randomly chosen states  $|k\rangle$ . The circles have been computed with Eq. (12) whereas the cross and plus symbols correspond to Eq. (27) with  $r = 0.4$  and  $r = 1$ , respectively. In all cases  $x = 3$ .

### B. Microcanonical observables

The very same simplified program can be applied to Eqs. (17), as we can also choose different values of  $r$  to make the procedure much more efficient. In Fig. 3 we plot Eq. (17a) for model (13) with  $g = 1$ ,  $h = 2$ , and we choose the energy as an observable, i.e.,  $A = H/N$ . Note that the microcanonical expectation value of  $A = H$  at  $E$  is only trivial (i.e., equal to  $E$  exactly) in the limit  $\delta \rightarrow 0$ . Here, our purpose is to investigate the convergence of  $H_{\delta,\psi}(E)$  with respect to  $\delta$ . We again choose 50 random states  $|k\rangle$  and subtract their mean energy  $E_k/N$  to optimize the visualization, since  $H_{\delta,k}(E_k) \sim E_k$ . We take  $\delta = 0.1$  for the triangles, and  $\delta = 1$  for the plus symbols and circles. Only for the latter do we choose the more efficient version (27) with  $r = 1$ . We observe a clear difference (at the percent level) for different values of  $\delta$ ; however, the value  $r = 1$  is sufficient to obtain reliable results.

An interesting question in this context is to what extent one can recover the microcanonical expectation value by decreasing  $\delta$ . This makes sense if the system is sufficiently large and fulfills the eigenstate thermalization hypothesis (ETH) [77,78], which ensures the convergence of the procedure in the thermodynamic limit. The question of how narrow the energy support of a pure state needs to be in order to recover microcanonical expectation values has been analyzed in Refs. [36,79] for one-dimensional systems. Dymarsky and Liu [79] concluded that, for generic systems,  $\delta$  needs to decrease with  $N$ , while Bañuls *et al.* [36] provided evidence that  $\delta \sim 1/\log(N)$  may suffice for local observables. In Appendix D we analyze this question for a nonintegrable model and up to  $N = 28$  spins, and give further evidence for the need to decrease  $\delta$  with  $N$ . This can be qualitatively understood as follows. Let us denote by  $|E\rangle$  the eigenvectors of  $H$  with

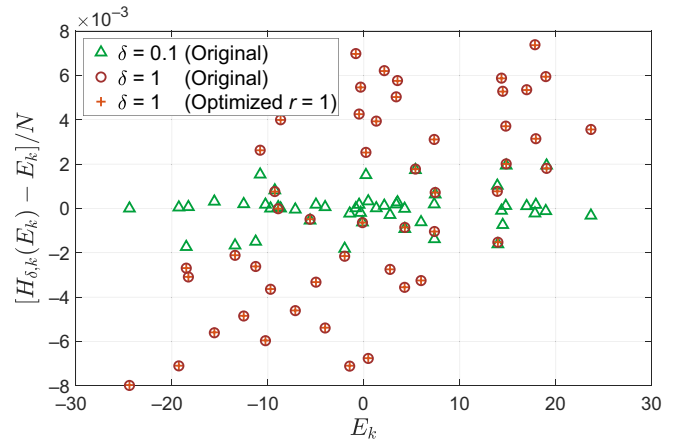


FIG. 3. Expectation value of the energy for the same model and parameters as in Fig. 2. The data show the results of Eq. (12) for  $\delta = 0.1$  (triangle) and 1 (circles), and Eq. (27) for  $\delta = 1$  with  $r = 1$  (plus symbols).

energy  $E$ . In the case where the ETH applies, the diagonal matrix elements of physical observables (which can also include, for instance, correlation functions) in the energy basis rapidly converge to the microcanonical expectation value, while off-diagonal elements vanish exponentially fast [77]. However, when one has a superposition of an exponential number of eigenstates around some energy, the sum of the off-diagonal terms does not need to converge even though the diagonal terms do. This is also the reason why  $\delta$  must decrease with  $N$ , since  $P_\delta(E)|\psi\rangle$  contains superpositions of  $|E\rangle$  and thus the expectation value depends on off-diagonal elements.

### V. QUANTUM ASSISTED MONTE CARLO ALGORITHMS FOR MICROCANONICAL AND CANONICAL OBSERVABLES

The considerations at the end of the previous section suggest an alternative strategy that can be used for microcanonical and canonical observables. At a high level, the algorithms we propose in this section can be seen as a quantum version of the classical quantum Monte Carlo algorithms: they use classical Monte Carlo to sample different initial states, while the quantum device assists with the computation of sampling probabilities and the measurements of observables.

Different to the algorithm presented in Sec. III, the quantum assisted sampling discussed here is not proven to be efficient. However, when compared to the first algorithm from an experimental point of view, these methods offer the potential advantage of requiring shorter coherence times, at the cost of an increased number of measurements in the quantum device. Moreover, the sampling probabilities are always positive and the sign problem in classical quantum Monte Carlo [51] is circumvented. We

also present numerical experiments for a  $N = 100$  Ising Hamiltonian, to demonstrate the viability of these methods on near-term devices. The results show that physical quantities can be obtained accurately even in the presence of a certain level of noise.

### A. Microcanonical observables

The quantity

$$A_\delta(E) = \frac{\text{tr}[AP_\delta(E)]}{\text{tr}[P_\delta(E)]} \quad (28)$$

may converge to the microcanonical expectation value even for constant  $\delta$  since, by definition,  $P_\delta(E)$  is diagonal in the eigenbasis of  $H$  and thus no off-diagonal element appears in the expectation value. This is indeed observed in Ref. [80], where tensor network techniques are used in order to compute quantities closely related to Eq. (28) with energy resolution corresponding to larger values of  $\delta$  than those required by Eqs. (17). In fact, that a constant  $\delta$  suffices also follows from the fact that, if the thermodynamic limit  $N \rightarrow \infty$  exists for the observable  $A$ , then in that limit the expectation value of  $A$  in almost all eigenstates of  $H$  should coincide if the corresponding energies fulfill  $|E_N - E'_N|/N \rightarrow 0$ . The intuitive reason is that in that limit, only intensive quantities matter. Therefore, what is relevant is  $\delta/N$ , so that as long as it vanishes, one should obtain the thermodynamic value.

In principle, one could obtain Eq. (28) in the very same way as Eq. (17a). This can be seen by noting that

$$A_\delta(E) = A_{\delta,\Phi}(E), \quad (29)$$

where  $\Phi$  is a maximally entangled state of each spin with an auxiliary one spin,

$$|\Phi\rangle = \frac{1}{2^{N/2}}[|0,0\rangle + |1,1\rangle]^{\otimes N}. \quad (30)$$

That is, just one has to add an auxiliary qubit for each existing one and prepare an entangled state of each pair. Thus, from Eqs. (29) and (17a), it follows that one could compute the numerator and denominator independently with the help of the quantum simulator, and then the quotient. However, we face the problem that the denominator will typically decrease exponentially with  $N$ . For local Hamiltonians, the reason is that  $\sigma_\Phi \propto \sqrt{N}$  and therefore, for any extensive value of the energy  $E = eN$ ,  $\langle \Phi | [P_\delta(E) \otimes \mathbb{1}] | \Phi \rangle \sim \exp(-cN)$  for some  $c = O(1)$ . This makes this procedure impracticable for  $N \gtrsim 20$ .

In the following, we propose an algorithm to circumvent, at least in part, this issue. Let us reexpress Eq. (28)

with the help of an (over)complete basis of states fulfilling

$$\int d\mu_\psi |\psi\rangle\langle\psi| = \mathbb{1}, \quad (31)$$

where  $d\mu_\psi$  is a measure in the basis set. For instance, we can take an orthonormal basis of product states  $|p_1, p_2, \dots, p_N\rangle$  or all product states (5), in which case

$$d\mu_p = \prod_{n=1}^N d\Omega_n = \frac{1}{(4\pi)^N} \prod_{n=1}^N \sin(\theta_n) d\theta_n d\varphi_n. \quad (32)$$

Inserting Eq. (31) into Eq. (28) and using definitions (17a) and (28), we can rewrite Eq. (28) as

$$A_\delta(E) = \frac{\int d\mu_\psi D_{\delta,\psi}(E) A_{\delta,\psi}(E)}{\int d\mu_\psi D_{\delta,\psi}(E)}. \quad (33)$$

This quantity can be computed using Monte Carlo algorithms so long as one is able to compute  $D_{\delta,\psi}(E)$  and  $A_{\delta,\psi}(E)$ . These values can be computed using the quantum simulator. More concretely, one can implement importance sampling according to a distribution proportional to  $D_{\delta,\psi}(E)$ . For instance, using a basis of product states, as mentioned above, this can be achieved by a Metropolis-Hastings algorithm in which, once a move is proposed, the quantum simulator is used to estimate the corresponding value of  $D_{\delta,\psi}(E)$ , and thus determine the acceptance rate. The Monte Carlo algorithm circumvents not only the burden of summing over all states, but also the need to measure  $D_{\delta,\psi}(E)$  with an exponentially small accuracy, because  $D_{\delta,\psi}(E)$  needs to be evaluated only for states  $\psi$  for which it is not negligible. We also emphasize that if the observable  $A$  is chosen so that  $A|\psi\rangle = \lambda|\psi\rangle$ , and  $\lambda$  can be classically computed, then one only has to determine  $D_{\delta,\psi}(E)$  with the quantum simulator.

We have tested this algorithm with Hamiltonian (13) and the simplest Monte Carlo algorithm that changes one spin at a time for the sampling. Note that our goal is to demonstrate the viability of the approach, rather than the competitiveness of the algorithm, since that would require optimizing the sampling methods and other parameters. We choose the Hamiltonian so that we can compare the results with an exact calculation for  $N \sim 100$  spins. We implement a Metropolis algorithm that takes a random state  $|k\rangle$  (16), according to a probability proportional to  $D_{\delta,k}(E)$ . We then compute the magnetization  $A = M$  with

$$M = \frac{1}{2N} \sum_{n=1}^N (\sigma_{n,z} + 1) \quad (34)$$

as a function of  $E$ . Since the model is exactly solvable, we also compute  $A_\delta(E)$  directly using the method of Appendix B. This direct numerical calculation requires the



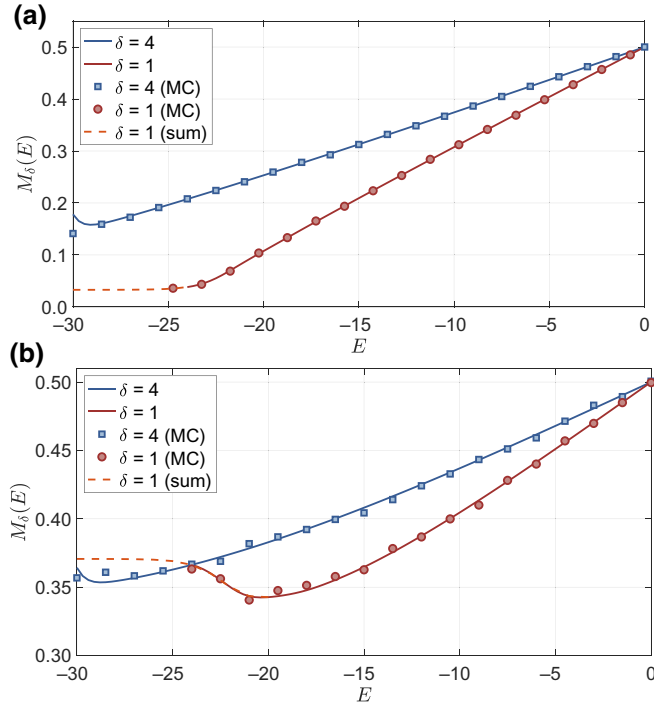


FIG. 4. Magnetization computed according to Eq. (28) for Hamiltonian (13) and  $N = 20$  spins computed exactly (solid lines) and via the Monte Carlo (MC) method described in the text with  $10^5$  samples per point (symbols). The Monte Carlo simulations are carried out for the full expression (12) with  $x = 3$  for the different values of  $\delta$  shown in the legend. In (a)  $g = 1$ ,  $h = 2$ , whereas in (b)  $g = 2$ ,  $h = 1$ .

computation of very large and small numbers, so one can easily run into precision problems. In fact, for some plots, we can only provide the exact result for some values of  $E$ , since otherwise our exact method does not give consistent values.

In Fig. 4(a) we plot  $M_\delta(E)$  for  $N = 20$  spins,  $g = 1$ ,  $h = 2$ , and  $\delta = 1, 4$  (red and blue lines) obtained with the numerical computation. The symbols are obtained with the Monte Carlo method with  $\delta = 1$  (circles) and  $\delta = 4$  (squares). We have checked that the results for  $\delta = 1$  and  $\delta = 0.1$  are almost indistinguishable, and this is why we only plot the results for  $\delta = 1$ . Each point is sampled  $10^5$  times. In Fig. 4(b) we plot the same for  $g = 1$  and  $h = 2$ , also showing very good results. Note that the lower curve terminates at around  $E_k \sim -25$ ; the reason is that the lowest energy that can be reached with states (16) is  $E_{p,\min} = -20, -14.39$  for Figs. 4(a) and 4(b), respectively, and thus at lower energies,  $D_{\delta,\psi}(E)$  becomes very small for any state  $\psi$ . Thus, we encounter convergence problems for such low energies not only with the Monte Carlo method but also with the exact method and this is why they are not included in the plot. We used an exact summation to obtain the dotted line. This is only possible because we only have  $N = 20$  spins in this figure.

In Fig. 5 we consider larger systems, with  $N = 100$ . In Fig. 5(a) we depict the magnetization for  $g = 1$  and  $h = 2$ . For a more convenient graphical representation of the data, we plot in the inset the modified quantity  $M'_\delta(E) = M_\delta(E) - 1/2 - 0.004E/N$  as a function of  $E$ . The solid lines indicate the exact results for  $\delta = 1, 4$  (red and blue), while the symbols show the Monte Carlo results (circles and crosses for  $\delta = 1$ , squares and plus symbols for  $\delta = 4$ ). As before, we sample  $10^5$  times for each point. The cross and plus symbols in the inset indicate the results when a cutoff of  $10^{-2}$  is set for  $D_{\delta,k}(E)$ . That is, in the Metropolis method, as soon as we compute it and obtain a smaller value, we set it to zero. In practice, the intention is to mimic an experiment where this quantity has been obtained to that precision. The solid lines are exact results, which have only been computed up to some value of  $E$ , since otherwise we encounter precision problems. As one can gather from the plots, the Monte Carlo results resemble the corresponding values well. For  $E < -60$ , the exact method encounters precision problems, while this happens for the Monte Carlo method only for  $E < -100 = E_{p,\min}$ . When decreasing  $\delta$  even further, the results are practically indistinguishable from those of  $\delta = 1$ . However, the program becomes unstable for energies  $E \lesssim E_{p,\min}$  since, as expected, the values of  $D_{\delta,k}(E)$  become extremely small due to the Gaussian dependence.

## B. Canonical observable

Now we show how, using similar ideas, one can also compute canonical observables, i.e.,

$$A(\beta) = \frac{\text{tr}(e^{-\beta H} A)}{\text{tr}(e^{-\beta H})}, \quad (35)$$

where  $\beta$  is the inverse temperature. We use the fact that, for sufficiently small  $\delta$ , we can approximate

$$e^{-\beta H} \approx \int_{E_0}^{E_1} dE e^{-\beta E} P_\delta(E), \quad (36)$$

where  $E_{rhs,0,1} = E'_{0,1} \mp y\delta$  with  $E'_0$  ( $E'_1$ ) the lowest (highest) eigenvalue of  $H$ , and  $y \gg 1$ . This motivates the definition

$$A_\delta(\beta) = \frac{\int_{E_0}^{E_1} dE e^{-\beta E} \text{tr}[P_\delta(E) A]}{\int_{E_0}^{E_1} dE e^{-\beta E} \text{tr}[P_\delta(E)]}, \quad (37)$$

which converges to the canonical value for  $\delta \rightarrow 0$ . Using Eq. (33), we have

$$A_\delta(\beta) = \frac{\int_{E_0}^{E_1} dE e^{-\beta E} \int d\mu_\psi D_{\delta,\psi}(E) A_{\delta,\psi}(E)}{\int_{E_0}^{E_1} dE e^{-\beta E} \int d\mu_\psi D_{\delta,\psi}(E)}. \quad (38)$$

The quantum algorithm proceeds in the same way as before, by using the quantum simulator to recover  $a_\psi(t_m)$

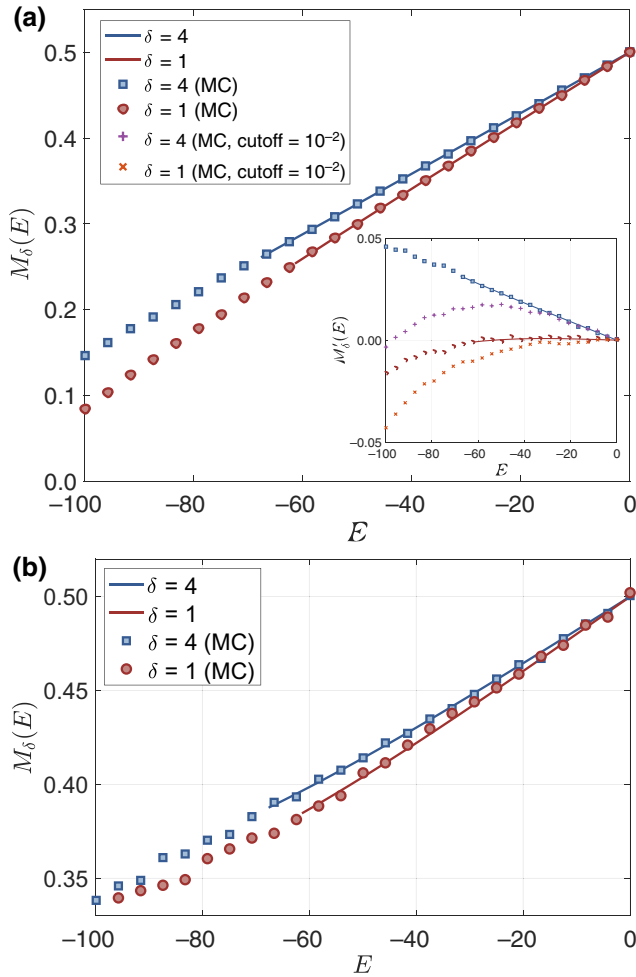


FIG. 5. Magnetization as a function of energy, as in Fig. 4 but with  $N = 100$ . For the inset plot, we subtracted  $1/2 + 0.004E/N$  in order to make the curves more visible. There, the crosses and plus symbols indicate the values obtained by setting a cutoff of  $10^{-2}$  in the sampling procedure to resemble an experiment. In (a)  $g = 1$ ,  $h = 2$ , whereas in (b)  $g = 2$ ,  $h = 1$ .

and  $a_{A,\psi}(t_m)$ , and then a classical computation to do the rest. In particular, Monte Carlo samples both the states  $\psi$  and the energies  $E$  with a probability proportional to  $e^{-\beta E} D_{\delta,\psi}(E)$ . For observables that are diagonal in the chosen basis, it is also possible to sample only over states, with the integrated probability  $\int dE e^{-\beta E} D_{\delta,\psi}(E)$ , which can be reconstructed using the quantum simulator in the same way as for the microcanonical observables, since the energy dependence can be integrated analytically.

We performed Monte Carlo computations and show the results in Fig. 6. We plot magnetization (34) as a function of the inverse temperature  $\beta$  for Hamiltonian (13) with  $N = 100$  spins and  $g = 0.3$ ,  $h = 0.8$  (lower curve) and  $g = 0.4$  and  $h = 0.4$  (upper curve). The solid line is the exact result (B14) computed with the formulas given in Appendix B. Note that here there is no problem with the precision, as most of the products in the numerator

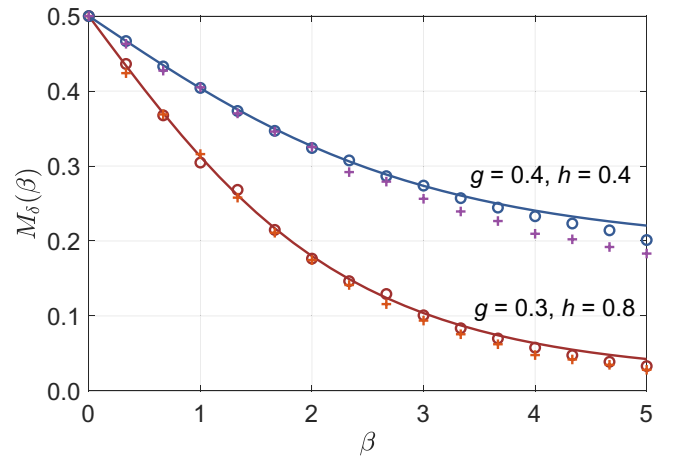


FIG. 6. Magnetization as a function of the inverse temperature for Hamiltonian (13) for  $N = 100$  spins,  $g = 0.3$ , and  $h = 0.8$  (lower curve) and  $g = h = 0.4$  (upper curve). The solid line represents the exact value  $M(\beta)$ , whereas the symbols are obtained with the Monte Carlo method, with  $10^5$  samples per point and  $\delta = 1$  and  $x = 3$ . We discretize the values of the energy from  $-N$  to  $N$  in intervals of 0.5. The plus symbols are obtained by setting a cutoff equal to  $10^{-2}$  in  $D_{\delta,k}(E)$ .

and denominator cancel and one ends up with a simple sum (B15), and that the result is independent of  $\delta$  [since definition (B14) is, too]. The symbols are obtained with the Monte Carlo simulation for  $\delta = 1$  and  $x = 3$ . We discretize the energy integral appearing in Eq. (38) by taking  $E$  from  $-N$  to  $N$  in steps of 0.5, and sampling only those values, although the Metropolis algorithm only takes values around certain energies,  $E \pm 5$ , for each value of  $\beta$ . For the Monte Carlo methods, we take  $E_{0,1} = \pm 3N/2$ , but these values are never reached in the Metropolis sampling. From the figure we conclude that the result converges very well for  $\delta = 1$ , and also that the performance of the Monte Carlo method is very good. One can observe that, for larger values of  $\beta$ , there is a little bias towards smaller values of the magnetization. In the inset of Fig. 5, we also display the results carried out with the Monte Carlo computation with a cutoff  $10^{-2}$ , still rendering competitive results. We note that the fact that the upper curve corresponds to the critical point  $g = h$  does not have relevant consequences at the temperatures considered here. For lower temperatures, the program does not give reliable results as it has to scan energies for which  $D_{\delta,\psi}(E)$  becomes very small.

## VI. SUMMARY AND OUTLOOK

We propose and analyze two types of quantum algorithm to characterize quantum many-body states in finite energy intervals and temperatures. They are based on the cosine filter, which when applied to a state, reduces its variance to a predetermined value around a given energy. However, instead of preparing the filtered state,

our algorithms compute different quantities that allow one to reconstruct expectation values of observables or other magnitudes. The algorithms we propose are quite flexible—they can either be implemented on digital quantum computers (e.g., based on superconducting qubits, trapped ions, or Rydberg atoms) and on analog quantum simulators (e.g., based on cold atoms, trapped ions, photons), as long as they can perform certain kinds of interferometric measurements. We also present in Appendix A more practical approaches that, in some cases, possess clear advantages with respect to such measurements.

We show how our first algorithm can be used to efficiently compute expectation values over filtered states, as long as we can prepare the initial state. The algorithm requires a time that scales polynomially with the system size,  $N$ , the inverse variance, and the inverse error. We also show how it can be simplified in practice, leading to a practicable method for present or planned quantum simulators. The simulator has to be run for times  $\leq 6/\delta$  (where we take  $x \sim 3$  and  $r \sim 1$ ) and therefore, for  $\delta = O(1)$ , this should be feasible for existing devices. The price one has to pay is that one has to perform many more measurements. This number can be estimated by taking into account the fact that we need to compute sums of the form (26), where we add  $2R$  terms so that to have a total error of the order of  $\epsilon$ , this will require of the order of  $R\epsilon^{-2} \approx 3\sqrt{N}/\delta\epsilon^2$  measurements. For  $N = 100$ ,  $\delta = 1$ , and  $\epsilon = 10^{-2}$ , this yields of the order of  $3 \cdot 10^5$  measurements. Actually, by taking into account the fact that  $c_m$  becomes small, using smaller  $x \sim 1$ , exploiting symmetries and other optimizations, we expect that it is possible to significantly reduce this number, to the order of  $10^4$ .

The second algorithm we introduce proposes a way to combine the first algorithm with Monte Carlo simulations in order to obtain the expectation value of microcanonical and canonical observables. The resulting methods also rely on the ability to prepare certain states (e.g., product states) efficiently, to run the quantum simulator for times of the order of  $6/\delta$  and to perform interferometric measurements. Our numerical simulations of this algorithm for a simple model indicate that with several tens of thousands of samplings one can obtain reliable results, at least in some energy regimes. This algorithm can also be significantly improved by using standard Monte Carlo strategies to speed up convergence. Such quantum assisted sampling algorithms open new possibilities to study thermal properties of quantum many-body systems with near-term quantum devices.

There are other modifications that may help to improve the algorithms. First, one could take a different expansion of the filter. For instance, instead of using the cosine filter, one can use Chebyshev expansions for a quantum computer [34,36,49,50], a low-pass filter, or choose  $c_m$  differently for an analog simulator to better adapt to the specific errors in which it may incur. Second, in the sampling, there

is information that can be collected to investigate other physical questions. For instance, the spins or energy samples that are used in the Monte Carlo methods can be used to estimate other quantities directly (like higher moments), without the need to make other computations. Another possibility is to use the adiabatic or variational algorithms to prepare states with small variance to start with, which would make the algorithms more efficient, or allow us to access energies that are out of the scope of product states.

We formulate most of our results for a many-body spin Hamiltonian with local interactions. However, they can be equally applied to fermionic systems, systems with longer-range interactions, or disordered setups, as long as the quantum device can emulate the corresponding Hamiltonians. Also, the ideas developed here can be easily adapted to measure other quantities of interest, like Green functions or structure factors.

## ACKNOWLEDGMENTS

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## APPENDIX A: WAYS TO RETRIEVE $a_{A,\psi}(t)$

In this appendix, we describe different methods to retrieve  $a_\psi(t)$  and  $a_{A,\psi}(t)$  from analog quantum simulators. Such methods typically possess severe limitations in the available operations as compared with quantum computers, and thus they may not be able to carry out the interferometric measurements that are needed to obtain those quantities. The methods adapt to different situations and have different requirements.

We start out by briefly reviewing the standard method based on conditional dynamics where, during the evolution, all the qubits have to be coupled to an extra one, called the control qubit. Then, we analyze a closely related method where the interaction with the control qubit only needs to occur at the beginning and at the end of the evolution, although at the expense of using other internal states. The third method does not require a control qubit but the possibility of creating certain catlike states, as well as an extra internal state; the latter can be avoided for certain kinds of Hamiltonian for which one can prepare one eigenstate efficiently, like, for instance, Hamiltonians of Heisenberg type. The fourth method builds on the previous one and does not need cat states. This is thus much simpler to implement in practice and may be more robust against decoherence. However, it requires more measurements. The last method applies to Hamiltonians with some

special symmetries, like  $XY$  or Hubbard models, and can be very practical.

### 1. Conditional dynamics

The conceptually simplest way to obtain  $a_{A,\psi}(t)$  is to carry out conditional dynamics [72] depending on the state of one of the qubits, the control qubit, which we call  $c$ . This corresponds to the operation

$$U|0\rangle_c \otimes |\psi\rangle = |0\rangle_c \otimes e^{-iHt}|\psi\rangle, \quad (\text{A1a})$$

$$U|1\rangle_c \otimes |\psi\rangle = |1\rangle_c \otimes |\psi\rangle. \quad (\text{A1b})$$

If we denote by  $H_c$  the Hadamard transformation on the control qubit then one can first produce the state

$$H_c^\dagger U H_c |0\rangle_c \otimes |\psi\rangle. \quad (\text{A2})$$

By then measuring the control qubit in the computational basis, and the observable  $A$  on the rest, one can retrieve Eqs. (2). This method requires the ability to couple all the qubits to the control one during the evolution in order to apply Eqs. (A1), which may be difficult in practice. However, it can be very naturally applied in trapped ion simulators [37,60], since all ions are coupled to the same phonon bus, and in Rydberg atoms in optical lattices, as one atom in a Rydberg state can influence the dynamics of the rest [81]. In fact, very efficient techniques have been proposed to perform this kind of dynamics and measurements using that implementation [82–84].

### 2. Additional internal states

The next method does not require coupling all the systems to the control qubit during the interaction, but makes use of extra levels in each of the systems. Instead of qubits, one uses four-level systems where, apart from the qubit states  $|0\rangle$  and  $|1\rangle$ , there are two other levels,  $|a_{0,1}\rangle$  (see Fig. 7), that are idle with respect to the action of the Hamiltonian; that is,  $\tilde{U} = e^{-iHt}$  only acts nontrivially in the subspace spanned by  $|0\rangle$  and  $|1\rangle$ . Let us assume that one has the possibility of adding extra two-body operations

$$W_n |0\rangle_c \otimes |i\rangle_n = |0\rangle_c \otimes |i\rangle_n, \quad (\text{A3a})$$

$$W_n |a_0\rangle_c \otimes |i\rangle_n = |a_0\rangle_c \otimes |a_i\rangle_n, \quad (\text{A3b})$$

as well as the Hadamard  $H_c$  between levels  $|0\rangle_c$  and  $|a_0\rangle_c$ . Then, one can implement

$$H_c^\dagger W^\dagger \tilde{U} W H_c |0\rangle_c \otimes |\psi\rangle, \quad (\text{A4})$$

where  $W = \bigotimes_{n=1}^N W_n$ . As before, by measuring in the basis  $|0\rangle_c, |a_0\rangle_c$  the control qubit, and the observable  $A$  on the rest, one can also obtain Eqs. (2).

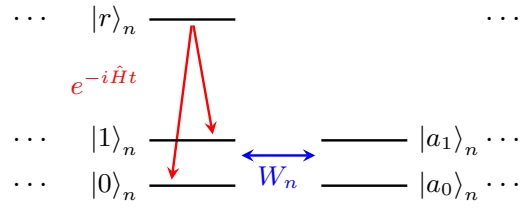


FIG. 7. Internal level scheme for the interferometric measurement: the simulation acts on levels  $|0\rangle$  and  $|1\rangle$ , while the other two levels are used to generate the catlike state.

### 3. Catlike states

We show now that neither an ancilla system nor additional internal states are required as long as one can prepare some catlike states. This has been achieved with different setups already [26,63–71].

Let us first assume that  $H$  has an eigenstate  $\varphi$ , i.e.,

$$H|\varphi\rangle = \lambda|\varphi\rangle, \quad (\text{A5})$$

such that, for a given state  $\psi$ , one can efficiently apply an operator  $V(\theta)$ , fulfilling

$$V(\theta)|0, \dots, 0\rangle = \frac{1}{\sqrt{2}}(|\varphi\rangle + e^{i\theta}|\psi\rangle). \quad (\text{A6})$$

In general,  $\psi$  and  $\varphi$  may be very different, in the sense that they can be distinguished by measuring few qubits [85], and thus the state on the right-hand side of Eq. (A6) will be a catlike state. In some cases,  $\varphi$  is just a product state. This occurs, for instance, for the Heisenberg Hamiltonian (in any dimension)

$$H = \sum_{n,m} J_{n,m} \vec{\sigma}_n \cdot \vec{\sigma}_m + \sum_n h_n \sigma_{n,z}. \quad (\text{A7})$$

Indeed, for arbitrary  $J_{n,m}$  and  $h_n$ ,  $|\varphi\rangle = |0, \dots, 0\rangle$  is an eigenstate. Here  $\vec{\sigma} = (\sigma_x, \sigma_y, \sigma_z)$  is a vector of Pauli operators and  $\sigma_{n,z}$  is the one that acts on the  $n$ th qubit.

Under the above assumption, one can prepare

$$|\Phi(\theta, \theta')\rangle = V(\theta')^\dagger e^{-iHt} V(\theta)|0, \dots, 0\rangle \quad (\text{A8})$$

and measure in the computational basis to obtain  $a_\psi(t)$ . In order to show this, without loss of generality, we set  $\lambda = 0$  as we can always subtract it from the Hamiltonian. The probability of obtaining the state  $|0, \dots, 0\rangle$  is

$$|\langle 0, \dots, 0 | \Phi(\theta, \theta') \rangle|^2 = \frac{1}{4} |e^{-i\alpha} + a_\psi(t) e^{i\alpha} + 2r \cos(\beta)|^2, \quad (\text{A9})$$

where  $\langle \psi | \varphi \rangle = r e^{-i\gamma}$ ,  $r \geq 0$ ,  $\alpha = (\theta - \theta')/2$ , and  $\beta = (\theta + \theta')/2 + \gamma$ . Note that, apart from  $a_\psi(t)$ ,  $r$  and  $\gamma$  are,

in principle, unknown. If they can be computed classically (e.g., if we deal with product states) then one can directly obtain  $a_\psi(t)$  by measuring this quantity for  $\beta = 0$  and  $\alpha = 0, \pi/2$ , and  $\pi/4$ . If not, one can proceed as follows. First, set  $\alpha = 0$ ; by changing  $\theta + \theta'$  one can find when  $\beta = 0$  or  $\pi$ , as the expression is maximal or minimal. Then, one can also identify when  $\beta = \pi/2$ . Using these three values, it is possible to recover the real part of  $a_\psi(t)$ . By choosing  $\alpha = \pi/2$  and proceeding in the same way, one can then get the imaginary part. The procedure to measure  $a_{A,\psi}(t)$  is the same, but (i) one has to apply the operator  $A$  before the measurement and (ii) one has to classically compute (or measure with the simulator)  $\langle \varphi | A | \varphi \rangle$ . The first task can be obtained as follows, as long as one can use the simulator to evolve according to the dynamics generated by  $A$ . One just applies the operator  $\exp(-iA\delta t) \approx \mathbb{1} - iA\delta t$  for a short time  $\delta t \ll \|A\|_\infty$  and measures as before. If  $A$  is not Hermitian, one can always write it as a sum  $A = A_r + iA_i$ , where  $A_r = A_r^\dagger$  and  $A_i = A_i^\dagger$  and measure these two independently. In practice, there may be simpler procedures that do not involve small times, for instance, if  $A^2 = \mathbb{1}$ .

So far we have requested that  $H$  has an eigenstate that can be easily prepared. This condition can always be satisfied so long as one has an extra level available on each system,  $|a_0\rangle$ , where the Hamiltonian does not act. In that case, taking  $|\varphi\rangle = |a_0, \dots, a_0\rangle$  one has  $H|\varphi\rangle = 0$  and thus it is an eigenstate.

#### 4. Product states

The procedure presented above can be further simplified if one wants to measure  $a_{A,\psi}(t)$  for states of the form

$$|\psi\rangle = \prod_{n=1}^N \sigma_{n,\alpha_n} |\varphi\rangle, \quad (\text{A10})$$

where  $\sigma_{\alpha_n}$  is a Pauli operator. For instance, if  $\varphi$  is a product state, this allows one to obtain Eqs. (2) for any product state. Also, it can be easily extended to states that are connected with products of other simple operators.

The idea is to sequentially measure  $a_{A,\varphi_m}(t)$  (for  $m = 1$  to  $m = N$ ), where

$$|\varphi_m\rangle = \prod_{n=1}^m \sigma_{n,\alpha_n} |\varphi\rangle = \sigma_{m,\alpha_m} |\varphi_{m-1}\rangle, \quad (\text{A11})$$

$\varphi_0 = \varphi$ , and  $\varphi_N = \psi$ . Once  $a_{\varphi_{m-1}}(t)$  is obtained, one can start out with  $\varphi_{m-1}$  and obtain  $a_{A,\varphi_m}(t)$  since the procedure explained immediately after Eq. (A8) can be applied if  $a_\varphi(t)$  is known (it does not have to be equal to one).

With single qubit operations acting only on the  $m$ th qubit, one can prepare

$$V_m(\theta)|\varphi_{m-1}\rangle = \frac{1}{\sqrt{2}}(|\varphi_{m-1}\rangle + e^{i\theta}|\varphi_m\rangle) \quad (\text{A12})$$

and follow the same procedure as before. This method does not require the preparation of catlike states, as the two states building the superposition at any given step differ just on one qubit. However, it requires many more measurements since one has to obtain  $a_{A,\varphi_n}(t)$  for  $n \leq m$ . Note that one has to obtain, at most,  $N$  of those quantities. Note also that this procedure may have some practical restrictions as the errors accumulate.

#### 5. Symmetric Hamiltonians

A special class of Hamiltonians for which one can use a simplified procedure contains those for which there exists a unitary operator  $R$  and a Hermitian operator  $T$  such that

$$R^\dagger H R = -H + T \quad (\text{A13})$$

with  $[T, H] = 0$  and for which  $\psi$  is an eigenstate, i.e.,

$$R|\psi\rangle = \lambda|\psi\rangle, \quad T|\psi\rangle = \mu|\psi\rangle. \quad (\text{A14})$$

Given such  $R$ ,  $T$ , and  $\psi$ ,

$$a_\psi(t) = \langle \psi | R^\dagger e^{-iHt} R | \psi \rangle = \langle \psi | e^{i(H-T)t} | \psi \rangle = \overline{a_\psi(t)} e^{-i\mu t}.$$

Thus,  $a_\psi(t)e^{i\mu t/2}$  is real. Moreover, if

$$R^\dagger A R = \pm A, \quad (\text{A15})$$

and additionally if either  $\psi$  is an eigenstate of  $A$  (with eigenvalue 1 for simplicity) or  $[A, H] = 0$ , then  $a_{A,\psi}(t)e^{i\mu t/2}$  is also real or purely imaginary, depending on the sign in Eq. (A15). To determine  $a_{A,\psi}(t)$  then, one only needs the absolute value  $|a_{A,\psi}(t)|$  and a sign. Since

$$|a_{A,\psi}(t)|^2 = |\langle \psi | A e^{-iHt} | \psi \rangle|^2, \quad (\text{A16})$$

the former can be found just by letting the system evolve and then measuring in a basis that contains  $A|\psi\rangle$ . The sign change can be inferred if we track when the absolute value becomes zero as follows. Because we have a finite system,  $a_{A,\psi}(t)$  will be an analytic function of  $t$  so that we can Taylor expand it when it is near zero. Suppose that  $a_{A,\psi}(t)$  approaches zero at a time  $t_0$ ; then one can expand it to the leading order:  $a_{A,\psi}(t) \approx \alpha(t - t_0)^n + O[(t - t_0)^{n+1}]$ . If  $n$  is odd then  $a_{A,\psi}(t)$  will change sign and if it is even, it will not. Since we are measuring the square of  $a_{A,\psi}(t)$ ,  $|\langle \psi | A e^{-iHt} | \psi \rangle|^2$ , we can determine the value  $2n$ , and thus  $n$ . In practice,  $n = 1$  or  $2$ , so that this can be obtained more easily. Hence, here we only require individual measurements, available in several labs worldwide that possess analog quantum computers [13,15,16,20,23,25]

In the rest of this section, we show physically relevant examples in which one can find the operators  $R$  and  $T$  with an especially simple structure, and a whole basis of

common product eigenstates  $\psi$  can be used to run the algorithms from Sec. IV. There are many textbook Hamiltonians and states fulfilling properties (A13) and (A14). For instance, if one has a bipartite lattice with sublattices  $A$  and  $B$ , and  $N$  sites in total, and a Hamiltonian of the form

$$H = \sum_{n \in A, m \in B} [J_{n,m,x} \sigma_{n,x} \sigma_{m,x} + J_{n,m,y} \sigma_{n,y} \sigma_{m,y}] + \sum_{n=1}^N h_n \sigma_{n,z}, \quad (\text{A17})$$

then

$$R = \bigotimes_{n \in A} \sigma_{n,x} \bigotimes_{m \in B} \sigma_{m,y} \quad (\text{A18})$$

fulfills Eq. (A13) with  $T = 0$ . Therefore, if one chooses  $\psi$  as an eigenstate of  $R$  (e.g., a product state), the requirements are satisfied. The operator  $R$  is, up to single qubit rotations, a parity operator.

The Hubbard model in a bipartite lattice

$$H = \sum_{n \in A, m \in B} \sum_{\sigma = \uparrow, \downarrow} J_{n,m,\sigma} [a_{n,\sigma}^\dagger a_{m,\sigma} + a_{m,\sigma}^\dagger a_{n,\sigma}] + \sum_{n=1}^N U a_{n,\uparrow}^\dagger a_{n,\uparrow} a_{n,\downarrow}^\dagger a_{n,\downarrow}, \quad (\text{A19})$$

with  $a^\dagger$  and  $a$  the fermionic creation and annihilation operators, also fulfills Eq. (A17). In such a case,

$$R = e^{i\pi \sum_{n \in A} a_{n,\downarrow}^\dagger a_{n,\downarrow}} \prod_n (a_{n,\uparrow} + a_{n,\uparrow}^\dagger), \quad (\text{A20a})$$

$$T = \sum_{n=1}^N U a_{n,\downarrow}^\dagger a_{n,\downarrow}. \quad (\text{A20b})$$

We can find a whole orthogonal basis of common eigenstates of  $R$  and  $T$  with simple structure in the following way. In particular, if the number of sites  $N$  is even, we can divide the lattice in  $N/2$  disjoint pairs of sites  $S_j = (n_j, m_j)$ . For each pair  $S_j$ , we choose a state  $|\phi_{0,j}\rangle$  which is the vacuum of spin up modes and an arbitrary Fock state of spin down modes, and define

$$|\varphi_j^\pm\rangle = (a_{n_j,\uparrow}^\dagger \pm i a_{m_j,\uparrow}^\dagger) |\phi_{0,j}\rangle, \quad (\text{A21})$$

$$|\phi_j^\pm\rangle = (1 \pm i a_{n_j,\uparrow}^\dagger a_{m_j,\uparrow}^\dagger) |\phi_{0,j}\rangle. \quad (\text{A22})$$

Then, an orthogonal basis of common eigenvectors  $|\psi\rangle$  can be formed as all possible products of one of these factors for each pair of sites.

## APPENDIX B: ISING MODEL

In this appendix we give some formulas that we have used in our numerical illustrations regarding the Ising model. Let us consider  $N$  fermionic modes and a Hamiltonian

$$H = \frac{g}{2} \sum_{n=1}^N (a_n + a_n^\dagger)(a_{n+1} - a_{n+1}^\dagger) + h \sum_n (a_n^\dagger a_n - 1/2), \quad (\text{B1})$$

where the  $a_n$  are annihilation operators of the vacuum  $|\text{vac}\rangle$  and we have taken periodic boundary conditions for the fermions,  $a_{N+1} = a_1$ . Through the Jordan-Wigner transformation, this corresponds to the Ising-like Hamiltonian

$$H = \frac{g}{2} \sum_{n=1}^N \sigma_{n,x} \sigma_{n+1,x} + \frac{h}{2} \sum_{n=1}^N \sigma_{n,z}, \quad (\text{B2})$$

where the  $\sigma_{x,z}$  are Pauli operators, with appropriate boundary conditions.

As usual, we first perform a Fourier transform with (we assume even  $N$ )

$$b_k = \frac{1}{\sqrt{N}} \sum_{n=1}^N e^{i2\pi nk/N} a_n, \\ a_n = \frac{1}{\sqrt{N}} \sum_{k=-N/2+1}^{N/2} e^{-i2\pi nk/N} b_k.$$

The new Hamiltonian is

$$H = \sum_{k=0}^{N/2} \tilde{H}_k, \quad (\text{B3})$$

where

$$\tilde{H}_k = x_k (b_k^\dagger b_k + b_{-k}^\dagger b_{-k} - 1) + i y_k (b_k b_{-k} - b_{-k}^\dagger b_k^\dagger), \\ \tilde{H}_0 = x_0 (b_0^\dagger b_0 - 1/2), \\ \tilde{H}_{N/2} = x_{N/2} (b_{N/2}^\dagger b_{N/2} - 1/2),$$

and

$$x_k = h + g \cos(2\pi k/N), \\ y_k = g \sin(2\pi k/N).$$

Note that  $\text{tr}(\tilde{H}_k) = 0$ .

We now proceed as follows. For each value of  $k = 1, \dots, N/2 - 1$ , we write  $H_k = \tilde{H}_k$  as a  $4 \times 4$  matrix in the

basis

$$\begin{aligned} |1\rangle &= |\text{vac}\rangle, \\ |2\rangle &= b_k^\dagger b_{-k}^\dagger |\text{vac}\rangle, \\ |3\rangle &= b_k^\dagger |\text{vac}\rangle, \\ |4\rangle &= b_{-k}^\dagger |\text{vac}\rangle, \end{aligned}$$

and also define  $H_0 = \tilde{H}_0 + \tilde{H}_{N/2}$  and write it as a  $4 \times 4$  matrix in the basis

$$\begin{aligned} |1\rangle &= |\text{vac}\rangle, \\ |2\rangle &= b_{N/2}^\dagger b_0^\dagger |\text{vac}\rangle \\ |3\rangle &= b_{N/2}^\dagger |\text{vac}\rangle, \\ |4\rangle &= b_0^\dagger |\text{vac}\rangle. \end{aligned}$$

We obtain

$$H = \sum_{k=0}^{N/2-1} H_k, \quad (\text{B4})$$

where

$$\begin{aligned} H_k &= [x_k \sigma_z + y_k \sigma_y] \oplus 0, \\ H_0 &= [x_{0,+} \sigma_z] \oplus [x_{0,-} \sigma_z], \end{aligned}$$

with

$$x_{0,\pm} = (x_0 \pm x_{N/2})/2 \quad (\text{B5})$$

and

$$\sigma_z = \begin{pmatrix} -1 & 0 \\ 0 & 1 \end{pmatrix}.$$

The direct sum structure corresponds to the subspaces  $\{|1\rangle, |2\rangle\}$  and  $\{|3\rangle, |4\rangle\}$ . Thus, the problem is reduced to  $N/2$  noninteracting four-level systems.

We can easily compute the eigenvalues of  $H_k$ . For  $k \neq 0$ , they are given by  $E_{k,\pm}^{1,2} = \pm z_k$  and  $E_{k,\pm}^{3,4} = 0$  (doubly degenerate), where

$$z_k = \sqrt{x_k^2 + y_k^2}, \quad (\text{B6})$$

and by  $E_{0,\pm}^{1,2} = \pm x_{0,+}$  and  $E_{0,\pm}^{3,4} = \pm x_{0,-}$  for  $k = 0$ .

The expectation value of the Hamiltonian in a product state  $|p\rangle = |p_1, \dots, p_{N/2}\rangle$  with  $p_k = 1, \dots, 4$  is

$$E_p = \langle p | H | p \rangle = \sum_{k=0}^{N/2-1} \langle p_k | H_k | p_k \rangle = \sum_{k=0}^{N/2-1} E_{p_k}, \quad (\text{B7})$$

where, for  $k = 1, \dots, N/2 - 1$ ,

$$\begin{aligned} E_{k,1} &= \langle 1 | H_k | 1 \rangle = -x_k, \\ E_{k,2} &= \langle 2 | H_k | 2 \rangle = x_k, \\ E_{k,3} &= \langle 3 | H_k | 3 \rangle = 0, \\ E_{k,4} &= \langle 4 | H_k | 4 \rangle = 0, \end{aligned}$$

with the rest zero, and

$$\begin{aligned} E_{0,1} &= \langle 1 | H_0 | 1 \rangle = -x_{0,+}, \\ E_{0,2} &= \langle 2 | H_0 | 2 \rangle = x_{0,+}, \\ E_{0,3} &= \langle 3 | H_0 | 3 \rangle = -x_{0,-}, \\ E_{0,4} &= \langle 4 | H_0 | 4 \rangle = x_{0,-}. \end{aligned}$$

In order to evaluate the expressions required for the simulation, we need to compute

$$\langle p | e^{iHt} | p \rangle = \prod_{k=1}^{N/2} \langle p_k | e^{iH_k t} | p_k \rangle. \quad (\text{B8})$$

We find that

$$\langle 1 | e^{iH_k t} | 1 \rangle = \cos(z_k t) - i \sin(z_k t) x_k / z_k, \quad (\text{B9a})$$

$$\langle 2 | e^{iH_k t} | 2 \rangle = \cos(z_k t) + i \sin(z_k t) x_k / z_k, \quad (\text{B9b})$$

$$\langle 3 | e^{iH_k t} | 3 \rangle = \langle 4 | e^{iH_k t} | 4 \rangle = 1, \quad (\text{B9c})$$

for  $k \neq 0$ , whereas

$$\langle n | e^{iH_0 t} | n \rangle = e^{iE_{0,n} t} \quad (\text{B10})$$

for  $n = 1, \dots, 4$ .

Although we are interested here in finite energies and temperatures, we now briefly discuss the zero temperature behavior (i.e., the ground state). In the basis introduced here, it can be written as

$$\prod_k (\alpha_k |1\rangle_k + \beta_k |2\rangle_k) \quad (\text{B11})$$

with eigenvalue  $E_0 = -\sum_k z_k$ . The coefficients  $\alpha_k$  and  $\beta_k$  depend on  $g$  and  $h$ . In particular, for  $h \gg g$ , we have  $\alpha_k \sim 1$  and  $\beta_k \sim 0$ , whereas for  $h \ll g$ , they change with  $k$ , and thus one gets superpositions of many configurations. At  $g = h$ ,  $z_{N/2-1} \rightarrow 0$  as  $N \rightarrow \infty$ , so that the gap closes and there is a quantum phase transition.

## 1. Microcanonical average

We now compute some of the quantities that are used in the numerical illustrations. For the sake of simplicity, let us assume that

$$A = \sum_{k=0}^{N/2-1} A_k, \quad (\text{B12})$$

where  $A_k$  only depends on  $b_{\pm k}$  for  $k \neq 0$  and on  $b_{0,N/2}$  for  $k = 0$ . Mapping back to qubits, operator  $A$  is such that it can be decomposed into a sum of local operators, of which the magnetization we used in Eq. (34) is an example.

Let us start with

$$\langle p | P_\delta(E) | p \rangle = \sum_{m=-R}^R c_m e^{i2mE/N} a_p(t_m),$$

where the  $c_m$  are defined with respect to  $\delta$  and  $R$  is given in Eq. (12) or (27). We have

$$a_p(t_m) = \langle p | e^{-i2mH/N} | p \rangle = \prod_k \langle p_k | e^{-i2mH_k/N} | p_k \rangle,$$

which can be readily computed using Eqs. (B9).

We can also compute Eq. (28) directly,

$$A(E) = \frac{\sum_{m=-R}^R c_m e^{i2mE/N} n_m}{\sum_{m=-R}^R c_m e^{i2mE/N} d_m}, \quad (\text{B13})$$

where

$$\begin{aligned} n_m &= \frac{1}{2^N} \text{tr}(e^{-i2mH/N} A) \\ &= \frac{1}{2^N} \sum_{k=0}^{N/2-1} \left[ \prod_{q=0, q \neq k}^{N/2-1} \text{tr}(e^{-i2mH_q/N}) \right] [\text{tr}(e^{-i2mH_k/N} A_k)], \end{aligned}$$

$$d_m = \frac{1}{2^N} \text{tr}(e^{-i2mH/N}) = \frac{1}{2^N} \prod_{k=0}^{N/2-1} \text{tr}(e^{-i2mH_k/N}).$$

Defining

$$\begin{aligned} r_{m,k} &= \frac{1}{4} \text{tr}(e^{-i2mH_k/N}), \\ s_{m,k}^{(A)} &= \frac{1}{4} \text{tr}(A_k e^{-i2mH_k/N}), \end{aligned}$$

we can write

$$\begin{aligned} n_m &= \sum_{k=0}^{N/2-1} \left[ \prod_{q=0, q \neq k}^{N/2-1} r_{m,q} \right] s_{m,k}, \\ d_m &= \prod_{k=0}^{N/2-1} r_{m,k}. \end{aligned}$$

Using Eqs. (B9) and (B10), we have

$$r_{m,k} = \cos^2 \left( \frac{mz_k}{N} \right)$$

for  $k = 1, \dots, N/2 - 1$  and

$$r_{m,0} = \frac{1}{2} [\cos(2mx_{0,+}/N) + \cos(2mx_{0,-}/N)],$$

while the values of  $s_{m,k}$  depend on the observable.

## 2. Canonical average

In a similar way we may compute the canonical average

$$A(\beta) = \text{tr}(e^{-\beta H} A) / \text{tr}(e^{-\beta H}). \quad (\text{B14})$$

We have

$$\text{tr}(e^{-\beta H}) = 2^N \prod_k \tilde{r}_k,$$

$$\text{tr}(e^{-\beta H} A) = 2^N \sum_k \left[ \prod_{q \neq k} \tilde{r}_q \right] \tilde{s}_k = \text{tr}(e^{-\beta H}) \sum_k \frac{\tilde{s}_k}{\tilde{r}_k},$$

and, thus,

$$A(\beta) = \sum_k \frac{\tilde{s}_k}{\tilde{r}_k}, \quad (\text{B15})$$

where

$$\begin{aligned} \tilde{r}_k &= \frac{1}{4} \text{tr}(e^{-\beta H_k}), \\ \tilde{s}_k^{(A)} &= \frac{1}{4} \text{tr}(A_k e^{-\beta H_k}). \end{aligned}$$

Using Eqs. (B9) and (B10),

$$\begin{aligned} \tilde{r}_0 &= \frac{1}{2} [\cosh(\beta x_{0,+}) + \cosh(\beta x_{0,-})], \\ \tilde{r}_k &= \cosh^2(\beta z_k/2). \end{aligned}$$

## APPENDIX C: EFFICIENT COMPUTATION

Here we prove the statement that was used to show that one can compute Eq. (17a) efficiently. Namely, we show that, for any state  $\psi$  and any  $\delta \leq N/\sqrt{2}$ , there exists an interval of width at least  $\delta^2/6N$  contained in an energy window  $|E - E_\psi| \leq r\sigma_\psi$  (20) such that, for any  $E$  in that interval,

$$n(E) := \langle \psi | P_\delta(E) | \psi \rangle \geq \frac{1}{4} \left( \frac{\delta^2}{\delta^2 + 2\sigma_\psi^2} \right)^{3/2} \quad (\text{C1})$$

and

$$r = \{3 \log[2(1 + 2\sigma_\psi^2/\delta^2)]\}^{1/2}. \quad (\text{C2})$$

We emphasize that, for our purposes, we do not need a tight bound, so we will be very rough when bounding different



quantities with the goal of obtaining simple expressions. Moreover, although  $n(E)$  involves cosine filter (7), it is simpler to bound the result of the Gaussian approximation (8). Specifically, we prove that there is an energy interval of radius at least  $\tilde{\delta}^2/3N$ , contained in window (20) and with its center inside the spectral limits of  $H$ , such that, for any energy inside this interval,

$$\tilde{n}(E) := \langle \psi | e^{-(H-E)^2/2\tilde{\delta}^2} | \psi \rangle \geq \frac{1}{4(1 + \sigma_\psi^2/\tilde{\delta}^2)^{3/2}}, \quad (\text{C3})$$

where  $\tilde{\delta} = \delta/\sqrt{2}$ . Since  $\cos(x) > e^{-x^2}$  for  $|x| < 1.3$ , if  $E$  fulfills

$$\|(H - E)/N\|_\infty < 1.3, \quad (\text{C4})$$

we have  $n(E) \geq \tilde{n}(E)$ . Because  $\|H\|_\infty = N/2$ , and the center of the interval is within the limits of the spectrum of  $H$ , we can ensure that  $\|H - E\| \leq N + \delta^2/6N \leq 13N/12$ , and thus Eq. (C4) is satisfied for  $E$ , so that bounding the Gaussian approximation is enough for the desired result.

Let

$$T(r) = \int_{E_\psi - r\sigma_\psi}^{E_\psi + r\sigma_\psi} dE \tilde{n}(E) \frac{e^{-(E-E_\psi)^2/2\sigma_\psi^2}}{\sqrt{2\pi}\sigma_\psi}. \quad (\text{C5})$$

First, we show that, for Eq. (C2),  $T(r)$  is upper and lower bounded by some quantity, which will indicate that there exists some  $E$  in the interval such that  $\tilde{n}(E)$  is sufficiently large. Then, by upper bounding the derivative of  $\tilde{n}(E)$ , we conclude that there is a neighborhood of that  $E$  where  $\tilde{n}(E)$  fulfills the required condition.

We write  $T(r) = T_0 - T_1(r)$ , where

$$T_0 = \int_{-\infty}^{\infty} dE \tilde{n}(E) \frac{e^{-(E-E_\psi)^2/2\sigma_\psi^2}}{\sqrt{2\pi}\sigma_\psi} \quad (\text{C6})$$

and  $T_1(r) = T(r) - T_0$ . Given that  $\tilde{n}(E) \leq 1$ , we have

$$\begin{aligned} |T_1(r)| &\leq \sqrt{\frac{2}{\pi}} \int_r^\infty dE e^{-E^2/2} \\ &= \text{erfc}(r/\sqrt{2}) \\ &\leq \sqrt{\frac{2}{\pi}} \frac{e^{-r^2/2}}{r} \\ &\leq e^{-r^2/2}, \end{aligned} \quad (\text{C7})$$

where the last step uses the fact that, for Eq. (C2),  $r \leq \sqrt{2/\pi}$ . We can perform the integration in Eq. (C6) explicitly using Eq. (C3) with Eq. (C2), to get

$$T_0 = \frac{\langle \psi | e^{-(H-E_\psi)^2/[2(\sigma_\psi^2 + \tilde{\delta}^2)]} | \psi \rangle}{\sqrt{1 + \sigma_\psi^2/\tilde{\delta}^2}}. \quad (\text{C8})$$

Using  $e^{-x^2} \geq 1 - 2x^2$ , we obtain

$$T_0 \geq (1 + \sigma_\psi^2/\tilde{\delta}^2)^{-3/2}. \quad (\text{C9})$$

Thus, putting things together we obtain the lower bound

$$T(r) \geq (1 + \sigma_\psi^2/\tilde{\delta}^2)^{-3/2} - e^{-r^2/2} \geq \frac{1}{2}(1 + \sigma_\psi^2/\tilde{\delta}^2)^{-3/2}, \quad (\text{C10})$$

where we have chosen Eq. (C2) and used the fact that  $r^2/2 > \log[2(1 + \sigma_\psi^2/\tilde{\delta}^2)^{3/2}]$ .

In order to bound  $T(r)$  from above, let us denote by  $E_0$  the value where  $\tilde{n}(E)$  attains its maximum within interval (20). Note that  $E_0$  must be within  $[E_{\min}, E_{\max}]$  (since, if it was outside, choosing the closest extreme eigenvalue of  $H$  would yield a larger value for  $\tilde{n}$ ). We then have

$$T(r) \leq \tilde{n}(E_0) \frac{1}{\sqrt{2\pi}\sigma} \int_{E_\psi - r\sigma_\psi}^{E_\psi + r\sigma_\psi} dE e^{-(E-E_\psi)^2/2\sigma_\psi^2} \leq \tilde{n}(E_0), \quad (\text{C11})$$

so that

$$\tilde{n}(E_0) \geq \frac{1}{2}(1 + \sigma_\psi^2/\tilde{\delta}^2)^{-3/2}. \quad (\text{C12})$$

It is now possible to show that in a neighborhood of  $E_0$  of radius  $\delta^2/3N$ ,  $\tilde{n}$  takes sufficiently large values,  $\tilde{n}(E) \geq \tilde{n}(E_0)/2$ . For that, we just have to bound the derivative of  $\tilde{n}$  within interval (20). Let us call  $\tilde{n}'_{\max}$  the maximal value of the derivative of  $\tilde{n}(E)$  with respect to  $E$  in that interval, which occurs at some  $E = E_1$ . Although  $E_1$  could lie outside the limits of the spectrum of  $H$ , if that is the case, and using the fact that  $xe^{-x^2/2}$  has a maximum at  $x = 1$ , it is easy to see that the maximum cannot occur more than  $\tilde{\delta}$  away from the edges, so that  $\|H - E_1\|_\infty \leq N + \tilde{\delta} \leq 3N/2$ . Then we can bound,

$$|\tilde{n}'_{\max}| \leq \frac{3N\tilde{n}(E_1)}{2\tilde{\delta}^2} \leq \frac{3N\tilde{n}(E_0)}{2\tilde{\delta}^2}. \quad (\text{C13})$$

Finally, for any value of energy such that

$$|E - E_0| \leq \frac{\tilde{\delta}^2}{3N}, \quad (\text{C14})$$

we have

$$\tilde{n}(E_0) - \tilde{n}(E) \leq |\tilde{n}'_{\max}| |E_0 - E| \leq \frac{\tilde{n}(E_0)}{2}, \quad (\text{C15})$$

and, thus,

$$\begin{aligned} \tilde{n}(E) &= \tilde{n}(E_0) - |\tilde{n}(E_0) - \tilde{n}(E)| \\ &\geq \frac{\tilde{n}(E_0)}{2} \\ &\geq \frac{1}{4}(1 + \sigma_\psi^2/\tilde{\delta}^2)^{-3/2}. \end{aligned} \quad (\text{C16})$$

We end this appendix with a remark on the difficulty of finding the state  $\psi$  given a prescribed energy. As mentioned at the end of Sec. II B, this can be analyzed, for instance, through mean field theory for product states, or with matrix product states in one-dimensional problems. Regarding rigorous bounds, for product states, it was shown by Lieb [86] that, for any local Hamiltonian with a spectrum contained in  $[-N, N]$ , one can efficiently find a product state with an energy  $-DN$ , where  $D > 1/9$ . While this is a theoretical bound, we expect that, for most relevant Hamiltonians,  $D$  will be much larger (in fact, one can find much tighter bounds for specific models, like those used in this paper). Furthermore, as emphasized in the main text, one does not necessarily have to use product states, which gives access to even larger values of  $D$ .

#### APPENDIX D: CONVERGENCE TO THE MICROCANONICAL AND CANONICAL VALUES FOR A NONINTEGRABLE MODEL

In this appendix we numerically investigate the convergence to the microcanonical and canonical values of the quantities defined in the main text using exact diagonalization. In particular, we show how a polynomially decreasing  $\delta \sim \text{poly}(1/N)$  seems to be enough for the quantity  $A'_{\delta, \psi}(E)$  (17b) to converge to the true microcanonical expectation value.

We consider the Ising model in a tilted field, described by the Hamiltonian

$$H_{\text{Ising}} = J \left[ \sum_{n=1}^{N-1} \sigma_{n,z} \sigma_{n+1,z} + h \sum_{n=1}^N \sigma_{n,z} + g \sum_{n=1}^N \sigma_{n,x} \right], \quad (\text{D1})$$

which is in general nonintegrable, except in the limits  $g = 0$  (classical) and  $h = 0$  (transverse field Ising model). In the following, we choose a strongly nonintegrable point  $h = 0.5, g = -1.05$  [87], and we take  $J = 1$ . Note that this corresponds to a different normalization for  $H/N$  than that used in the main text. However, it is enough to ensure that Eq. (7) filters out energies much farther than the width  $\delta$  for all the states analyzed here (see Appendix E), so that it allows us to study how the microcanonical values are approached as the width decreases.

We consider real translationally invariant product states, which can be parameterized as  $|\Psi(\theta)\rangle = |p(\theta)\rangle^{\otimes N}$ , where  $|p(\theta)\rangle = \cos\theta|0\rangle + \sin\theta|1\rangle$ . In the thermodynamic limit, these states have energy density  $E/N = \cos^2(2\theta) + h \cos(2\theta) + g \sin(2\theta)$ , ranging over most of the energy band (see Fig. 8).

We choose three values of  $\theta$  corresponding to states in the lower part of the energy band, namely  $\theta_1 = \pi/4$ , for which  $E_1/(JN) = g = -1.05$ ,  $\theta_2 = \pi/3$ , for which  $E_2/(JN) = -0.909$ , and  $\theta_3 = \pi/6$ , with  $E_3/(JN) = -0.409$ . For each of these states, and for several local

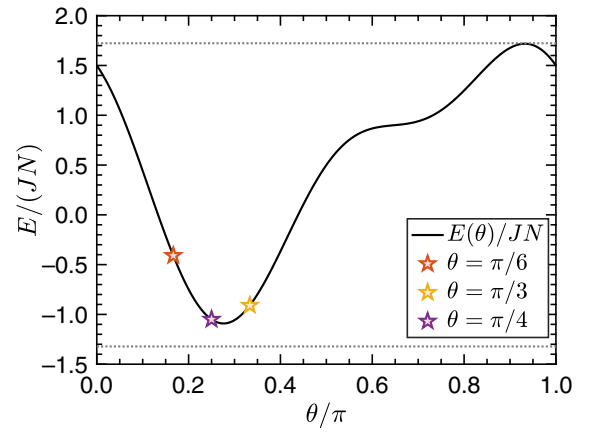


FIG. 8. Energy density in the nonintegrable Ising chain (D1) in the thermodynamic limit. The solid line indicates the energy density of translationally invariant real product states, while the dashed horizontal lines indicate the energy densities of the ground and maximally excited states (estimated numerically with MPS). The colored symbols indicate the states chosen in our numerical simulations.

observables, we compute exactly the expression  $A'_{\delta, \psi}(E = E_\psi)$  from Eq. (17b) for different values of  $\delta$ . In order to check that it is enough to decrease  $\delta$  polynomially with the system size, we run the calculations for system sizes  $10 \leq N \leq 28$ , and  $\delta \propto N^s$  for  $s = 0, -1, -2$ . The results, for observables  $A = \sigma_{[N/2],z} \otimes \sigma_{[N/2+1],z}$  and  $A = \sigma_{[N/2],x}$ , are shown in Fig. 9. As reference, we estimate the microcanonical expectation values in the thermodynamic limit using uniform matrix product state (MPS) [74] (more concretely, we approximate the canonical ensemble at the same energy density in the thermodynamic limit as a matrix product operator, in which the observables can be easily computed, and use the fact that in this limit, both ensembles are equivalent).

Our results indicate that, although a constant value of  $\delta$  is not enough for  $A'_{\delta, \psi}$  to approximate the microcanonical value, when  $\delta$  decreases as  $1/N^2$ , the expectation values indeed converge. For  $\delta \propto 1/N$ , we observe that the values are reasonably close, and they would be compatible with a slower convergence.

#### APPENDIX E: APPROXIMATING THE COSINE FILTER AS GAUSSIAN

We want to bound the absolute value of the difference

$$f_M(x) = e^{-Mx^2/2} - \cos^M x. \quad (\text{E1})$$

Since both terms are even, we can consider only  $x \geq 0$ . If  $|x| < \pi/2$ ,  $e^{-x^2/2} \geq \cos x$ , and both terms are positive, so also  $0 \leq \cos^M x \leq e^{-Mx^2/2}$ , and

$$|f_M(x)| = f_M(x) \leq e^{-Mx^2/2}. \quad (\text{E2})$$

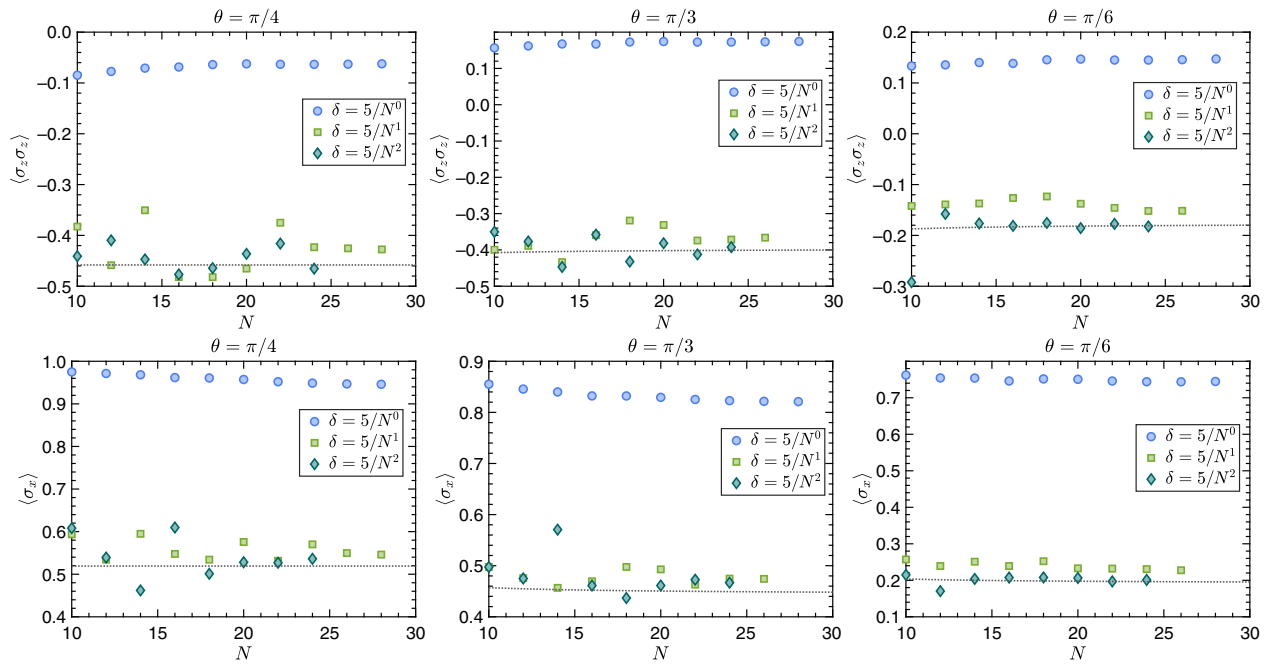


FIG. 9. Convergence of Eq. (17b) to the microcanonical values for several local observables, in the nonintegrable Ising model (D1). Each column corresponds to a real translationally invariant product state (in order of increasing energy density), determined by  $\theta = \pi/4$  (left),  $\pi/3$  (center), and  $\pi/6$  (right). The upper row illustrates the results for the two-site observable  $\sigma_z \otimes \sigma_z$ , and the lower row for  $\sigma_x$ , measured, in both cases, in the center of the chain. The dashed lines indicate the microcanonical values in the thermodynamic limit corresponding to the same energy density. Our results show that  $\delta \propto 1/N^2$  converges fast to the microcanonical expectation value, and for  $\delta \propto 1/N$ , the values are reasonably close.

The bound actually holds for slightly larger  $x$ , as long as  $e^{-x^2/2} \geq |\cos x|$ , which is true up to  $x_1 \approx 0.566\pi$ .

Actually, the Gaussian form also approximates the cosine beyond this value, since, for  $x_1 \leq x \leq \pi - \mu$ , it holds that

$$|f_M(x)| = \cos^M x - e^{-Mx^2/2} \leq \cos^M x \leq \cos^M \mu, \quad (\text{E3})$$

where we have assumed that  $M$  is even, as in the text. Thus, the difference decreases exponentially with  $M$  up to  $x = \pi - \mu$ .

At very small  $x$  it is more useful to use the Taylor expansion. There we can note that at very small  $|x|$  the difference vanishes as  $f_M(x) \approx Mx^4/12 + O(x^6)$ , and, in fact,  $f_M(x) \leq Mx^4/12$ .

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