Wall modes and the transition to bulk convection in rotating Rayleigh-Bénard convection

Xuan Zhang ^(D), ¹ Philipp Reiter, ¹ Olga Shishkina, ¹ and Robert E. Ecke ^(D), ², ³

¹Max Planck Institute for Dynamics and Self-Organization, 37077 Göttingen, Germany ²Center for Nonlinear Studies, Los Alamos National Laboratory, Los Alamos, New Mexico 87545, USA ³Department of Physics, University of Washington, Seattle, Washington 98195, USA

(Received 16 October 2023; accepted 26 April 2024; published 21 May 2024)

We investigate states of rapidly rotating Rayleigh-Bénard convection in a cylindrical cell over a range of Rayleigh numbers $3 \times 10^5 \le \text{Ra} \le 5 \times 10^9$ and Ekman numbers $10^{-6} \leq \text{Ek} \leq 10^{-4}$ for Prandtl number Pr = 0.8 and aspect ratios $1/5 \leq \Gamma \leq 5$ using direct numerical simulations. We characterize, for perfectly insulating sidewall boundary conditions, the first transition to convection via wall mode instability and the nonlinear growth and instability of the resulting wall mode states, including a secondary transition to time dependence. We show how the radial structure of the vertical velocity u_{z} and the temperature T is captured well by the linear eigenfunctions of the wall mode instability where the radial width of u_z is $\delta_{u_z} \sim Ek^{1/3}r/H$ whereas $\delta_T \sim e^{-kr}$ (k is the wave number of a laterally infinite wall mode state). The disparity in spatial scales for $Ek = 10^{-6}$ means that the heat transport is dominated by the radial structure of u_z since T varies slowly over the radial scale δ_{u} . We further describe how the transition to a state of bulk convection is influenced by the presence of the wall mode states. We use temporal and spatial scales as measures of the local state of convection and the Nusselt number Nu as representative of global transport. Our results elucidate the evolution of the wall state of rotating convection and confirm that wall modes are strongly linked with the boundary zonal flow being the robust remnant of nonlinear wall mode states. We also show how the heat transport (Nu) contributions of wall modes and bulk modes are related and discuss approaches to disentangling their relative contributions.

DOI: 10.1103/PhysRevFluids.9.053501

I. INTRODUCTION

Thermal convection driven by buoyancy and subject to rotation [1,2] is a phenomenon of great relevance in many physical disciplines, especially in geo- and astrophysics and also in engineering applications. In the Rayleigh-Bénard convection (RBC) geometry a fluid with thermal-expansion coefficient α , kinematic viscosity ν , and thermal diffusivity κ is bounded above and below by isothermal boundaries separated by a height H and heated from below so as to produce a temperature difference Δ . In a gravitational field with acceleration g, the strength of thermal forcing is determined by the Rayleigh number $\text{Ra} \equiv \alpha g \Delta H^3/(\kappa \nu)$, and the fluid type by the Prandtl number $\text{Pr} = \nu/\kappa$. In the presence of rotation about an axis parallel to gravity, the influence of the Coriolis force is a balance of angular rotation rate Ω_d and viscous diffusion time H^2/ν . There are many ways to represent this balance, namely, nondimensional $\Omega \equiv \Omega_d H^2/\nu$ and the Taylor

Published by the American Physical Society under the terms of the Creative Commons Attribution 4.0 International license. Further distribution of this work must maintain attribution to the author(s) and the published article's title, journal citation, and DOI. Open access publication funded by Max Planck Society.

number Ta $\equiv (2\Omega)^2$. Another measure popular in geophysics that we use here is the Ekman number Ek $\equiv \nu/(2\Omega_D H^2)$ that nicely captures the limit of rapid rotation as a small parameter, i.e., Ek $\rightarrow 0$. A similar parameter that measures the balance of buoyancy and rotation is the convective Rossby number Ro $\equiv \sqrt{\text{Ra}/(\text{PrTa})} = (\text{Ra}/\text{Pr})^{1/2}\text{Ek} = \sqrt{g\alpha\Delta/H}/(2\Omega_d)$ where the latter is the ratio of buoyancy frequency and rotation frequency. Note that Ro is independent of dissipation parameters. Unless otherwise specified, we use length scale *H*, timescale $\sqrt{H/(g\alpha\Delta)}$, and temperature scale Δ for nondimensional variables.

In physical realizations of convection, there are solid horizontal boundaries, usually of cylindrical geometry with aspect ratio $\Gamma \equiv D/H$ (*D* is the cylinder diameter), that have a finite thermal conductivity and diffusivity. In direct numerical simulations (DNSs), these sidewall boundaries can be perfectly insulating or, less frequently, perfectly conducting. Experimental boundaries are somewhere in between these limits and tend towards the insulating case so as to minimize heat transport through the sidewalls. We consider here the case of perfectly insulating sidewalls; results for perfectly conducting boundaries will be presented elsewhere.

The global response of the system is measured by the molecular-diffusion-normalized heat transport, Nusselt number $\text{Nu} \equiv (\langle u_z T \rangle_z - \kappa \partial_z \langle T \rangle_z)/(\kappa \Delta/H)$. Here, T denotes the temperature, **u** is the velocity field with component u_z in the vertical direction, and $\langle \cdot \rangle_z$ denotes the average in time and over a horizontal cross section at height z from the bottom.

The multitude of states of rotating Rayleigh-Bénard convection (RRBC) in a fluid layer bounded above and below by isothermal boundaries have been elucidated over many decades going back to early linear stability calculations [3,4] and preliminary experiments [5] (see also Ref. [2]). They found that the onset of bulk convection was strongly suppressed by the rotation-induced Coriolis force, which inhibits the vertical motions associated with thermal convection via mechanisms associated with the Taylor-Proudman constraint on slow, inviscid flows [6]. In particular, the critical Rayleigh number Ra_c was found to vary approximately as $Ek^{-4/3}$ for rapid rotation. Whereas the form of convection without rotation is in the form of convection rolls with characteristic length $\lambda \approx 2H$ (for a roll pair), rotation induces vortical motions near onset [7] with a corresponding characteristic length $\lambda \sim Ek^{1/3}$ [3]. Experimental heat transport measurements [8–11] confirmed the suppression of the convective onset by rotation, noted that some manner of convection ensued prior to strong bulk flow of linear theory, and revealed that rotation could, over some ranges in Ra and Pr, enhance heat transport over nonrotating turbulent RBC. The nature of the "subcritical instability" identified in the heat transport measurements was partially explained by linear stability calculations of stationary wall-mode states [11,12] but simultaneous heat transport and shadowgraph visualization experiments [13,14] together with theory [15,16] showed that these wall-mode states precess in the rotating frame and have precession frequency ω and integer (resulting from the cylindrically periodic azimuthal symmetry) mode number m. Further experiments [17-19] and linear stability calculations [16,20,21] showed beautiful correspondence between theory and experiment—here we denote $Ra_w \sim Ek^{-1}$ as the critical value for the transition from the diffusive conduction state to one of wall-mode convection. A phase diagram showing the boundaries of different convective states in RRBC is presented in Fig. 1 where we plot $Ra/Ra_c(Ek = 0)$ versus Ek.

The enhancement of heat transport (Nu) in RRBC compared with its value in nonrotating RBC (Nu₀) has been observed in many experimental studies [8,14,18,25–29] and DNS [30–34] and has been attributed to Ekman pumping from the thermal boundary layers. This enhancement occurs over a range of Ra, Pr, and Ek, with states at low Pr or small Ek not showing an enhancement but rather a decrease in Nu for all Ra [9,11,24,35–38]. We denote the value Ra_t at fixed Ek as the transition from buoyancy-dominated flow to rotation-influenced behavior.

Another well-defined region of rapidly rotating convection is the geostrophic regime where the flow maintains geostrophic balance and is rotation-dominated. The importance of this regime to geophysical and astrophysical situations was articulated through the derivation [39] and simulation [40–42] of a reduced set of equations valid in the limit $Ek \rightarrow 0$ and $Ra \rightarrow \infty$ such that the scaled variable $\widetilde{Ra} \equiv RaEk^{4/3}$ remains finite. The evolution of states from the onset of bulk convection at $\widetilde{Ra}_c \approx 8.7$ depends on Pr and includes cellular patterns, convective Taylor vortices, plume



FIG. 1. Schematic phase diagram of states of rotating convection (perfectly insulating sidewall boundary conditions): RaEk^{4/3}/A₀ (dashed black line) where $A_0 = 8.7$ is the asymptotic Ek \rightarrow 0 value versus Ek. The lines for Ra_c (dashed blue line) [4,22,23] and Ra_w (solid blue line) [20] are from linear stability analysis with idealized boundary conditions on laterally infinite domains. Ra_g marks the approximate boundary of the geostrophic, rotation-dominated region and Ra_t denotes the transition to buoyancy-dominated flow [1,2,24]. The precise details of this diagram depend on parameters including Pr and Γ and are not included here. The data points show the parameter values reported in this paper.

states, and geostrophic turbulence. The range $1 \leq Ra/Ra_c \leq 20$ was explored using these reduced equations and a prediction was made that in the geostrophic turbulence regime Nu $\sim \widetilde{Ra}^{3/2}$ where it is controlled by the interior quasigeostrophic (QG) flow rather than being boundary layer controlled, a result that was demonstrated in the simulations only for Pr = 1 over a short range in RaEk^{4/3}. The experimental and DNS test of these predicted states [26,28,29,36] and their Nu scalings is challenging [43] owing to the need for very small Ek < 10⁻⁶, large Ra, and the experimental resolution to measure rather small deviations from Nu \approx 1, i.e., over a range 1 \leq Nu \leq 20.

The flow structure of rotating convection for sufficiently small Ek ≤ 0.02 starts with wall modes that arise as a supercritical bifurcation from the no-flow base state [13–16,20,21]. These initial flow states persist to higher Ra and serve as the nonlinear base state for the onset of fully three-dimensional bulk convection. In the idealized case usually considered in DNS, the sidewall is perfectly adiabatic. In experiments, however, this is not the case. The details of the onset and evolution of wall modes depend on the ratio of thermal conductivity of the sidewalls to that of the nonconvecting fluid. Often the sidewalls are acrylic or polycarbonate with thermal conductivity of order 0.2 W/(m K) compared with, for example, water with 0.6 W/(m K). These conditions yield small (order 10%) modification of the asymptotic critical parameters for insulating sidewall boundary conditions [21] with higher Rayleigh number for the onset of the wall modes Ra_w (74.4 versus 63.6), lower procession frequency at the onset $\omega_{\kappa_c} = \omega_c H^2/\kappa = (57.7 \text{ versus 66.1})$, and lower wave number k_w (5.55 versus 6.07). Here we consider the idealized case of perfectly insulating sidewalls.

A. Direct numerical simulations

We consider rotating Rayleigh-Bénard convection (RRBC) in a cylindrical container with radius R and height H ($\Gamma \equiv 2R/H$). The governing equations for RRBC in the Oberbeck–Boussinesq

approximation include the continuity equation $\nabla \cdot \mathbf{u} = 0$ and the momentum equation and heat equations

$$\partial_t \mathbf{u} + (\mathbf{u} \cdot \nabla) \mathbf{u} = -\nabla p + \nu \nabla^2 \mathbf{u} - 2\Omega_d \mathbf{e}_z \times \mathbf{u} + \alpha (T - T_0) g \mathbf{e}_z, \tag{1}$$

$$\partial_t T + (\mathbf{u} \cdot \nabla)T = \kappa \nabla^2 T. \tag{2}$$

Here, *p* denotes the reduced kinematic pressure, \mathbf{e}_z is the unit vector pointing upward, and $T_0 \equiv (T_+ + T_-)/2$. We do not include any centrifugal acceleration because, in laboratory realizations of RRBC, it is negligible compared with the gravitational acceleration, i.e., the Froude number is very small, $Fr \equiv \Omega_d^2 R/g \ll 1$. We apply the standard boundary conditions for RBC which are no-slip for the velocity ($\mathbf{u} = 0$) at all walls, isothermal at the bottom plate ($T = T_+$ at z = 0) and top plate ($T = T_-$ at z = H, where $T_- < T_+$), and adiabatic ($\partial T/\partial \mathbf{n} = 0$) at the sidewall for the temperature. Here \mathbf{n} denotes a unit vector orthogonal to the boundary surface.

Taking as the reference quantities $\tau = H/\sqrt{\alpha g \Delta H}$ for time, Δ for temperature, H for length, $\sqrt{\alpha g \Delta H}$ for velocity, and $\alpha g \Delta H$ for the reduced kinematic pressure, from equations (1) and (2) we obtain the following dimensionless governing equations:

$$\partial_t \mathbf{u} + (\mathbf{u} \cdot \nabla) \mathbf{u} = -\nabla p + (\Pr/\operatorname{Ra})^{1/2} \nabla^2 \mathbf{u} - \operatorname{Ro}^{-1} \mathbf{e}_z \times \mathbf{u} + T \mathbf{e}_z,$$
(3)

$$\partial_t T + (\mathbf{u} \cdot \nabla)T = (\Pr \operatorname{Ra})^{-1/2} \nabla^2 T.$$
 (4)

The resulting set of equations (3) and (4) together with $\nabla \cdot \mathbf{u} = 0$ is solved numerically using the direct numerical solver GOLDFISH [44,45] in its latest version [46,47], which is optimized for efficient computation with massive parallelization. The computational code uses a fourth-order finite-volume discretization on staggered grids and a third-order Runge–Kutta time integration scheme. The direct numerical solver GOLDFISH utilizes a two-dimensional pencil decomposition employing the 2DECOMP library [48] and a HDF5 file management.

Unless otherwise specified, all quantities are made dimensionless by τ , H, and Δ (velocity by H/τ). The exception is the radial distance, which we explicitly scale with the radius R, i.e., $r/R = (2/\Gamma)r/H$.

B. Linear stability theory

One can establish using linear stability calculations the onset of RRBC of a laterally infinite fluid layer [4,22,23] and of precessing wall modes [16,20,21,25] that arise owing to the existence of sidewalls. Here we summarize the asymptotic results (see the Appendix for empirical fits for larger Ek) for $Pr \gtrsim 0.68$ and $Ek \rightarrow 0$. For the bulk instability the onset Rayleigh number Ra_c, critical wave number k_c and characteristic length λ_c are

$$Ra_{c} \approx 8.70 \text{ Ek}^{-4/3} (1 - 1.11 \text{Ek}^{1/6} + 0.153 \text{Ek}^{1/3}),$$
(5)

$$k_{\rm c} \approx 1.31 \,{\rm Ek}^{-1/3} (1 - 0.554 {\rm Ek}^{1/6} - 0.345 {\rm Ek}^{1/3}),$$
 (6)

$$\lambda_{\rm c} \approx 4.80 \,{\rm Ek}^{1/3} (1 + 0.554 {\rm Ek}^{1/6} + 0.345 {\rm Ek}^{1/3}),$$
(7)

where k_c and λ_c are normalized by H^{-1} and H, respectively, and $\lambda_c k_c = 2\pi$.

For the precessing wall modes with perfectly insulating boundaries, one has for a planar wall (no curvature) and to second order in Ek (see also the Appendix) [20,21,25]

$$Ra_{w} \approx \pi^{2} \sqrt{6\sqrt{3}} Ek^{-1} + 46.5 Ek^{-2/3} \approx 31.8 Ek^{-1} + 46.5 Ek^{-2/3},$$
(8)

$$k_{\rm w} \approx \pi \sqrt{2 + \sqrt{3}} - 35.0 \,{\rm Ek}^{1/3} \approx 6.07 - 35.0 \,{\rm Ek}^{1/3},$$
 (9)

$$m_{\rm w} = (\Gamma/2)k_{\rm w},\tag{10}$$

053501-4

$$\omega_{\kappa_{w}} \equiv \omega_{w} H^{2} / \kappa = \omega_{ff_{w}} \text{Ra}^{1/2} \text{Pr}^{1/2} \approx 2\pi^{2} \sqrt{6} + 3\sqrt{3} - 732 \text{ Ek}^{1/3}$$

$$\approx 66.0 - 732 \text{ Ek}^{1/3}, \qquad (11)$$

$$\omega_{d_w} \equiv \omega_w / \Omega = 2\omega_{\kappa_w} \text{Ek} \operatorname{Pr}^{-1} \approx \text{Ek} \operatorname{Pr}^{-1} (4\pi^2 \sqrt{6} + 3\sqrt{3} - 1465 \text{Ek}^{1/3})$$

$$\approx 132.1 \text{ Ek} \operatorname{Pr}^{-1} - 1465 \text{ Ek}^{4/3} \operatorname{Pr}^{-1}, \qquad (12)$$

where the spatial scale is H and the precession frequency ω_w is made dimensionless with the timescale H^2/κ for ω_{κ_w} , $\sqrt{H/(g\alpha\Delta)}$ for ω_{ff_w} , and Ω^{-1} for ω_{κ_c} and ω_{d_w} , respectively. Recall that for realistic sidewall boundary conditions [21] these asymptotic values are modified by about 10%; something that needs to be considered when comparing experimental and DNS results. For finite cylindrical geometries, especially for small Γ , the values also shift slightly with smaller Ra_w for m = 1, 2 [16,20] compared with their planar wall values.

The paper is organized to generally track the evolution of the RRBC state for $Ek = 10^{-6}$ and Pr = 0.8 over the range $3 \times 10^7 \le Ra \le 5 \times 10^9$. In Sec. II B, we describe the properties of wall modes including (Sec. II A) effects of aspect ratio, (Sec. II B) their spatial and temporal structure in terms of linear state eigenfunctions, and (Sec. II C) the evolution of the nonlinear state through a subcritical bifurcation. We also consider (Sec. II D) the transition to the bulk state and (Sec. II E) the associated spatial and temporal properties of the coexisting wall mode and bulk mode states where we use the term boundary zonal flow (BZF) to label the wall-localized state in the presence of bulk flow. In Sec. II F, we focus on the heat transport (Nu). We provide conclusions in Sec. III and end with the acknowledgments. The Appendix has details of the data presented here, some extra features of wall modes, and the empirical fits of critical parameters for larger Ek.

II. WALL MODES

The goal of this paper is to address the evolution of the states of rotating convection at small Ek from the onset of wall modes through the onset of bulk convection and up to the upper threshold of the geostrophic regime (Fig. 1). We begin with a description of pure wall modes for Pr = 0.8, Ek = 10^{-6} and 10^{-4} , and for a range of $1/5 \le \Gamma \le 5$ (details of these parameters are presented in Table II in the Appendix). Within this range we identify the critical onset Ra_w , the mode number m (and corresponding azimuthal wave number k), and different measures of temporal and spatial degrees of freedom. Owing to the previous lack of data for wall modes at small Ek, we systematically consider aspects of the onset of wall modes with a particular emphasis on the effects of small $\Gamma < 1$. We then address the nonlinear evolution of the wall modes including a subcritical transition to time dependence and the mechanisms for that transition. We identify three distinct regions. First, one has steady (in the precessing frame) wall mode states for $3 \times 10^7 \lesssim \text{Ra} \lesssim 3 \times 10^8$. For $4 \times 10^8 \lesssim \text{Ra} \lesssim$ 9×10^8 , there is a subcritical bifurcation to time dependence followed by increasingly nonlinear states. The mechanism for this bifurcation is the lateral ejection of plumes originating near the azimuthal zero crossings of temperature and u_z and propagating into the interior, see also Ref. [49]. Finally, one gets a transition at $Ra \approx 10^9$ to bulk convection where the wall mode states are in a dynamic balance with the bulk states in the form of a BZF. We characterize this later state in the range $10^9 \leq \text{Ra} \leq 5 \times 10^9$. Some aspects of these states emphasizing the connection between wall modes and BZF were reported earlier by us [50].

A. Effect of aspect ratio Γ

In Fig. 2, we show *m* for a number of datasets with different Γ and Ek with corresponding values of Γ of 2 [14], 5 [17], 10 [19], 2 [51], 1/5 [52], 3/2 [49], and {1/3, 1/2, 3/4, 1, 2} [53]. The first four examples are for Pr = 6.4 and primarily involve the properties of the wall-mode state. Most of the other values are for the BZF state where the wall localization coexists with a turbulent



FIG. 2. Azimuthal mode number *m* vs Γ . The shaded region between dashed lines corresponds to the dependence for $m(\Gamma)$ between the critical values for Ek = 0 (upper, asymptotic, blue dashed curve) and for Ek = 10^{-5} (lower, red dashed curve). Blue circles: Pr = 0.8 [53,54], black squares: Pr = 5.2 [52], red diamonds: Pr = 1 [49], magenta upward triangles: Pr = 6.4 [14,17,19], cyan downward triangles: Pr = 0.8 [50], open circles: Pr = 6.4 [51].

interior bulk mode. Because of the azimuthal periodicity, *m* is constrained to integer values so that, for small $\Gamma < 1$, one has m = 1. The BZF state has systematically smaller values of $m \approx 2\Gamma$ than the asymptotic (Ek \rightarrow 0) condition $m_{\rm w} \approx 3.03\Gamma$ although the Ek correction in (1) is substantial in the range $10^{-6} \leq \text{Ek} \leq 10^{-4}$, corresponding to $2.1 \leq m_{\rm w} \leq 2.83$. (For Ek $> 10^{-4}$ the expansion to second order in Ek is inaccurate, see the Appendix). In Fig. 2, we show as a shaded region the predicted range of $m_{\rm w}$ over the range $0 < \text{Ek} < 10^{-5}$. The trends and approximate range of *m* are consistent with a strong connection between wall mode states and the BZF. We next consider the onset of wall modes in more detail.

The critical Ra_w for $\Gamma = 1/5$, 1, 5 are determined from the DNS with Pr = 0.8 at Ek = 10⁻⁴ and $\Gamma = 1/2$, 2 at Ek = 10⁻⁶, see Table I. One can understand the dependence of Ra_w on Γ by considering the discrete wave number $k_{\rm m}$ of the corresponding mode number m ($k_{\rm m} = 2m/\Gamma$) which is selected owing to azimuthal periodicity and comparing it to the critical wave number $k_{\rm w}$ for a planar (flat) sidewall boundary. The marginal stability curve, Fig. 3, is defined by the k dependence of the onset Ra designated as the marginal stability Rayleigh number Ra_M(k). For small differences $k - k_{\rm w}$ the dependence of Ra_M is quadratic, i.e., $\epsilon_{\rm M} \equiv \text{Ra}_{\rm M}(m)/\text{Ra}_{\rm w} - 1 = \xi_0^2 (k - k_{\rm w})^2$ [17,19,21]. The quadratic approximation is valid only for small $k - k_{\rm w}$ and ϵ whereas for larger values there are terms higher order in $k - k_{\rm w}$ (see, e.g., Ref. [55]). Nevertheless, this representation gives a

Ek	Г	Ra _w	Ro _w	m
10 ⁻⁶	1/2	2.8×10^{7}	0.07	1
10^{-6}	2	$3. \times 10^{7}$	0.07	3
10 ⁻⁴	1/5	8.3×10^{5}	0.10	1
	1	$3. \times 10^{5}$	0.03	2
	5	3.3×10^{5}	0.06	14

TABLE I. Data from DNS with Pr = 0.8 showing Ek, Γ , Ra_w , Ro_w , and *m*.



FIG. 3. $\operatorname{Ra}_M/\operatorname{Ra}_w$ versus k, where Ra_M is the value of Ra along the marginal stability curve given by (solid black) $\operatorname{Ra}_M/\operatorname{Ra}_w = 1 + \xi_0^2 (k - k_w)^2$ where, for periodic azimuthal conditions, one has discrete $k = 2m/\Gamma$. (a) $\operatorname{Ek} = 10^{-6}$ with corresponding $k_w = 5.75$ and $\xi_0 = 0.18$ [21], $\operatorname{Pr} \approx 1$ and data points (Γ, m) = (1/5, 1), red; (1/3, 1), green; (1/2, 1) and (1/2, 2), blue; (3/2, 3) to (3/2, 6), black [49]; (2, 4) to (2, 8), magenta. (b) $\operatorname{Ek} = 10^{-4}$ with corresponding $k_w = 5.75$ and $\xi_0 = 0.18$ [21] and data points (Γ, m) = (1/5, 1), red; (1, 2) to (1, 4), blue; (5, 11) to (5, 15), magenta. Dashed black line [55] shows computed effects of higher-order terms at $\operatorname{Ek} \approx 2 \times 10^{-3}$ and corresponding data points (Γ, m) = (1/5, 1), red; (1, 4), blue.

reasonable semiquantitative description of the variation of Ra_M for different parameter values. The difference $k - k_w$ is particularly important for small Γ as we see below. The coefficient ξ_0^2 was evaluated for experimentally realized [14,17] sidewall thermal conductivity as a function of Ek [21] (see the Appendix). In Fig. 3(a), the curve is evaluated for $Ek = 10^{-6}$ with $k_w = 5.75$ and $\xi_0 = 0.18$, and the labeled data points are the discrete values k_m for $\Gamma = 1/5, 1/3, 1/2, 3/2$ [49], and 2 which have the lowest values of Ra_M/Ra_w . One can see the large and unintuitive dependence on $\Gamma \lesssim 1/2$ where m = 1: for $\Gamma = 1/5$, $Ra_M/Ra_w = 1.6$ whereas $\Gamma = 1/3$ yields $Ra_M/Ra_w = 1.003$ and $\Gamma = 1/2$ has $Ra_M/Ra_w = 1.1$. The increase in Ra_M/Ra_w is nonmonotonic in Γ and depends sensitively on k_w . Another observation concerns the predicted lowest value for $\Gamma = 3/2$ which is m = 4 whereas m = 6 was observed in DNS [49] upon increasing Ra but m = 4 when decreasing Ra. This would be consistent with the m = 6 mode resulting from starting the DNS for Ra $\gtrsim Ra_w$ from random initial conditions at Ra = 0 so that the state would evolve to an $m \neq m_w$, where m_w is the value corresponding to the smallest $k_{\rm m} - k_{\rm w}$. The effect of a sudden increase in Ra was used in experiments on wall modes [14,17,19] to access a wide range of available m. In Fig. 3(b), we consider Ek = 10^{-4} where $k_w = 4.9$ and $\xi_0 = 0.2$ for $\Gamma = 1/5, 1, 5$. Here the shift in Ra_M/Ra_w is even larger for $\Gamma = 1/5$ than for Ek = 10^{-6} but as Γ increases the density of discrete k_m increases because $\Delta k \equiv k_{m+1} - k_m = 2/\Gamma$ so that they increasingly cluster near Ra_M/Ra_w ≈ 1 . For example, for $\Gamma = 5$, we observe modes m = 14, 15 compared with the predicted values with the lowest Ra_M, m = 12, 13 (a small shift in k_w would account for that difference). More importantly, we again appear to find that the selected m can be different than the one closest to critical m_w owing to starting from random initial conditions; m = 14 near onset whereas m = 15 above onset and similarly for $\Gamma = 1, m = 2$ near onset and m = 3 above onset as shown in Fig. 4. When we slowly raise Ra from near onset, the selected *m* near onset is preserved.

To directly evaluate the observed and predicted onset values Ra_M , we plot in Fig. 5(a) $Ra_M/Ra_w = \epsilon_M + 1$ for different observed values of *m* at the corresponding Γ . The dramatic effect for $\Gamma = 1/5$ where m = 1 noted above is well captured by the marginal stability model with an



FIG. 4. Instantaneous temperature field at mid-height for $\text{Ek} = 10^{-4}$, Pr = 0.8. (a) $\Gamma = 1$, $\text{Ra} = 3 \times 10^5$, $\text{Nu} \approx 1$, m = 2, and (b) $\Gamma = 5$, $\text{Ra} = 3 \times 10^5$, Nu = 1.03, m = 14. Horizontal black bars show scale R/Γ .



FIG. 5. $\text{Ra}_{\text{M}}/\text{Ra}_{\text{w}}$ versus Γ . (a) The lines are calculations using values for m, $k_{\text{w}}(\text{Ek})$, and $\xi_{0}(\text{Ek})$. Different m = 1, 2, 4, 6, 14 are shown with corresponding colors (blue, red, orange, magenta, black); the dashed blue curve is evaluated for $\text{Ek} = 10^{-6}$ whereas the solid blue is for $\text{Ek} = 10^{-4}$. The data points are solid circle: $\text{Ek} = 10^{-4}$, Pr = 0.8; solid square: $\text{Ek} = 10^{-6}$, Pr = 0.8; triangle: $\text{Ek} = 10^{-6}$, Pr = 1 [49]; diamond: $\text{Ek} = 2 \times 10^{-3}$, Pr = 6.7 [16]. (b) Dotted lines show calculations [16] for m = 1, 2 taking into account the finite curvature of the cylindrical container for specific Γ .



FIG. 6. Linear eigenfunctions (discretized from data in Figs. 4 and 5 of Ref. [20] for $Ek = 10^{-6}$ and Pr = 7): (a) Eigenfunction of u_z versus $Ek^{-1/3}x$, where x reflects the distance from a flat sidewall. Characteristic lengths are indicated by the location of the peak $\delta_{u_{z_p}}$, the first zero crossing δ_{u_z} , and the minimum $\delta_{u_{z_m}}$. (b) Eigenfunction of T versus x. The dashed line is an exponential fit to the data $1.05e^{-x/0.13}$ that fits extremely well over the whole range except very near the sidewall on the order of the u_z radial width [see the inset where short red (long) dashed vertical line is r_{max} (r_0) of u_z]. (c) The geometric length scale $L_R = R = \Gamma/2$ is plotted versus values of Γ considered (solid black circles) and compared with the temperature length scale δ_T and the vertical velocity length scale δ_{u_z} evaluated at $Ek = 10^{-4}$ (dashed blue) and $Ek = 10^{-6}$ (solid blue).

observed factor of 2.6 in the onset Ra_M compared with a predicted ratio of about 2.1. For $\Gamma = 1/2$, 1, there is actually a decrease relative to Ra_w which is captured in calculations for finite Γ [16], see Fig. 5(b). As Γ is further increased, the observed *m* values yield $k \approx k_w$ so there is little or no shift in the onset Ra_M . Given that ϵ_M is not precisely quadratic except for small ϵ , ξ_0 is computed for imperfectly insulating sidewall boundary conditions, and the computations for Ra_w and k_w do not include rigid top or bottom boundary conditions, the agreement is quite satisfactory.

B. Wall mode spatial and temporal structure

We now turn to the spatial and temporal structure of the wall mode states. In previous work [50,53,54], we considered the radial boundary length scales of wall modes and BZF states defined by $\delta_{u_z} = \max \langle u_z^2 \rangle_{\phi,t}^{1/2}$ and the zero crossing δ_0 of $\langle u_\phi \rangle_{\phi,t}$, i.e., δ_0 is the distance to the sidewall where $\langle u_\phi \rangle_{\phi,t} = 0$, where $\langle \cdot \rangle_{\phi,t}$ means averaging in the azimuthal direction and in time. As we discuss below, δ_0 may not be a good measure of radial extent owing to the midplane symmetry of u_{ϕ} . Here we consider characteristics of the stationary and subcritical wall modes as represented in the fluid fields: temperature *T*, vertical velocity u_z , azimuthal velocity u_{ϕ} , radial velocity u_r , and vertical vorticity ω_z .

One feature of wall modes that is perhaps not appreciated is the decoupling of the radial length scales of the temperature field *T* from the vertical velocity u_z and vertical vorticity ω_z fields with decreasing Ek [20]. In Figs. 6(a) and 6(b), we show, respectively, the radial eigenfunctions of $u_z(r)$ and T(r) (interpolated from Ref. [20] for Ek = 10⁻⁶) where a Fourier–Bessel expansion is an excellent representation of u(z) (anticipating the results in a cylindrical cell) and *T* is well fit by an exponential function. The inset of Fig. 6(b) shows that the necessary condition at the wall, $\partial T/\partial r(r = R) = 0$, relaxes quickly with *r* on length scales of order δ_{u_z} . The decoupling of length scales results from the *T* length scale $\delta_T \sim k_w^{-1}$ (1/*e* decay length) being dependent on Ek only through its weak dependence $k_w(Ek)$ whereas the u_z length scale (first zero crossing) $\delta_{u_z} \sim Ek^{1/3}$ has an explicit Ek scaling. This separation of scales has important implications for RRBC in small aspect ratio containers with rapid rotation and for understanding the heat transport properties of wall modes and the BZF.



FIG. 7. Positive (red) and negative (blue) linear eigenfunctions of temperature T as in Fig. 6(b) and their sum (black) for planar wall modes at $Ek = 10^{-6}$ and Pr = 7 [20] plotted to approximate a radial profile at mid-height in a confined cylindrical geometry with an odd m so that T(r) is radially asymmetric about r = 0. (a) $\Gamma = 1/2$, (b) $\Gamma = 1$, (c) $\Gamma = 2$ and $\Gamma = 5$ (magenta).

In Fig. 6(c), we consider the different length scales associated with the eigenfunctions of u_z and T. We take the geometric length scale L_R of the container as R so that $L_R = \Gamma/2$ and compare with the length scales δ_T , and $\delta_{u_{z0}}$ evaluated at $Ek = 10^{-6}$ and 10^{-4} . One obtains $\delta_T \approx 0.13$ (consistent with Ref. [20], which predicts scaling of order $k_w^{-1} = 0.17$), and $\delta_{u_z} = 2.5 \text{Ek}^{1/3}$ which yields values of 0.12 for $Ek = 10^{-4}$ and 0.025 for $Ek = 10^{-6}$. Thus, for $Ek = 10^{-4}$, $\delta_T \approx \delta_{u_z}$ whereas for $Ek = 10^{-6}$, $\delta_T \approx 5\delta_{u_z}$. For $\Gamma \approx 1/3$, $\delta_T \approx L_R$. Thus, the temperature field is highly constrained by the geometry for small $\Gamma \leq 1$. On the other hand, for $Ek = 10^{-6}$, $\delta_{u_{z0}} \ll \delta_T$ and only becomes of order R for very small $\Gamma \approx 1/20$. Thus, u_z should not be affected by geometry and should have a form in cylindrical geometry similar to that of a planar boundary.

We expand on the disparity in length scales by considering in Figs. 7(a)–7(c), the wall-mode linear temperature eigenfunction of a planar wall mode state with perfectly insulating sidewall boundary conditions plotted versus r/R for different Γ : Fig. 7(a) 1/2, Fig. 7(b) 1, and Fig. 7(c) 2, 5. The positive and negative representations corresponding to opposite sides of the cylinder and the sum of the two as an approximation of the effects of finite geometry (see also Ref. [16]) are illustrated. For $\Gamma = 1/2$, there is considerable overlap of the two functions. For $\Gamma = 1$, one has just about achieved separation into individual contributions but one needs $\Gamma \gtrsim 2$ for a clean separation of the thermal field. For comparison, we have plotted vertical dashed lines showing δ_{u_c} .

Whereas the analysis above was for a planar wall [20], our finite cylindrical convection cell suggests that the appropriate representation in that geometry are radial eigenfunctions using sums of Bessel functions of the first kind, $J_m(kr)$ of order m [16]. It is therefore natural to consider $u_z(r)$ fields as Fourier–Bessel transforms with m = 1:

$$c_n = 2/[J_2(j_{1n})]^2 \int_0^1 r u_z(r) J_1(j_{1n}r) dr,$$
$$u_z(r) \approx \sum_{j=1}^N c_n J_1(j_{1n}r),$$

where $u_z(r)$ is the radial component of the vertical velocity field and j_{1n} are the *n*th order zeros of $J_1(x)$. Empirically, we find that N = 50 yields excellent fits to $u_z(r)$ which is primarily consequential in the advection of heat and the determination of the global heat transport Nu. In Fig. 8(a), we show the radial dependence of mean u_z for Pr = 0.8, $\Gamma = 1/2$, and $Ek = 10^{-6}$ evaluated at its azimuthal maximum (in the precessing frame) and at the midplane (z = 1/2) height for



FIG. 8. Midplane u_z (averaged in precessing frame at maximum in ϕ) versus r/R for (a) Ra/10⁷ = 3, 4, 5, 7, 10, 20, 30 where dark (light) blue corresponds to the smallest (largest) Ra. The inset of panel (a) shows Fourier–Bessel coefficients c_n normalized by $c_{n_{max}}$ (Ra) and averaged over the set of Ra in panel (a). (b) Expanded version of panel (a) with $0.7 \le r/R \le 1.0$. Inset of (b) is $\tilde{u}_z(r)$ reconstructed from the average of $c_n/c_{n_{max}}$ in the inset of panel (a). (c) u_z versus r/R of time-dependent wall-mode state for $4 \times 10^8 \le \text{Ra} \le 9 \times 10^8$ (colors per legend) with saturating peak amplitude and wider spread in δ_{u_z} . (d) Ra = 3×10^7 average radial structure (solid line) compared with linear eigenfunctions (dashed) [20] for u_{ϕ} (red) u_z (blue), and u_r (black). The first zero crossing of u_z is at about the same r as the maximum of u_r . All profiles are scaled such that the maximum values are one. (e) Variation of radial maximum values $u_{z_{max}}$ (blue solid circles), $u_{z_{ms,max}}$ (light blue inverted triangles), $u_{\phi_{max}}$ (red solid squares), and $u_{r_{max}}$ (black solid triangles) vs Ra – Ra_w. Linear scaling $u \sim (\text{Ra} - \text{Ra}_w)^{1/2}$ (blue, red, and black dashed lines, respectively). Vertical dashed lines from smaller to larger Ra: end of quasilinear regime (long-dashed) Ra $\approx 9 \times 10^7$ —note the decrease in u_{ϕ} , the subcritical instability to lateral jet ejection (dashed), the transition to aperiodic (chaotic) jet ejection (long-dashed), and the bulk onset (dashed).

 $3 \times 10^7 \leq \text{Ra} \leq 3 \times 10^8$. The inset is the normalized Fourier–Bessel decomposition coefficients c_n which has a simple form and is quite independent (modulo an overall scale factor that increases with Ra for Ra $< 3 \times 10^8$, i.e., for steady wall modes). An expanded view in Fig. 8(b) shows the details of the profiles near the sidewall and the definition of the radial extent of the main wall peak—here we take the first zero crossing of $u_z(r)$ as the characteristic width δ_{u_z} . The inset shows the radial profile obtained by averaging the normalized c_n values over the Ra range. The data points are from the amplitude-normalized planar eigenfunction of u_z shown in Fig. 6(a) and match almost perfectly, indicating again that cylindrical curvature is not important here for u_z . For larger



FIG. 9. Midplane T fields for Ra: (a) 3×10^7 , (b) 2×10^8 , and (c) 5×10^8 . Dashed lines show contours of the arithmetic mean of the top and bottom temperatures. Directions of rotation and precession are indicated in panel (a).

Ra $\geq 4 \times 10^8$, the profiles change slightly as shown in Fig. 8(b) with a larger spread in δ_{u_c} . This change is associated with a secondary bifurcation of the wall modes as discussed below. It is quite surprising that the radial profile of u_z maintains a close correspondence with the linear eigenfunction over such a large range of Ra. In Fig. 8(d), we compare the linear eigenfunctions (dashed lines) for u_z , u_{ϕ} , and u_r with data for Ra = 3×10^7 (solid lines) where one sees very close correspondence between the eigenfunctions and the data for u_z and u_{ϕ} with a slightly smaller first zero crossing for u_{ϕ} . u_r has a very different shape with a maximum value at a radius r_{max} (we define the characteristic length of u_r as $\delta_{u_r} = 1 - r_{\text{max}}$) near the first zero crossing of u_z , and there is a distinct difference between the eigenfunction and the data indicating that u_r is more affected by finite wall curvature than the other convective fields. Figure 8(e) shows the peak amplitudes of u_z , u_{ϕ} , and u_r plotted as a function of Ra – Ra_w. Close to onset $u \sim (\text{Ra} - \text{Ra}_w)^{1/2}$, as expected for a supercritical bifurcation. The monotonic increase in u_z and u_r appears to saturate near the onset of bulk convection at Ra $\approx 10^9$ whereas u_{rmax} drops rapidly after the onset of the subcritical instability described below. Finally, there is a small drop in $u_{\phi_{\text{max}}}$ at Ra $\approx 9 \times 10^7$ which signals the end of the weakly nonlinear wall mode regime. This signature shows up in measures of T.

The other dominant field contributing to the heat transport is the temperature field which, as noted above, has a much broader radial distribution and is strongly affected by small Γ as was suggested by considerations of the temperature eigenfunction [20] shown in Fig. 7(a). To get an overall flavor of the wall mode states for $Ra \leq 3 \times 10^8$, we show in Fig. 9 some representative azimuthal midplane contours and corresponding fields; in all of them T extends significantly over most of the radial domain. Figure 9(a) shows the nearly sinusoidal mean contour for Ra = 3×10^7 ($\epsilon = \text{Ra}/\text{Ra}_{w} - 1 =$ 0.07) with counterclockwise rotation and clockwise precession. For Ra = 2×10^8 ($\epsilon = 6.1$) in Fig. 9(b), the contour is more nonlinear but the precession remains steady whereas for Ra = 5×10^8 (ϵ = 17) there is increased distortion of the mean contour and the state is time dependent. By going to the precessing frame one can obtain well-converged statistical averages of the radial profiles of different fields. We show in Fig. 10, the radial distribution of the max value of T (averaged in precessing frame) in a midplane (z = 1/2) cross section. For Ra $\leq 10^8$, the profiles are remarkably linear, more so than the eigenfunction, indicating that the finite geometry is having a significant effect on T(r). For larger Ra > 10⁸, T(r) becomes increasingly nonlinear with radial undulations owing to the transition to time-dependent states and bulk-BZF convection described below. In Fig. 10(c), we show the Ra variation of the max value of T which indicates the weakly nonlinear growth of the wall mode amplitude but for higher Ra $\approx 9 \times 10^7$ this reverses and T decreases slowly owing to the decreasing vertical gradient of T in the cell center as discussed below. This feature is correlated with the small drop in $u_{\phi_{max}}$ in Fig. 8(e). Also shown is T_{rms} for Ra $\gtrsim 10^9$ owing to the unsteady and intermittent precession once the bulk mode is present which makes averaging in the precessing frame much slower to converge.



FIG. 10. Midplane temperature field maximum (averaged in precessing frame at maximum with respect to ϕ) T(r) versus r/R for (a) Ra/10⁷ = 3, 10, and 30. Dashed colored lines indicate the temperature eigenfunction for a planar wall [20] [as in Fig. 7(a)] adjusted for approximate sidewall mean temperature. Mean width δ_{u_z} and its variation (lateral error bars) in our computations are shown. (b) Ra/10⁷ = 40, 60, and 80. The corresponding temperature eigenfunction for Ra = 4×10^8 and the mean width δ_{u_z} are shown. (c) maximum radial value (at r/R = 1) T_{max} (black circles) and T_{rms} (red squares) versus Ra/10⁷ with different regions labeled by vertical dashed lines. The long-dashed lines denote transitions: At Ra $\approx 8.5 \times 10^7$ from weakly nonlinear growth of T, where solid black line is 6.0×10^{-5} (Ra – Ra_w)^{1/2} – 7.6×10^{-10} (Ra – Ra_w), to decreasing amplitude resulting from steepening of the vertical temperature profile; at Ra $\approx 6.5 \times 10^8$ from periodic to aperiodic (chaotic) time dependence of Nu.

With the insights provided by the structure of $u_z(r)$ and T(r), one can understand the radial structure of the heat transport Nu(r). In Fig. 11(a), the normalized (peak values set to one) fields $\tilde{u}_z(r)$, $\tilde{T}(r)$, $\tilde{u}_z(r)\tilde{T}(r)$, and $\tilde{N}u(r)$ are shown for an average over time-independent wall modes $3 \leq \text{Ra}/10^7 < 30$. T(r) varies slowly with r and the fields are time independent in the precessing frame. If the azimuthal phase difference between u_z and T is small, i.e., the place where u_z is maximally upward (downward) is the same as where T is maximally positive (negative) with respect to the mean, then one has Nu(r) ~ $u_z(r)$ to within about 10%. Taking Nu ~ $\langle T \rangle \langle u_z \rangle$ improves the agreement to about $\pm 5\%$. Averaging over the time-dependent wall mode states in the range $40 \leq \text{Ra}/10^7 \leq 90$ yields similar results shown in Fig. 11(b). More surprising is that averaging over



FIG. 11. $\widetilde{Nu}(r)$ (blue), $\widetilde{u_z}(r)$ (red), $\widetilde{T}(r)$ (black), and $\widetilde{u_z}(r)\widetilde{T}(r)$ for $Ek = 10^{-6}$, Pr = 0.8. Each profile for a given Ra is the radial profile evaluated at the max value in ϕ in the mean precessing frame and then normalized by the maximum value. Then different ranges of Ra are averaged. (a) Averaged over $3 \leq \text{Ra}/10^7 \leq 30$, (b) Averaged over $40 \leq \text{Ra}/10^7 \leq 90$. (c) Averaged over $100 \leq \text{Ra}/10^7 \leq 500$ where the bulk contribution is subtracted out.



FIG. 12. (a)–(d) Vertical temperature field $T(r = 0.98R, \phi, z)$ for Ek = 10^{-6} , Ra: (a) 3×10^{7} , (b) 5×10^{7} , (c) 5×10^8 , (d) 10^9 . Precession direction is right to left as indicated by the arrow. (e)–(h) Averaged angular profiles (temporally averaged in the precessing frame, i.e., $\phi_p = \phi_0 + \omega t$) for Ra = 5×10⁸ (arbitrary amplitude scaled approximately to $\pm 1/2$). The first Fourier mode corresponding to the linear eigenfunction in ϕ is black dashed line whereas different z = 0.5, 0.8, and 0.95 are indicated by black, red, and blue solid lines, respectively. The insets of panels (e)–(g) are the approximate amplitudes ΔT , Δu_z , Δu_{ϕ} of the shock-like feature as a function of z near the upper boundary (z = 1), respectively, with the horizontal arrows indicating the minimum and maximum extent of the shock. The dashed blue lines are (e) to guide the eye with local maximum at $1 - z_{\text{max}} = 0.03$, (f) an exponential saturation fit to $\sim 1 - e^{-(1-z)/z_0}$ with $z_0 = 0.04$, and (g) an exponential growth fit to $\sim e^{(1-z)/z_1}$ with $z_1 = 0.003$. (e) $\langle T(r = 0.98R, \phi, z) \rangle_{\phi_p}$, (f) $\langle u_z(r = 0.98R, \phi, z = H/2) \rangle_{\phi_p}$, (g) $\langle u_{\phi}(r=0.98R, \phi, z=H/2) \rangle_{\phi_p}$, and (h) $\langle u_r(r=0.98R, \phi, z=H/2) \rangle_{\phi_p}$. For T(z=1/2) and $u_z(z=1/2)$, the profiles are close to the linear eigenfunction as is $u_{\phi}(z=0.8)$ ($u_{\phi} \approx 0$ at z=1/2) and display very sharp nonlinear fronts at their advancing edge for z = 0.95, i.e., close to the upper boundary. On the other hand, u_r has a pulse-like shape with very little of the profile being described by the first Fourier mode. The inset shows a expanded view at the advancing edge (region indicated by black arrow) of the narrow peak at $\phi \approx 0.3\pi$ with a spatial oscillation with length about $2\lambda_c$.

the bulk-BZF states $100 \le \text{Ra}/10^7 \le 500$, Fig. 11(c), also produces a fairly good correspondence where here we subtract out the bulk contribution in the cell interior. There is a widening zone of boundary zone influence in the BZF but the form of Nu near the wall is driven predominately by the vertical velocity u_z of the wall mode which remains highly localized to within a radial width of order Ek^{1/3} characteristic of the linear wall-mode eigenfunctions. A more quantitative test of this comparison requires detailed amplitude averaging of the $u_z(\phi)$ and $T(\phi)$ for which we would need better converged statistical quantities in the bulk phase with Ra > Ra_c; we consider this in future work.

In addition to the radial structure of the fields, we consider their azimuthal structure. In Figs. 12(a)–12(d), we show instantaneous vertical temperature fields $T(r = 0.98R, \phi, z)$ for values

of Ra in the interval $3 \times 10^7 \le \text{Ra} \le 5 \times 10^9$. A very nonlinear profile develops with increasing Ra with only the smallest $Ra = 3 \times 10^7$ being well approximated by a single-mode sinusoidal function. To characterize this nonlinearity for $Ra = 5 \times 10^8$, we show in Figs. 12(e)–12(h) azimuthal profiles of T, u_z , u_{ϕ} , and u_r , respectively, averaged in the precessing frame ($\phi_p = \phi_0 + \omega t$) at r = 0.98Rand at three different z = 0.5, 0.8, and 0.95 [horizontal dashed lines in Fig. 12(c)]. For T and u_z , the midplane profiles are rather close to the linear eigenfunction, i.e., to the lowest Fourier mode. The eigenfunction solutions for the ϕ dependence of T and u are discrete Fourier series $\sum_{n=1}^{N} a_n \sin(nm\phi) e^{i\omega t}$ with the first term being $\sin(m\phi + \phi_p)$. For u_{ϕ} and u_r , the midplane is problematic because it corresponds to a zero crossing. Considering z = 0.8, one also sees reasonable correspondence with the lowest Fourier mode for u_{ϕ} but u_r is very different with almost no weight in the lowest Fourier mode. Close to the top (or bottom) boundary with z = 0.95, the nonlinearity of the wall mode is revealed where there is a sharp front-like feature appearing in T, u_z , and u_{ϕ} and a pulse-like structure for u_r . The emerging contribution of bulk-like excitations are visible in the azimuthal oscillations of order $2\lambda_c$ at the leading edge of u_r shown in the inset of Fig. 12(h). The shock-like features for T, u_z , and u_ϕ vanish at the upper boundary consistent with rigid and isothermal boundary conditions as shown in the insets of Figs. 12(f) and 12(g): T approaches zero consistent with global heat transport constraint, u_z because the zero normal velocity BC, and u_{ϕ} because of the zero tangential BC. The latter should be controlled by an Ekman boundary layer condition of order $\delta_E = \text{Ek}^{1/2} = 0.001$ but more quantitatively there is a coefficient relating δ_E to the viscous boundary layer of order 3 [56]. The blue dashed line in Fig. 12(g) is proportional to $e^{(1-z)/0.003}$, consistent with an Ekman boundary layer. More detail on the z dependence is provided in the Appendix (Fig. 28) where we compute for Ra = 5×10^8 : $\langle X \rangle_{\rm rms}$ and $\langle X - X^{(1)} \rangle_{\rm rms}$, where X is the convective field and $X^{(1)}$ is its lowest Fourier mode.

C. Evolution of wall modes: Nonlinearity and instability

Wall modes for $\text{Ek} > 10^{-4}$ [14,17,19] have a narrow Ra-window between their onset and the onset of bulk convection $\text{Ra}_c/\text{Ra}_w < 4$ whereas for $\text{Ek} = 10^{-6}$ one has $\text{Ra}_c/\text{Ra}_w \approx 30$. Thus, the wall mode state becomes increasingly nonlinear and eventually time-dependent over that interval, as noted previously [49,57]. In Figs. 12(a)–12(d), representative instantaneous temperature fields $T(r = 0.98R, \phi, z)$ are shown for increasing Ra at $\text{Ek} = 10^{-6}$. As Ra increases, the vertical profile is initially sinusoidal near onset [Fig. 12(a), Ra = 3×10^7] but becomes highly nonlinear—square-wave like and nonsingle-valued at the leading (left) and trailing (right) edge [Fig. 12(b), Ra = 5×10^7 ; Fig. 12(c), Ra = 5×10^8). Above Ra = 7×10^8 , there are strong lateral spatial striations [Fig. 12(d) Ra = 10^9 , near bulk onset] superimposed on the overall single-mode structure.

We characterize the transition to an oscillatory state through the time dependence of the heat transport which is time-independent for precessing wall modes. For $10^8 \leq \text{Ra} \leq 3 \times 10^8$, Nu(*t*) is a decaying oscillation of the form Nu \approx Nu_{∞} + Nu₀ cos $[\omega_2(t - t_0)]e^{-t/\tau_0}$ starting from some random initial condition, as shown in Figs. 13(a) and 13(b). For $4 \times 10^8 \leq \text{Ra} \leq 6 \times 10^8$, the waveform is approximately stationary (or slightly increasing saturation, see Appendix A 3) with amplitude Δ Nu (mean peak-to-peak) and an oscillation frequency ω_2 as in Fig. 13(c). The nature of the onset of time dependence is a subcritical bifurcation as shown in Fig. 14 where we plot the max-min envelope of Nu(*t*), the difference Nu_{max} - Nu_{min} = Δ Nu, the oscillation frequency ω_{d_2}/ω_d , and the decay frequency ω_{d_0}/ω_d versus Ra, where ω_d is the precession frequency in units of Ω^{-1} .

We show in Fig. 14(a) the Nu(t) envelope defined by its maximum and minimum values as a function of Ra. For Ra $\leq 3 \times 10^8$, Nu is time-independent and Nu_{max} = Nu_{min}. There is a rapid increase in Nu_{max} for Ra $\geq 4 \times 10^8$ whereas the minima stay roughly constant at Nu_{min} ≈ 6.3 , very close to its value at Ra = 3×10^8 . The solid (dashed) blue line is an three-parameter adjustable fit to the stable (unstable) portions of a subcritical bifurcation. This schematic illustration of the nonlinear time-dependent amplitude (solid line) and the unstable subcritical region (dashed line) is presented in a conventional bifurcation diagram in Fig. 14(b). In Fig. 14(c), we show ω_{d_2}/ω_d where the black circles and corresponding black dashed line indicate the oscillatory part of the decaying



FIG. 13. Time dependence of the heat transport Nu(t) for Ra: (a) 10^8 , (b) 2×10^8 , and (c) 5×10^8 . The dashed lines in panels (a) and (b) have the form Nu(∞) + Nu₀ cos $[\omega_2(t - t_0)]e^{-t/\tau_0}$, where $\omega_2 = 2\pi/\tau_2$ is the oscillatory frequency and τ_0 is the decay time. For panel (c) one has a more complicated relationship that accounts for large or small asymmetry $\sim \cos^4[\omega_2(t - t_0)]$ and a transient growth $\sim 1 - e^{-t/\tau_0}$. The dashed lines are fits to the data using these forms, fitted values are indicated, and horizontal (blue or black) bars indicate τ_2 and τ_0 , respectively.

solutions associated with the time-independent state whereas the blue circles and dashed line show $\omega_{d_2}/\omega_d \sim \text{Ra} \approx 1.7$. The decay frequency $\omega_{d_0}/\omega_d \sim (\text{Ra}_{w_2} - \text{Ra})$ (magenta diamonds and dashed line), where its zero intercept indicates the Rayleigh number Ra_{w_2} at which the wall mode state becomes unstable to infinitesimal perturbations. For $\text{Ra} \ge 7 \times 10^8$, the waveform becomes chaotic with increasing fluctuations and irregular oscillation frequency. The inference from Fig. 14(a) is that the minimum value Nu_{min} is when the system is close to the pure wall mode state whereas Nu_{max} occurs when the lateral jet makes the maximal contribution to the heat transport. We discuss this in more detail below when we consider the full range of Nu.



FIG. 14. (a) Extreme maximum and minimum values of Nu(t) (minima: black solid circles, maxima: blue solid circles and solid red square for chaotic state), where the error bars depict the range of less extreme maximum or minimum values, (b) Δ Nu (mean peak-to-peak), and (c) normalized oscillation frequency ω_{d_2}/ω_d (solid circles: black, stable state and blue, unstable state) and normalized decay frequency (scaled by 1/2 for comparison) $(1/2)(2\pi/\tau_{d_0})/\omega_d$ (magenta diamonds) vs Ra/10⁸, where ω_d is the wall mode precession frequency. Error bars denote variability of Δ Nu, ω_{d_2}/ω_d , and $(2\pi/\tau_{d_0})/\omega_d$ owing to shorter or longer lengths of time series and/or unsteady oscillations. The blue shaded region denotes an approximate zone of subcritical instability. The solid (dashed) lines are schematic suggestions for stable (unstable) behavior. The square (red) indicates a state where the oscillations have become chaotic with increased fluctuations and variable frequency. The dashed magenta line represents the stable state transient inverse time constant which is expected to vary as $\tau_0^{-1} \sim (\text{Ra}_{w_2} - \text{Ra})$.



FIG. 15. For Ra = 5×10^8 , Ek = 10^{-6} , and Pr = 0.8: horizontal cross sections of (a) $T(r, \phi)$ and (b)–(d) u_z (dashed lines are at radii r/R = 0.98 and 0.90, corresponding to the peak and the first zero crossing, respectively), u_{ϕ} , and u_r , respectively, for z = 0.8. Red (blue) corresponds to hot (cold) and up (down), AC—anticyclonic (C—cyclonic), and radially out (radially in), respectively. Vertical profiles (e)–(h) for corresponding fields (ϕ , z) evaluated at r/R = 0.98 [see panel (b), outer dashed line]. Arrows and \pm symbols indicate qualitative motion directions. Spatial oscillations of order $2\lambda_c$ are visible in horizontal cross sections of u_z and u_r and in vertical azimuthal profiles of u_z , u_{ϕ} , and u_r . Note that the major ejections of jets (and return flow) represented in u_r happen near the top and bottom boundaries.

The time dependence of the wall mode was noted earlier [49] for Pr = 1, $\Gamma = 3/2$, and Ek = 10^{-6} where the emission of horizontally propagating thermal plumes originating within the wallmode region and moving into the interior were observed for Ra $\gtrsim 5 \times 10^8$, consistent with our results (see also Refs. [57,58]). The mechanism for these fluctuations was hypothesized to result from a shear instability of a Stewartson layer of width $\sim Ek^{1/4}$ [49,59] that develops a net mean flow for $Ra \gg Ra_w$. The criterion for instability in a differentially sheared rotating layer (not a wall-bounded flow) is Re ≈ 10 to 20 [60] which is qualitatively consistent with the computational results [49]. For comparison we get Re values of 6 and 12 for Ra = 2×10^8 and 5×10^8 , respectively, based on the wall-bounded layer and 15 and 30 based on the outer free shear layer. Although the qualitative Re are about right for shear instability of a zonal barotropic mode [61], a more detailed analysis was beyond their scope [49]. Our results show that small scale fluctuation structures appear in the wall-bounded zone of order $Ek^{1/3}$, see Figs. 15(d), 15(g), and 15(h), rather than in the outer shear layer of order $Ek^{1/4}$ (see also below for length scale of striations). Further analysis and characterization of this instability is also beyond the scope of our work. Here we elucidate the onset of this time-dependent degree of freedom for Pr = 0.8, $\Gamma = 1/2$, and $Ek = 10^{-6}$. In Fig. 15, we show instantaneous fields of horizontal and vertical profiles, respectively: T [Figs. 15(a) and 15(e)], u_z [Figs. 15(b) and 15(f)], u_{ϕ} [Figs. 15(c) and 15(g)], and u_r [Figs. 15(d) and 15(h)] for Ra = 5×10⁸, above the subcritical onset of time dependence. Figs. 15(a)-15(d) show horizontal cross sections at z = 1/2 [Fig. 15(a)] and z = 0.8 [Figs. 15(b)–15(d)] whereas Figs. 15(e) and 15(f) illustrate the wall mode region r/R =0.98 with vertical $0 \le z \le 1$ and horizontal $0 \le \phi \le 2\pi$. The horizontal fields show the lateral plume emission into the interior near the zero crossing between positive and negative u_{ϕ} and u_{z} whereas the vertical fields show a spatially oscillatory signature of order $2\lambda_c$ that generates in the build-up to the emission. The oscillations are strongest for u_{ϕ} and u_r compared with less obvious



FIG. 16. Horizontal cross sections (a), (c) *T* (red—warmer, blue—colder than the arithmetic mean of the top and bottom temperatures); (b), (d) ω_z (red—cyclonic, blue—anticyclonic); and (e)–(g) u_z (red—up, blue—down), u_{ϕ} (red—cyclonic, blue—anticyclonic), u_r (red—outwards, blue—inwards), respectively. Ra: (a), (b) 10⁹; (c)–(g) 4×10⁹ at the same time. (a), (c) Length scale depicted is bulk λ_c .

features in T and u_z . In the bulk region represented in Figs. 15(a)–15(d), u_z is quite small compared with its value near the wall whereas the other fields have significant contributions in the interior. Thus, although the contribution to Nu $\propto u_z$ seems small; the perturbations in the interior region probably act as the foundation for the nucleation of bulk convection. At the same time, the vertical striations of order $2\lambda_c$ seem to be related to bulk instability implying that the underlying wall mode acts as the foundation for the nucleation of bulk instability. Thus, the jet instability and the $2\lambda_c$ bulk structures seem inextricably linked. The striation length scale (proportional to Ek^{1/3}) suggests that the origins of the instability arise within the wall shear zone rather than in the outer free shear zone of order Ek^{1/4} [49]. Other characteristics of the subcritical state include unsteady precession and the qualitative change in the shape of the contours of T: compare Figs. 9(b) and 9(c).

D. Transition to bulk modes

The transition to finite amplitude bulk convection occurs for $\text{Ra} \approx 10^9$ as compared with the theoretical prediction of $\text{Ra}_c \approx 7.8 \times 10^8$ for an infinite layer without sidewall effects. For comparison with the theoretical predictions of the nonhydrostatic quasigeostrophic (NHQGS) equations (see also below), the range $8 \times 10^8 < \text{Ra} \leq 5 \times 10^9$ corresponds to $1 \leq \text{Ra}/\text{Ra}_c \leq 6$ $(8 \leq \text{RaEk}^{4/3} \leq 60)$ where the plume state is predicted to be prevalent for Pr = 1 [40,41]. In Fig. 16, representative horizontal cross sections for instantaneous fields of T, ω_z , u_z , u_{ϕ} , and u_r are shown (z = 1/2 for T and ω_z and z = 0.8 for u). At $\text{Ra} = 10^9$, there are strong lateral jets of contrasting thermal signature for T in Fig. 16(a) which are generated near the intersection of cold and hot zones, i.e., where $T \approx \langle T \rangle$ and $u_z \approx 0$ (see Fig. 12). There are corresponding regions of cyclonic vorticity generation as shown for ω_z in Fig. 16(b). Around the sidewall boundary, there are azimuthal variations of T and ω_z with some spatially intermittent structure. When bulk convection is well established at $\text{Ra} = 5 \times 10^9$ as in Figs. 16(c) and 16(d) for T and ω_z , respectively, the BZF remains visible in the T field but there are significant small-scale temperature and vorticity fluctuations with spatial scales of the order of the linear length scale $\lambda_c = 4.8 \text{Ek}^{1/3}$ [0.051 for Ek = 10⁻⁶, see Eq. (7)]. In Figs. 16(e)–16(g), u_z, u_{ϕ} , and u_r , evaluated for $\text{Ra} = 4 \times 10^9$ and at the same time, show that u_z displays little remnant of the horizontal jet flow but u_{ϕ} and u_r show that even at



FIG. 17. Vertical azimuthal slice at r/R = 0.98 with fields indicated and z = 0.5 (top) and z = 0.8 (bottom) as labeled on the far right. (a), (c) *T* (red—warmer, blue—colder); (b), (d) ω_z , and (e) u_z , (f) u_{ϕ} , (g) u_r . Red (blue): warmer (colder), cyclonic (anticyclonic), up (down), cyclonic (anticyclonic), out (in), respectively. Ra: (a), (b) 1×10^9 at the same time; (c)–(g) 4×10^9 at the same time; (a), (c) length scale depicted is bulk $2\lambda_c$.

 $Ra/Ra_c \approx 4$, there are large-scale motions suggesting residual emanations from the BZF region note in particular the in and out radial velocity structure in Fig. 16(g).

The vertical profiles of the same fields evaluated for r/R = 0.98, Figs. 17(a)-17(g) further characterize interactions of the boundary flow and the bulk interior flow. The vertically coherent structures of T and ω_z in Figs. 17(a) and 17(b), respectively, at Ra = 10⁹ have a characteristic length of order $2\lambda_c$ shown in both plots and the m = 1 BZF remains quite coherent. At Ra = 5×10^9 , the BZF is instantaneously less coherent in T [Fig. 17(c)] but with vertical structures with remaining vertical correlation. In the nonhydrostatic, quasigeostrophic description [40,41], the state at $Ra = 10^9$ corresponds to the cellular regime of bulk convection whereas $2 \times 10^9 \le Ra \le 5 \times 10^9$ corresponds to the plume regime (ending at about Ra/Ra_c \approx 6, i.e., for Ra = 6×10⁹ for our data at $Ek = 10^{-6}$). The vertical profiles of the velocity fields [Figs. 17(e)-17(g)] for $Ra = 4 \times 10^9$ show similar spatial structure with some vertical coherence and a range of horizontal spatial scales. The BZF structure is again most visible in u_z [Fig. 17(e)] compared with u_{ϕ} [Fig. 17(f)] and u_r [Fig. 17(g)]. The spatial structure of u_{ϕ} is similar to u_z whereas u_r has much finer spatial structures. The u_r field in particular shows the strong radially inward and outward exchange with the bulk interior flow. To make these statements involving spatial length scales more quantitative, we compute the different measures of spatial length scale covering the range from the wall modes up to full bulk convection.

E. Length scales of wall modes and boundary zonal flow

There are multiple different length scales to consider when spanning the regimes of steady wall modes, $3 \times 10^7 \le \text{Ra} \le 3 \times 10^8$, unsteady subcritical wall modes $4 \times 10^8 \le \text{Ra} \le 8 \times 10^8$, and bulk modes $9 \times 10^8 \le \text{Ra} \le 5 \times 10^9$. For our $\Gamma = 1/2$, the largest length scale is the m = 1 wall mode state with $\lambda_1 = \pi/2$ which persists over the full range considered here. The radial length scale of the wall-mode or BZF states can be found in different ways [53,54]. Here we take a radial length defined either by the first zero crossing r_0 of each field in the radial direction starting from the sidewall boundary (r/R = 1), for u_r , the radial position of the peak maximum [see Fig. 8(d)], and for ω_z we take the second zero crossing. Note that $\delta_u/H = (1 - r_0/R)\Gamma/2 = (1 - r_0/R)/4$. We compute this in the precessing frame at the maximum of the field in the azimuthal direction; the



FIG. 18. Radial widths $\delta \text{Ek}^{-1/3}$ versus Ra/10⁷ for Ek = 10⁻⁶: $\Gamma = 1/2$ (closed symbols) and $\Gamma = 2$ (open symbol); δ_{u_z} (blue circles), $\delta_{u_{\phi}}$ (red squares), δ_{u_r} (magenta diamonds), and δ_{ω_z} (orange triangles), and $\delta_{u_{zms}}$ (blue triangles). The three regions are steady wall modes (yellow), subcritical unsteady wall modes (blue), and BZF with bulk modes (pink). The lateral (blue) dashed line is the length scale of the eigenfunction of u_z of 2.5, the lateral red dashed line is the length scale of the eigenfunction of u_{ϕ} of 2.24, and the orange lateral line is for the radial width of ω_z of 3.3. The values for δ_{u_z} for Ra $\geq 9 \times 10^8$ are the rms values and the blue dashed curve is the function $\delta_{u_r} = 2.5 + 0.16(\text{Ra}/10^7 - 80)^{1/2}$.

approach we used above for comparing with linear eigenfunctions [20]. Previously, we measured the radial length scale of the wall-mode or BZF states by the first zero crossing of $\langle u_{\phi} \rangle_{\phi,t}$ at the midplane [53,54]. This does not work well for the wall mode state because $u_{\phi}(z)$ has an approximate zero at the midplane, a radial length scale $\delta_{u_{\phi}}$ is about 1/2 of its value at, for example, z = 0.8. This explains the apparent scaling $\sim Ek^{2/3}$ of δ_0 found earlier for a range of Ek [53,54] instead of the expected $Ek^{1/3}$ scaling. This approach works well even for unsteady precession in the subcritical regime but there are insufficient data to get statistical convergence for the highly intermittent BZF state in this study where strong bulk flow is present. In that case, we use the first minimum of the root-mean-square (rms) field $(\langle u_z^2 \rangle_{\phi,t})^{1/2}$ which should be at the approximate location of the first zero crossing used for the wall-mode regime. The results are shown in Fig. 18 for u_z , u_{ϕ} , u_r , and ω_z where the widths δ are normalized by Ek^{1/3} for comparison with the wall-mode linear eigenfunctions. In the weakly nonlinear regime with Ra $\leq 7 \times 10^7$, $\delta_{u_{\phi}} \approx 2.25$ Ek^{1/3} consistent with the linear eigenfunction whereas for Ra $\geq 10^8$, we have $\delta_{u_{\phi}} \approx \delta_{u_z} \approx 2.5 \text{Ek}^{1/3}$. In contrast, $\delta_{\omega_z} \approx$ $3.3 \text{Ek}^{1/3} \approx 1.3 \delta_{u_z}$ for Ra $\leq 10^9$. Both δ_{u_ϕ} and δ_{u_z} start to increase slightly in the region of subcritical instability. For $\Gamma = 2$ (blue, open circle), $\delta_{u_z} \approx \delta_{u_\phi}$ at $Ra = 5 \times 10^8$. For $Ra \ge 10^9$, $\delta_{u_{zms}}$ increases rapidly with Ra. δ_{u_r} shows more variability, perhaps owing to the smaller amplitude of u_r which might require better statistical averaging.

Other length scales are the horizontal length scales in the interior of the cell as measured in horizontal cross sections and the horizontal and vertical length scales in the azimuthal and vertical direction around the circumference in the sidewall boundary region. The first is characteristic of bulk convection whereas the second is reflective of BZF–bulk interactions. For horizontal cross sections, we compute the two-dimensional (2D) autocorrelation function of the field *F* as $C_{FF}(r) \equiv \langle F(\mathbf{x})F(\mathbf{x}+\mathbf{r})\rangle_{A,t}/\langle F^2(\mathbf{x})\rangle_{A,t}$ in a centered rectangular domain with dimensions $0.6D \times 0.6D$, where $\langle \cdot \rangle_{A,t}$ means averaging over a horizontal cross-section and in time. We define the length scale δ such that $C_{FF}(\delta) = 0.25$ with uncertainty defined by $d\delta = \pm [dC(r)/dr]^{-1}\Delta C = \pm 0.05[dC(r)/dr]^{-1}$ [as compared with Ref. [57] which uses $\delta = \int_0^\infty C(r)dr$ so our lengths are a bit larger]. In Fig. 19(a),



FIG. 19. (a) Bulk length scales $\delta_T \text{Ek}^{-1/3}$ versus Ra/10⁷ for horizontal cross sections of *T* at z = 1/2 in a square domain with 46% of the circular area [shown in images in panel (b)]. The labels (1), (2), (3) correspond to Ra values of images in panel (b). (b) False color images of *T*: (left) *T*(*x*, *y*, z = 1/2) and (right) *T*($r/R = 0.98, \phi, z$). (1) Ra = 2×10⁸, (2) Ra = 1×10⁹, and (3) Ra = 5×10⁹. (c) Wall-mode or BZF scales $\delta_T \text{Ek}^{-1/3}$ vs Ra/10⁷ for vertical sections *T*($r/R = 0.98, \phi, z$): *C*_{TT}(0.98*R* ϕ) (horizontal) and *C*_{TT}(*z*) (vertical). Unless labeled in the figure, dashed lines (with corresponding colors) are guides to the eye.

we show $\delta_T \text{Ek}^{-1/3}$ versus Ra with corresponding representative images (left) in Fig. 19(b) for (1) Ra = 2×10⁸, (2) Ra = 1×10⁹, and (3) Ra = 5×10⁹. The size of the interior domain of area 0.36 D^2 is shown. In the steady wall-mode region, $\delta_T \approx 13$ which corresponds roughly to the exponential decay shown in Fig. 7(a). In the unsteady subcritical regime, δ_T decreases owing to an increasing fraction of lateral spatial structures with smaller length scale. In the bulk regime for Ra $\geq 1 \times 10^9$, $\delta_T \text{Ek}^{-1/3} \approx \lambda_c \text{Ek}^{-1/3} = 5.1$, characteristic of bulk rotating convection.

We also compute the one-dimensional (1D) correlation functions for the anisotropic $T(r/R = 0.98, \phi, z)$: $C_{TT}(R\phi) = \langle \langle T(r/R) \rangle$ directions in a vertical azimuthal slice $(0.98, \theta, z)T(r/R = 0.98, (\theta + \phi, z))\rangle_z/\langle T^2 \rangle_z\rangle_t$ and $C_{TT}(z) = \langle \langle T(r/R = 0.98, \phi, y) T(r/R =$ $(0.98, \phi, y+z)\rangle_{\phi}/\langle T^2\rangle_{\phi}\rangle_t$. The results are shown in Fig. 20(c) with the corresponding representative images in Fig. 20(b) (right). The azimuthal length scale is dominated by the m = 1 sinusoidal character of the wall mode and is about 30 up to the onset of bulk convection—a m = 1 cosine function has a correlation length of about 20 in these units. In the bulk regime, it drops to about 17. The vertical length scale is roughly independent of Ra which is consistent with the main vertical variation of T being a linear gradient with a corresponding slope independent correlation length of about 25 (the correlation length of a linear gradient is independent of its slope). Instead, the value is somewhat lower, about 18–20.

To obtain a more complete description of the length scales defined above, we apply the autocorrelation analysis to the velocity and vertical vorticity fields for horizontal cross sections and for the horizontal and vertical directions in a azimuthal slice near the sidewall. In particular, we use for the horizontal sections, z = 0.8 for u_z , u_ϕ , u_r , and ω_z to avoid the approximate vertical zero crossing of u_ϕ at z = 1/2. We show in Figs. 20(a)–20(c) the correlation lengths $\delta Ek^{-1/3}$ versus Ra/10⁷ for horizontal cross sections [Fig. 20(a)], horizontal variations in a vertical azimuthal profile [Fig. 20(b)], and vertical variations in a vertical azimuthal profile [Fig. 20(c)]. The length scales for horizontal cross sections all show the same overall trend with a decrease from about ten starting at Ra $\approx 5 \times 10^8$ to between two and five corresponding to $\lambda_c/2$ and λ_c , respectively, for higher Ra. Whereas δ_T decreases gradually over the subcritical regime, the velocities and ω_z abruptly decrease and maintain roughly constant values through the subcritical and bulk regimes $5 \times 10^8 \leq \text{Ra} \leq 5 \times 10^9$. This suggests that the subcritical wall-mode instability is related to the bulk instability with structures of the order of the linear stability length scale. The slower decline of δ_T



FIG. 20. Autocorrelation lengths $\delta \text{Ek}^{-1/3}$ vs Ra/10⁷: δ_T (black triangles), δ_{u_z} (blue circles), $\delta_{u_{\phi}}$ (red squares), δ_{u_r} (magenta diamonds), δ_{ω_z} (orange inverted triangles). (a) Bulk length scales from horizontal cross sections of T and ω_z at z = 1/2 and u_z , u_{ϕ} , and u_r at z = 0.8 in a square domain with 46% of the circular area, (b) wall-mode or BZF horizontal autocorrelation lengths for vertical sections at r/R = 0.98 and (c) wall-mode or BZF vertical correlation lengths for vertical sections r/R = 0.98. Averaging is as in Fig. 19. Unless labeled in the figure, the dashed lines (with corresponding colors) are guides to the eye.

is related to the longer (exponential) width of the temperature field relative to the radially confined velocity and ω_z fields.

The horizontal and vertical lengths in the sidewall boundary region, Figs. 20(b) and 20(c), are more complicated to interpret. The length scales of T, u, and ω_z are consistently about 28 for Ra \leq Ra_c $\approx 1 \times 10^9$ with the exception of $\delta_{u_r} \approx 8$ in the subcritical regime. For $1 \times 10^9 \leq$ Ra $\leq 5 \times 10^9$, $\delta_T \approx \delta_{u_z} \approx \delta_{u_\varphi} \approx \delta_{\omega_z} \sim F(\text{Ra}/\text{Ra}_c - 1)$ (assuming a dependence on Ra/Ra_c - 1). An empirical approximation of the data is $F(x) \sim x^{1/3}$. The exception is $u_r \approx 1.5$ which reflects the small scale nature of the boundary-interior interactions, see Fig. 17(g). The vertical correlations are somewhat different although δ_{u_z} and δ_{u_φ} are reasonably described by $F(x) \sim x^{\gamma}$ with $\gamma = 1.2, 0.7$, respectively. Comparatively, δ_{u_r} starts to decrease in the subcritical regime, similar to the horizontal component. The quantitative trends shown here are consistent with the qualitative impressions provided by representative images in Fig. 17.

F. Heat transport Nu

There are heat transport contributions from the wall modes and their remnants in the BZF and from the bulk flow. As we have seen above, the two states are not independent. The wall-mode jet instability for $Ra \ge 4 \times 10^8$ injects temperature and velocity perturbations into the interior prior to and continuing into the region of bulk instability, see Fig. 16. In the other direction, the finite convective amplitude near the sidewall boundary appears to be the nucleation site for bulk-like perturbations with spatial separation of order λ_c and with strong vertical coherence, see Fig. 17. Thus, cleanly separating the individual contributions, especially for convection cells with $\Gamma < 1$, is a complicated proposition. In Fig. 14(a), we noted that the time series of Nu in the subcritical regime $4 \times 10^8 \leq \text{Ra} \lesssim 10^9$ had minima corresponding to a more quiescent state and maxima corresponding to the lateral jet ejection process. Extending the range of this analysis to include bulk convection up to $Ra = 5 \times 10^9$ yields Fig. 21(a) where both the maxima and minima of Nu vary linearly with Ra (the red and black dashed lines are of the form a + bRa) so the difference ΔNu also varies linearly with Ra. What then is the base state for the onset of bulk interior convection? In Fig. 21(b), T fields corresponding to minimum and maximum Nu for Ra = 5×10^8 and 1×10^9 are shown. For $Ra = 5 \times 10^8$, the minimum Nu state corresponds quite closely to a nonlinear time-independent state whereas the maximum Nu state involves the emission of lateral jet structures.



FIG. 21. (a) Nu_{max} (blue solid circles, red solid square and diamonds) and Nu_{min} (black, solid circles and triangles) vs Ra/10⁸. Regions are labeled with vertical dashed lines separating them. Red and black dashed lines are linear fits (of the form a + bRa) to Nu vs Ra for Ra $\ge 10^9$. (b) T fields for (top) Ra = 5×10^8 and (bottom) Ra = 10^9 showing example at the minimum (left image) and maximum (right image) of Nu. Color bar of T is shown.

The situation near the onset of bulk instability at $Ra = 1 \times 10^9$ is more complex with lateral jets visible in both images; the timing and strength of the jets is probably important and not characterized here.

It appears from the analysis presented in Figs. 21(a) and 21(b) that the lateral jets from the wall region contribute substantially to the heat transport at the onset of bulk convection at Ra $\approx 1 \times 10^9$. With that in mind, we look at the average Nu and consider how best to separate the wall and bulk contributions. Previously, we have done so in several ways. First, owing to the rather unique radial structure of the wall modes and their remnants, we defined the BZF contribution as $2\pi \int_{r_0}^{t} Nu(r) r dr$, where r_0 is the first zero crossing of Nu(r) [53,54]. But this misses the negative contribution in Nu(r) as discussed above, see Figs. 11(a) and 11(b). Recently, we generalized this approach by dividing the total Nu into inner and outer portions $(2/r_s^2) \int_0^{r_s} \text{Nu}(r) r dr$ and $2/(1-r_s^2) \int_{r_s}^1 \text{Nu}(r) r dr$, respectively, where r_s was varied [50]. Here we choose $r_s = 0.8$ (in units of R) which is a good approximation to where the radial signature of the wall-mode or BZF seems to have become small, see Fig. 11(b). So from this perspective there are three contributions to the heat transport: pure wall mode (with jet instability), BZF, and bulk. In Fig. 22(a), these different pieces are plotted versus Ra/Ra_c. For $r_s = 0.8$, Nu contributions of BZF and bulk are about the same and both vary quite linearly with Ra/Ra_c. For comparison, Nu from the nonhydrostatic quasigeostrophic model (NHQGS model) [40,41] for Ek \rightarrow 0 and from Nu from a laterally periodic DNS for Pr = 1 and Ek = 10^{-7} [42] which includes Ekman pumping contributions (see also Refs. [62,63]) are plotted. The NHQGS is smaller owing to ignoring the contributions of Ekman pumping whereas the DNS is a bit higher, perhaps owing to larger Ek, i.e., Ek > 0.

Another possible separation strategy is to define $r_s(\text{Ra})$ based on some feature of, for example, $u_z(r)$ or Nu(r) rather than having it be a constant value. Varying r_s slightly for different Ra introduces a net upward (downward) curvature to Nu versus Ra for BZF (bulk) pieces. We choose instead to evaluate the separation for several constant r_s , recognizing that there is no unambiguous, testable separation strategy using the data we have here. In Fig. 11(c), the features of Nu, \tilde{u}_z , and \tilde{T} suggest $0.7 < r_s/R < 0.8$. In Fig. 22(b), we show results using $r_s/R = 0.7$ to contrast with $r_s/R = 0.8$ in Fig. 22(a). The trends are both consistent with a linear increase of Nu with Ra/Ra_c but with a greater (lesser) slope for BZF (bulk) contributions. Here the bulk contribution is quite comparable to the NHQGS data with little Ekman pumping correction whereas the seeming correspondence of the BZF contribution with the DNS results is purely coincidental.



FIG. 22. Nu versus Ra/Ra_c: (a) $r_s/R = 0.8$, Nu_{>rs} – Nu_{wm} BZF (red circles), Nu_{<rs} Bulk (blue squares), Nu_{NHQGS} Pr = 1 [41] (magenta diamonds), and Nu_{DNS} Pr = 1, Ek = 10⁻⁷ (black triangles) [42]. Horizontal dashed line for Ra/Ra_c \leq 1 is the wall-mode contribution Nu_{wm}. The blue dashed line is a linear approximation to Nu_{>rs} – Nu_{wm} \approx Nu_{<rs}. (b) Same data with $r_s/R = 0.7$. The blue and red dashed lines are linear fits to the bulk (blue squares) and BZF (red circles), respectively.

Although the procedure above is a tempting one that offers a more or less clean separation, there are concerns about this approach that can be elucidated by considering the vertical temperature profile $\langle T(r = r_{\text{max}}, \phi, z) \rangle_{\phi,t}$, where r_{max} is the value of, and its variation with, Ra. In Fig. 23(a), profiles of $\langle T \rangle$ are shown as functions of z for $3 \times 10^7 \leq \text{Ra} \leq 7 \times 10^8$ and in the inset for $10^9 \leq \text{Ra} \leq 5 \times 10^9$. There is a continual increasing slope at the top and bottom boundaries but the interior gradient is not monotonic with a smaller slope at $\text{Ra} = 5 \times 10^8$ compared with that at $\text{Ra} = 7 \times 10^8$. This reversal is also observed in the inset where higher Ra have steeper interior slope. These variations are features associated with the different states of wall modes for $\text{Ra} \leq \text{Ra}_c \approx 9 \times 10^8$ and the development of the bulk instability. An important point here is that the slope is continuously varying (see Fig. 23 in the Appendix) so that the definition of a thermal BL thickness is quite problematic. Perhaps there is localized Ekman pumping associated with the BZF—a topic for further investigation.



FIG. 23. (a) Temperature profiles $\langle T(r = 0.98R) \rangle_{\phi,t}$ for $3 \times 10^7 \leq \text{Ra} \leq 7 \times 10^8$ as indicated. The diagonal black dashed line is the linear conductivity profile of the thermally conducting state. Inset shows $\langle T(r = r_{\text{max}}) \rangle_{\phi,t}$ for $10^9 \leq \text{Ra} \leq 5 \times 10^9$, where r_{max} is the radial position of the maxima of $\langle u_z(r) \rangle_{\phi,t}$. (b) $-dT/dz|_{z=0} - 1$ (blue, solid circles) and $1 + dT/dz|_{z=1/2}$ (red, solid squares) at the same *r* as in panel (a). (c) Comparison of $-dT/dz|_{z=0}$ at the same *r* as in panel (a) (blue, solid circles) and Nu (red, solid squares), showing close correspondence.

In Fig. 23(b), we present the slope at the lower boundary -dT/dz as $-dT/dz(z = 0, r = r_{max}) - 1$ for which there are three distinct regions: weakly nonlinear growth \sim (Ra – Ra_w), the nonlinear state including the periodic time-dependent state up to Ra = 6×10^8 , and the combined BZF and bulk state with increasing slope \sim Ra. If one averaged over r as well as over ϕ and t, one would expect an exact correspondence between global heat transport and local temperature slope Nu – $1 = -dT/dz(z = 0, r = r_{max}) - 1$. In Fig. 23(c), there is very close agreement between these two quantities except in the weakly nonlinear regime. Returning to the interior slope, we plot in Fig. 23(b) dT/dz(z = 1/2) + 1 which is zero at Ra_w and would approach one with increasing Ra for nonrotating convection but has a finite slope for bulk rotating convection [50]. Here we have between 0.8 and 0.9 compared with a value of 0.6 for pure bulk geostrophic convection [40]. The difference is not surprising given the coexistence of bulk fluctuations on top of the wall-localized state as previously documented above.

Perhaps the most interesting feature of these profiles of $\langle T \rangle$ is the close agreement of the localized (in *r*) measure of Nu_{local} = $-dT/dz(z = 0, r_{max})$ and the global heat transport Nu that averages over the whole domain. If there was a clean separation between wall and bulk modes, Nu_{local} would not increase rapidly and in close correspondence to Nu. One can then infer that the bulk-like fluctuations in the wall zone continue to increase in strength and dominate the heat transport with the average wall-mode foundation playing a minor role. One cannot escape the realization that there are strong interactions among the bulk and coexisting BZF that plays a major role in the heat transport in some transition region from wall modes to bulk rotating convection.

III. CONCLUSION

We employed DNS using the Goldfish code of RRBC in a cylindrical geometry with aspect ratio $\Gamma = 1/2$, insulating sidewall boundary conditions, $Ek = 10^{-6}$, and Pr = 0.8 to characterize the progression of states over a range $3 \times 10^7 \le \text{Ra} \le 5 \times 10^9$. We also computed a smaller set of properties for the same parameters but with $\Gamma = 1, 2$ and for $Ek = 10^{-4}$ with $\Gamma = 0.3, 1/2, 1,$ 2, 5, primarily to determine the mode number m. The set of values for $Ek = 10^{-6}$ are tabulated in the Appendix, Table II. We elucidated how the mode number m observed for different Γ was consistent between wall modes and the BZF state, how the observed wall mode critical Ra_w varies with Ek, Γ , and *m*, and how small $\Gamma \leq 0.7$ always has m = 1 owing to the azimuthal periodicity. We further demonstrate the decoupling of radial length scales for $T \sim (k_w)^{-1}$ and $u \sim \text{Ek}^{-1/3}$ in the steady wall-mode regime $3 \times 10^7 \le \text{Ra} \le 3 \times 10^8$ and show that the eigenfunctions of the planar wall linear solutions [20] provide an excellent representation of the data, particularly for the radial dependence of u(r) where, for Ek = 10^{-6} , the radial localization near the sidewall of about 0.1 makes the sidewall curvature effects very small, i.e., $0.1/(2\pi) \approx 0.02$. On the other hand, for T the linear eigenfunction extends over a significant fraction of R and deviations owing to finite curvature and growing nonlinearity are more keenly felt. Similarly, the azimuthal and vertical eigenfunctions are also more effected by nonlinearity and curvature but give insight into the development of the steady wall-mode state.

In the steady wall-mode regime, Nu is constant despite the traveling-wave nature of the wall mode. For $4 \times 10^8 \leq \text{Ra} \leq 8 \times 10^8$, the steady wall mode undergoes a subcritical bifurcation to a state of time-dependent Nu through a nonlinear mechanism of lateral jet ejection from the wall mode into the bulk interior (see Refs. [49,57]). Throughout the linear (steady) and nonlinear (time-dependent) regimes for Ek = 10^{-6} , Nu(r) ~ $\langle u_z(r) \rangle \langle T(r) \rangle$ owing to the rapid radial variation of u_z compared with the very slow spatial variation of T and the fixed small phase difference between them. The jet instability strengthens as Ra \rightarrow Ra_c $\approx 1 \times 10^9$ and fine scale spatial (and temporal oscillations, see Appendix) structures form around the azimuth in the wall localized region. We show that the radial length scale δ_{u_z} and $\delta_{u_{\phi}}$ defined by first zero crossing is about 2.5Ek^{1/3} whereas $\delta_{\omega_z} \approx 3.3$ Ek^{1/3} using its second zero crossing. We also use autocorrelation analysis to extract correlation lengths for different cross sections of fields, namely, horizontal cross sections and vertical surfaces of { ϕ , z} at constant r/R = 0.98 [the maximum value of $u_z(r)$]. The correlation lengths are fairly

constant in the wall-mode regime and decrease rapidly once the bulk mode sets in for Ra $\gtrsim 10^9$. In horizontal regions in the cell center and Ra $\gtrsim 10^9$, δ_T , $\delta_{u_{\phi}} \approx \lambda_c$ whereas δ_{u_c} , δ_{u_r} , $\delta_{\omega_z} \approx \lambda_c/2$. In the wall localized regime, the azimuthal correlation lengths are roughly constant at about a half wavelength of the m = 1 wall mode but decrease to $2\lambda_c - 4\lambda_c$ in the bulk mode region. The vertical correlation lengths are more varied among the different fields but systematically decrease in the bulk mode region.

When considered in the context of heat transport scaling, the analysis of the structure of wall modes and their remnant robust features in the presence of bulk convection—the BZF—play a key role in attempting to separate bulk influence from that of the wall-localized portion. The negative feature in Nu is understood as arising from the oscillatory radial structure of $u_z(r)$ and the much slower radial dependence of T. Once bulk convection begins in the interior, the BZF spreads out radially and different features of Nu, u_z , and T including the lateral jet instability makes separation of the two components complicated. The resulting separation using $r_s/R = 0.8$ and 0.7 and subtracting the wall-mode contribution [50,53], yields Nu ~ Ra/Ra_c but with different slopes: one that gives agreement with the NHQGS theory and the other with Ekman pumping corrections. There are also subtle changes in the Ra dependence for the two r_s values. The average vertical temperature profiles reinforce the understanding that the BZF carries strong bulk fluctuations that contribute heavily to the total heat transport. There seems to be no unambiguous way to disentangle the two components of Nu in small aspect ratios such as studied here. A closer inspection of the interplay of thermal boundary layers in the transitioning wall mode and for bulk convection with coexisting BZF states is certainly warranted.

ACKNOWLEDGMENTS

The authors acknowledge the support from the German Research Foundation (DFG), Grants No. Sh405/20 and No. Sh405/22, and the Leibniz Supercomputing Centre (LRZ) for providing computing time. R.E.E. acknowledges support from the Los Alamos National Laboratory LDRD Program. We thank R. Kunnen for helpful comments and suggestions.

APPENDIX

In this section, we provide tabulated data from our DNS, empirically expand the set of critical parameters to higher order in Ek, and show some additional features of wall modes in the BZF regime. In addition, we tabulate movies that we present in Supplemental Material [64].

1. Data

In Table II, the computed values of control parameters $\text{Ek} = 10^{-6}$, Pr = 0.8, Γ , Ra, and $\epsilon = \text{Ra}/\text{Ra}_{\text{w}} - 1$ are listed with the resultant values of Nu, ω_d , ω_{d_2} , and ω_{d_0}/ω_d .

2. Wall-mode critical parameters and aspect ratio effects

The asymptotic approximations for the critical values of Ra_w , ω_w , and k_w are valid over varying ranges of Ek but the approximations are not adequate to accurately estimate these critical values for larger Ek. Using the numerical data presented in Ref. [20], one can empirically extend the asymptotic results to larger Ek using fits to the data and expansions in higher powers of Ek^{1/3}. In Fig. 24, we fit the data for the critical values assuming insulating sidewall boundary conditions at Pr = 0.7, 7 (see Fig. 3 [20]). Only Ra_w has significant variation with Pr. Finite wall curvature ($\Gamma = 1$ [16]) is shown to slightly decrease Ra_w whereas finite wall conductivity slight increases Ra_w [19] as shown in Fig. 24(a). We also fit the curvature parameter ξ_0 of the marginal stability boundary computed for Pr = 7 and for partially insulating sidewalls [21]. The results are shown in Fig. 24.

Ek	Г	Ra	ϵ	Ro	Nu	ω_d	ω_{d_2}	$ au_0 \omega_d / (2\pi)$
10 ⁻⁶	1/2	3.0×10^{7}	0.07	0.0061	1.1	1.9×10^{-4}		
	,	4.0×10^{7}	0.43	0.0071	1.6	$2.2 imes 10^{-4}$		
		5.0×10^{7}	0.79	0.0079	2.0	$2.4 imes 10^{-4}$		
		7.0×10^{7}	1.5	0.0094	2.5	$2.8 imes 10^{-4}$	$1.7 imes 10^{-4}$	-0.25
		1.0×10^{8}	2.6	0.011	4.2	$3.7 imes 10^{-4}$	3.3×10^{-4}	-0.25
		2.0×10^{8}	6.1	0.016	5.4	$6.0 imes 10^{-4}$	$8.9 imes 10^{-4}$	-0.42
		3.0×10^{8}	9.7	0.019	6.5	$8.1 imes 10^{-4}$	1.4×10^{-3}	-0.56
		4.0×10^{8}	13.3	0.022	7.9	$9.7 imes 10^{-4}$	1.5×10^{-3}	
		5.0×10^{8}	16.9	0.025	8.4	1.2×10^{-3}	1.8×10^{-3}	0.21
		6.0×10^{8}	20.4	0.027	8.7	1.4×10^{-3}	2.4×10^{-3}	0.34
		7.0×10^{8}	24.	0.030	9.1	2.9×10^{-3}	$1.7(5) \times 10^{-3}$	
		8.0×10^{8}	27.6	0.032	10.3	2.7×10^{-3}		
		9.0×10^{8}	27.6	0.034	10.8	3.4×10^{-3}		
		1.0×10^{9}	34.7	0.035	11.2	4.4×10^{-3}		
		2.0×10^{9}	70	0.05	24.1	7.4×10^{-3}		
		3.0×10^{9}	106	0.061	36.4	9.6×10^{-3}		
		4.0×10^{9}	142	0.071	47.1	1.2×10^{-3}		
		5.0×10^9	178	0.079	55.8	$1.5 imes 10^{-2}$		

TABLE II. Data from DNS with Pr = 0.8 indicating Ek, Γ , Ra, ϵ , Ro, Nu, ω_d , ω_d , and $\tau_0 \omega_d/(2\pi)$.

With the resulting fit parameters:

$$\begin{aligned} & \operatorname{Ra_{w}} = 31.82 \mathrm{Ek^{-1}} + 46.5 \mathrm{Ek^{-2/3}} - 465 \mathrm{Ek^{-1/3}} + 1564 \quad (\mathrm{Pr} = 0.7) \\ & = 31.82 \mathrm{Ek^{-1}} + 46.5 \mathrm{Ek^{-2/3}} - 194 \mathrm{Ek^{-1/3}} + 845 \quad (\mathrm{Pr} = 7), \\ & k_{w} = 6.07 - 35 \mathrm{Ek^{1/3}} + 250 \mathrm{Ek^{2/3}} - 990 \mathrm{Ek} \quad (\mathrm{Pr} = 0.7) \\ & = 6.07 - 35 \mathrm{Ek^{1/3}} + 250 \mathrm{Ek^{2/3}} - 870 \mathrm{Ek} \quad (\mathrm{Pr} = 7), \\ & \omega_{\kappa_{w}} = 66.05 - 732 \mathrm{Ek^{1/3}} + 5700 \mathrm{Ek^{2/3}} - 2300 \mathrm{Ek} \\ & \xi_{0} = 0.142 + 0.447 \mathrm{Ek^{1/6}} - 1.08 \mathrm{Ek^{1/3}} + 1.34 \mathrm{Ek^{1/2}} \quad (\mathrm{Pr} = 7). \end{aligned}$$

3. Transients

Whereas the form of the time dependence of Nu is very precisely sinusoidal [e.g., Fig. 25(b)], the subcritical nonlinear state has some interesting additional features. At Ra = 4×10^8 , there is a very clear large versus small Nu asymmetry in Nu(*t*) as shown in Fig. 25(a), i.e., the system prefers to be in a larger Nu state relative to a smaller one. The solid line is a fit with $a - b\{\cos^4 [(\omega_2/2)t - \phi_0]\}$ to the normalized Nu(*t*), where Nu_n is its approximate mean value. In Fig. 13(c) (Ra = 5×10^8), there is some remaining asymmetry with both $\cos (\omega t)$ and $\cos [(\omega_2/2)t]^4$ contributions. Whereas the amplitude Δ Nu at Ra = 4×10^8 is very constant over several oscillation periods, the amplitude increases slowly for Ra = 6×10^8 as a saturation of the form $1 - e^{-t/\tau_0}$; this is also observed for Ra = 5×10^8 , see Fig. 13(c), but the single oscillation period makes the saturation difficult to measure accurately. The time series for Ra = 7×10^8 in Fig. 25(c) is quite chaotic with an aperiodic oscillation and an intermittent amplitude.

4. $\Gamma = 1/3$ prograde dynamics

Most of the BZF states that we explored had retrograde (anticyclonic) precession. The exception is for $\Gamma = 1/3$ with Ra = 10⁸, and Ek = 1.1×10^{-5} , where we find prograde (cyclonic) precession and alternations between prograde and retrograde directions, see Figs. 26(a) and 26(b). At early



FIG. 24. Critical parameters of a planar wall mode with perfectly insulating sidewall boundary conditions versus Ek using digitized data from Fig. 3 of Ref. [20]. Blue and red data points and corresponding fits in powers of Ek^{1/3} are for Pr = 7 and Pr = 0.7, respectively. (a) Ra_w where black [16] and magenta [19] data points for Pr \approx 7 [below and above the Pr = 7 (blue) curve] show the effects, respectively, of finite curvature ($\Gamma = 1$) and finite wall conductivity (plexiglass walls). (b) ω_{k_w} (recall that $\omega_d = 2\text{Ek}/\text{Pr}\,\omega_k$), and (c) k_w . (d) The curvature parameter ξ_0 of the marginal stability boundary $\epsilon_M = \xi_0^2 (k - k_w)^2$ (schematically shown in the inset) from data at Pr = 7 and weakly conducting sidewall boundary conditions [21].



FIG. 25. Time dependence of the heat transport Nu(*t*) for (a) Ra = 4×10^8 showing large or small Nu asymmetry $\sim \cos^4(\omega_2/2t)$ and (b) Ra = 6×10^8 where there is a long transient relaxation of the form $(1 - e^{-t/\tau_0})$. Dashed lines in panels (a) and (b) are fits with these functional forms with fitted values for oscillation period τ_2 and τ_0 indicated as horizontal (blue or black) bars, respectively. (c) Ra = 7×10^8 chaotic time series with some time intervals from peak (blue) and zero crossing (red): mean is 110 ± 30 .



FIG. 26. Angle-time plot for Ra = 10^8 , $\Gamma = 1/3$, and Ek = 1.1×10^{-5} of $T(r = 0.98R, \phi(t), z = 1/2)$ and (a) the whole time series of 1000 time units, (b) the last 200 time units. Hotter (red), cooler (blue).

 $t \leq 400$ [Fig. 26(b) shows an expanded interval] and late times $t \gtrsim 700$, the precession is prograde but for intermediate times, an interval of retrograde precession is observed. It is not determined from our data whether the prograde state is stable at long times. Uniform cyclonic precession dynamics for wall-mode states was predicted theoretically for small Γ and $\Pr < 1$ [16] and observed in previous simulations [51] for $\Gamma = 1/2$, $\Pr = 0.8$, $\operatorname{Ra} = 10^5$, and $\operatorname{Ek} = 1.4 \times 10^{-3}$. The prograde direction of precession depended sensitively on a combination of Γ and Ek as we also find here whereas retrograde precession is observed in almost all cases; we see prograde precession for $\Gamma = 1/3$ but not for $\Gamma = 1/5$, 1/2, 3/4, 1 over similar ranges of Ra and Ek.

5. Eigenfunction root mean square

The qualitative features for the field shapes at $Ra = 5 \times 10^8$ seen in Figs. 12(e)-12(h) show that the development of the nonlinearity of the wall mode is near regions close to the horizontal isothermal top or bottom boundaries. To make this more quantitative, we compute $X(\phi, z, r = 0.98)$ and its first Fourier mode $X^{(1)}(\phi, z, r = 0.98)$ at fixed z. We then find the rms value with respect to ϕ of X and of $X - X^{(1)}$ and average in time. Figures 27(a)-27(d) show the z variation of the rms values for T, u_z , u_{ϕ} , and u_r , respectively, for $Ra = 5 \times 10^8$ whereas [Figs. 27(f)-27(h)] show u fields for $Ra = 3 \times 10^7$; the differences with Ra are not large. The difference $\Delta_X(z)$ [as shown in Fig. 27(a) for T] indicates the degree to which the $X^{(1)}(z)$ captures the main features of its ϕ dependence. Figure 27(e) shows the normalized ratio $\Delta_X/X - 1$. Except for u_r , over much of the vertical extent almost 80% of the full rms value is captured by $X^{(1)}(z)$ with only a small enhancement for $Ra \approx Ra_w$. On the other hand, u_r has very little weight—about 25%—in the first Fourier mode.

For the vertical profiles T(z), we computed the rms analysis of the deviation of T(z) from a linear profile, i.e., $\theta(z) = T(z) - (1 - z)$, with respect to its first Fourier mode for different Ra as shown in Fig. 28(a) where one sees the weakly nonlinear growth $\sim (\text{Ra} - \text{Ra}_w)^{1/2}$ (blue dashed) for the θ_{rms} . Subtracting out the dominant linear eigenfunction one gets second-order scaling of $\sim \text{Ra} - \text{Ra}_w$ (red dashed) for $\text{Ra} \leq 1 \times 10^8$. The saturation at higher Ra results from increasing gradients near the top



FIG. 27. $\langle X(\phi, z) \rangle_{\phi_{rmsr}}$ (blue solid circles) and $\langle X(\phi, z) - X^{(1)}(z) \rangle_{\phi_{rmsr}}$ (red solid squares vs z, where $X^{(1)}$ is first Fourier mode of field X where solid and dashed lines are guides to the eye unless otherwise specified. (a) T, (b) u_z , (c) u_{ϕ} , and (d) u_r for Ra = 5×10^8 and (f) u_z , (g) u_{ϕ} , and (h) u_r for Ra = 5×10^7 . ($u_{\phi_{rms}}$ and u_{rms} go to zero at r/R = 1 for 0.99 $\lesssim z < 1.0$ consistent with zero tangential BC—not shown). (e) The normalized difference $\langle X - X^{(1)} \rangle_{rms} / \langle X \rangle_{rms} - 1$. For most of the height (excluding the center z = 1/2 for u_{ϕ} and u_r), $X^{(1)}$ captures about 80% of the ϕ variation of T (blue solid circles), u_z (red solid squares), and u_{ϕ} (magenta solid diamonds) with little variation for Ra \approx Ra_w (solid blue, red, and magenta lines). u_r (black solid triangles and black line) is quite different with only about 20% for the higher Ra and 50%–70% for lower Ra.



FIG. 28. (a) $\langle \langle T(z) \rangle_{\phi,t} \rangle_{\text{rms}}$ (blue, solid circles) and $\langle \langle T(z) \rangle_{\phi,t} - T^{(1)}(z) \rangle_{\text{rms}}$ (red, solid squares) vs Ra – Ra_w; $T^{(1)(z)}$ is the first Fourier component of $\langle T(z) \rangle_{\phi,t}$. Inset shows the ratio of $\langle T - T^{(1)(z)} \rangle_{\text{rms}} / \langle T \rangle_{\text{rms}}$. Vertical dashed lines show boundaries of different regions. (b) (dT/dz)/[dT/dz(z=0)] vs z for values of Ra/10⁷ in the range of bulk rotating convection: 100, 200, 300, 400, 500. There is no clear demarcation of a thermal boundary layer with an approximately constant slope only for z values within an Ekman boundary layer $\delta_{\text{Ek}} = \text{Ek}^{1/2} =$ 0.001 indicated by dashed black vertical line. The dashed orange and black curves are the corresponding profiles for nonrotating convection with the same Nu as for the corresponding solid orange and black curves.



FIG. 29. Representative images of movies presented in Supplemental Material [64]. Ek = 10^{-6} , Pr = 0.8, and $\Gamma = 1/2$. Ra = 5×10^8 : (a) $T(r, \phi, z = 1/2)$ (b) $T(r = 0.98R, \phi, z)$; Ra = 2×10^9 : (c) $T(r, \phi, z = 1/2)$, (d) $T(r = 0.98R, \phi, z)$; Ra = 10^9 : (e) $u_z(r, \phi, z = 0.8)$, (f) $u_z(r = 0.98R, \phi, z)$, (g) $u_\phi(r, \phi, z = 0.8)$, (h) $u_\phi(r = 0.98R, \phi, z)$, (i) $u_r(r, \phi, z = 0.8)$, (j) $u_r(r = 0.98R, \phi, z)$, (k) $\omega_z(r, \phi, z = 0.8)$, (l) $\omega_z(r = 0.98R, \phi, z)$.

or bottom plates but saturating amplitude of θ . Finally, we show in Fig. 26(b) the boundary layer profiles dT/dz(z) near the bottom boundary in the bulk or BZF region $1 \le \text{Ra}/10^9 \le 5$. The curve (blue) for Ra = 1×10^9 has a form still dominated by the wall-mode whereas the higher Ra curves fall off much faster. For comparison the boundary layer profiles of dT/dz for nonrotating convection for the same total Nu are shown for Ra/ $10^9 = 4$ and 5. The RRBC profiles are quite different than those of nonrotating RBC.

6. Movies

The dynamics and structure of the wall modes and the bulk/BZF states are hard to fully characterize through static single frame images. In the Supplemental Material [64], we provide representative movies that illustrate a variety of states. Figure 29 shows a representative frame for each movie with details for each in the caption.

^[1] R. P. J. Kunnen, The geostrophic regime of rapidly rotating turbulent convection, J. Turbul. 22, 267 (2021).

^[2] R. E. Ecke and O. Shishkina, Turbulent rotating Rayleigh–Bénard convection, Annu. Rev. Fluid Mech. 55, 603 (2023).

- [3] S. Chandrasekhar, The instability of a layer of fluid heat below and subject to Coriolis forces, Proc. Roy. Soc. London A 217, 306 (1953).
- [4] S. Chandrasekhar, Hydrodynamic and Hydromagnetic Stability (Clarendon Press, Oxford, 1961).
- [5] Y. Nakagawa and P. Frenzen, A theoretical and experimental study of cellular convection in rotating fluids, Tellus 7, 1 (1955).
- [6] H. P. Greenspan, The Theory of Rotating Fluids (Cambridge University Press, Cambridge, 1968).
- [7] G. Veronis, Cellular convection with finite amplitude in a rotating fluid, J. Fluid Mech. 5, 401 (1959).
- [8] T. H. Rossby, A study of Bénard convection with and without rotation, J. Fluid Mech. 36, 309 (1969).
- [9] P. G. J. Lucas, J. M. Pfotenhauer, and R. J. Donnelly, Stability and heat transfer of rotating cryogens. Part 1. Influence of rotation on the onset of convection in liquid ⁴He, J. Fluid Mech. 129, 251 (1983).
- [10] J. M. Pfotenhauer, P. G. J. Lucas, and R. J. Donnelly, Stability and heat transfer of rotating cryogens. Part 2. Effects of rotation on heat-transfer properties of convection in liquid ⁴He, J. Fluid Mech. 145, 239 (1984).
- [11] J. M. Pfotenhauer, J. J. Niemela, and R. J. Donnelly, Stability and heat transfer of rotating cryogens. part 3. effects of finite cylindrical geometry and rotation on the onset of convection, J. Fluid Mech. 175, 85 (1987).
- [12] J. C. Buell and I. Catton, The effect of wall conduction on the stability of a fluid in a right circular cylinder heated from below, J. Heat Transfer 105, 255 (1983).
- [13] F. Zhong, R. Ecke, and V. Steinberg, Asymmetric modes and the transition to vortex structures in rotating Rayleigh-Bénard convection, Phys. Rev. Lett. 67, 2473 (1991).
- [14] F. Zhong, R. Ecke, and V. Steinberg, Rotating Rayleigh–Bénard convection: Asymmetric modes and vortex states, J. Fluid Mech. 249, 135 (1993).
- [15] R. Ecke, F. Zhong, and E. Knobloch, Hopf bifurcation with broken reflection symmetry in rotating Rayleigh–Bénard convection, Europhys. Lett. 19, 177 (1992).
- [16] H. F. Goldstein, E. Knobloch, I. Mercader, and M. Net, Convection in a rotating cylinder. part 1 linear theory for moderate Prandtl numbers, J. Fluid Mech. 248, 583 (1993).
- [17] L. Ning and R. E. Ecke, Rotating Rayleigh–Bénard convection: Aspect-ratio dependence of the initial bifurcations, Phys. Rev. E 47, 3326 (1993).
- [18] Y. Liu and R. E. Ecke, Heat transport scaling in turbulent Rayleigh–Bénard convection: Effects of rotation and Prandtl number, Phys. Rev. Lett. 79, 2257 (1997).
- [19] Y. Liu and R. E. Ecke, Nonlinear traveling waves in rotating Rayleigh–Bénard convection: Stability boundaries and phase diffusion, Phys. Rev. E 59, 4091 (1999).
- [20] J. Herrmann and F. H. Busse, Asymptotic theory of wall-attached convection in a rotating fluid layer, J. Fluid Mech. 255, 183 (1993).
- [21] E. Y. Kuo and M. C. Cross, Traveling-wave wall states in rotating Rayleigh–Bénard convection, Phys. Rev. E 47, R2245 (1993).
- [22] P. P. Niiler and F. E. Bisshopp, On the influence of coriolis force on onset of thermal convection, J. Fluid Mech. 22, 753 (1965).
- [23] G. M. Homsy and J. L. Hudson, The asymptotic stability of a bounded rotating fluid heated from below: The conductive basic state, J. Fluid Mech. 45, 353 (1971).
- [24] M. Wedi, D. P. M. van Gils, S. Weiss, and E. Bodenschatz, Rotating turbulent thermal convection at very large Rayleigh numbers, J. Fluid Mech. 912, A30 (2021).
- [25] K. Zhang and X. Liao, The onset of convection in rotating circular cylinders with experimental boundary conditions, J. Fluid Mech. 622, 63 (2009).
- [26] E. M. King, S. Stellmach, J. Noir, U. Hansen, and J. M. Aurnou, Boundary layer control of rotating convection systems, Nature (London) 457, 301 (2009).
- [27] S. Weiss, P. Wei, and G. Ahlers, Heat-transport enhancement in rotating turbulent Rayleigh–Bénard convection, Phys. Rev. E 93, 043102 (2016).
- [28] J. S. Cheng, S. Stellmach, A. Ribeiro, A. Grannan, E. M. King, and J. M. Aurnou, Laboratory-numerical models of rapidly rotating convection in planetary cores, Geophys. J. Int. 201, 1 (2015).
- [29] J. S. Cheng, M. Madonia, A. J. Aguirre Guzmán, and R. P. J. Kunnen, Laboratory exploration of heat transfer regimes in rapidly rotating turbulent convection, Phys. Rev. Fluids 5, 113501 (2020).

- [30] K. Julien, S. Legg, J. McWilliams, and J. Werne, Rapidly rotating turbulent Rayleigh–Bénard convection, J. Fluid Mech. 322, 243 (1996).
- [31] R. P. J. Kunnen, H. J. H. Clercx, and B. J. Geurts, Heat flux intensification by vortical flow localization in rotating convection, Phys. Rev. E 74, 056306 (2006).
- [32] R. J. A. M. Stevens, H. J. H. Clercx, and D. Lohse, Optimal Prandtl number for heat transfer in rotating Rayleigh–Bénard convection, New J. Phys. 12, 075005 (2010).
- [33] S. Horn and O. Shishkina, Rotating non-Oberbeck–Boussinesq Rayleigh–Bénard convection in water, Phys. Fluids 26, 055111 (2014).
- [34] R. P. J. Kunnen, R. Ostilla-Mónico, E. Van der Poel, R. Verzicco, and D. Lohse, Transition to geostrophic convection: The role of the boundary conditions. J. Fluid Mech. 799, 413 (2016).
- [35] J. J. Niemela, S. Babuin, and K. R. Sreenivasan, Turbulent rotating convection at high Rayleigh and Taylor numbers, J. Fluid Mech. 649, 509 (2010).
- [36] R. E. Ecke and J. J. Niemela, Heat transport in the geostrophic regime of rotating Rayleigh–Bénard convection, Phys. Rev. Lett. 113, 114301 (2014).
- [37] S. Horn and O. Shishkina, Toroidal and poloidal energy in rotating Rayleigh–Bénard convection, J. Fluid Mech. 762, 232 (2015).
- [38] H.-Y. Lu, G.-Y. Ding, J.-Q. Shi, K.-Q. Xia, and J.-Q. Zhong, Heat-transport scaling and transition in geostrophic rotating convection with varying aspect ratio, Phys. Rev. Fluids 6, L071501 (2021).
- [39] M. Sprague, K. Julien, E. Knobloch, and J. Werne, Numerical simulation of an asymptotically reduced system for rotationally constrained convection, J. Fluid Mech. 551, 141 (2006).
- [40] K. Julien, E. Knobloch, A. M. Rubio, and G. M. Vasil, Heat transport in low-Rossby-number Rayleigh-Bénard convection, Phys. Rev. Lett. 109, 254503 (2012).
- [41] K. Julien, A. M. Rubio, I. Grooms, and E. Knobloch, Statistical and physical balances in low Rossby number Rayleigh–Bénard convection, Geophys. Astrophys. Fluid Dyn. 106, 392 (2012).
- [42] S. Stellmach, M. Lischper, K. Julien, G. Vasil, J. S. Cheng, A. Ribeiro, E. M. King, and J. M. Aurnou, Approaching the asymptotic regime of rapidly rotating convection: boundary layers versus interior dynamics, Phys. Rev. Lett. 113, 254501 (2014).
- [43] J. S. Cheng, J. M. Aurnou, K. Julien, and R. P. J. Kunnen, A heuristic framework for next-generation models of geostrophic convective turbulence, Geophys. Astrophys. Fluid Dyn. 112, 277 (2018).
- [44] O. Shishkina, S. Horn, S. Wagner, and E. S. C. Ching, Thermal boundary layer equation for turbulent Rayleigh–Bénard convection, Phys. Rev. Lett. 114, 114302 (2015).
- [45] G. L. Kooij, M. A. Botchev, E. M. A. Frederix, B. J. Geurts, S. Horn, D. Lohse, E. P. van der Poel, O. Shishkina, R. J. A. M. Stevens, and R. Verzicco, Comparison of computational codes for direct numerical simulations of turbulent Rayleigh–Bénard convection, Comput. Fluids 166, 1 (2018).
- [46] P. Reiter, O. Shishkina, D. Lohse, and D. Krug, Crossover of the relative heat transport contributions of plume ejecting and impacting zones in turbulent Rayleigh–Bénard convection, Europhys. Lett. 134, 34002 (2021).
- [47] P. Reiter, X. Zhang, and O. Shishkina, Flow states and heat transport in Rayleigh–Bénard convection with different sidewall boundary conditions, J. Fluid Mech. 936, A32 (2022).
- [48] N. Li and S. Laizet, 2DECOMP&FFT A Highly Scalable 2D Decomposition Library and FFT Interface (2010).
- [49] B. Favier and E. Knobloch, Robust wall states in rapidly rotating Rayleigh–Bénard convection, J. Fluid Mech. 895, R1 (2020).
- [50] R. E. Ecke, X. Zhang, and O. Shishkina, Connecting wall modes and boundary zonal flows in rotating Rayleigh–Bénard convection, Phys. Rev. Fluids 7, L011501 (2022).
- [51] S. Horn and P. J. Schmid, Prograde, retrograde, and oscillatory modes in rotating Rayleigh–Bénard convection, J. Fluid Mech. 831, 182 (2017).
- [52] X. M. de Wit, A. J. Aguirre Guzmán, M. Madonia, J. S. Cheng, H. J. H. Clercx, and R. P. J. Kunnen, Turbulent rotating convection confined in a slender cylinder: The sidewall circulation, Phys. Rev. Fluids 5, 023502 (2020).
- [53] X. Zhang, R. E. Ecke, and O. Shishkina, Boundary zonal flows in rapidly rotating turbulent thermal convection, J. Fluid Mech. 915, A62 (2021).

- [54] X. Zhang, D. P. M. van Gils, S. Horn, M. Wedi, L. Zwirner, G. Ahlers, R. E. Ecke, S. Weiss, E. Bodenschatz, and O. Shishkina, Boundary zonal flow in rotating turbulent Rayleigh–Bénard convection, Phys. Rev. Lett. **124**, 084505 (2020).
- [55] J. D. Scheel, M. R. Paul, M. C. Cross, and P. F. Fischer, Traveling waves in rotating Rayleigh–Bénard convection: Analysis of modes and mean flow, Phys. Rev. E 68, 066216 (2003).
- [56] A. J. Aguirre Guzmán, M. Madonia, J. S. Cheng, R. Ostilla-Mónico, H. J. H. Clercx, and R. P. J. Kunnen, Flow- and temperature-based statistics characterizing the regimes in rapidly rotating turbulent convection in simulations employing no-slip boundary conditions, Phys. Rev. Fluids 7, 013501 (2022).
- [57] M. Madonia, A. J. Aguirre Guzmán, H. J. H. Clercx, and R. P. J. Kunnen, Velocimetry in rapidly rotating convection: Spatial correlations, flow structures and length scales, Europhys. Lett. 135, 54002 (2021).
- [58] X. M. de Wit, W. J. M. Boot, M. Madonia, A. J. Aguirre Guzmán, and R. P. J. Kunnen, Robust wall modes and their interplay with bulk turbulence in confined rotating Rayleigh–Bénard convection, Phys. Rev. Fluids 8, 073501 (2023).
- [59] R. P. J. Kunnen, R. J. A. M. Stevens, J. Overkamp, C. Sun, G. F. van Heijst, and H. J. H. Clercx, The role of Stewartson and Ekman layers in turbulent rotating Rayleigh–Bénard convection, J. Fluid Mech. 688, 422 (2011).
- [60] W.-T. Früh and P. L. Read, Experiments on a barotropic rotating shear layer. Part 1. Instability and steady vortices, J. Fluid Mech. 383, 143 (1999).
- [61] F. H. Busse, Shear flow instabilities in rotating systems, J. Fluid Mech. 33, 577 (1968).
- [62] M. Plumley, K. Julien, P. Marti, and S. Stellmach, The effects of Ekman pumping on quasi-geostrophic Rayleigh–Bénard convection, J. Fluid Mech. 803, 51 (2016).
- [63] M. Plumley, K. Julien, P. Marti, and S. Stellmach, Sensitivity of rapidly rotating Rayleigh–Bénard convection to Ekman pumping, Phys. Rev. Fluids 2, 094801 (2017).
- [64] See Supplemental Material at http://link.aps.org/supplemental/10.1103/PhysRevFluids.9.053501 for animations of cross sections and vertical azimuthal slices of fields of temperature, vertical vorticity, and vertical, azimuthal, and radial velocity as described in Fig. 29.