

PAPER • OPEN ACCESS

## A rederivation of the conformal anomaly for spin- $\frac{1}{2}$

To cite this article: Hadi Godazgar and Hermann Nicolai 2018 *Class. Quantum Grav.* **35** 105013

View the [article online](#) for updates and enhancements.

### Related content

- [One-loop quantum gravitational corrections to the scalar two-point function at fixed geodesic distance](#)  
Markus B Fröb
- [Setting the boundary free in AdS/CFT](#)  
Geoffrey Compère and Donald Marolf
- [Interactions in higher-spin gravity: a holographic perspective](#)  
Charlotte Sleight

# A rederivation of the conformal anomaly for spin- $\frac{1}{2}$

Hadi Godazgar<sup>1</sup>  and Hermann Nicolai

Max-Planck-Institut für Gravitationsphysik, (Albert-Einstein-Institut),  
Mühlenberg 1, D-14476 Potsdam, Germany

E-mail: [hadi.godazgar@aei.mpg.de](mailto:hadi.godazgar@aei.mpg.de) and [hermann.nicolai@aei.mpg.de](mailto:hermann.nicolai@aei.mpg.de)

Received 13 February 2018, revised 27 March 2018

Accepted for publication 29 March 2018

Published 16 April 2018



CrossMark

## Abstract

We rederive the conformal anomaly for spin- $\frac{1}{2}$  fermions by a genuine Feynman graph calculation, which has not been available so far. Although our calculation merely confirms a result that has been known for a long time, the derivation is new, and thus furnishes a method to investigate more complicated cases (in particular concerning the significance of the quantum trace of the stress tensor in non-conformal theories) where there remain several outstanding and unresolved issues.

Keywords: gravitational anomaly, Weyl anomaly, conformal symmetry

## 1. Introduction

Conformal anomalies have been studied for a long time, see [1–12] for original references and [13–18] for reviews and further references. In four dimensions the gravitational part of the conformal anomaly takes the form

$$\mathcal{A} = a E_4 + b \square R + c C^{\mu\nu\rho\sigma} C_{\mu\nu\rho\sigma} \quad (1)$$

where  $C_{\mu\nu\rho\sigma}$  is the Weyl tensor and  $E_4$  the Euler invariant. Unlike the first and last term the middle contribution can in principle be removed by a local counterterm ( $\sim R^2$ ), but we will keep it here for later purposes. These three terms are the only local expressions which satisfy the Wess–Zumino consistency condition, while an  $R^2$  contribution would require a non-local completion of the anomaly for the consistency condition to be obeyed (see e.g. [19] and references therein).

<sup>1</sup> Author to whom any correspondence should be addressed.



Original content from this work may be used under the terms of the [Creative Commons Attribution 3.0 licence](https://creativecommons.org/licenses/by/3.0/). Any further distribution of this work must maintain attribution to the author(s) and the title of the work, journal citation and DOI.

In this paper we give a new derivation of the coefficients  $a, b, c$  for spin- $\frac{1}{2}$  (Majorana) fermions, by directly calculating

$$\mathcal{A} = g^{\mu\nu} \langle T_{\mu\nu} \rangle \quad (2)$$

up to second order in the metric fluctuations, thus extending the  $\mathcal{O}(\hbar)$  calculation of Capper and Duff [1]. We note that the  $b$  and  $c$  coefficients were originally determined from the two-point correlator of stress tensors in [1] because the two-point function is renormalised by the same counterterm as the 3-point function [3], but this calculation does not yield the  $a$  coefficient. In this paper, by going to  $\mathcal{O}(\hbar^2)$ , we find all coefficients ‘in one go’; there is thus no need to distinguish between type  $A$  and  $B$  anomalies [9], as both appear on an equal footing. Of course, the coefficients of the spin- $\frac{1}{2}$  conformal anomaly have been known for a long time and have been determined by various different methods, via one-loop divergences and heat kernel expansions [4, 5, 13–15, 17], conformal higher spins [12], path integral methods [20, 21], or by QFT in curved spacetime methods [22, 23]. The trace anomaly appears in the TTT correlator in CFT; the general structure of this correlator has been studied in [24–27]—moreover, in [24] the 3-point function of stress tensors for free fields is calculated in  $x$ -space. Curiously, however, to the best of our knowledge, this computation has never been done *à la* Capper–Duff up to  $\mathcal{O}(\hbar^2)$ . In fact, a derivation closest in spirit to the present one is in recent work by Bonora *et al* [28, 29], where, however, only the simpler parity odd contribution (related to the Pontryagin invariant) was considered. Our rederivation is, in principle, a straightforward calculation, very much like the standard textbook derivation of the axial anomaly via triangle graphs, though far more cumbersome in practice. Notably, and in contrast to several other derivations, it does not rely on kinematic choices, such as special gauges for the external graviton  $h_{\mu\nu}$ , nor special values for external momenta, nor on-shell conditions. It thus also provides a toolkit for a similar ‘textbook calculation’ of the (again known)  $s = 0, 1$  anomalies that still remains to be done in this way.

To be sure, we basically regard the present derivation as just a ‘warm-up’ exercise for investigating the conformal anomaly in *non*-conformal theories, in particular for  $s = \frac{3}{2}$  (that is, Poincaré supergravity) where there remain several open issues. These concern for instance the occurrence (or not) of  $R^2$  and/or non-local contributions to the anomaly; a full clarification of non-local terms will probably require the full machinery of scalar  $n$ -point integrals that we review and further develop in section 4 of this paper. The dependence of the  $a$  and  $c$  coefficients on the choice of gauge for the external gravity fluctuations that has been observed for  $s \geq \frac{3}{2}$  [6, 8, 12, 30, 31] is a very strange feature, as it would seem to indicate a breakdown of general covariance—whereas a proper definition of the conformal anomaly should result in a gauge invariant answer also for non-conformal theories (this question is relevant for the possible cancellation of the  $c$  coefficients for  $N \geq 5$  Poincaré supergravities [32]). Another open issue is to see precisely *why* the result for spin- $\frac{3}{2}$  comes out to be negative (this is the only field that contributes with a negative  $c$  coefficient, and is thus indispensable for any cancellation), a feature that is probably related to the absence of a gauge invariant stress tensor and a positive definite Hilbert space of states for spin- $\frac{3}{2}$ . Understanding these issues will hopefully lead to an understanding of whether and what type of non-local terms have to be added to the effective action, with potential cosmological implications as in, for example, [33].

The organisation of this paper is as follows: in section 2.1, we give the Weyl transformation of the curvature, Ricci tensor and scalar and review the Weyl transformation properties of the actions for scalar, Dirac, Maxwell and gravitino fields. In section 3, we consider the action for a massless Majorana field and the expectation value of the stress tensor for such a theory. We present the Feynman rules and calculate the expectation value at first order, section 3.1,

and calculate the  $\square R$  anomaly. We then consider the expectation value of the stress tensor at second order, section 3.2, and show that it is also conserved. We review and develop methods for calculating scalar 3-point loop integrals in section 4, which are then used to find the trace anomaly at second order in the metric perturbation in section 5. We provide a list of the expansion of some relevant quantities under metric perturbations, appendix A; give the result of scalar 2-point integrals, appendix B and list some useful gamma matrix and integral identities in appendix C. We also relegate some technical calculations to appendices D and E.

A final word on our conventions. Lest our multiple use of Greek indices may raise confusion let us state once and for all the convention that we will follow throughout this paper: contractions with the full metric  $g_{\mu\nu}$  are always fully covariant, whereas the *flat* metric  $\eta_{\mu\nu}$  is to be used for all contractions involving the metric fluctuations  $h_{\mu\nu}$  or any quantities appearing inside Feynman diagrams. For instance, when writing out a contraction like  $g^{\mu\nu}T_{\mu\nu}$  in terms of the metric fluctuations  $h_{\mu\nu}$  the result will be an infinite series in terms of the latter where now all contractions are w.r.t. the Minkowski metric  $\eta_{\mu\nu}$ . Where appropriate we will also use flat (Lorentz) indices  $a, b, \dots$  in the fully covariant context, whereas the distinction between flat and curved indices becomes void in terms of the fluctuation expansion.

## 2. Preliminaries

In this section we summarize some general results concerning Weyl transformations so as to make our presentation self-contained, and for reference in future work. We use mostly positive metric signature and positive curvature conventions, namely the conventions of [34].

### 2.1. Weyl transformations

We collect a list of the transformations of some tensors under a Weyl transformation

$$g_{\mu\nu} \longrightarrow \Omega^2 g_{\mu\nu} = e^{2\sigma} g_{\mu\nu}, \quad (3)$$

where all quantities depend on  $x$ . The curvature tensor in  $d$ -dimensions

$$R_{\mu\nu\rho\sigma} = C_{\mu\nu\rho\sigma} + \frac{2}{d-2}g_{\mu[\rho}R_{\sigma]\nu} - \frac{2}{d-2}g_{\nu[\rho}R_{\sigma]\mu} - \frac{2}{(d-1)(d-2)}g_{\mu[\rho}g_{\sigma]\nu}R, \quad (4)$$

and its contractions transform as follows:

$$\begin{aligned} R^\mu{}_{\nu\rho\sigma} &\longrightarrow R^\mu{}_{\nu\rho\sigma} - 2\delta_{[\rho}^\mu \nabla_{\sigma]} \nabla_\nu \sigma + 2g^{\mu\alpha} g_{\nu[\rho} \nabla_{\sigma]} \nabla_\alpha \sigma + 2\delta_{[\rho}^\mu \partial_{\sigma]} \sigma \partial_\nu \sigma \\ &\quad - 2g^{\mu\alpha} g_{\nu[\rho} \partial_{\sigma]} \sigma \partial_\alpha \sigma - 2\delta_{[\rho}^\mu g_{\sigma]\nu} g^{\alpha\beta} \partial_\alpha \sigma \partial_\beta \sigma. \end{aligned} \quad (5)$$

$$R_{\mu\nu} \longrightarrow R_{\mu\nu} - (d-2)\nabla_\mu \nabla_\nu \sigma + (d-2)\partial_\mu \sigma \partial_\nu \sigma - g_{\mu\nu} \square \sigma - (d-2)g_{\mu\nu} g^{\rho\sigma} \partial_\rho \sigma \partial_\sigma \sigma, \quad (6)$$

$$R \longrightarrow \Omega^{-2} \left[ R - 2(d-1)\square \sigma - (d-1)(d-2)g^{\mu\nu} \partial_\mu \sigma \partial_\nu \sigma \right]. \quad (7)$$

The covariant derivative also transforms under a Weyl transformation. In particular, the Christoffel symbol transforms as

$$\Gamma_{\mu\nu}^\rho \longrightarrow \Gamma_{\mu\nu}^\rho + 2\delta_{(\mu}^\rho \partial_{\nu)} \sigma - g^{\rho\sigma} g_{\mu\nu} \partial_\sigma \sigma, \quad (8)$$

while the spin connection transforms as

$$\omega_\mu^{ab} \longrightarrow \omega_\mu^{ab} + 2e_\mu^{[a} e_\nu^{b]} g^{\nu\rho} \partial_\rho \sigma. \quad (9)$$

## 2.2. Weyl invariant actions for spins $s \leq 1$

Given the transformation property of the quadratic operator

$$\begin{aligned} \sqrt{-g} \left( -\square + \frac{d-2}{4(d-1)} R \right) &\longrightarrow \Omega^{d-2} \sqrt{-g} \left( -\square + \frac{d-2}{4(d-1)} R \right) - \frac{d-2}{2} \Omega^{d-2} \sqrt{-g} \square \sigma \\ &\quad - (d-2) \Omega^{d-2} \sqrt{-g} g^{\mu\nu} \left( \frac{d-2}{4} \partial_\mu \sigma \partial_\nu \sigma + \partial_\mu \sigma \partial_\nu \right) \end{aligned} \quad (10)$$

this operator is Weyl covariant if it acts on a scalar  $\phi$  of conformal weight  $-\frac{d-2}{2}$ ,

$$\phi \longrightarrow \Omega^{-\frac{d-2}{2}} \phi. \quad (11)$$

Furthermore, it is then clear that

$$\sqrt{-g} \phi \left( -\square + \frac{d-2}{4(d-1)} R \right) \phi \quad (12)$$

is Weyl invariant.

For a spinor  $\chi$  of conformal weight  $-\frac{d-1}{2}$ ,

$$\chi \longrightarrow \Omega^{-\frac{d-1}{2}} \chi, \quad (13)$$

the Dirac Lagrangian

$$\bar{\chi} \gamma^\mu D_\mu \chi \equiv \bar{\chi} \gamma^\mu \left( \partial_\mu + \frac{1}{4} \omega_{\mu ab} \gamma^{ab} \right) \chi, \quad (14)$$

is already Weyl-invariant by itself without any modification, and for any  $d$ . This can be seen using the transformation of the spin connection, (9), and noting that

$$\gamma_\mu \gamma^{\mu\nu} = (d-1) \gamma^\nu. \quad (15)$$

In four dimensions, the invariance of the Yang–Mills action is anyhow clear because of the invariance of the factor  $\sqrt{-g} g^{\mu\rho} g^{\nu\sigma}$  multiplying  $\text{Tr}(F_{\mu\nu} F_{\rho\sigma})$  under Weyl transformations (where the vector field  $A_\mu$  is assigned Weyl weight zero). In arbitrary dimensions the Yang–Mills action is not, however, Weyl-invariant.

For completeness and later applications let us also display the action of a Weyl transformation on the Rarita–Schwinger action, which is not invariant. It transforms as

$$\epsilon^{\mu\nu\rho\sigma} \bar{\psi}_\mu \gamma_5 \gamma_\nu \nabla_\rho \psi_\sigma \longrightarrow \Omega^{-4} \epsilon^{\mu\nu\rho\sigma} \bar{\psi}_\mu \gamma_5 \gamma_\nu \nabla_\rho \psi_\sigma - 2i g^{\mu\rho} \partial_\rho \sigma \bar{\psi}_{[\mu} \gamma^\nu \psi_{\nu]}, \quad (16)$$

where

$$\psi_\mu \longrightarrow \Omega^{-1/2} \psi_\mu. \quad (17)$$

Hence we see that Weyl invariance is already broken at the classical level. Indeed, it is known that for spin- $\frac{3}{2}$  one needs an action cubic in derivatives for conformal invariance [7].

### 3. Majorana fermions

In this paper we will consider only spin- $\frac{1}{2}$  fermions as they appear to provide the simplest context in which to perform the analysis up to  $\mathcal{O}(h^2)$ . Accordingly, we start with the Dirac action for a Majorana fermion<sup>2</sup>:

$$S = \frac{i}{2} \int e \bar{\chi} \gamma^\mu D_\mu \chi = S^{(0)} + S^{(1)} + S^{(2)} + \dots, \quad (18)$$

where  $D_\mu$  is the covariant derivative with spin-connection  $\omega$  and  $S^{(k)}$  is the action at order  $k$  in the metric fluctuation  $h_{\mu\nu}$ , from

$$g_{\mu\nu}(x) = \eta_{\mu\nu} + h_{\mu\nu}(x). \quad (19)$$

Using the expansions in appendix A, we find that, up to second order in  $h$ ,

$$\begin{aligned} S^{(0)} &= \frac{i}{4} \int \bar{\chi} \overleftrightarrow{\not{D}} \chi, \\ S^{(1)} &= -\frac{i}{8} \int \left( h^{\mu\nu} \bar{\chi} \gamma_\mu \overleftrightarrow{\partial}_\nu \chi - h \bar{\chi} \overleftrightarrow{\not{D}} \chi \right), \\ S^{(2)} &= \frac{i}{32} \int \left( 3 h^{\mu\rho} h_\rho{}^\nu \bar{\chi} \gamma_\mu \overleftrightarrow{\partial}_\nu \chi - 2 h h^{\mu\nu} \bar{\chi} \gamma_\mu \overleftrightarrow{\partial}_\nu \chi - 2 h^{\mu\nu} h_{\mu\nu} \bar{\chi} \overleftrightarrow{\not{D}} \chi \right. \\ &\quad \left. + h^2 \bar{\chi} \overleftrightarrow{\not{D}} \chi + h^\sigma{}_\mu \partial_\nu h_{\rho\sigma} \bar{\chi} \gamma^{\mu\nu\rho} \chi \right), \end{aligned} \quad (20)$$

where  $h \equiv \eta^{\mu\nu} h_{\mu\nu}$ ,  $\overleftrightarrow{\partial}_\mu = \overrightarrow{\partial}_\mu - \overleftarrow{\partial}_\mu$ , and where the left action of the differential operator is only on the fermion  $\bar{\chi}$ . Also, we use lower case Latin letters for tangent space indices, we use Greek indices for tensors after perturbatively expanding the metric. In both cases the position of indices is raised/lowered with the Minkowski metric. Moreover, the fermionic stress tensor admits a similar expansion<sup>3</sup>,

$$T_{\mu\nu} = \frac{2}{e} g_{\mu\rho} g_{\nu\sigma} \frac{\delta S}{\delta g_{\rho\sigma}} = -\frac{1}{e} e_{\mu a} \frac{\delta S}{\delta e_{a\nu}} = -\frac{i}{4} \left( \bar{\chi} \gamma_{(\mu} \overleftrightarrow{D}_{\nu)} \chi - g_{\mu\nu} \bar{\chi} \overleftrightarrow{\not{D}} \chi \right) = T_{\mu\nu}^{(0)} + T_{\mu\nu}^{(1)} + \dots, \quad (21)$$

where to first order in  $h$ ,

$$T_{\mu\nu}^{(0)} = -\frac{i}{4} \left( \bar{\chi} \gamma_{(\mu} \overleftrightarrow{\partial}_{\nu)} \chi - \eta_{\mu\nu} \bar{\chi} \overleftrightarrow{\not{D}} \chi \right) \quad (22)$$

$$T_{\mu\nu}^{(1)} = -\frac{i}{8} \left( h_{\rho(\mu} \bar{\chi} \gamma^{\rho} \overleftrightarrow{\partial}_{\nu)} \chi + \eta_{\mu\nu} h^{\rho\sigma} \bar{\chi} \gamma_\rho \overleftrightarrow{\partial}_\sigma \chi - 2 h_{\mu\nu} \bar{\chi} \overleftrightarrow{\not{D}} \chi - \partial_\rho h_{\sigma(\mu} \bar{\chi} \gamma_{\nu)}{}^{\rho\sigma} \chi \right). \quad (23)$$

In the Majorana representation  $\bar{\chi} \gamma^\mu \chi = 0$ , hence terms containing such contractions do not contribute. However, even for Dirac fermions for which  $\bar{\chi} \gamma^\mu \chi \neq 0$  terms with such contractions cancel in the final result, and the expansion is, up to an overall factor of 2, given by

<sup>2</sup> Up to an overall factor of  $\frac{1}{2}$  this action is the same for Dirac and Majorana fermions. The action for a massless Majorana fermion is also classically the same as the action for a Weyl fermion up to a total derivative term. There are recent claims that they are different at the quantum level and that there is, in particular, an odd parity anomaly for Weyl fermions [28, 29]. We will not address this claim here, but we just note that there is no issue for Majorana fermions as (18) is real.

<sup>3</sup> The symmetrisation of the stress–energy tensor comes from the variation of the spin connection in a second-order formalism.

**Figure 1.** Feynman rules for graviton-fermion interactions; the crossed vertex comes from the expansion of  $T_{\mu\nu}$  to  $\mathcal{O}(h)$ .

the very same expression (20). Hence the anomaly for a Dirac fermion is twice the one for a Majorana fermion. From the Lagrangian density above it is then straightforward to read off the Feynman rules with up to two external graviton lines. The relevant expressions are given in figure 1<sup>4</sup>.

We are interested in the expectation value of the stress tensor at first and second order in the metric perturbation,

$$\begin{aligned}
 \langle T_{\mu\nu}(x) \rangle &= \left\langle T_{\mu\nu}(x) e^{i(S^{(1)} + S^{(2)} + \dots)} \right\rangle_0, \\
 &= \left\langle \left( T_{\mu\nu}^{(0)}(x) + T_{\mu\nu}^{(1)}(x) + \dots \right) \left( 1 + iS^{(1)} + \left( iS^{(2)} - \frac{1}{2}S^{(1)}S^{(1)} \right) + \dots \right) \right\rangle_0, \\
 &= i \left\langle T_{\mu\nu}^{(0)}(x) S^{(1)} \right\rangle_0 + i \left\langle T_{\mu\nu}^{(1)}(x) S^{(1)} \right\rangle_0 + i \left\langle T_{\mu\nu}^{(0)}(x) S^{(2)} \right\rangle_0 \\
 &\quad - \frac{1}{2} \left\langle T_{\mu\nu}^{(0)}(x) S^{(1)} S^{(1)} \right\rangle_0 + \dots
 \end{aligned} \tag{24}$$

where  $\langle \dots \rangle_0$  denotes the free expectation value (to be evaluated in the spin- $\frac{1}{2}$  Fock space). Note that at zeroth order,  $\langle T_{\mu\nu}^{(0)}(x) \rangle_0$ , we only have tadpole diagrams, which vanish in dimensional regularisation. Furthermore, there is no  $\langle T_{\mu\nu}^{(1)}(x) \rangle_0$  contribution at first order in  $h$ , since these also contribute only tadpole diagrams.

<sup>4</sup> Since we are working with Lorentzian signature it should be understood that we are using the usual  $i\epsilon$  prescription for the propagator, although we do not write this out explicitly.

### 3.1. Expectation value of the stress tensor at $\mathcal{O}(h)$

In this section we briefly summarise the old  $\mathcal{O}(h)$  result of [1]. At first order, from equation (24) the expectation value of the stress tensor is

$$\langle T_{\mu\nu}(x) \rangle|_{\mathcal{O}(h)} = i \langle T_{\mu\nu}^{(0)}(x) S^{(1)} \rangle_0 = \int d^d y \int \frac{d^d p}{(2\pi)^d} e^{-ip \cdot (x-y)} T_{\mu\nu\rho\sigma}(p) h^{\rho\sigma}(y), \quad (25)$$

which defines the two-point function  $T_{\mu\nu\rho\sigma}(p)$  in momentum space. Using the Feynman rules we have

$$T_{\mu\nu\rho\sigma}(p) = \frac{i}{16} \int \frac{d^d k}{(2\pi)^d} \text{tr} \left( \frac{\not{k}}{k^2} (2k-p)_{(\mu} \gamma_{\nu)} \frac{\not{k}-\not{p}}{(k-p)^2} (2k-p)_{(\rho} \gamma_{\sigma)} \right), \quad (26)$$

where we have neglected all terms proportional to  $\eta_{\mu\nu}$  and  $\eta_{\rho\sigma}$ , since, using the identity

$$\frac{\not{k}(2\not{k}-\not{p})(\not{k}-\not{p})}{k^2(k-p)^2} = \frac{(\not{k}-\not{p})}{(k-p)^2} + \frac{\not{k}}{k^2}, \quad (27)$$

these terms reduce to tadpole integrals which vanish. Note also the simple identities

$$T_{\mu\nu\rho\sigma}(p) = T_{\mu\nu\rho\sigma}(-p), \quad T_{\mu\nu\rho\sigma}(p) = T_{\rho\sigma\mu\nu}(p). \quad (28)$$

As shown in appendix D, equation (D.1), the explicit symmetrisation of the  $\mu\nu$  indices in the integral (26) is not required, as the antisymmetric part vanishes, a fact that we will exploit to simplify some of the subsequent calculations.

The conservation of the stress tensor

$$\nabla^\mu \langle T_{\mu\nu} \rangle = 0 \quad (29)$$

and the tracelessness

$$\langle g^{\mu\nu} T_{\mu\nu} \rangle = 0 \quad (30)$$

at order  $h$ , translate to the following Ward identities

$$p^\mu T_{\mu\nu\rho\sigma} = 0, \quad (31)$$

$$\eta^{(d)\mu\nu} T_{\mu\nu\rho\sigma} = 0, \quad (32)$$

where it is important that the trace is taken in  $d$  dimensions (indicated in the notation by putting the trace *inside* the brackets in (30) and superscript  $(d)$  on the  $\eta$ ). In order to verify the conservation Ward identity, we note that

$$p^\mu (2k-p)_\mu = k^2 - (k-p)^2. \quad (33)$$

Hence  $p^\mu T_{\mu\nu\rho\sigma}$  reduces to a tadpole integral which vanishes. Similarly, the  $d$  dimensional trace reduces to a tadpole integral upon using identity (27). This is in accord with the fact that the Dirac Lagrangian density is classically Weyl invariant in all dimensions with a  $d$ -dependent scaling of the fermions.



Evaluating the 2-point function integral, (26), using the integral identities (B.1)–(B.4), we obtain

$$-i \frac{1}{(2\sqrt{\pi})^d} \cdot \frac{2^{d/2} I(p)}{32(d^2 - 1)} \left[ (d-2) p_\mu p_\nu p_\rho p_\sigma - 2(d-1) p^2 p_{(\mu} \eta_{\nu)(\rho} p_{\sigma)} + p^2 (\eta_{\mu\nu} p_\rho p_\sigma + \eta_{\rho\sigma} p_\mu p_\nu) + (p^2)^2 ((d-1) \eta_{\mu(\rho} \eta_{\sigma)\nu} - \eta_{\mu\nu} \eta_{\rho\sigma}) \right]. \quad (34)$$

where the extra factor of  $1/(2\sqrt{\pi})^d$  in front is due to our normalisation of the integral  $I(p)$  in (B.5). It is now straightforward to verify that the contraction of the above expression with  $p^\mu$  vanishes, confirming that the Ward identity for general covariance is satisfied. Furthermore, we can verify again that the contraction of the  $\mu\nu$  indices in  $d$  dimensions is also zero.

However, contracting the  $\mu\nu$  indices in four dimensions we obtain

$$\eta^{(4)\rho\sigma} T_{\rho\sigma\mu\nu} = \frac{p^2}{60(4\pi)^2} (p_\mu p_\nu - \eta_{\mu\nu} p^2), \quad (35)$$

from which we find the  $\square R$  anomaly at  $\mathcal{O}(\hbar)$ , to wit,

$$g^{\mu\nu} \langle T_{\mu\nu} \rangle|_{\mathcal{O}(\hbar)} = \frac{1}{60(4\pi)^2} \square R|_{\mathcal{O}(\hbar)} \quad (36)$$

where we now put the  $g^{\mu\nu}$  *outside* the bracket to indicate that the trace is to be taken in four dimensions, after regularisation and renormalisation. We again stress that the  $\square R$  anomaly is scheme-dependent in the sense that its coefficient can be changed by adding a local counterterm. However, within dimensional regularisation, we choose to calculate this coefficient without adding any counterterms to the action.

We stress that this  $\mathcal{O}(\hbar)$  calculation can *not* give the  $a$  and  $c$  coefficients as these require at least  $\mathcal{O}(\hbar^2)$ . However, with an extra assumption on the counterterm it is possible to derive the  $c$  coefficient at least by indirect arguments [3]. This can be seen as follows: introducing the counterterm  $\epsilon^{-1} \Delta W$ , where  $\Delta W \equiv \int d^d x \sqrt{-g} C^2$  and  $C$  is the 4-dimensional Weyl tensor, and functionally differentiating, we get

$$2e^{-1} g_{\mu\nu} \frac{\delta}{\delta g_{\mu\nu}} \Delta W = (d-4) \left( C^2 + \frac{2}{3} \square R \right) \quad (37)$$

which shows that

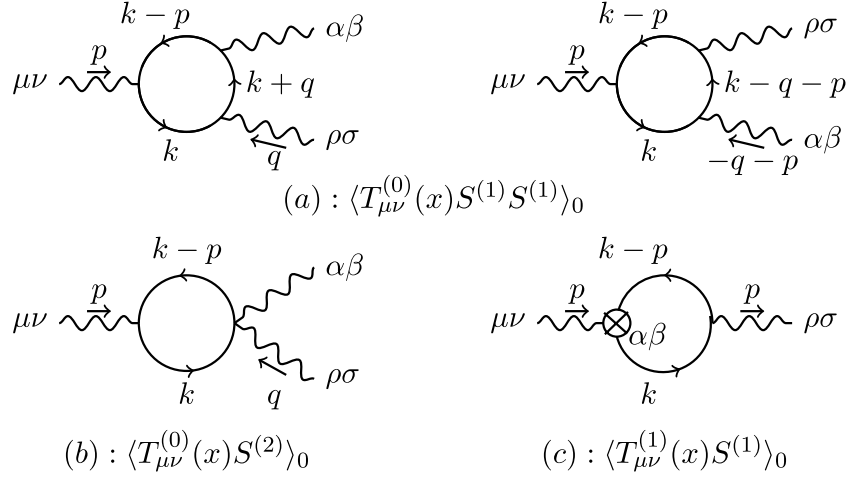
$$b = \frac{2}{3} c, \quad (38)$$

a relation which we shall later verify explicitly at  $\mathcal{O}(\hbar^2)$ . By contrast, there is no such indirect and labor saving argument for the coefficient  $a$ .

### 3.2. Expectation value of the stress tensor at $\mathcal{O}(\hbar^2)$

The evaluation of the expectation value of the stress tensor to second order in the metric fluctuations is far more involved than at first order because there are many more contributions. In particular we must now consider the 3-point functions which are given by Feynman diagrams labelled (a) and (b) in figure 2<sup>5</sup>. These diagrams, as well as a new diagram labelled (c)

<sup>5</sup>The labelling in the Feynman diagrams should be seen as mnemonics for writing down the integrals that contribute to the expectation value of the stress tensor. The object that is calculated is  $x$ -dependent and the momenta  $p$  and  $q$  are integrated over.



**Figure 2.** Feynman diagrams for 3-point function of stress tensor insertions. The diagrams at the top, labelled (a), contribute to  $T_{\mu\nu\rho\sigma\alpha\beta}^{(1)}$ ; diagram (b) and (c) contribute to  $T_{\mu\nu\rho\sigma\alpha\beta}^{(2)}$  and  $T_{\mu\nu\rho\sigma\alpha\beta}^{(3)}$ , respectively.

in figure 2, contribute to the the expectation value of the stress tensor at second order in the metric perturbation,

$$\begin{aligned} \langle T_{\mu\nu}(x) \rangle|_{\mathcal{O}(h^2)} = \int d^d y d^d z \int \frac{d^d p}{(2\pi)^d} \frac{d^d q}{(2\pi)^d} e^{-ip \cdot (x-y) - iq \cdot (z-y)} h^{\rho\sigma}(y) \left( T_{\mu\nu\rho\sigma\alpha\beta}^{(1)}(p, q) h^{\alpha\beta}(z) \right. \\ \left. + T_{\mu\nu\rho\sigma\alpha\beta}^{(2)}(p, q) h^{\alpha\beta}(z) + T_{\mu\nu\rho\sigma\alpha\beta}^{(3)}(p) h^{\alpha\beta}(x) \right). \end{aligned} \quad (39)$$

Here  $T_{\mu\nu\rho\sigma\alpha\beta}^{(1)}$ ,  $T_{\mu\nu\rho\sigma\alpha\beta}^{(2)}$  and  $T_{\mu\nu\rho\sigma\alpha\beta}^{(3)}$  are the 3-point diagrams corresponding to  $\langle T_{\mu\nu}^{(0)}(x)S^{(1)}S^{(1)} \rangle_0$ ,  $\langle T_{\mu\nu}^{(0)}(x)S^{(2)} \rangle_0$  and  $\langle T_{\mu\nu}^{(1)}(x)S^{(1)} \rangle_0$  in equation (24), respectively. Using the Feynman rules given in figure 1, the Feynman diagrams labelled (a) in figure 2 give

$$\begin{aligned} T_{\mu\nu\rho\sigma\alpha\beta}^{(1)}(p, q) = \frac{i}{128} \int \frac{d^d k}{(2\pi)^d} \text{tr} \left\{ \frac{k}{k^2} \left[ (2k-p)_{(\mu\gamma\nu)} - \eta_{\mu\nu}(2k-p) \right] \frac{k-p}{(k-p)^2} \right. \\ \times \left( \left[ (2k-p+q)_{(\rho\gamma\sigma)} - \eta_{\rho\sigma}(2k-p+q) \right] \frac{k+q}{(k+q)^2} \right. \\ \times \left. \left[ (2k+q)_{(\alpha\gamma\beta)} - \eta_{\alpha\beta}(2k+q) \right] \right. \\ \left. + \left[ (2k-2p-q)_{(\alpha\gamma\beta)} - \eta_{\alpha\beta}(2k-2p-q) \right] \frac{k-p-q}{(k-p-q)^2} \right. \\ \left. \times \left. \left[ (2k-p-q)_{(\rho\gamma\sigma)} - \eta_{\rho\sigma}(2k-p-q) \right] \right) \right\}. \end{aligned} \quad (40)$$

Letting  $k \rightarrow -k+p$  in the second term and using the gamma matrix identity (C.2), we can show that the second term is identical to the first term, viz the contribution from the two

diagrams labelled (a) is identical. Furthermore the terms involving  $\eta_{\alpha\beta}$  or  $\eta_{\rho\sigma}$ , can be written as two-point function integrals, defined in equation (26), using identities analogous to (27). Terms with more than one  $\eta$  reduce to tadpole integrals, which vanish, or integrals of the form

$$\int \frac{d^d k}{(2\pi)^d} \text{tr} \left( \frac{\not{k}}{k^2} (2\not{k} - \not{p} + \not{q}) \frac{\not{k} - \not{p}}{(k-p)^2} (2k-p)_{(\rho}\gamma_{\sigma)} \right), \quad (41)$$

which using identity (27) reduces to

$$\int \frac{d^d k}{(2\pi)^d} \text{tr} \left( \frac{\not{k}}{k^2} \not{q} \frac{\not{k} - \not{p}}{(k-p)^2} (2k-p)_{(\rho}\gamma_{\sigma)} \right) = 0 \quad (42)$$

by identity (C.5). Hence, we can rewrite  $T^{(1)}$  as

$$\begin{aligned} T_{\mu\nu\rho\sigma\alpha\beta}^{(1)}(p, q) &= \frac{1}{64} T_{\mu\nu\rho\sigma\alpha\beta}(p, q) - \frac{1}{4} \eta_{\mu\nu} [T_{\rho\sigma\alpha\beta}(p+q) + T_{\rho\sigma\alpha\beta}(q)] \\ &\quad - \frac{1}{4} \eta_{\rho\sigma} [T_{\mu\nu\alpha\beta}(p) + T_{\mu\nu\alpha\beta}(q)] - \frac{1}{4} \eta_{\alpha\beta} [T_{\mu\nu\rho\sigma}(p+q) + T_{\mu\nu\rho\sigma}(p)], \end{aligned} \quad (43)$$

where we define

$$T_{\mu\nu\rho\sigma\alpha\beta} \equiv i \int \frac{d^d k}{(2\pi)^d} \text{tr} \left\{ \frac{\not{k}}{k^2} (2k-p)_{(\mu}\gamma_{\nu)} \frac{\not{k} - \not{p}}{(k-p)^2} (2k-p+q)_{(\rho}\gamma_{\sigma)} \frac{\not{k} + \not{q}}{(k+q)^2} (2k+q)_{(\alpha}\gamma_{\beta)} \right\} \quad (44)$$

that is, the original expression but *without* the trace terms.

Moreover, the Feynman diagrams labelled (b) and (c), respectively, give

$$\begin{aligned} T_{\mu\nu\rho\sigma\alpha\beta}^{(2)}(p, q) &= -\frac{i}{128} \int \frac{d^d k}{(2\pi)^d} \text{tr} \left\{ \frac{\not{k}}{k^2} (2k-p)_{(\mu}\gamma_{\nu)} \frac{\not{k} - \not{p}}{(k-p)^2} \right. \\ &\quad \times \left( 3 (\eta_{(\rho|\alpha}(2k-p)_{\beta)}\gamma_{|\sigma)} + \eta_{(\alpha|(\rho}(2k-p)_{\sigma)}\gamma_{|\beta)}) \right. \\ &\quad \left. - 2 (\eta_{\rho\sigma}(2k-p)_{(\alpha}\gamma_{\beta)} + \eta_{\alpha\beta}(2k-p)_{(\rho}\gamma_{\sigma)}) \right. \\ &\quad \left. \left. + \frac{1}{2} (2q+p)^\tau (\gamma_{\tau\alpha(\rho}\eta_{\sigma)\beta} + \gamma_{\tau\beta(\rho}\eta_{\sigma)\alpha}) \right) \right\}, \end{aligned} \quad (45)$$

$$\begin{aligned} T_{\mu\nu\rho\sigma\alpha\beta}^{(3)}(p) &= \frac{i}{32} \int \frac{d^d k}{(2\pi)^d} \text{tr} \left\{ \frac{\not{k} - \not{p}}{(k-p)^2} (2k-p)_{(\rho}\gamma_{\sigma)} \frac{\not{k}}{k^2} \right. \\ &\quad \left. \times \left[ (2k-p)_{(\mu}\eta_{\nu)(\alpha}\gamma_{\beta)} + \eta_{\mu\nu}(2k-p)_{(\alpha}\gamma_{\beta)} + \frac{1}{2} p^\tau (\gamma_{\tau\alpha(\mu}\eta_{\nu)\beta} + \gamma_{\tau\beta(\mu}\eta_{\nu)\alpha}) \right] \right\}, \end{aligned} \quad (46)$$

where, as in the two-point function evaluation, we have used the fact that some terms lead to tadpole integrals which vanish. It is also straightforward to show that the terms proportional to  $\gamma^{\tau\alpha\beta}$  in both  $T^{(2)}$  and  $T^{(3)}$  vanish. Therefore, we can express both contributions solely in terms of the two-point function integral, (26),

$$T_{\mu\nu\rho\sigma\alpha\beta}^{(2)}(p) = -\frac{3}{4} \eta_{(\alpha|(\rho} T_{\sigma)|\beta)\mu\nu}(p) + \frac{1}{4} \eta_{\rho\sigma} T_{\mu\nu\alpha\beta}(p) + \frac{1}{4} \eta_{\alpha\beta} T_{\mu\nu\rho\sigma}(p), \quad (47)$$

$$T_{\mu\nu\rho\sigma\alpha\beta}^{(3)}(p) = \frac{1}{2}\eta_{(\alpha|\mu}T_{\nu)|\beta\rho\sigma}(p) + \frac{1}{2}\eta_{\mu\nu}T_{\rho\sigma\alpha\beta}(p). \quad (48)$$

### 3.3. Conservation

The conservation of the expectation value of the stress tensor can be expressed as follows:

$$\nabla^\mu \langle T_{\mu\nu} \rangle = g^{\mu\rho} \left( \partial_\rho \langle T_{\mu\nu} \rangle - \Gamma_{\rho\mu}^\sigma \langle T_{\sigma\nu} \rangle - \Gamma_{\rho\nu}^\sigma \langle T_{\mu\sigma} \rangle \right) = 0. \quad (49)$$

Using the expansion of the metric and Christoffel symbol components in appendix A, at second order in the metric perturbation the above identity reduces to

$$\begin{aligned} & \partial^\mu \langle T_{\mu\nu} \rangle|_{\mathcal{O}(h^2)} - h^{\mu\rho} \partial_\rho \langle T_{\mu\nu} \rangle|_{\mathcal{O}(h)} \\ & - \frac{1}{2} (2\partial_\mu h^{\mu\rho} - \partial^\rho h) \langle T_{\rho\nu} \rangle|_{\mathcal{O}(h)} - \frac{1}{2} \partial_\nu h^{\mu\rho} \langle T_{\mu\rho} \rangle|_{\mathcal{O}(h)} = 0, \end{aligned} \quad (50)$$

where  $\langle T_{\mu\nu} \rangle|_{\mathcal{O}(h)}$  and  $\langle T_{\mu\nu} \rangle|_{\mathcal{O}(h^2)}$  are defined in equations (25) and (39), respectively. Equation (49) is fully covariant. However, in equation (50), and for the rest of the section, the indices are raised and lowered with the flat metric.

We first consider

$$\begin{aligned} \partial^\mu \langle T_{\mu\nu} \rangle(x)|_{\mathcal{O}(h^2)} &= \int d^d y d^d z \int \frac{d^d p}{(2\pi)^d} \frac{d^d q}{(2\pi)^d} e^{-ip \cdot (x-y) - iq \cdot (z-y)} h^{\rho\sigma}(y) \\ & \times \left\{ (-ip^\mu) \left( T_{\mu\nu\rho\sigma\alpha\beta}^{(1)}(p, q) h^{\alpha\beta}(z) + T_{\mu\nu\rho\sigma\alpha\beta}^{(2)}(p, q) h^{\alpha\beta}(z) \right. \right. \\ & \left. \left. + T_{\mu\nu\rho\sigma\alpha\beta}^{(3)}(p) h^{\alpha\beta}(x) \right) + T_{\mu\nu\rho\sigma\alpha\beta}^{(3)}(p) \partial^\mu h^{\alpha\beta}(x) \right\}. \end{aligned} \quad (51)$$

Using equation (43) and the conservation Ward identity at first order in  $h$ , (31),

$$\begin{aligned} p^\mu T_{\mu\nu\rho\sigma\alpha\beta}^{(1)}(p, q) &= \frac{1}{64} p^\mu T_{\mu\nu\rho\sigma\alpha\beta} - \frac{1}{4} p_\nu [T_{\rho\sigma\alpha\beta}(p+q) + T_{\rho\sigma\alpha\beta}(q)] \\ & - \frac{1}{4} p^\mu [\eta_{\rho\sigma} T_{\mu\nu\alpha\beta}(q) + \eta_{\alpha\beta} T_{\mu\nu\rho\sigma}(p+q)]. \end{aligned} \quad (52)$$

Furthermore, using equations (47) and (48),

$$p^\mu T_{\mu\nu\rho\sigma\alpha\beta}^{(2)}(p) = 0, \quad (53)$$

$$p^\mu T_{\mu\nu\rho\sigma\alpha\beta}^{(3)}(p) = \frac{1}{4} p_{(\alpha} T_{\beta)\nu\rho\sigma}(p) + \frac{1}{2} p_\nu T_{\rho\sigma\alpha\beta}(p). \quad (54)$$

We have expressed all the terms in terms of the two-point function integral except the first term on the rhs of equation (52), which we would like to also rewrite in terms of two-point function integral,

$$\begin{aligned}
& p^\mu T_{\mu\nu\rho\sigma\alpha\beta} \\
&= \frac{i}{2} \int \frac{d^d k}{(2\pi)^d} \text{tr} \left\{ \left( \not{k} \gamma_\nu + (2k - p)_\nu \right) \frac{\not{k} - \not{p}}{(k - p)^2} (2k - p + q)_{(\rho} \gamma_\sigma) \frac{\not{k} + \not{q}}{(k + q)^2} (2k + q)_{(\alpha} \gamma_\beta) \right\} \\
&\quad - \frac{i}{2} \int \frac{d^d k}{(2\pi)^d} \text{tr} \left\{ \frac{\not{k}}{k^2} \left( \gamma_\nu (\not{k} - \not{p}) + (2k - p)_\nu \right) (2k - p + q)_{(\rho} \gamma_\sigma) \frac{\not{k} + \not{q}}{(k + q)^2} (2k + q)_{(\alpha} \gamma_\beta) \right\},
\end{aligned} \tag{55}$$

where we have used (33) and  $\not{p} = \not{k} - (\not{k} - \not{p})$  to cancel a propagator factor. We redefine  $k \rightarrow -k + p$  in the first term and  $k \rightarrow -k$  in the second term, whereupon we obtain

$$p^\mu T_{\mu\nu\rho\sigma\alpha\beta} = \frac{1}{2} \left( \tilde{T}_{\nu\rho\sigma\alpha\beta}(p, q) - \tilde{T}_{\nu\alpha\beta\rho\sigma}(-p, p + q) \right), \tag{56}$$

where

$$\tilde{T}_{\nu\rho\sigma\alpha\beta}(p, q) \equiv i \int \frac{d^d k}{(2\pi)^d} \text{tr} \left\{ \frac{\not{k}}{k^2} \left( \gamma_\nu (\not{k} + \not{p}) + (2k + p)_\nu \right) (2k + p - q)_{(\rho} \gamma_\sigma) \frac{\not{k} - \not{q}}{(k - q)^2} (2k - q)_{(\alpha} \gamma_\beta) \right\} \tag{57}$$

and we have used the identity (C.2). In appendix E, we simplify this integral and derive equation (E.5). Using this result, we arrive at

$$\begin{aligned}
p^\mu T_{\mu\nu\rho\sigma\alpha\beta} &= 8 \left[ 3 p_{(\rho} T_{\sigma)\nu\alpha\beta}(q) + 2 (p + q)_\nu T_{\rho\sigma\alpha\beta}(q) - p^\tau \eta_{\nu(\rho} T_{\sigma)\tau\alpha\beta}(q) \right. \\
&\quad \left. + 3 p_{(\alpha} T_{\beta)\nu\rho\sigma}(p + q) - 2 q_\nu T_{\rho\sigma\alpha\beta}(p + q) - p^\tau \eta_{\nu(\alpha} T_{\beta)\tau\rho\sigma}(p + q) \right].
\end{aligned} \tag{58}$$

Hence, from equation (52),

$$\begin{aligned}
p^\mu T_{\mu\nu\rho\sigma\alpha\beta}^{(1)}(p, q) &= \frac{1}{8} \left[ 3 p_{(\alpha} T_{\beta)\nu\rho\sigma}(p + q) - 2 (p + q)_\nu T_{\rho\sigma\alpha\beta}(p + q) - 2 p^\mu \eta_{\alpha\beta} T_{\mu\nu\rho\sigma}(p + q) \right. \\
&\quad \left. - p^\tau \eta_{\nu(\alpha} T_{\beta)\tau\rho\sigma}(p + q) + 3 p_{(\rho} T_{\sigma)\nu\alpha\beta}(q) + 2 q_\nu T_{\rho\sigma\alpha\beta}(q) \right. \\
&\quad \left. - 2 p^\mu \eta_{\rho\sigma} T_{\mu\nu\alpha\beta}(q) - p^\tau \eta_{\nu(\rho} T_{\sigma)\tau\alpha\beta}(q) \right].
\end{aligned} \tag{59}$$

Integrating the above equation over  $p$  and letting  $p \rightarrow p - q$ ,

$$\begin{aligned}
\int \frac{d^d p}{(2\pi)^d} e^{-ip \cdot (x-y)} p^\mu T_{\mu\nu\rho\sigma\alpha\beta}^{(1)}(p, q) &= \frac{1}{8} \int \frac{d^d p}{(2\pi)^d} e^{-i(p-q) \cdot (x-y)} \left[ 3 (p - q)_{(\alpha} T_{\beta)\nu\rho\sigma}(p) \right. \\
&\quad \left. - 2 p_\nu T_{\rho\sigma\alpha\beta}(p) + 2 q^\mu \eta_{\alpha\beta} T_{\mu\nu\rho\sigma}(p) + q^\tau \eta_{\nu(\alpha} T_{\beta)\tau\rho\sigma}(p) \right. \\
&\quad \left. + 3 (p - q)_{(\rho} T_{\sigma)\nu\alpha\beta}(q) + 2 q_\nu T_{\rho\sigma\alpha\beta}(q) \right. \\
&\quad \left. - 2 p^\mu \eta_{\rho\sigma} T_{\mu\nu\alpha\beta}(q) - p^\tau \eta_{\nu(\rho} T_{\sigma)\tau\alpha\beta}(q) \right].
\end{aligned} \tag{60}$$

Therefore, using also equations (53) and (54) and (48) and reparametrising the integration variables,

$$\begin{aligned}
\partial^\mu \langle T_{\mu\nu} \rangle(x)|_{\mathcal{O}(h^2)} &= \int d^d y d^d z \int \frac{d^d p}{(2\pi)^d} \frac{d^d q}{(2\pi)^d} e^{-ip \cdot (x-y) - iq \cdot (z-x)} h^{\rho\sigma}(y) \\
&\times \left\{ -\frac{i}{8} h^{\alpha\beta}(z) \left[ (5p - 3q)_\alpha T_{\beta\nu\rho\sigma}(p) + 2p_\nu T_{\rho\sigma\alpha\beta}(p) \right. \right. \\
&\quad + 2q^\mu \eta_{\alpha\beta} T_{\mu\nu\rho\sigma}(p) + q^\tau \eta_{\nu\alpha} T_{\beta\tau\rho\sigma}(p) + 3(p - q)_\rho T_{\sigma\nu\alpha\beta}(q) \\
&\quad + 2q_\nu T_{\rho\sigma\alpha\beta}(q) - 2p^\mu \eta_{\rho\sigma} T_{\mu\nu\alpha\beta}(q) - p^\tau \eta_{\nu\rho} T_{\sigma\tau\alpha\beta}(q) \left. \right] \\
&\quad + \frac{1}{4} \left[ T_{\rho\sigma\nu\alpha}(p) \partial_\beta h^{\alpha\beta}(z) + T_{\rho\sigma\mu\alpha} \eta_{\beta\nu}(p) \partial_\mu h^{\alpha\beta}(z) \right. \\
&\quad \left. + 2T_{\rho\sigma\alpha\beta}(p) \partial_\nu h^{\alpha\beta}(z) \right] \left. \right\}. \tag{61}
\end{aligned}$$

The terms in the integrand that are proportional to two-point function integrals with arguments  $q$  can be replaced by terms proportional to those with arguments  $p$ <sup>6</sup>. Whereupon,

$$\begin{aligned}
\partial^\mu \langle T_{\mu\nu} \rangle|_{\mathcal{O}(h^2)} &= -\frac{1}{4} \int d^d y d^d z \int \frac{d^d p}{(2\pi)^d} \frac{d^d q}{(2\pi)^d} e^{-ip \cdot (x-y) - iq \cdot (z-x)} h^{\rho\sigma}(y) \\
&\times \left( i h^{\alpha\beta}(z) \left[ (4p - 3q)_\alpha T_{\beta\nu\rho\sigma}(p) + 2q^\mu \eta_{\alpha\beta} T_{\mu\nu\rho\sigma}(p) + q^\tau \eta_{\nu\alpha} T_{\beta\tau\rho\sigma}(p) \right] \right. \\
&\quad \left. - T_{\rho\sigma\nu\alpha}(p) \partial_\beta h^{\alpha\beta}(z) - \eta^{\alpha\beta} T_{\rho\sigma\mu\alpha}(p) \partial^\mu h_{\beta\nu}(z) - 2T_{\rho\sigma\alpha\beta}(p) \partial_\nu h^{\alpha\beta}(z) \right). \tag{62}
\end{aligned}$$

Integrating by parts over the  $y$  and  $z$  integrals,

$$\begin{aligned}
\partial^\mu \langle T_{\mu\nu} \rangle|_{\mathcal{O}(h^2)} &= \int d^d y d^d z \int \frac{d^d p}{(2\pi)^d} \frac{d^d q}{(2\pi)^d} e^{-ip \cdot (x-y) - iq \cdot (z-x)} h^{\rho\sigma}(y) \left\{ -ip_\alpha T_{\beta\nu\rho\sigma}(p) h^{\alpha\beta}(z) \right. \\
&\quad \left. + \partial_\alpha h^{\alpha\beta}(z) T_{\beta\nu\rho\sigma}(p) - \frac{1}{2} \partial^\mu h(z) T_{\mu\nu\rho\sigma}(p) + \frac{1}{2} \partial_\nu h^{\alpha\beta}(z) T_{\rho\sigma\alpha\beta}(p) \right\}, \\
&= h^{\alpha\beta} \partial_\alpha \langle T_{\beta\nu} \rangle|_{\mathcal{O}(h)} + \partial_\alpha h^{\alpha\beta} \langle T_{\beta\nu} \rangle|_{\mathcal{O}(h)} - \frac{1}{2} \partial^\mu h \langle T_{\mu\nu} \rangle|_{\mathcal{O}(h)} + \frac{1}{2} \partial_\nu h^{\alpha\beta}(z) \langle T_{\alpha\beta} \rangle|_{\mathcal{O}(h)}, \tag{63}
\end{aligned}$$

where in the last line we have integrated over  $q$  and  $z$ , which sets  $z = x$ , and used definition (25). We have, therefore, verified equation (50) and hence

$$\nabla^\mu \langle T_{\mu\nu} \rangle = 0 \tag{64}$$

up to and including second order in the metric perturbation.

#### 4. Scalar 3-point loop integrals

We have seen in the preceding chapters that the evaluation of the expectation value of the stress tensor to second order in the metric fluctuations requires the computation of certain 3-point

<sup>6</sup> More precisely, the relabelling of the integration variables implies that the integrand must be invariant under  $p \leftrightarrow -q$  and  $(\rho\sigma) \leftrightarrow (\alpha\beta)$ .

Feynman loop integrals (and correspondingly the evaluation of  $(n + 1)$ -point loop integrals if one expands to  $n$ -th order in the metric fluctuations). Such integrals have been much investigated in the literature, see e.g. [35, 36] for recent reviews and references. Nevertheless, and also with regard to possible future applications, we here collect some formulae needed for our computation that to the best of our knowledge have not been given in fully explicit form in the literature, although the general procedure for their derivation is of course known, see in particular [37–39].

The relevant integrals are of the form

$$J_{\mu_1 \dots \mu_M}(d | p, q) = \int \frac{d^d k}{\pi^{d/2}} \frac{k_{\mu_1} \cdots k_{\mu_M}}{k^2 (k-p)^2 (k+q)^2}, \quad (65)$$

or more generally

$$J_{\mu_1 \dots \mu_M}(d; m_1, m_2, m_3 | p, q) \equiv \int \frac{d^d k}{\pi^{d/2}} \frac{k_{\mu_1} \cdots k_{\mu_M}}{k^{2m_1} (k-p)^{2m_2} (k+q)^{2m_3}} \quad (66)$$

with (not necessarily integer) exponents  $m_1, m_2, m_3$ <sup>7</sup>. For the computation of the conformal anomaly we are in particular interested in the pole part of these integrals for  $d \rightarrow 4$ . Note that we normalise the loop integrals (65) and (66) with the factor  $\pi^{-d/2}$ , different from the normalisation adopted in the rest of this paper. This we do only for convenience in order to simplify the subsequent calculations: because

$$\frac{1}{(2\pi)^d} = \frac{1}{(2\sqrt{\pi})^d} \cdot \frac{1}{\pi^{d/2}} \quad (67)$$

we then only need to multiply the final results by  $(2\sqrt{\pi})^{-d}$  to revert to the normalisation conventions used in the rest of this article.

To evaluate the integrals we will follow a method developed by Davydychev [37, 38], whereby the above integrals can be reduced to the basic scalar 3-point loop integral

$$J(d; 1, 1, 1 | p, q) = \int \frac{d^d k}{\pi^{d/2}} \frac{1}{k^2 (k-p)^2 (k+q)^2} \quad (68)$$

which is again a special case of the more general integral

$$J(d; m_1, m_2, m_3 | p, q) = \int \frac{d^d k}{\pi^{d/2}} \frac{1}{k^{2m_1} (k-p)^{2m_2} (k+q)^{2m_3}} \quad (69)$$

and so-called boundary integrals for which one of the exponents  $m_i$  vanishes (see appendix B)

$$I(d; m_1, m_2 | p) \equiv J(d; m_1, m_2, 0 | p) = \int \frac{d^d k}{\pi^{d/2}} \frac{1}{k^{2m_1} (p-k)^{2m_2}} \quad (70)$$

up to explicit factors which are rational functions of the external momenta. The final result will be completely explicit because for (70) we have the explicit formula

$$\int \frac{d^d k}{\pi^{d/2}} \frac{1}{k^{2m_1} (p-k)^{2m_2}} = i(p^2)^{d/2-m} \frac{\Gamma(m - \frac{d}{2}) \Gamma(\frac{d}{2} - m_1) \Gamma(\frac{d}{2} - m_2)}{\Gamma(d-m) \Gamma(m_1) \Gamma(m_2)} \quad (71)$$

where  $m \equiv m_1 + m_2$  and the factor of  $i$  comes from Wick rotating from Lorentzian space to Euclidean signature. A further advantage of our choice is the simple normalisation

<sup>7</sup>In the remainder we will usually not write out all arguments displayed on the lhs of (66).

$$I(p) \equiv I(d; 1, 1|p) = \frac{i}{\epsilon} + \mathcal{O}(1). \quad (72)$$

As we said, our derivation relies largely on the general formalism developed in [37, 38] but we will spell out the formulae given there in more detail for the cases of interest. The final result will thus express (65) directly in terms of explicitly known functions, where all the UV divergences (needed for the determination of the conformal anomaly) are contained in the boundary integrals. The extension of our results to higher  $n$ -point scalar loop integrals is straightforward, though increasingly tedious for higher values of  $n$ .

In the remainder we will assume the external momenta  $p$  and  $q$  to assume generic values, for which  $p^2 q^2 \neq (p \cdot q)^2$ , so as to avoid IR or kinematical singularities—the latter can then be easily and explicitly extracted from our final expressions. First we note that, in the Feynman parametrisation, the scalar integral (69) is given by

$$J(d; m_1, m_2, m_3) = \frac{i \Gamma(m - \frac{d}{2})}{\Gamma(m_1) \Gamma(m_2) \Gamma(m_3)} \int_0^1 d\xi_1 d\xi_2 d\xi_3 \frac{\xi_1^{m_1-1} \xi_2^{m_2-1} \xi_3^{m_3-1} \delta(1 - \xi_1 - \xi_2 - \xi_3)}{[\xi_1 \xi_2 p^2 + \xi_1 \xi_3 q^2 + \xi_2 \xi_3 (p + q)^2]^{m-d/2}}, \quad (73)$$

where  $m \equiv m_1 + m_2 + m_3$ . Differentiating the lhs of (69) with respect to  $p_\mu$  we find

$$2m_2 \left( J_\mu(d; m_1, m_2 + 1, m_3) - p_\mu J(d; m_1, m_2 + 1, m_3) \right). \quad (74)$$

On the other hand, differentiating the rhs of equation (73) gives

$$-2m_2 \left( m_1 p_\mu J(d + 2; m_1 + 1, m_2 + 1, m_3) + m_3 (p_\mu + q_\mu) J(d + 2; m_1, m_2 + 1, m_3 + 1) \right). \quad (75)$$

Equating expressions (74) and (75), we obtain an equation for  $J_\mu(d; m_1, m_2, m_3)$  in terms of scalar integrals [37]. We can further simplify this expression by noting the identity

$$J(d; \{m_i\}) = \sum_{j=1}^3 m_j J(d + 2; \{m_i + \delta_{ij}\}), \quad (76)$$

which can be proved directly from (73). Using the above identity and equating expressions (74) and (75), we obtain [37]

$$J_\mu(d; m_1, m_2, m_3) = m_2 p_\mu J^{(3)}(d + 2; m_1, m_2 + 1, m_3) - m_3 q_\mu J^{(3)}(d + 2; m_1, m_2, m_3 + 1). \quad (77)$$

This method can be inductively implemented, by further differentiating with respect to  $p_\mu$ , to find similar identities for  $J_{\mu_1 \dots \mu_M}$  in terms of scalar integrals, see [38] for the general formulae. We list the relevant identities for  $M$  up to  $M = 6$ , found using the method outlined above

$$J_\mu(d; 1, 1, 1) = p_\mu J(d + 2; 1, 2, 1) - q_\mu J(d + 2; 1, 1, 2), \quad (78)$$

$$J_{\mu\nu}(d; 1, 1, 1) = \frac{1}{2} \eta_{\mu\nu} J(d + 2; 1, 1, 1) + 2 \left[ p_\mu p_\nu J(d + 4; 1, 3, 1) - p_{(\mu} q_{\nu)} J(d + 4; 1, 2, 2) + q_\mu q_\nu J(d + 4; 1, 1, 3) \right], \quad (79)$$



$$\begin{aligned}
J_{\mu\nu\rho}(d; 1, 1, 1) &= \frac{3}{2}\eta_{(\mu\nu} [p_\rho] J(d+4; 1, 2, 1) - q_\rho) J(d+4; 1, 1, 2)] \\
&\quad + 6 [p_\mu p_\nu p_\rho J(d+6; 1, 4, 1) - p_{(\mu} p_\nu q_\rho) J(d+6; 1, 3, 2) \\
&\quad\quad + p_{(\mu} q_\nu q_\rho) J(d+6; 1, 2, 3) - q_\mu q_\nu q_\rho J(d+6; 1, 1, 4)], \\
J_{\mu\nu\rho\sigma}(d; 1, 1, 1) &= \frac{3}{4}\eta_{(\mu\nu} \eta_{\rho\sigma)} J(d+4; 1, 1, 1) + 6\eta_{(\mu\nu} [p_\rho p_\sigma] J(d+6; 1, 3, 1) \\
&\quad - p_\rho q_\sigma) J(d+6; 1, 2, 2) + q_\rho q_\sigma) J(d+6; 1, 1, 3)] \\
&\quad + 24 [p_\mu p_\nu p_\rho p_\sigma J(d+8; 1, 5, 1) - p_{(\mu} p_\nu p_\rho q_\sigma) J(d+8; 1, 4, 2) \\
&\quad\quad + p_{(\mu} p_\nu q_\rho q_\sigma) J(d+8; 1, 3, 3) - p_{(\mu} q_\nu q_\rho q_\sigma) J(d+8; 1, 2, 4) \\
&\quad\quad + q_\mu q_\nu q_\rho q_\sigma J(d+8; 1, 1, 5)],
\end{aligned} \tag{80}$$

$$\begin{aligned}
J_{\mu\nu\rho\sigma\alpha}(d; 1, 1, 1) &= \frac{15}{4}\eta_{(\mu\nu} \eta_{\rho\sigma} [p_\alpha] J(d+6; 1, 2, 1) - q_\alpha) J(d+6; 1, 1, 2)] \\
&\quad + 30\eta_{(\mu\nu} [p_\rho p_\sigma p_\alpha] J(d+8; 1, 4, 1) - p_\rho p_\sigma q_\alpha) J(d+8; 1, 3, 2) \\
&\quad\quad + p_\rho q_\sigma q_\alpha) J(d+8; 1, 2, 3) - q_\rho q_\sigma q_\alpha) J(d+8; 1, 1, 4)] \\
&\quad + 120 [p_\mu p_\nu p_\rho p_\sigma p_\alpha J(d+10; 1, 6, 1) - p_{(\mu} p_\nu p_\rho p_\sigma q_\alpha) J(d+10; 1, 5, 2) \\
&\quad\quad + p_{(\mu} p_\nu p_\rho q_\sigma q_\alpha) J(d+10; 1, 4, 3) - p_{(\mu} p_\nu q_\rho q_\sigma q_\alpha) J(d+10; 1, 3, 4) \\
&\quad\quad + p_{(\mu} q_\nu q_\rho q_\sigma q_\alpha) J(d+10; 1, 2, 5) - q_\mu q_\nu q_\rho q_\sigma q_\alpha J(d+10; 1, 1, 6)],
\end{aligned} \tag{81}$$

$$\begin{aligned}
J_{\mu\nu\rho\sigma\alpha\beta}(d; 1, 1, 1) &= \frac{15}{8}\eta_{(\mu\nu} \eta_{\rho\sigma} \eta_{\alpha\beta)} J(d+6; 1, 1, 1) + \frac{45}{2}\eta_{(\mu\nu} \eta_{\rho\sigma} [p_\alpha p_\beta] J(d+8; 1, 3, 1) \\
&\quad - p_\alpha q_\beta) J(d+8; 1, 2, 2) + q_\alpha q_\beta) J(d+8; 1, 1, 3)] \\
&\quad + 180\eta_{(\mu\nu} [p_\rho p_\sigma p_\alpha p_\beta] J(d+10; 1, 5, 1) - p_\rho p_\sigma p_\alpha q_\beta) J(d+10; 1, 4, 2) \\
&\quad\quad + p_\rho p_\sigma q_\alpha q_\beta) J(d+10; 1, 3, 3) - p_\rho q_\sigma q_\alpha q_\beta) J(d+10; 1, 2, 4) \\
&\quad\quad + q_\rho q_\sigma q_\alpha q_\beta) J(d+10; 1, 1, 5)] + 720 [p_\mu p_\nu p_\rho p_\sigma p_\alpha p_\beta J(d+12; 1, 7, 1) \\
&\quad\quad - p_{(\mu} p_\nu p_\rho p_\sigma p_\alpha q_\beta) J(d+12; 1, 6, 2) + p_{(\mu} p_\nu p_\rho p_\sigma q_\alpha q_\beta) J(d+12; 1, 5, 3) \\
&\quad\quad - p_{(\mu} p_\nu p_\rho q_\sigma q_\alpha q_\beta) J(d+12; 1, 4, 4) + p_{(\mu} p_\nu q_\rho q_\sigma q_\alpha q_\beta) J(d+12; 1, 3, 5) \\
&\quad\quad - p_{(\mu} q_\nu p_\rho q_\sigma q_\alpha q_\beta) J(d+12; 1, 2, 6) + q_\mu q_\nu q_\rho q_\sigma q_\alpha q_\beta J(d+12; 1, 1, 7)],
\end{aligned} \tag{82}$$

where the integrals on the lhs are all in  $d$  dimensions, whereas the dimension varies on the rhs. Here, as elsewhere in this paper, all symmetrisations are with strength one. The scalar integrals on the rhs are now of type (69), but they still involve different dimensions  $D = d, d + 2, \dots$  and different exponents  $m_1, m_2, m_3$ . To further simplify the above expressions we exploit the basic result [37, 38] that for integer  $m_i$  all integrals of the form (69) can be expressed in terms of  $J(d; 1, 1, 1)$  and boundary integrals of the type (70) and (71).

The first part in this reduction procedure is to decrease the values  $m_1, m_2, m_3$  in integer steps while leaving the dimension unchanged; this is done by noting that

$$\int \frac{d^d k}{\pi^{d/2}} \frac{\partial}{\partial k_\mu} \left\{ \frac{k_\mu}{k^{2m_1}(k-p)^{2m_2}(k+q)^{2m_3}} \right\} = 0, \quad (83)$$

which gives a relation between  $J(D; m_1, m_2, m_3)$  with  $\sum m_i = m$  and  $J(D; m_1, m_2, m_3)$  with  $\sum m_i = m - 1$ . Two more relations can be found by changing the numerator in the integrand in (83) to  $k_\mu - p_\mu$  and  $k_\mu + q_\mu$ . These three equations can be solved [38] to obtain

$$\begin{aligned} J(m_1, m_2, m_3 + 1) = & \frac{1}{2m_3(p+q)^2 q^2} \left( (2m_1 + m_2 + m_3 - d)(p+q)^2 \right. \\ & + (m_1 + 2m_2 + m_3 - d)q^2 - (m_1 + m_2 + 2m_3 - d)p^2 \Big) J(m_1, m_2, m_3) \\ & + m_2(p+q)^2 J(m_1 - 1, m_2 + 1, m_3) + m_3(p+q)^2 J(m_1 - 1, m_2, m_3 + 1) \\ & + m_1 q^2 J(m_1 + 1, m_2 - 1, m_3) + m_3 q^2 J(m_1, m_2 - 1, m_3 + 1) \\ & \left. - m_1 p^2 J(m_1 + 1, m_2, m_3 - 1) - m_2 p^2 J(m_1, m_2 + 1, m_3 - 1) \right), \quad (84) \end{aligned}$$

$$\begin{aligned} J(m_1 + 1, m_2, m_3) = & \frac{1}{2m_1 p^2 q^2} \left( (m_1 + 2m_2 + m_3 - d)q^2 + (m_1 + m_2 + 2m_3 - d)p^2 \right. \\ & \left. - (2m_1 + m_2 + m_3 - d)(p+q)^2 \right) J(m_1, m_2, m_3) \\ & + m_2 p^2 J(m_1, m_2 + 1, m_3 - 1) + m_1 p^2 J(m_1 + 1, m_2, m_3 - 1) \\ & + m_3 q^2 J(m_1, m_2 - 1, m_3 + 1) + m_1 q^2 J(m_1 + 1, m_2 - 1, m_3) \\ & - m_2(p+q)^2 J(m_1 - 1, m_2 + 1, m_3) \\ & \left. - m_3(p+q)^2 J(m_1 - 1, m_2, m_3 + 1) \right), \quad (85) \end{aligned}$$

$$\begin{aligned}
 J(m_1, m_2 + 1, m_3) = & \frac{1}{2m_2(p+q)^2p^2} \left( (2m_1 + m_2 + m_3 - d)(p+q)^2 \right. \\
 & + (m_1 + m_2 + 2m_3 - d)p^2 - (m_1 + 2m_2 + m_3 - d)q^2 \Big) J(m_1, m_2, m_3) \\
 & + m_3(p+q)^2 J(m_1 - 1, m_2, m_3 + 1) + m_2(p+q)^2 J(m_1 - 1, m_2 + 1, m_3) \\
 & + m_1p^2 J(m_1 + 1, m_2, m_3 - 1) + m_2p^2 J(m_1, m_2 + 1, m_3 - 1) \\
 & \left. - m_1q^2 J(m_1 + 1, m_2 - 1, m_3) - m_3q^2 J(m_1, m_2 - 1, m_3 + 1) \right).
 \end{aligned} \tag{86}$$

We repeat that the dimension  $D$  is the *same* in all these integrals, whence

$$J(m_1, m_2, m_3) \equiv J(d; m_1, m_2, m_3)$$

with the same  $d$  in the three equations above.

Having reduced the integrals  $J_{\mu_1 \dots \mu_M}$  of the form given in equation (65) to a sum of scalar integrals  $J(D; 1, 1, 1)$  and boundary integrals, where  $D = d, d + 2, \dots, d + 2M$  we next require a further identity which lowers the values of  $D$  by relating  $J(D + 2; 1, 1, 1)$  to  $J(D; 1, 1, 1)$  so that finally all integrals can be reduced to  $J(d; 1, 1, 1)$ , where now  $d = 4 - 2\epsilon$ . The relevant identity is found by contracting the indices in equation (79), whereupon the lhs of (79) reduces to a boundary integral, and we get

$$\begin{aligned}
 J(d; 0, 1, 1) = & \frac{d}{2} J(d + 2; 1, 1, 1) + 2 \left[ p^2 J(d + 4; 1, 3, 1) \right. \\
 & \left. - (p \cdot q) J(d + 4; 1, 2, 2) + q^2 J(d + 4; 1, 1, 3) \right].
 \end{aligned} \tag{87}$$

Then, using the reduction formulae (84)–(86), we express  $J(d + 4; 1, 3, 1), J(d + 4; 1, 2, 2)$  and  $J(d + 4; 1, 1, 3)$  in terms of  $J(d + 4; 1, 1, 1)$  and boundary integrals. Substituting, these expression in the equation above, and replacing  $d \rightarrow d - 2$ , we obtain

$$\begin{aligned}
 & 2(d - 2) ((p \cdot q)^2 - p^2q^2) J(d + 2; 1, 1, 1) - p^2q^2(p + q)^2 J(d; 1, 1, 1) \\
 = & (p \cdot q)(p + q)^2 J(d; 0, 1, 1) - p^2((p \cdot q) + q^2) J(d; 1, 1, 0) - q^2(p^2 + (p \cdot q)) J(d; 1, 0, 1),
 \end{aligned} \tag{88}$$

or

$$\begin{aligned}
 J(d + 2; 1, 1, 1 | p, q) = & \frac{1}{2(d - 2)} \frac{1}{(p \cdot q)^2 - p^2q^2} \left[ p^2q^2(p + q)^2 J(d; 1, 1, 1 | p, q) + \right. \\
 & \left. + (p \cdot q)(p + q)^2 I(d | p + q) - p^2((p \cdot q) + q^2) I(d | p) - q^2(p^2 + (p \cdot q)) I(d | q) \right]
 \end{aligned} \tag{89}$$

where we have now substituted the boundary integrals from the appendix B. This formula seems to be a new result: it allows us to reduce any given integral of type (65) for even  $d$  to sums involving the convergent integral  $J(4; 1, 1, 1)$  and various boundary integrals which contain all the UV divergences as  $d \rightarrow 4$ . Using the formula (B.5) from appendix B, the latter can be exhibited explicitly:

$$J(d+2; 1, 1, 1) = -2i \frac{d-1}{d-2} \cdot \frac{\Gamma(1-\frac{d}{2})\Gamma(\frac{d}{2})^2}{\Gamma(d)} + \text{finite terms.} \quad (90)$$

The factor  $(p^2q^2 - (p \cdot q)^2)^{-1}$  is cancelled, whence the UV divergence does not depend on the external momenta, as expected. Furthermore, formula (89) makes the kinematical singularities completely explicit. We note that the above formulae, (89) and (90) cannot be used for  $d=2$  because of IR singularities and the factor  $(d-2)^{-1}$ .

Using equations (84)–(86) and (89), the integrals  $J_{\mu_1 \dots \mu_M}$  can thus be reduced to boundary integrals and  $J(d, 1, 1, 1)$ . Since  $J(4, 1, 1, 1)$  is finite, the  $1/\epsilon$  poles in  $J_{\mu_1 \dots \mu_M}$  can easily be found by expanding  $d = 4 - 2\epsilon$  and using the result for the poles of the boundary integral, (B.5). The  $1/\epsilon$  expansion of  $J_{\mu_1 \dots \mu_M}$  up to  $M=6$  is:

$$J_\mu(d|p, q) = O(1), \quad (91)$$

$$J_{\mu\nu}(d|p, q) = \frac{i}{4\epsilon} \eta_{\mu\nu} + O(1), \quad (92)$$

$$J_{\mu\nu\rho}(d|p, q) = \frac{i}{4\epsilon} \eta_{(\mu\nu} (p_\rho) - q_\rho) + O(1), \quad (93)$$

$$J_{\mu\nu\rho\sigma}(d|p, q) = -\frac{i}{32\epsilon} (p^2 + q^2 + (p+q)^2) \eta_{(\mu\nu} \eta_{\rho\sigma)} + \frac{i}{4\epsilon} \eta_{(\mu\nu} (p_\rho p_\sigma) - p_\rho q_\sigma + q_\rho q_\sigma) + O(1), \quad (94)$$

$$\begin{aligned} J_{\mu\nu\rho\sigma\alpha}(d|p, q) &= -\frac{i}{32\epsilon} \eta_{(\mu\nu} \eta_{\rho\sigma} [(2p^2 + q^2 + 2(p+q)^2) p_\alpha] - (p^2 + 2q^2 + 2(p+q)^2) q_\alpha] \\ &\quad + \frac{i}{4\epsilon} \eta_{(\mu\nu} (p_\rho p_\sigma p_\alpha) - p_\rho p_\sigma q_\alpha + p_\rho q_\sigma q_\alpha - q_\rho q_\sigma q_\alpha) + O(1), \end{aligned} \quad (95)$$

$$\begin{aligned} J_{\mu\nu\rho\sigma\alpha\beta}(d|p, q) &= \frac{i}{192\epsilon} \left[ (p^2 + q^2 + (p+q)^2)^2 - (p+q)^2(p^2 + q^2) - p^2q^2 \right] \eta_{(\mu\nu} \eta_{\rho\sigma} \eta_{\alpha\beta)} \\ &\quad - \frac{i}{32\epsilon} \eta_{(\mu\nu} \eta_{\rho\sigma} [(3p^2 + q^2 + 3(p+q)^2) p_\alpha p_\beta] - 2(p^2 + q^2 + 2(p+q)^2) p_\alpha q_\beta \\ &\quad + (p^2 + 3q^2 + 3(p+q)^2) q_\alpha q_\beta] + \frac{i}{4\epsilon} \eta_{(\mu\nu} (p_\rho p_\sigma p_\alpha p_\beta) - p_\rho p_\sigma p_\alpha q_\beta \\ &\quad + p_\rho p_\sigma q_\alpha q_\beta - p_\rho q_\sigma q_\alpha q_\beta + q_\rho q_\sigma q_\alpha q_\beta) + O(1). \end{aligned} \quad (96)$$

These coefficients and polynomials in the external momenta are what we need for the evaluation of the conformal anomaly.

If one is just interested in the divergent parts, this result can also be arrived at without invoking the full machinery of  $n$ -point loop integrals and in a much simpler way as follows: first of all, one notes that the divergence must be polynomial in the external momenta  $p$  and  $q$ . Secondly the resulting polynomial must be symmetric under interchange of  $p$  and  $-q$ . Thirdly, by shifting the integration variable as  $k \rightarrow -k + p$  one obtains a relation constraining the polynomials by replacing the external momenta  $(p, q)$  by  $(p, -p - q)$ . When applying this trick to the above integrals, one first notes that the integrals  $J$  and  $J_\mu$  are convergent, whence the first divergence arises in  $J_{\mu\nu}$ ; the latter divergence is proportional to  $\eta_{\mu\nu}$  and can thus be extracted by contracting with  $\eta_{\mu\nu}$ , thereby cancelling one propagator and reducing the

determination of the pole term to that of a 2-point integral. Likewise the divergence in  $J_{\mu\nu\rho}$  can only appear in the term linear in the external momenta, which by symmetry must appear in the combination  $(p - q)_\mu$ ; again the result can be read off from the corresponding 2-point integral after contraction, and so on for the integrals with more momenta in the numerator.

It is easy to see that this procedure can also be applied inductively to  $n$ -point integrals for  $n > 3$  by successively reducing them to  $(n - 1)$ -loop integrals, etc. In other words, the determination of the pole parts at any order in  $h_{\mu\nu}$  does not require the actual evaluation of  $n$ -point integrals. However, this shortcut may no longer be available for classically non-conformal theories where there could arise extra non-local contributions.

## 5. The conformal anomaly at $\mathcal{O}(\hbar^2)$

The anomaly is given by the trace of  $\langle T_{\mu\nu}(x) \rangle$  after regularisation. If we calculate the trace before finding the regularised expression, the trace vanishes by the Ward identities as the Dirac action is scale-invariant in all dimensions. At second order in the external graviton, the anomaly is given by

$$\begin{aligned} g^{\mu\nu} \langle T_{\mu\nu}(x) \rangle |_{\mathcal{O}(\hbar^2)} &= \eta^{\mu\nu} \langle T_{\mu\nu} \rangle |_{\mathcal{O}(\hbar^2)} - h^{\mu\nu}(x) \langle T_{\mu\nu} \rangle |_{\mathcal{O}(\hbar)} \\ &= \int d^d y d^d z \int \frac{d^d p}{(2\pi)^d} \frac{d^d q}{(2\pi)^d} e^{-ip \cdot (x-y) - iq \cdot (z-y)} h^{\rho\sigma}(y) \\ &\quad \times \left\{ \eta^{\mu\nu} \left( T_{\mu\nu\rho\sigma\alpha\beta}^{(1)}(p, q) h^{\alpha\beta}(z) + T_{\mu\nu\rho\sigma\alpha\beta}^{(2)}(p) h^{\alpha\beta}(z) + T_{\mu\nu\rho\sigma\alpha\beta}^{(3)}(p) h^{\alpha\beta}(x) \right) \right. \\ &\quad \left. - T_{\mu\nu\rho\sigma}(p) h^{\mu\nu}(x) \right\}. \end{aligned} \quad (97)$$

Using equations (43), (47) and (48), and rewriting

$$\begin{aligned} \int d^d y \int \frac{d^d p}{(2\pi)^d} e^{-ip \cdot (x-y)} h^{\mu\nu}(x) h^{\rho\sigma}(y) T_{\mu\nu\rho\sigma}(p) \\ = \int d^d y d^d z \int \frac{d^d p}{(2\pi)^d} \frac{d^d q}{(2\pi)^d} e^{-ip \cdot (x-y) - iq \cdot (z-y)} h^{\rho\sigma}(y) h^{\mu\nu}(z) T_{\mu\nu\rho\sigma}(p + q), \end{aligned}$$

$$\begin{aligned} g^{\mu\nu} \langle T_{\mu\nu}(x) \rangle &= \frac{1}{4} \eta^{\mu\nu} \int d^d y d^d z \int \frac{d^d p}{(2\pi)^d} \frac{d^d q}{(2\pi)^d} e^{-ip \cdot (x-y) - iq \cdot (z-y)} h^{\rho\sigma}(y) h^{\alpha\beta}(z) \\ &\quad \times \left( \frac{1}{16} T_{\mu\nu\rho\sigma\alpha\beta} + \eta_{\mu\nu} T_{\rho\sigma\alpha\beta}(p + q) - \eta_{\mu\nu} T_{\rho\sigma\alpha\beta}(q) - \eta_{\rho\sigma} T_{\mu\nu\alpha\beta}(q) \right. \\ &\quad \left. - \eta_{\alpha\beta} T_{\mu\nu\rho\sigma}(p + q) - 3 \eta_{\alpha\rho} T_{\mu\nu\sigma\beta}(p) - 2 \eta_{\alpha\mu} T_{\nu\beta\rho\sigma}(p + q) \right). \end{aligned}$$

By redefining the integration variables  $p$  and  $q$ , we can write the above expression in a more symmetric way,

$$\begin{aligned}
g^{\mu\nu}\langle T_{\mu\nu}(x)\rangle &= \frac{1}{4}\eta^{\mu\nu}\int d^d y d^d z \int \frac{d^d p}{(2\pi)^d} \frac{d^d q}{(2\pi)^d} e^{-ip\cdot(x-y)-iq\cdot(x-z)} h^{\rho\sigma}(y) h^{\alpha\beta}(z) \\
&\quad \times \left( \frac{1}{16} \hat{T}_{\mu\nu\rho\sigma\alpha\beta} - \eta_{\alpha\mu} T_{\rho\sigma\nu\beta}(p) - \eta_{\rho\mu} T_{\nu\sigma\alpha\beta}(q) \right. \\
&\quad \left. - 2\eta_{\alpha\beta} T_{\mu\nu\rho\sigma}(p) - 3\eta_{\alpha\rho} T_{\mu\nu\sigma\beta}(p+q) \right), \tag{98}
\end{aligned}$$

where we have exploited the symmetry under  $p \leftrightarrow q$  and  $\rho\sigma \leftrightarrow \alpha\beta$ , to simplify the integrand and where  $\hat{T}_{\mu\nu\rho\sigma\alpha\beta}$  is the expression by letting  $p \rightarrow p+q$  and  $q \rightarrow -q$ , viz

$$\hat{T}_{\mu\nu\rho\sigma\alpha\beta} \equiv i \int \frac{d^d k}{(2\pi)^d} \text{tr} \left\{ \frac{\not{k}}{k^2} (2k+q)_{(\alpha\gamma\beta)} \frac{\not{k}+\not{q}}{(k+q)^2} (2k-p+q)_{(\mu\gamma\nu)} \frac{\not{k}-\not{p}}{(k-p)^2} (2k-p)_{(\rho\gamma\sigma)} \right\}. \tag{99}$$

The trace of the expectation in  $d$ -dimensions should be zero as the Dirac action is Weyl-invariant in all dimensions. The anomaly appears because the expectation value of the regularised 4-dimensional stress tensor is evaluated in  $d = 4 - 2\epsilon$  dimensions, which gives rise to a non-zero 4-dimensional trace. As a consistency check we show that the  $d$ -dimensional trace of the expression of the rhs of equation (98) vanishes.

First consider,

$$\begin{aligned}
\eta^{(d)\mu\nu} \hat{T}_{\mu\nu\rho\sigma\alpha\beta} &= i \int \frac{d^d k}{(2\pi)^d} \text{tr} \left\{ \frac{\not{k}}{k^2} (2k-p)_{(\rho\gamma\sigma)} \frac{(\not{k}-\not{p})}{(k-p)^2} (2k+q)_{(\alpha\gamma\beta)} \right. \\
&\quad \left. + \frac{\not{k}}{k^2} (2k+p)_{(\rho\gamma\sigma)} \frac{(\not{k}-\not{q})}{(k-q)^2} (2k-q)_{(\alpha\gamma\beta)} \right\}, \\
&= 16(T_{\rho\sigma\alpha\beta}(p) + T_{\rho\sigma\alpha\beta}(q)), \tag{100}
\end{aligned}$$

where we have used an identity similar to identity (27) and reparametrised the variable of integration  $k$  in the first equality, and equation (C.5) and the definition of the two-point function integral, (26), in the second equality. Therefore, substituting into equation (98) and using the Weyl-invariance of the 2-point function in  $d$ -dimensions, (32),

$$\langle g^{\mu\nu} T_{\mu\nu}(x) \rangle|_{d-\text{dim}} = 0. \tag{101}$$

Therefore, from equation (98), the anomaly at second order in  $h$  is given by

$$\begin{aligned}
\mathcal{A}(x)|_{\mathcal{O}(h^2)} &= \frac{1}{4} \int d^d y d^d z \int \frac{d^d p}{(2\pi)^d} \frac{d^d q}{(2\pi)^d} e^{-ip\cdot(x-y)-iq\cdot(z-y)} h_{\rho\sigma}(y) h_{\alpha\beta}(z) \\
&\quad \times \left( \frac{1}{16} \eta^{(4)\mu\nu} \hat{T}_{\mu\nu\rho\sigma\alpha\beta} - T_{\rho\sigma\alpha\beta}(p) - T_{\rho\sigma\alpha\beta}(q) \right. \\
&\quad \left. - 2\eta_{\alpha\beta} \eta^{(4)\mu\nu} T_{\mu\nu\rho\sigma}(p) - 3\eta_{\rho\alpha} \eta^{(4)\mu\nu} T_{\mu\nu\sigma\beta}(p+q) \right), \tag{102}
\end{aligned}$$

where  $\eta^{(4)\mu\nu}\hat{T}_{\mu\nu\alpha\beta\rho\sigma}$  is the 4-dimensional trace of the regularised 3-point function in momentum space.

The 4-dimensional trace of the regularised 2-point function integral is already known and given in equation (35). Therefore, it remains to consider the terms on the second line of the rhs of equation (102). We write,

$$\frac{1}{16}\hat{T}_{\mu\nu\rho\sigma\alpha\beta} - \eta_{\alpha\mu}T_{\rho\sigma\nu\beta}(p) - \eta_{\rho\mu}T_{\nu\sigma\alpha\beta}(q) = \frac{A_{\mu\nu\rho\sigma\alpha\beta}(p,q)}{2\epsilon} + B_{\mu\nu\rho\sigma\alpha\beta}(p,q) + \mathcal{O}(\epsilon). \quad (103)$$

The terms on the lhs are regularised integrals in  $d$ -dimensions and we denote the pole terms in the expression by  $A_{\mu\nu\rho\sigma\alpha\beta}$  and the finite terms by  $B_{\mu\nu\rho\sigma\alpha\beta}$ . We are interested in the 4-dimensional trace of the expression on the lhs, which gives the terms on the second line of the rhs of equation (102). Namely, we are interested in

$$\eta^{(4)\mu\nu}B_{\mu\nu\rho\sigma\alpha\beta} = \frac{1}{16}\eta^{(4)\mu\nu}\hat{T}_{\mu\nu\rho\sigma\alpha\beta} - T_{\rho\sigma\alpha\beta}(p) - T_{\rho\sigma\alpha\beta}(q). \quad (104)$$

Note that the 4-dimensional trace of  $A_{\mu\nu\rho\sigma\alpha\beta}$  necessarily vanishes, since the anomaly is finite.

The tensor  $A_{\mu\nu\rho\sigma\alpha\beta}$  is local in the momenta  $p$  and  $q$  and can be found using equations (91)–(96). Meanwhile, the tensor  $B_{\mu\nu\rho\sigma\alpha\beta}$  is given by the terms labelled  $\mathcal{O}(1)$  in equations (91)–(96) and is in general non-local in the momenta. The 4-dimensional trace of  $B_{\mu\nu\rho\sigma\alpha\beta}$  can nevertheless be found from  $A_{\mu\nu\rho\sigma\alpha\beta}$  by taking a trace in  $D$  dimensions, where  $D$  is arbitrary (but remember that the 2-point and 3-point functions above are computed in  $d$ -dimensions, so  $D$  is just an auxiliary variable here).

From equation (100), we know that

$$\eta^{(D)\mu\nu}\left(\frac{1}{16}\hat{T}_{\mu\nu\rho\sigma\alpha\beta} - \eta_{\alpha\mu}T_{\rho\sigma\nu\beta}(p) - \eta_{\rho\mu}T_{\nu\sigma\alpha\beta}(q)\right) = (D-d)\left(\frac{C_{\rho\sigma\alpha\beta}}{2\epsilon} + D_{\rho\sigma\alpha\beta} + \mathcal{O}(\epsilon)\right), \quad (105)$$

where  $C_{\rho\sigma\alpha\beta}$  and  $D_{\rho\sigma\alpha\beta}$  are tensorial functions of the momenta. Substituting equation (103) on the lhs of equation (105) and expanding the rhs in  $\epsilon$ , we find

$$\eta^{(D)\mu\nu}A_{\mu\nu\rho\sigma\alpha\beta} = (D-4)C_{\rho\sigma\alpha\beta}, \quad (106)$$

$$\eta^{(D)\mu\nu}B_{\mu\nu\rho\sigma\alpha\beta} = C_{\rho\sigma\alpha\beta} + (D-4)D_{\rho\sigma\alpha\beta}, \quad (107)$$

at order  $1/\epsilon$  and order 1. Letting  $D = 4$  in equation (107), and using (104), we find that

$$C_{\rho\sigma\alpha\beta} = \frac{1}{16}\eta^{(4)\mu\nu}\hat{T}_{\mu\nu\rho\sigma\alpha\beta} - T_{\rho\sigma\alpha\beta}(p) - T_{\rho\sigma\alpha\beta}(q). \quad (108)$$

However,  $C_{\rho\sigma\alpha\beta}$  can also be found by taking the  $D$ -dimensional trace of  $A_{\mu\nu\rho\sigma\alpha\beta}$ , (106).

After a lengthy calculation (that involves collecting several hundred terms!) we determine  $A_{\mu\nu\rho\sigma\alpha\beta}$ , defined in equation (103), and identify  $C_{\rho\sigma\alpha\beta}$  by taking an arbitrary  $D$ -dimensional trace, (106). This gives, (108), an expression for the terms on the second line of the rhs of equation (102) and, as we have already mentioned, the other terms on the rhs of equation (102) are given by equation (35). The final result is

$$\begin{aligned}
 \mathcal{A}|_{\mathcal{O}(h^2)} = & -\frac{1}{720(4\pi)^2} \int d^4y d^4z \int \frac{d^4p}{(2\pi)^4} \frac{d^4q}{(2\pi)^4} e^{-ip \cdot (x-y) - iq \cdot (x-z)} h^{\rho\sigma}(y) h^{\alpha\beta}(z) \\
 & \times \left\{ \begin{aligned}
 & (4(p \cdot q)^2 + p^2(12p \cdot q + 5q^2)) \eta_{\rho\sigma} \eta_{\alpha\beta} \\
 & - (12(p^2)^2 + 25(p \cdot q)^2 + 14p^2(3p \cdot q + q^2)) \eta_{\alpha(\rho} \eta_{\sigma)\beta} \\
 & - 8(3p^2 + 3p \cdot q + 2q^2) \eta_{\rho\sigma} p_{\alpha} p_{\beta} - 2(3p \cdot q + 5q^2) \eta_{\alpha\beta} p_{\rho} p_{\sigma} \\
 & - 4(3p^2 + 4p \cdot q + 6q^2) \eta_{\alpha\beta} p_{(\rho} q_{\sigma)} + 8(3p^2 + 6p \cdot q + 4q^2) p_{(\rho} \eta_{\sigma)(\alpha} p_{\beta)} \\
 & + 4(6p^2 + 5p \cdot q) p_{(\rho} \eta_{\sigma)(\alpha} q_{\beta)} + 2(6p^2 + 13p \cdot q) q_{(\rho} \eta_{\sigma)(\alpha} p_{\beta)} \\
 & + 12p_{\rho} p_{\sigma} p_{\alpha} p_{\beta} + 12p_{\rho} p_{\sigma} p_{(\alpha} q_{\beta)} + 5p_{\rho} p_{\sigma} q_{\alpha} q_{\beta} - 4p_{(\rho} q_{\sigma)} p_{(\alpha} q_{\beta)} - 7q_{\rho} q_{\sigma} p_{\alpha} p_{\beta} \end{aligned} \right\}. \tag{109}
 \end{aligned}$$

The expression is, in particular, polynomial in  $p$  and  $q$ —the dependence on inverse powers or logarithms of the external momenta, which are in higher order terms in  $\epsilon$ , has dropped out, hence the anomaly is local in  $x$ -space, as expected.

Note that when comparing with the anomalies the terms quadratic in curvature must all have the structure  $\partial\partial h\partial\partial h$ , which in Fourier space is equivalent to having two  $p$  and two  $q$  in each term, whereas all other terms with a different distribution of derivatives must originate from  $\square R$ . Therefore, we can use the term proportional to  $p_{\rho} p_{\sigma} p_{\alpha} p_{\beta}$  (see equation (A.14)), for example, to fix the coefficient of  $\square R$ ,

$$\mathcal{A}|_{\mathcal{O}(h^2)} = \frac{1}{60(4\pi)^2} \square R|_{\mathcal{O}(h^2)} + \dots \tag{110}$$

Furthermore, from equation (A.11)–(A.13), we note that  $q_{\rho} q_{\sigma} p_{\alpha} p_{\beta}$ ,  $p_{(\rho} q_{\sigma)} p_{(\alpha} q_{\beta)}$ ,  $p_{\rho} p_{\sigma} q_{\alpha} q_{\beta}$  only appear in Riemann-squared, Ricci-squared and scalar-squared, respectively. Hence terms containing these expressions can be used to fix the coefficient of all the terms in the anomaly. Altogether we have thus shown that

$$\begin{aligned}
 \mathcal{A}|_{\mathcal{O}(h^2)} = & \left[ \frac{7}{720(4\pi)^2} \text{Riem}^2 + \frac{1}{90(4\pi)^2} \text{Ric}^2 - \frac{1}{144(4\pi)^2} R^2 + \frac{1}{60(4\pi)^2} \square R \right]_{\mathcal{O}(h^2)} \\
 = & \left[ \frac{1}{40(4\pi)^2} C^{\mu\nu\rho\sigma} C_{\mu\nu\rho\sigma} - \frac{11}{720(4\pi)^2} E_4 + \frac{1}{60(4\pi)^2} \square R \right]_{\mathcal{O}(h^2)}. \tag{111}
 \end{aligned}$$

Note that the coefficient of  $\square R$  at second order in  $h$  matches the coefficient at first order, (36), as it must do for consistency. Furthermore, this explicit calculation confirms the relation (38), and agrees with the values for  $a, b, c$  in the literature.

## 6. Outlook

In this paper we have given a new and direct derivation of the spin- $\frac{1}{2}$  anomaly, along the lines of the textbook derivation of the axial anomaly. Although at this point the calculation merely confirms a known result, our derivation based on standard Feynman diagram techniques has



brought out several subtleties, and we expect similar subtleties for a rederivation of the (again known) results for  $s = 0, 1$ .

However, as we already said in the introduction, the present work should be regarded as only preparatory for what we are really after, namely a proper computation of and a better understanding of the conformal anomaly in *non-conformal* theories, where the anomaly can be *defined* by

$$\mathcal{A} := g^{\mu\nu} \langle T_{\mu\nu} \rangle - \langle g^{\mu\nu} T_{\mu\nu} \rangle \quad (112)$$

and where the second term subtracts the terms due to the classical violation of Weyl invariance. Most significantly we will be interested in the cases  $s = \frac{3}{2}$  and  $s = 2$ , where there remain several issues (dependence of  $a$  and  $c$  coefficients on gauge choices for external gravitons, appearance of  $R^2$  contributions for non-conformal theories, etc) that remain open even after many years. Future directions are thus:

- A computation of conformal anomaly for  $s = \frac{3}{2}$  along the lines of this paper.
- Understanding the appearance of  $R^2$  and possible non-local contributions that may be required to satisfy WZ consistency condition.
- Understanding the dependence of  $a$  and  $c$  coefficients on the choice of gauge for metric fluctuation  $h_{\mu\nu}$ . Such a gauge dependence should not exist, as the anomaly coefficients should be *gauge invariant* with the (natural) assumption of unbroken general covariance.
- Understanding the appearance of negative anomaly coefficients for  $s = \frac{3}{2}$ , which is in apparent conflict with positivity theorems. However, the latter rely on unitarity (positive definite) Hilbert space, and the existence of a gauge invariant stress tensor, whereas both these assumptions are violated for  $s \geq \frac{3}{2}$ .

## Acknowledgments

We would like to thank T Bautista, L Casarin, M Duff, K Meissner and A Schwimmer for discussions related to this work, and L Bonora for correspondence and explanations concerning [28, 29].

## Appendix A. Expansions

In this appendix we collect all necessary formulae for the expansion of curvature-squared quantities to second order in  $h$ :

$$g_{\mu\nu} = \eta_{\mu\nu} + h_{\mu\nu}, \quad g^{\mu\nu} = \eta^{\mu\nu} - h^{\mu\nu} + h^{\mu\rho} h^{\nu\sigma} \eta_{\rho\sigma}, \quad (A.1)$$

$$e_\mu^a = \delta_\mu^a + \frac{1}{2} h_\mu^a - \frac{1}{8} h_{\mu\nu} h^{\nu a}, \quad e^\mu_a = \delta_a^\mu - \frac{1}{2} h_a^\mu + \frac{3}{8} h^{\mu\nu} h_{\nu a}, \quad (A.2)$$

$$e = 1 + \frac{1}{2} h - \frac{1}{4} h^{\mu\nu} h_{\mu\nu} + \frac{1}{8} h^2, \quad e^{-1} = 1 - \frac{1}{2} h + \frac{1}{4} h^{\mu\nu} h_{\mu\nu} + \frac{1}{8} h^2, \quad (A.3)$$

$$\omega_{\mu ab} = \partial_{[b} h_{a]\mu} + \frac{1}{4} h^\nu_{[b} \partial_\mu h_{a]\nu} - \frac{1}{2} h^\nu_{[b} (\partial_\nu h_{a]\mu} - \partial_a h_{\mu\nu}), \quad (A.4)$$

$$\Gamma_{\rho\sigma}^{\mu} = \frac{1}{2}(\eta^{\mu\nu} - h^{\mu\nu})(2\partial_{(\rho}h_{\sigma)\nu} - \partial_{\nu}h_{\rho\sigma}), \quad (\text{A.5})$$

$$\text{Riem}^2 = \partial^{\mu}\partial^{\nu}h^{\rho\sigma}(\partial_{\mu}\partial_{\nu}h_{\rho\sigma} - 2\partial_{\mu}\partial_{\rho}h_{\nu\sigma} + \partial_{\rho}\partial_{\sigma}h_{\mu\nu}), \quad (\text{A.6})$$

$$\begin{aligned} \text{Ric}^2 &= \frac{1}{2}\partial_{\mu}\partial^{\rho}h^{\mu\sigma}(\partial^{\nu}\partial_{\rho}h_{\nu\sigma} - 2\partial^2h_{\rho\sigma} - 2\partial_{\rho}\partial_{\sigma}h + \partial^{\nu}\partial_{\sigma}h_{\nu\rho}) \\ &+ \frac{1}{4}\partial^2h^{\rho\sigma}(\partial^2h_{\rho\sigma} + 2\partial_{\rho}\partial_{\sigma}h) + \frac{1}{4}\partial^{\rho}\partial^{\sigma}h\partial_{\rho}\partial_{\sigma}h, \end{aligned} \quad (\text{A.7})$$

$$R^2 = \partial^{\rho}\partial^{\sigma}h_{\rho\sigma}\partial^{\mu}\partial^{\nu}h_{\mu\nu} - 2\partial_{\rho}\partial_{\sigma}h_{\rho\sigma}\partial^2h + \partial^2h\partial^2h, \quad (\text{A.8})$$

$$\begin{aligned} \square R &= \partial^2\partial^{\rho}\partial^{\sigma}h_{\rho\sigma} - \partial^2\partial^2h \\ &+ h^{\rho\sigma}(2\partial^2\partial_{\rho}\partial_{\sigma}h - 2\partial^2\partial^{\mu}\partial_{\rho}h_{\mu\sigma} + \partial^2\partial^2h_{\rho\sigma} - \partial_{\rho}\partial_{\sigma}\partial^{\mu}\partial^{\nu}h_{\mu\nu}) \\ &+ \partial^{\mu}h^{\rho\sigma}\left(2\partial_{\mu}\partial_{\rho}\partial_{\sigma}h - 4\partial_{\mu}\partial_{\rho}\partial^{\nu}h_{\nu\sigma} + \frac{7}{2}\partial^2\partial_{\mu}h_{\rho\sigma} - \partial^2\partial_{\rho}h_{\sigma\mu}\right) \\ &- \left(\partial_{\mu}h^{\mu\rho} - \frac{1}{2}\partial^{\rho}h\right)(2\partial^2\partial^{\nu}h_{\rho\nu} - 2\partial_{\rho}\partial^2h + \partial_{\rho}\partial^{\nu}\partial^{\sigma}h_{\nu\sigma}) \\ &+ \frac{1}{2}\partial^{\mu}\partial^{\nu}h^{\rho\sigma}(3\partial_{\mu}\partial_{\nu}h_{\rho\sigma} - 2\partial_{\mu}\partial_{\rho}h_{\nu\sigma}) - 2\partial^{\rho}\partial_{\mu}h^{\mu\sigma}(\partial_{\rho}\partial^{\nu}h_{\nu\sigma} - \partial_{\rho}\partial_{\sigma}h + \partial^2h_{\rho\sigma}) \\ &+ \partial^2h^{\rho\sigma}(\partial_{\rho}\partial_{\sigma}h + \partial^2h_{\rho\sigma}) - \frac{1}{2}\partial^{\rho}\partial^{\sigma}h\partial_{\rho}\partial_{\sigma}h, \end{aligned} \quad (\text{A.9})$$

$$\begin{aligned} T_{\mu\nu} &= \bar{\chi}\gamma_{(\mu}D_{\nu)}\chi \\ &= \bar{\chi}\gamma_{(\mu}\partial_{\nu)}\chi + \frac{1}{2}h_{\rho(\mu}\bar{\chi}\gamma^{\rho}\partial_{\nu)}\chi + \frac{1}{4}\partial_{\sigma}h_{\rho(\nu}\bar{\chi}\gamma_{\mu)}\gamma^{\rho\sigma}\chi + \frac{1}{8}h_{\tau(\mu}\partial_{\sigma}h_{\rho|\nu)}\bar{\chi}\gamma^{\tau}\gamma^{\rho\sigma}\chi, \end{aligned} \quad (\text{A.10})$$

where on the rhs all indices are lowered and raised with the Minkowski metric and  $\gamma_{\mu}$  is the flat gamma-matrix.

At second-order in the metric perturbation, the curvature-squared quantities can also be written in momentum space as:

$$\begin{aligned} \text{Riem}^2 &= \int d^4y d^4z \int \frac{d^4p}{(2\pi)^4} \frac{d^4q}{(2\pi)^4} e^{-ip\cdot(x-y) - iq\cdot(x-z)} h^{\rho\sigma}(y) h^{\alpha\beta}(z) \\ &\times \left\{ (p\cdot q)^2 \eta_{\alpha(\rho}\eta_{\sigma)\beta} - 2p\cdot q q_{(\rho}\eta_{\sigma)(\alpha}p_{\beta)} + q_{\rho}q_{\sigma}p_{\alpha}p_{\beta} \right\}, \end{aligned} \quad (\text{A.11})$$

$$\begin{aligned}
\text{Ric}^2 = & \int d^4y d^4z \int \frac{d^4p}{(2\pi)^4} \frac{d^4q}{(2\pi)^4} e^{-ip \cdot (x-y) - iq \cdot (x-z)} h^{\rho\sigma}(y) h^{\alpha\beta}(z) \\
& \times \left\{ \frac{1}{4} (p \cdot q)^2 \eta_{\rho\sigma} \eta_{\alpha\beta} + \frac{1}{4} p^2 q^2 \eta_{\alpha(\rho} \eta_{\sigma)\beta} + \frac{1}{2} q^2 \eta_{\rho\sigma} p_{\alpha} p_{\beta} - p \cdot q \eta_{\alpha\beta} p_{(\rho} q_{\sigma)} \right. \\
& \left. - q^2 p_{(\rho} \eta_{\sigma)(\alpha} p_{\beta)} + \frac{1}{2} p \cdot q p_{(\rho} \eta_{\sigma)(\alpha} q_{\beta)} + \frac{1}{2} p_{(\rho} q_{\sigma)} p_{(\alpha} q_{\beta)} \right\}, \quad (\text{A.12})
\end{aligned}$$

$$\begin{aligned}
R^2 = & \int d^4y d^4z \int \frac{d^4p}{(2\pi)^4} \frac{d^4q}{(2\pi)^4} e^{-ip \cdot (x-y) - iq \cdot (x-z)} h^{\rho\sigma}(y) h^{\alpha\beta}(z) \\
& \times \left\{ p^2 q^2 \eta_{\rho\sigma} \eta_{\alpha\beta} - 2 q^2 \eta_{\alpha\beta} p_{\rho} p_{\sigma} + p_{\rho} p_{\sigma} q_{\alpha} q_{\beta} \right\}, \quad (\text{A.13})
\end{aligned}$$

$$\begin{aligned}
\Box R = & \int d^4y d^4z \int \frac{d^4p}{(2\pi)^4} \frac{d^4q}{(2\pi)^4} e^{-ip \cdot (x-y) - iq \cdot (x-z)} h^{\rho\sigma}(y) h^{\alpha\beta}(z) \\
& \times \left\{ -\frac{1}{2} ((p \cdot q)^2 + 2 p^2 p \cdot q) \eta_{\rho\sigma} \eta_{\alpha\beta} \right. \\
& + \frac{1}{2} (2 (p^2)^2 + 3 (p \cdot q)^2 + 7 p^2 p \cdot q + 2 p^2 q^2) \eta_{\alpha(\rho} \eta_{\sigma)\beta} \\
& + (2 p^2 + 2 p \cdot q + q^2) \eta_{\rho\sigma} p_{\alpha} p_{\beta} + \frac{1}{2} p \cdot q \eta_{\alpha\beta} p_{\rho} p_{\sigma} \\
& + (p^2 + 2 p \cdot q + 2 q^2) \eta_{\alpha\beta} p_{(\rho} q_{\sigma)} - 2 (p^2 + 2 p \cdot q + q^2) p_{(\rho} \eta_{\sigma)(\alpha} p_{\beta)} \\
& - 2 (p^2 + p \cdot q) p_{(\rho} \eta_{\sigma)(\alpha} q_{\beta)} - (p^2 + p \cdot q) q_{(\rho} \eta_{\sigma)(\alpha} p_{\beta)} \\
& \left. - p_{\rho} p_{\sigma} p_{\alpha} p_{\beta} - p_{\rho} p_{\sigma} p_{(\alpha} q_{\beta)} \right\}. \quad (\text{A.14})
\end{aligned}$$

## Appendix B. Boundary integrals

We here present some well-known results for scalar 2-point integrals, see for example [40] which are also referred to as boundary integrals,

$$\int \frac{d^d k}{\pi^{d/2}} \frac{k_{\mu}}{k^2 (p-k)^2} = \frac{1}{2} p_{\mu} I(p), \quad (\text{B.1})$$

$$\int \frac{d^d k}{\pi^{d/2}} \frac{k_{\mu} k_{\nu}}{k^2 (p-k)^2} = \frac{1}{4(d-1)} (d p_{\mu} p_{\nu} - p^2 \eta_{\mu\nu}) I(p), \quad (\text{B.2})$$

$$\int \frac{d^d k}{\pi^{d/2}} \frac{k_\mu k_\nu k_\rho}{k^2(p-k)^2} = \frac{1}{8(d-1)} ((d+2)p_\mu p_\nu p_\rho - 3p^2 \eta_{(\mu\nu} p_{\rho)}) I(p), \quad (\text{B.3})$$

$$\int \frac{d^d k}{\pi^{d/2}} \frac{k_\mu k_\nu k_\rho k_\sigma}{k^2(p-k)^2} = \frac{1}{16(d^2-1)} ((d+4)(d+2)p_\mu p_\nu p_\rho p_\sigma - 6(d+2)p^2 \eta_{(\mu\nu} p_\rho p_{\sigma)} + 3(p^2)^2 \eta_{(\mu\nu} \eta_{\rho\sigma)}) I(p), \quad (\text{B.4})$$

where  $I(p) \equiv I(d|p) \equiv J(d; 1, 1, 0)$ , or more explicitly

$$\begin{aligned} I(p) &= \int \frac{d^d k}{\pi^{d/2}} \frac{1}{k^2(p-k)^2} \\ &= -2i(d-1)(p^2)^{d/2-2} \frac{\Gamma(1-\frac{d}{2})\Gamma(\frac{d}{2})^2}{\Gamma(d)} = \frac{i}{\epsilon} + i \left( 2 + \gamma - \log \frac{p^2}{\mu^2} \right) + \mathcal{O}(\epsilon), \end{aligned} \quad (\text{B.5})$$

where  $\gamma$  is the Euler–Mascheroni constant and  $\mu$  is the renormalisation scale. Note the overall factor of  $i$  which is a result of Wick rotating from Lorentzian signature. The normalisation in the definition of  $I(p)$  (and in (65) and (69)) has thus been chosen in order to eliminate all factors of  $\pi^{d/2}$  in the final expression.

### Appendix C. Useful formulae and identities

$$\text{tr}(\gamma_\mu \gamma_\nu \gamma_\rho \gamma_\sigma) = \text{tr}(\gamma_\mu \gamma_\sigma \gamma_\rho \gamma_\nu). \quad (\text{C.1})$$

$$\text{tr}(\gamma_\mu \gamma_\nu \gamma_\rho \gamma_\sigma \gamma_\alpha \gamma_\beta) = \text{tr}(\gamma_\rho \gamma_\nu \gamma_\mu \gamma_\beta \gamma_\alpha \gamma_\sigma). \quad (\text{C.2})$$

$$\text{tr}(\gamma_\mu \gamma_\nu \gamma_\rho (\gamma_\sigma \gamma_\alpha \gamma_\beta + \gamma_\beta \gamma_\alpha \gamma_\sigma)) = 2 \text{tr}(\gamma_\mu \gamma_\nu \gamma_\rho (\eta_{\sigma\alpha} \gamma_\beta - \eta_{\sigma\beta} \gamma_\alpha + \eta_{\alpha\beta} \gamma_\sigma)). \quad (\text{C.3})$$

$$\text{tr}(\gamma_\mu \gamma_\nu \gamma_\rho (\gamma^\sigma \gamma^\alpha \gamma^\beta - \gamma^\beta \gamma^\alpha \gamma^\sigma)) = -12 \delta_{\mu\nu\rho}^{\sigma\alpha\beta} \text{tr} \mathbf{1}. \quad (\text{C.4})$$

Making use of the redefinition of the integration variable, letting  $k \rightarrow -k + p$ , and the gamma matrix identity (C.1), it can be shown that

$$\int \frac{d^d k}{(2\pi)^d} \text{tr} \left[ \frac{\not{k}}{k^2} \gamma_\alpha \frac{(\not{k} - \not{p})}{(k-p)^2} (2k-p)_{(\rho} \gamma_{\sigma)} \right] = 0. \quad (\text{C.5})$$

Using identity (C.1),

$$\begin{aligned} &\int \frac{d^d k}{(2\pi)^d} \frac{1}{k^2(k-p)^2} \text{tr} \left\{ (2\not{k} - \not{p})(2k-p)_\mu \gamma_\nu (2\not{k} - \not{p})(2k-p)_{(\rho} \gamma_{\sigma)} \right\}, \\ &= -64i T_{\mu\nu\rho\sigma}(p) + \int \frac{d^d k}{(2\pi)^d} \frac{1}{k^2(k-p)^2} \text{tr} \left\{ \not{p} (2k-p)_\mu \gamma_\nu \not{p} (2k-p)_{(\rho} \gamma_{\sigma)} \right\}. \end{aligned} \quad (\text{C.6})$$

Using identities (C.2) and (C.3),

$$\begin{aligned} & \int \frac{d^d k}{(2\pi)^d} \frac{1}{k^2(k-p)^2} \text{tr} \left\{ (2\mathbb{k} - \not{p})(2k-p)_{(\alpha\gamma\beta)} (2\mathbb{k} - \not{p})(2k-p)_{\tau\gamma\rho\gamma\sigma\gamma\nu} \right\}, \\ &= -64i [\eta_{\rho\sigma} T_{\alpha\beta\tau\nu}(p) - \eta_{\rho\nu} T_{\alpha\beta\tau\sigma}(p) + \eta_{\sigma\nu} T_{\alpha\beta\tau\rho}(p)] \\ &+ \int \frac{d^d k}{(2\pi)^d} \frac{1}{k^2(k-p)^2} \text{tr} \left\{ \not{p}(2k-p)_{(\alpha\gamma\beta)} \not{p}(2k-p)_{\tau\gamma\rho\gamma\sigma\gamma\nu} \right\}, \end{aligned} \quad (\text{C.7})$$

where we have also used equation (C.6).

Furthermore, using identities (C.2) and (C.3) and

$$\text{tr} [(2\mathbb{k} - \not{p})\gamma_\nu(2\mathbb{k} - \not{p})\not{p}] = \text{tr} [\not{p}\gamma_\nu\not{p}\not{p}] + 2(k^2 - (k-p)^2)\text{tr} [\gamma_\nu(2\mathbb{k} - \not{p})] - 2(k^2 + (k-p)^2)\text{tr} [\gamma_\nu\not{p}], \quad (\text{C.8})$$

we also have the following identity:

$$\begin{aligned} & \int \frac{d^d k}{(2\pi)^d} \frac{1}{k^2(k-p)^2} \text{tr} \left\{ (2\mathbb{k} - \not{p})\gamma_\nu(2\mathbb{k} - \not{p})(2k-p)_{(\alpha\gamma\beta)} \not{p}(2k-p)_{(\rho\gamma\sigma)} \right\}, \\ &= -64i [p_{(\alpha} T_{\beta)\nu\rho\sigma}(p) + p_{(\rho} T_{\sigma)\nu\alpha\beta}(p)] \\ &+ \int \frac{d^d k}{(2\pi)^d} \frac{1}{k^2(k-p)^2} \text{tr} \left\{ \not{p}\gamma_\nu\not{p}(2k-p)_{(\alpha\gamma\beta)} \not{p}(2k-p)_{(\rho\gamma\sigma)} \right\}. \end{aligned} \quad (\text{C.9})$$

## Appendix D. Symmetrisation of the two-point function

We show that in the integral expression for the two-point function, equation (26), the antisymmetrised part of the integral in indices  $\mu\nu$  vanishes,

$$\int \frac{d^d k}{(2\pi)^d} \text{tr} \left\{ \frac{\mathbb{k}(2k-p)_{[\mu}\gamma_\nu](\mathbb{k} - \not{p})(2k-p)_{(\rho\gamma\sigma)}}{k^2(k-p)^2} \right\} = 0. \quad (\text{D.1})$$

Expanding out the trace and rewriting

$$k \cdot (k-p) = \frac{1}{2}k^2 + \frac{1}{2}(k-p)^2 - \frac{1}{2}p^2,$$

the integral on the lhs of equation (D.1) becomes

$$2^{d/2-1} \int \frac{d^d k}{(2\pi)^d} \frac{1}{k^2(k-p)^2} k_{[\mu} (\eta_{\nu](\sigma} - 2p_{\nu]}k_{(\sigma)} (2k-p)_{\rho)}. \quad (\text{D.2})$$

Using equations (B.1)–(B.3), equation (D.1) is established.

## Appendix E. Evaluation of $\tilde{T}$

In this appendix we evaluate  $\tilde{T}_{\nu\rho\sigma\alpha\beta}(p, q)$ , defined in equation (57),

$$i \int \frac{d^d k}{(2\pi)^d} \text{tr} \left\{ \frac{\mathbb{k}}{k^2} (\gamma_\nu(\mathbb{k} + \not{p}) + (2k+p)_\nu) (2k+p-q)_{(\rho\gamma\sigma)} \frac{\mathbb{k} - \not{q}}{(k-q)^2} (2k-q)_{(\alpha\gamma\beta)} \right\}. \quad (\text{E.1})$$

We first note that each factor can be written as a sum of a term proportional to  $(2k - q)$  and another  $k$ -independent term. For example,

$$k_\mu = \frac{1}{2}(2k - q)_\mu + \frac{1}{2}q_\mu. \quad (\text{E.2})$$

Now, again using the trick in appendix D that under a reparametrisation of the integration variable  $k \rightarrow -k + q$  we have

$$\frac{1}{k^2(k - q)^2} \rightarrow \frac{1}{k^2(k - q)^2}, \quad (2k - q) \rightarrow -(2k - q), \quad (\text{E.3})$$

we rewrite integral (E.1) in terms of an even number of  $(2k - q)$  factors,

$$\begin{aligned} \frac{i}{8} \int \frac{d^d k}{(2\pi)^d} \frac{1}{k^2(k - q)^2} \text{tr} \left\{ (2\cancel{k} - \cancel{q}) \gamma_\nu (2\cancel{k} - \cancel{q}) (2k - q)_{(\alpha\gamma\beta)} (2\cancel{k} - \cancel{q}) P_{(\rho\gamma\sigma)} \right. \\ - (2\cancel{k} - \cancel{q}) \gamma_\nu (2\cancel{k} - \cancel{q}) (2k - q)_{(\alpha\gamma\beta)} \cancel{q} (2k - q)_{(\rho\gamma\sigma)} \\ + (2\cancel{k} - \cancel{q}) \gamma_\nu \cancel{q} (2k - q)_{(\alpha\gamma\beta)} (2\cancel{k} - \cancel{q}) (2k - q)_{(\rho\gamma\sigma)} \\ + (2\cancel{p} + \cancel{q}) \gamma_\nu (2\cancel{k} - \cancel{q}) (2k - q)_{(\alpha\gamma\beta)} (2\cancel{k} - \cancel{q}) (2k - q)_{(\rho\gamma\sigma)} \\ - (2\cancel{k} - \cancel{q}) \gamma_\nu \cancel{q} (2k - q)_{(\alpha\gamma\beta)} \cancel{q} P_{(\rho\gamma\sigma)} \\ - (2\cancel{p} + \cancel{q}) \gamma_\nu (2\cancel{k} - \cancel{q}) (2k - q)_{(\alpha\gamma\beta)} \cancel{q} P_{(\rho\gamma\sigma)} \\ + (2\cancel{p} + \cancel{q}) \gamma_\nu \cancel{q} (2k - q)_{(\alpha\gamma\beta)} (2\cancel{k} - \cancel{q}) P_{(\rho\gamma\sigma)} \\ - (2\cancel{p} + \cancel{q}) \gamma_\nu \cancel{q} (2k - q)_{(\alpha\gamma\beta)} \cancel{q} (2k - q)_{(\rho\gamma\sigma)} \\ + 2(2k - q)_\nu (2\cancel{k} - \cancel{q}) (2k - q)_{(\alpha\gamma\beta)} (2\cancel{k} - \cancel{q}) P_{(\rho\gamma\sigma)} \\ + 2(p + q)_\nu (2\cancel{k} - \cancel{q}) (2k - q)_{(\alpha\gamma\beta)} (2\cancel{k} - \cancel{q}) (2k - q)_{(\rho\gamma\sigma)} \\ - 2(2k - q)_\nu \cancel{q} (2k - q)_{(\alpha\gamma\beta)} \cancel{q} P_{(\rho\gamma\sigma)} \\ \left. - 2(p + q)_\nu \cancel{q} (2k - q)_{(\alpha\gamma\beta)} \cancel{q} (2k - q)_{(\rho\gamma\sigma)} \right\}, \quad (\text{E.4}) \end{aligned}$$

cancelling some terms using the gamma matrix identity (C.1). Using identity (C.4), we can show that the terms on the sixth and seventh line in the expression above cancel. Furthermore, we can use identities (C.6), (C.7) and (C.9) to simplify the expressions on the ninth and tenth lines; first, third and fourth lines and the second line of above expression, respectively. Collecting all the terms we find that

$$\tilde{T}_{\nu\rho\sigma\alpha\beta}(p, q) = 16 [3p_{(\rho} T_{\sigma)\nu\alpha\beta}(q) + 2(p + q)_\nu T_{\rho\sigma\alpha\beta}(q) - p^\tau \eta_{\nu(\rho} T_{\sigma)\tau\alpha\beta}(q)]. \quad (\text{E.5})$$

## ORCID iDs

Hadi Godazgar  <https://orcid.org/0000-0002-9263-2325>

## References

- [1] Capper D M and Duff M J 1974 *Nuovo Cimento* **23A** 173  
Capper D M and Duff M J 1974 *Nucl. Phys. B* **82** 147

- [2] Deser S, Duff M J and Isham C 1976 *Nucl. Phys. B* **111** 45
- [3] Duff M J 1977 *Nucl. Phys. B* **125** 334
- [4] Tsao H S 1977 *Phys. Lett.* **68B** 79
- [5] Perry M J 1978 *Nucl. Phys. B* **143** 114
- [6] Fradkin E S and Tseytlin A 1982 *Nucl. Phys. B* **203** 157
- [7] Fradkin E S and Tseytlin A A 1982 *Phys. Lett.* **110B** 117
- [8] Fradkin E S and Tseytlin A 1984 *Phys. Lett. B* **134** 187
- [9] Deser S and Schwimmer A 1993 *Phys. Lett. B* **309** 279
- [10] Barvinsky A O, Gusev Y V, Vilkovisky G A and Zhytnikov V V 1995 *Nucl. Phys. B* **439** 561
- [11] Schwimmer A and Theisen S 2011 *Nucl. Phys. B* **847** 590
- [12] Tseytlin A 2013 *Nucl. Phys. B* **877** 598
- [13] Parker L 1978 *Recent Developments in Gravitation (Cargèse, 1978)* ed M E Lévy and S Deser (New York: Plenum)
- [14] Eguchi T, Gilkey P B and Hanson A J 1980 *Phys. Rep.* **66** 213
- [15] Birrell N D and Davies P C W 1982 *Quantum Fields in Curved Space* (Cambridge: Cambridge University Press)
- [16] Duff M J 1994 *Class. Quantum Grav.* **11** 1387
- [17] Vassilevich D 2003 *Phys. Rep.* **388** 279
- [18] Barvinsky A O 2012 *Phys. Rev. D* **85** 104018
- [19] Boulanger N 2007 *Phys. Rev. Lett.* **98** 261302
- [20] Endo R 1984 *Prog. Theor. Phys.* **71** 1366
- [21] Bastianelli F and van Nieuwenhuizen P 2009 *Path Integrals and Anomalies in Curved Space* (Cambridge: Cambridge University Press)
- [22] Dowker J S and Critchley R 1977 *Phys. Rev. D* **16** 3390
- [23] Brown L S and Cassidy J P 1977 *Phys. Rev. D* **15** 2810
- [24] Osborn H and Petkou A C 1994 *Ann. Phys.* **231** 311
- [25] Cappelli A, Guida R and Magnoli N 2001 *Nucl. Phys. B* **618** 371
- [26] Coriano C, Delle Rose L, Mottola E and Serino M 2015 *Field theory J. High Energy Phys.* JHEP08(2012)147
- [27] Bzowski A, McFadden P and Skenderis K 2017 arXiv:1711.09105 [hep-th]
- [28] Bonora L and Giaccari S 2013 *J. High Energy Phys.* JHEP08(2013)116
- [29] Bonora L, Giaccari S and Lima de Souza B 2014 *J. High Energy Phys.* JHEP07(2014)117
- [30] Christensen S M and Duff M J 1978 *Phys. Lett.* **76B** 571
- [31] Christensen S M and Duff M J 1979 *Nucl. Phys. B* **154** 301
- [32] Meissner K A and Nicolai H 2017 *Phys. Lett. B* **772** 169
- Meissner K A and Nicolai H 2017 *Phys. Rev. D* **96** 041701
- [33] Antoniadis I and Mottola E 1992 *Phys. Rev. D* **45** 2013
- [34] Wald R M 1984 *General Relativity* (Chicago, IL: University of Chicago Press)
- [35] Smirnov V A 2004 *Evaluating Feynman Integrals (Springer Tracts in Modern Physics vol 211)* (Berlin: Springer)
- [36] Weinzierl S 2010 *Introduction to Feynman integrals 6th Summer School on Geometric and Topological Methods for Quantum Field Theory (6–23 Jul 2009, Villa de Leyva, Colombia)* C09-07-06.8 pp 144–87 (arXiv:1005.1855 [hep-ph])
- [37] Davydychev A I 1991 *Phys. Lett. B* **263** 107
- [38] Davydychev A I 1992 *J. Phys.: Math. Gen. A* **25** 5587
- [39] Boos E E and Davydychev A I 1991 *Theor. Math. Phys.* **89** 1052
- [40] 't Hooft G and Veltman M J G 1972 *Nucl. Phys. B* **44** 189