HELICITY, SPIN, AND INFRA-ZILCH OF LIGHT: A LORENTZ COVARIANT FORMULATION

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Abstract. The spin part of angular momentum of the electromagnetic field is known since the 1990’s to be a separately conserved quantity. Cameron et al. [1] introduced the helicity array, a non-covariant analog of Lipkin’s zilch tensor, which expresses the hierarchy of conservation laws including helicity, spin, as well as the spin-flux or infra-zilch. In this paper, a novel conserved Lorentz covariant tensor, termed the helicity tensor, is introduced. The conservation laws arising from the helicity array can be obtained from the helicity tensor. The Lorentz covariance of the helicity tensor is in contrast to previous formulations of the helicity hierarchy of conservation laws, which required the non-Lorentz covariant transverse gauge. The helicity tensor is shown to arise as a Noether current for a variational symmetry of a duality-symmetric Lagrangian for Maxwell theory.

1. Introduction

The physical relevance of the spin and intrinsic orbital parts of angular momentum of the Maxwell field was first demonstrated by Allen et al. [2] and van Enk and Nijenhuis [3] in the early 1990’s. These and subsequent works have led to the understanding that these two parts of angular momentum are separately conserved. The related conservation laws have recently been re-examined by several groups, including I. and Z. Bialynicki-Birula [4], Bliokh et al. [5] and Barnett et al. [6]. The two parts of the angular momentum are physically meaningful observable quantities despite the fact that they do not satisfy the commutation relations of angular momenta. Among the phenomena related to the just mentioned conservation laws is the experimentally verified spin Hall effect of light, cf. Bliokh et al. [7] and references therein.

The notion of electromagnetic (or optical) helicity was introduced by Trueba and Rañada based on similar notions defined in the context of magnetohydrodynamics and fluid dynamics [8]. The conservation of helicity is due to the fact that it is the density of a Noether current associated to the duality symmetry of the free Maxwell equations, while the spin part of the angular momentum turns out to be the flux of helicity [9]. The electromagnetic helicity has been used in describing the interaction of light beams with chiral molecules and nano-particles [10]. These quantities also generalize to linearized gravitational fields and acoustic waves, cf. e.g. [11]–[14].

The zilch tensor, a third rank tensor constructed from the product of the Maxwell field and its first derivative was introduced by Lipkin [15], see also Morgan [16] and Kibble [17]. The zilch tensor is conserved for the Maxwell field in the absence of sources. The hierarchy of conservation laws expressed by the zilch tensor, like the above mentioned helicity and related conservation laws, constitute new conserved quantities, that are not equivalent to the classical conservation laws related to Poincaré and conformal symmetries of Minkowski space, cf. [18] and references therein.
The zilch conservation laws were generalized by Morgan to a hierarchy of quadratic tensors with an arbitrary number of derivatives. Morgan also found another such hierarchy generalizing the Maxwell stress-energy tensor. It has been proposed that the purely temporal component of the zilch tensor describes the chirality of the electromagnetic fields [19]. Electromagnetic helicity and chirality are proportional for monochromatic waves but are in general distinct notions. It is still under debate which one of these quantities is more appropriate for describing relevant properties of light in its interaction with chiral matter, cf. [20] for discussion.

Motivated by the structure of the zilch tensor, the conserved currents associated to the helicity density \( H_{abc} \), spin density \( S_{a} \), and infra-zilch \( \Sigma_{ab} \) have been arranged by Cameron et al. [1] in what they named the helicity array, \( N_{abc} \). Here, the indices label the spacetime components of the array, taking the values \{0, 1, 2, 3\}. The array is by construction symmetric in the first two indices, \( N_{abc} = N_{(ab)c} \).

The corresponding conservation laws could then be represented in the compact form \( \partial_{c} N_{abc} = 0 \), valid provided the field equations hold and the transverse gauge condition is imposed. However, as stated in [1], the array \( N_{abc} \) is not a tensor.

Further, the analysis in [1] as well as earlier papers on the subject including the papers cited above makes use of the transverse gauge condition, i.e. a combination of temporal gauge \( A_{0} = 0 \) and Coulomb gauge. Due to this elliptic gauge condition, the helicity array is neither tensorial nor Lorentz covariant.

As was noted in [1], the structure of the helicity array is closely related to the zilch tensor but differs in the number of derivatives. As mentioned above, the zilch tensor is of the third order in derivatives of the Maxwell potential, while the helicity array is of the first order. The similarities in the structures is illustrated by the formal transformation \((A, C) \rightarrow (\nabla \times A, \nabla \times C)\) of the electric and magnetic potentials mapping the helicity array to the zilch tensor, \( N_{abc} \rightarrow Z_{abc} \), cf. [1]. Here \( A, C \) are the 3-vector parts of the electric and magnetic potentials for the Maxwell field. The fact that this mapping involves both potentials is an indication that a tensor description of the helicity array is related to a self-dual formulation of Maxwell theory.

Let \( A_{a} = A_{a} + iC_{a} \) be the complex potential formed from the electric and magnetic potentials \( A_{a}, C_{a} \), and let

\[
\mathcal{Y}_{abc} = \frac{i}{2} \tilde{A}_{a} \partial_{b} A_{c}. \tag{1.1}
\]

where \( \tilde{A}_{a} \) denotes complex conjugation, and we have used the using the bidirectional-arrow derivative notation, which is often used in the quantum theory literature. This is defined in the present case by \( \tilde{A}_{a} \partial_{b} A_{c} = \tilde{A}_{a} \partial_{c} A_{b} - A_{b} \partial_{c} \tilde{A}_{a} \) and extends to general fields in the same manner. In section 3 we shall define the helicity tensor by

\[
\mathcal{H}_{abc} = (\hat{P} \mathcal{Y})_{abc}, \tag{1.2}
\]

where \( \hat{P} \) is a linear involution, which transforms the components of the auxiliary tensor \( \mathcal{Y}_{abc} \) with even parity to odd parity. By this construction, \( \mathcal{H}_{abc} \) is a real tensor, which like the helicity array, and the classical angular momenta, has odd parity. The helicity, spin and infra-zilch conservation laws are consequences of

\[
\partial_{c} \mathcal{H}_{abc} = 0, \tag{1.3}
\]

which is valid provided that the vacuum Maxwell equations \( \partial_{a} F^{a}_{b} = 0 \), and the Lorenz gauge condition \( \partial_{a} A_{a} = 0 \) hold, where \( F_{ab} = \partial_{a} A_{b} - \partial_{b} A_{a} \) is the complex field strength.

The helicity tensor \( \mathcal{H}_{abc} \) is a quadratic tensor constructed from the electric and magnetic potentials and their first derivatives and contains the same information...
as that contained in the components of the helicity array. For details, see section 3.

We further show in section 4 that the helicity tensor arises as the Noether current of a symmetry of the gauge extended, duality-symmetric Maxwell action.

**Overview of this paper.** In section 2 we review the construction of helicity array and the interrelated conservation laws of its components. Then, in section 3 we introduce the covariant helicity tensor. It is a Lorentz covariant tensor that is shown to contain the same information as the helicity array. In this analysis we make use of the 1 + 3 decompositions discussed in appendix B, in terms of the notation introduced in appendix A. In section 4 we demonstrate that the helicity tensor arises as a Noether current for a symmetry of the duality-symmetric action for the Maxwell field, and find the generator of the corresponding symmetry. In particular this is a symmetry of the duality-symmetric Lagrangian with Lorenz gauge fixing terms.

## 2. The helicity array in Maxwell theory

We shall consider fields on Minkowski spacetime with metric $g_{ab}$. We use the index notation with indices $a, b, c, \ldots$ taking values $0, \ldots, 3$, using $0$ for a temporal index and overlines indices $\bar{a}, \bar{b}, \ldots$ for spatial indices, cf. appendix A. We shall sometimes use the standard 3-dimensional vector notation, with bold-face capital letters $\mathbf{A}, \ldots$ denoting 3-vectors. In this context, div and $\times$ denote the divergence and cross-product. Note that we also use the notation div for the 4-dimensional divergence.

In this section we shall review the helicity array and the associated conservation laws. Let $A_a, C_a$ be the electric and magnetic potentials and let $F_{ab} = \partial_a A_b - \partial_b A_a$, $G_{ab} = \partial_a C_b - \partial_b C_a$ be the field strength and its dual. In this section, unless otherwise stated, we shall use the transverse gauge condition

$$A_0 = 0 = C_0, \quad \text{div} \, A = 0 = \text{div} \, C. \quad (2.1)$$

The helicity current

$$J^a = \frac{1}{2} (G_{ab} A^b - F_{ab} C^b) \quad (2.2)$$

satisfies the conservation law

$$\partial_a J^a = \frac{1}{2} A^a G_{a,b} - \frac{1}{2} C^a F_{a,b} \cong 0 \quad (2.3)$$

The temporal part of the helicity current is referred to as the helicity density and has the form

$$\mathcal{H} = J^0 = \frac{1}{2} (\mathbf{A} \cdot \mathbf{B} - \mathbf{C} \cdot \mathbf{E}) \quad (2.4)$$

where $\mathbf{E}, \mathbf{B}$ are the electric and magnetic field vectors expressed in traditional spatial vector notation and $\mathbf{A}, \mathbf{C}$ are the spatial vectors with components $A_\alpha, C_\alpha$, and $J^0 = - J^a u_a$ is the helicity density. The spatial part of the helicity current is the spin density given by

$$S_a \equiv J_a = \frac{1}{2} (\mathbf{E} \times \mathbf{A} + \mathbf{B} \times \mathbf{C}) + \frac{1}{2} (C_0 \mathbf{E} - A_0 \mathbf{B}), \quad (2.5)$$

usually presented in transverse gauge \[.\] Here $J_a = h_{a,b} J_b$, where $h_{ab} = g_{ab} + u_a u_b$ is the spatial projection, cf. appendix A. Written in terms of these quantities, the conservation law \[2.3\] has the form

$$\dot{\mathcal{H}} + \partial_a S^a \cong 0 \quad (2.6)$$

---

1 Here and below $\cong$ signifies equality modulo the field equations $F_{a,b}^a = 0$ and $G^a_{a,b} = 0$. 

where here and in the following the overdot denotes time derivative, \( \dot{\mathcal{H}} = u^a \partial_a \mathcal{H} = \partial_b \mathcal{H} \). This relation is the first level of the helicity hierarchy, a three level hierarchy of conservation laws

\[
\dot{\mathcal{H}} + \partial_a S^a \simeq 0 \tag{2.7}
\]

\[
\dot{S}_a + \partial_b \Sigma^b_a \simeq 0 \tag{2.8}
\]

\[
\dot{\Sigma}_{ab} + \partial_c N_{ab}^c \simeq 0 \tag{2.9}
\]

where the spin flux (or infra-zilch \([1]\)) is the symmetric spatial tensor

\[
\Sigma_{ab} = h_{ab} \mathcal{H} + C_{(a} E_{b)} - A_{(a} B_{b)} \tag{2.10}
\]

and the flux of infra-zilch is the spatial tensor

\[
N_{abc} = N_{(ab)c} = h_{ab} S_c + C_{(a} \partial_{c]} A_{b)} = h_{ab} S_c + C_{(a} A_{b),c} - A_{(a} C_{b),c} \tag{2.11}
\]

using the bidirectional-arrow notation. Cameron et al. \([1]\) defined a conserved rank-three object \(N_{abc} = N^{(ab)c}\) that they called the helicity array having three spacetime indices. The array was given the components

\[
N^{000} = \mathcal{H} \tag{2.12}
\]

\[
N^{0a0} = N^{00a} = S^a \tag{2.13}
\]

\[
N^{a00} = N^{00a} = \Sigma^{ab} \tag{2.14}
\]

\[
N^{abc} = N^{abc} \tag{2.15}
\]

leading to the conservation relation \(\partial_c N^{abc} \simeq 0\). They also showed that the array, while not being a tensor, could be mapped to the zilch tensor \([15–17]\) by the mapping

\[
A \to \nabla \times A, \quad C \to \nabla \times C. \tag{2.16}
\]

Despite this close relation to the zilch tensor, the helicity array has up to now not been reformulated in a Lorentz covariant form. In the next section we introduce a Lorentz covariant tensor that contains the same information as the helicity array.

3. Helicity tensor in complex duality-symmetric formulation

Recall the formulation of Maxwell theory in terms of a complex vector potential \(A_a = A_a + i C_a\) with field strength \(F_{ab} = -2 A_{[a} A_{b]}\), cf. \([21–23]\) and references therein. We shall now define a Lorentz covariant complex tensor \(\mathcal{H}_{abc}\), which we call the helicity tensor, that contains the same information as the helicity array. Let

\[
\mathcal{H}_{abc} = \frac{i}{2} \bar{A}_a \bar{\partial}_b A_c. \tag{3.1}
\]

This tensor is manifestly duality invariant and is conserved with respect to the third index as we will show in the following.

Taking the divergence of \((3.1)\) we find

\[
\mathcal{H}_{ab} \cdot c = \frac{i}{2} (A_a \Box A_b - A_b \Box A_a) = \frac{i}{2} A_a F^b_{a,b} - \frac{i}{2} A_b F^b_{a,b} + \frac{i}{2} A_a (\text{div} A),_b - \frac{i}{2} A_b (\text{div} A),_a, \tag{3.2}
\]

where now \(\text{div} A\) is the spacetime divergence \(\partial_a A^a\). This shows that \(\mathcal{H}_{abc}\) is conserved in Lorenz gauge,

\[
\text{div} A = \text{div} \bar{A} = 0, \tag{3.3}
\]

provided that the field equations \(F^b_{a,b} = 0\) and \(\bar{F}^b_{a,b} = 0\) are satisfied.

Let

\[
P_{abcd} = g_{(a} g_{b) d - \frac{i}{2} \varepsilon_{abcd}, \tag{3.4}
\]
where $g_{ab}$ is the Minkowski metric, and $\varepsilon_{abcd}$ is the volume element, and define the operator $\hat{P}$ by its action on 2-tensors, $(\hat{P} t)_{ab} = P_{ab}^{\ \cd} t_{cd}$. Due to $P_{abcd}P^{cdfe} = g_a^{\ d}g_b^{\ f}$, we have that $(\hat{P})^2$ is the identity operator, and hence $\hat{P}$ is an involution. We now define the helicity tensor by

$$\mathcal{H}_{abc} = (\hat{P} \mathcal{Y})_{abc}$$

(3.5)

where $\hat{P}$ acts on the first two indices of $\mathcal{Y}_{abc}$. Taking into account the definition of $\hat{P}$ we have that

$$\partial_a \mathcal{H}^{abc} \equiv 0$$

(3.6)

provided the vacuum Maxwell equations are satisfied and $A_a$ is in Lorenz gauge. As we shall see, $\mathcal{H}_{abc}$ has odd parity. Since $\hat{P}$ is invertible, $\mathcal{H}_{abc}$ contains the same information as $\mathcal{Y}_{abc}$. Therefore, in order to demonstrate that $\mathcal{H}_{abc}$ contains the same information as $N_{abc}$, it is sufficient to consider $\mathcal{Y}_{abc}$.

We now note that the helicity tensor has the form

$$\mathcal{H}_{abc} = Y_{abc} + iX_{abc}$$

(3.7)

We have

$$Y_{abc} = \Re\{\mathcal{Y}_{abc}\} = \mathcal{Y}_{(ab)c} = \frac{i}{2} \tilde{A}_a \tilde{\partial}_c |A_b|, \quad (3.8)$$

$$X_{abc} = \Im\{\mathcal{Y}_{abc}\} = -i \mathcal{Y}_{(ab)c} = \frac{1}{2} \tilde{A}_a \tilde{\partial}_c |A_b|. \quad (3.9)$$

In terms of real potentials $A_a$ and $C_a$ we have

$$Y_{abc} = C_{(a} |\tilde{\partial}_c| A_{b)}, \quad (3.10a)$$

$$X_{abc} = \frac{1}{2}(A_a \tilde{\partial}_a A_b + C_a \tilde{\partial}_a C_b) = A_{[a} A_{b]}^{c} + C_{[a} C_{b]}^{c}. \quad (3.10b)$$

We now note that the helicity array and corresponding conservation laws are incorporated in the helicity tensor $\mathcal{H}_{abc}$.

**Remark 3.1.** In previous work on the helicity array and related conservation laws, the transverse gauge (2.1) has been assumed. Here we use the Lorenz gauge (3.3) which is Lorentz invariant. On Minkowski spacetime, transverse gauge is consistent with Lorenz gauge, but breaks Lorentz invariance.

The helicity current $J^a$, which has a covariant form (2.2), is expected to have a covariant relationship with the helicity tensor. In fact, it can be written as

$$J^a = \frac{1}{4}(\mathcal{F}^{ab} A_b - \mathcal{F}^{ab} \tilde{A}_b)$$

$$= -\frac{i}{2} \mathcal{Y}^b_{\ b} + \frac{1}{4}(\tilde{A}^a \text{ div } A - A^a \text{ div } \tilde{A}) + \frac{1}{4}(\tilde{A}^b A^a - A^b \tilde{A}^a)_b, \quad (3.12)$$

which shows that the helicity current $J^a$ is equivalent to the trace of the helicity tensor $\mathcal{H}_{ba}^b = \mathcal{Y}_{ba}^b$, up to the Lorenz gauge and a trivial current (i.e. having a vanishing divergence). To be concrete, we introduce the equivalent helicity current

$$\tilde{J}_a = \frac{1}{4} \tilde{A}^b \tilde{\partial}_a A_b = -\frac{i}{2} \mathcal{Y}^b_{\ ba} = -\frac{1}{2} \mathcal{H}^b_{\ ba}. \quad (3.13)$$

From (3.12), we find that

$$J^a - \tilde{J}^a = \frac{1}{4}(\tilde{A}^a \text{ div } A - A^a \text{ div } \tilde{A}) + \frac{1}{2} K^a, \quad (3.14)$$
from while the second conservation law in (2.8) (i.e. the conservation of sp in) is obtained

where

\[ K^a = i (\tilde{A}^b A^a)_b = 2 (C^{[b} A^{a]})_b \]  

(3.15)

has a vanishing divergence. Finally, we note that the trace of the helicity tensor \( \mathcal{H}_{abc} \) is given by the trace of \( Y_{abc} \), because \( *X_{abc} \) is antisymmetric in its first two indices.

Using (3.13) and performing the 1 + 3 decomposition of \( \tilde{J}_a \) (see appendix B) the helicity density in transverse gauge can be obtained as:

\[ \mathcal{H} \equiv -\frac{1}{2} Y_a^a = -\frac{1}{2} \mathcal{H}^{a0} \]  

(3.16)

The second level of the helicity array turns out to be related to \( *X_{abc} \). In particular the spin density is related to the helicity tensor through the 1 + 3 decomposition (see appendix B) by

\[ S^a \equiv *X^{0a0} = \mathcal{H}^{0a0} \]  

(3.17)

while the second conservation law in (2.8) (i.e. the conservation of spin) is obtained from

\[ \tilde{S}^{a0} + \Sigma^{bc, c} \equiv *X^{000} = 0. \]  

(3.18)

Finally, from (3.6) the infra-zilch can be constructed as

\[ \Sigma^{ab} \equiv h^{ab} \mathcal{H} + Y^{ab0} = -\frac{1}{2} h^{ab} \mathcal{H}_{c}^{c0} + \mathcal{H}^{a00} , \]  

(3.19)

and its conservation law (i.e. the third level of conservation of the helicity array) is obtained from

\[ \Sigma^{ab} + \mathcal{N}^{abc, c} \equiv -\frac{1}{2} h^{ab} \mathcal{H}_{c}^{cd} + \mathcal{H}^{abc, c} = 0. \]  

(3.20)

Finally, the flux of infra-zilch can be written as

\[ \mathcal{N}^{abc, c} \equiv -\frac{1}{2} h^{ab} (\mathcal{H}_{d}^{dc} - K^{c}) + \mathcal{H}^{abc} , \]  

(3.21)

where \( K^a \) is the null current (3.15). A summary of these relations between the helicity array and the helicity tensor is given in table 1.

### Table 1. Summary of the relations between the helicity tensor and the helicity array. Transverse gauge is assumed.

<table>
<thead>
<tr>
<th>symbol</th>
<th>helicity array</th>
<th>helicity tensor</th>
</tr>
</thead>
<tbody>
<tr>
<td>( \mathcal{H} )</td>
<td>( N^{000} )</td>
<td>( -\frac{1}{2} \mathcal{H}_{c}^{a0} )</td>
</tr>
<tr>
<td>spin density</td>
<td>( S^a )</td>
<td>( N^{0a0} = N^{a00} )</td>
</tr>
<tr>
<td>infra-zilch density</td>
<td>( \Sigma^{ab} )</td>
<td>( N^{a00} = N^{0a0} - \frac{1}{2} h^{ab} \mathcal{H}<em>{c}^{c0} + \mathcal{H}^{a00} = -\frac{1}{2} h^{ab} (\mathcal{H}</em>{c}^{cd} - K^{c}) + \mathcal{H}^{abc} )</td>
</tr>
<tr>
<td>infra-zilch flux</td>
<td>( \mathcal{N}^{abc} )</td>
<td>( -\frac{1}{2} h^{ab} (\mathcal{H}_{c}^{cd} - K^{c}) + \mathcal{H}^{abc} )</td>
</tr>
</tbody>
</table>

### 4. Noether analysis

Using a complex vector potential \( A_a = A_a + i C_a \), Maxwell’s equations can be derived from the duality-symmetric Lagrangian (cf. 22)

\[ \mathcal{L} = -\frac{i}{2} F_{ab} \tilde{F}^{ab} = -\frac{1}{2} \mathcal{H}_{abcd} A_{a,b} \tilde{A}_{c,d} \]  

(4.1)
where $F_{ab} = -2A_{[a,b]}$ and the bar denotes complex conjugation. Here, $\kappa_{abcd} = g_{[a[c}g_{b]d]}$ is the antisymmetry projector $^3$.

In $^2$, the real and imaginary parts of $A_a$ were taken as the basic field variables. Here, we will instead use $A_a$ and $\bar{A}_a$ as the basic fields. Since they represent linearly independent combinations of $A_a$ and $C_a$, they can be used as alternative field variables. They carry the same number of degrees of freedom (4 each) and are treated as formally independent. Defining $M^a = E^a_\bar{\alpha}(\mathcal{L})$, where $E$ is the Euler operator for a variable $X$, the Euler-Lagrange expressions are given by

$$M^a = -\frac{\partial}{\partial x^b} \frac{\partial \mathcal{L}}{\partial A_{a,b}} = -\frac{1}{2} \bar{g}^{ab} \omega$$

(4.2)

and $E^a_\bar{\alpha}(\mathcal{L}) = \bar{M}^\alpha$. Expressed in terms of the potentials they become

$$M_a = \frac{1}{4} \square A_a - \frac{1}{2} (\text{div} \bar{A})_a.$$ 

(4.3)

Here $\text{div} A = \partial_a A^a$ is the divergence.

$$\mathcal{B}_{ab}^c = \frac{1}{2} \left( \bar{A}_a \square A_b - A_b \square \bar{A}_a \right) = 2i \bar{A}_a M_b - 2i A_b M_a + \frac{1}{2} \bar{A}_a (\text{div} \bar{A})_b - \frac{1}{2} A_b (\text{div} A)_a.$$ 

(4.4)

The presence of the gauge dependent terms, the equation (4.4) does not as it stands lead to a conservation law in the characteristic form needed for correspondence with Noether theory (see $^2$, $^3$ for details of our use of the Noether formalism). One way to find such a form is to modify the Lagrangian by adding terms which take care of the gauge constraints. To this end we define a modified Lagrangian depending on an auxiliary complex dependent variable $\omega$.

$$\mathcal{L} = \mathcal{L} + \omega \text{div} A + \bar{\omega} \text{div} \bar{A}.$$ 

(4.5)

Since the additional parts are linear in $A_{a,b}$ and $\bar{A}_{a,b}$, no new terms enter in the Euler-Lagrange expressions (4.2). The modified Lagrangian (4.5) leads to the complex Euler-Lagrange expression

$$\Omega = E_\omega(\mathcal{L}) = \frac{\partial \mathcal{L}}{\partial \omega} = \text{div} A$$

(4.6)

and its complex conjugate in addition to those in (4.2). This shows that the Lagrangian $\mathcal{L}$ gives the duality-symmetric Maxwell equations in Lorenz gauge. The relation (4.4) can then be written in the form

$$\mathcal{P}_{ab}^c = 2i \bar{A}_a M_b - 2i A_b M_a + \frac{i}{2} \bar{A}_a \Omega_{b.a} - \frac{i}{2} A_b \Omega_{a.a}.$$ 

(4.7)

As in the case of the zilch conservation law (see $^2$, $^3$), there are Euler-Lagrange expressions $(\Omega$ and $\bar{\Omega})$ that are differentiated in (4.7). Therefore to arrive at a conservation law in characteristic form we need to first perform partial integrations leading to

$$\mathcal{P}_{ab}^c = 2i \bar{A}_a M_b - 2i A_b M_a + \frac{i}{2} \bar{A}_a \Omega + \frac{i}{2} (\bar{A}_a \Omega)_b - \frac{i}{2} (A_b \Omega)_a.$$ 

(4.8)

Defining a tensor which is equivalent$^1$ to $\mathcal{P}_{ab}$ by

$$\mathcal{P}_{abc} = \mathcal{P}_{abc} - \frac{i}{2} g_{bc} \bar{A}_a \text{div} A + \frac{i}{2} g_{ab} A_b \text{div} \bar{A}$$

(4.9)

the conservation law can be written in the characteristic form

$$\mathcal{P}^a_{abc} = 2i \bar{A}_a M_b - 2i A_b M_a - \frac{i}{2} \bar{A}_a \Omega + \frac{i}{2} A_b \Omega.$$

(4.10)

$^2$This tensor projects out the antisymmetric part of any index pair, e.g. $M_{[ab]} = \kappa_{ab}^{cd} M_{cd}$. It has the same index symmetries as the Riemann tensor, $\kappa_{abcd} = \kappa_{dabc} = \kappa_{[ab]cd}$ and $\kappa_{a[bcd]} = 0$.

$^3$In the sense of being equal for solutions of the $\mathcal{L}$ field equations.
From equation (4.10) we can read off the characteristic functions, which are the coefficients of the Euler-Lagrange expressions:

\[ Q_{abc} = -2 i A_b g_{ac}, \quad Q_{abc}^\dagger = 2 i \tilde{A}_a g_{bc}, \quad Q^\dagger = -i \tilde{A}_{a,b} \quad \text{and} \quad Q^\dagger = i A_{b,a} \]

The corresponding symmetry generator then becomes

\[ v_{ab} = Q_{abc}^\dagger \frac{\partial}{\partial \tilde{A}_c} + Q_{abc} \frac{\partial}{\partial A_c} + Q_{abc}^\dagger \frac{\partial}{\partial \omega^b} + Q_{abc} \frac{\partial}{\partial \omega^b} \]

\[ = -2 i A_b g_{ac} \frac{\partial}{\partial \tilde{A}_c} + 2 i \tilde{A}_a g_{bc} \frac{\partial}{\partial A_c} - \frac{i}{2} \tilde{A}_{a,b} \frac{\partial}{\partial \omega^b} + \frac{i}{2} A_{b,a} \frac{\partial}{\partial \omega^b}. \quad (4.12) \]

The first prolongation of the generator (4.12) is given by

\[ \text{pr} v_{ab} = v_{ab} + Q_{abc,d}^\dagger \frac{\partial}{\partial \tilde{A}_{c,d}} + Q_{abc,d} \frac{\partial}{\partial A_{c,d}} + Q_{abc,c}^\dagger \frac{\partial}{\partial \omega^b} + Q_{abc,c} \frac{\partial}{\partial \omega^b} \]

\[ = v_{ab} - 2 i A_{b,d} g_{ac} \frac{\partial}{\partial \tilde{A}_{c,d}} + 2 i \tilde{A}_a g_{bc} \frac{\partial}{\partial A_{c,d}} + \frac{i}{2} \tilde{A}_{a,b} (\text{div} A) + \frac{i}{2} A_{b,a} (\text{div} \tilde{A}) \]

\[ = \mathcal{U}_{abc} \quad (4.14) \]

where

\[ \mathcal{U}_{abc} = \frac{i}{2} \tilde{A}_a A_{c,b} - \frac{i}{2} A_b \tilde{A}_{c,a} - \frac{i}{2} g_{bc} \tilde{A}_a \text{ div } A + \frac{i}{2} g_{ac} A_b \text{ div } \tilde{A}. \quad (4.15) \]

This leads to the conserved tensor

\[ -Q_{abc} \frac{\partial \mathcal{L}}{\partial A_{c,d}} - Q_{abc}^\dagger \frac{\partial \mathcal{L}}{\partial \tilde{A}_{c,d}} + \mathcal{U}_{abc} = \mathcal{Y}_{abc} - \frac{i}{2} g_{bc} \tilde{A}_a \text{ div } A + \frac{i}{2} g_{ac} A_b \text{ div } \tilde{A} \]

\[ = \mathcal{Y}_{abc}, \quad (4.16) \]

as given in (4.11). We note that \( \mathcal{Y}_{abc} \) and \( \tilde{\mathcal{Y}}_{abc} \) are equal modulo Lorenz gauge as expected. In the present version of the Noether formalism, this corresponds to the vanishing of their difference on-shell. Such a quantity is itself considered as a trivial conservation law of the first kind (see [24]). The conservation laws of \( \mathcal{Y}_{abc} \) and \( \tilde{\mathcal{Y}}_{abc} \) are then said to be equivalent.

It is instructive to notice the symmetric and antisymmetric parts of the generator (4.12), which are, respectively, responsible for the conservation of real and imaginary parts of the helicity tensor in (4.13). The symmetric part is

\[ v_{(ab)} = -2 i A_{a} g_{bc} \frac{\partial}{\partial A_{c}} - \frac{i}{2} \tilde{A}_{a,b} \frac{\partial}{\partial \omega} + \text{c.c.} \]

\[ = 2 C_{a} g_{bc} \frac{\partial}{\partial A_{c}} - 2 A_{a} g_{bc} \frac{\partial}{\partial \tilde{C}_{c}} + \frac{i}{2} C_{a,b} \frac{\partial}{\partial \omega^b} - \frac{i}{2} A_{a,b} \frac{\partial}{\partial \omega^b}, \quad (4.17) \]

while the antisymmetric part is

\[ v_{[ab]} = 2 i A_{a} g_{bc} \frac{\partial}{\partial A_{c}} - \frac{i}{2} \tilde{A}_{[a,b]} \frac{\partial}{\partial \omega} - \text{c.c.} \]

\[ = [2 A_{a} g_{bc} \frac{\partial}{\partial A_{c}} + 2 C_{a} g_{bc} \frac{\partial}{\partial \tilde{C}_{c}} - \frac{i}{2} A_{a,b} \frac{\partial}{\partial \omega^b} - \frac{i}{2} C_{a,b} \frac{\partial}{\partial \omega^b}] \quad (4.18) \]
5. Concluding Remarks

We have introduced a new Lorentz covariant tensor, $\mathcal{H}_{abc}$, which we call the helicity tensor, that is conserved in Lorenz gauge, and contains the same information as the helicity array of Cameron et al. [1]. In particular, the conserved currents expressing helicity, spin, and infra-zilch can be obtained from the helicity tensor by performing a 1 + 3 decomposition and specializing to transverse gauge. The fact that helicity, spin and infra-zilch are observer dependent parts of a Lorentz covariant object, the helicity tensor, is analogous to the fact that spin and orbital angular momentum are parts of the Lorentz covariant total angular momentum. With this in mind, it would be interesting to investigate if there are observer dependence properties of these quantities which are analogous to the spin-orbit interaction discussed by Bliokh et al. [25] and Smirnova et al. [26].

The construction of the helicity tensor is carried out in terms of the duality-symmetric formulation of Maxwell theory and is local with respect to the pair of electromagnetic potentials $A_a, C_a$. However, as in any other treatment of the conservation laws expressed in the helicity array, the construction is non-local with respect to the standard formulation in terms of a single electromagnetic potential. Indeed, the intrinsic non-locality of these conserved quantities is emphasized by the work of I. and Z. Bialynicki-Birula [4] who showed that the spin and orbital parts of the angular momentum of the electromagnetic field are given in a gauge invariant manner in terms of vector potentials constructed from the Maxwell field strength using a non-local transformation. It appears worthwhile to consider whether a similar construction can be carried out for the other conserved quantities contained in the helicity tensor.

We have further demonstrated that the helicity tensor and its associated conservation laws correspond to Noether currents associated to variational symmetries of the duality-symmetric Lagrangian for Maxwell theory, amended with gauge fixing terms. Symmetries giving rise to the helicity array components as Noether quantities were discussed in papers by Cameron et al. [1, 27] and Bliokh et al. [5]. The analysis in the just cited papers was restricted to the transverse gauge, and the variational aspect of the symmetries were not investigated. The symmetry generator given here is new, and differs from the non-local symmetries discussed in these papers. It would be interesting to understand the relation between these different symmetries. In particular, the physical interpretation of the symmetries presented here should be clarified.

Here we have defined the helicity tensor $\mathcal{H}_{abc}$ by applying the parity modifier operator $\hat{P}$ to the complex conserved tensor $\mathcal{Y}_{abc}$, obtaining a real tensor with the desired parity properties. It would be interesting to discuss the complex nature of $\mathcal{Y}_{abc}$ and its parity properties in the framework of geometrical algebra of the 4-dimensional spacetime, cf. [28]. We plan to return to this subject in a future publication.

Finally, we mention that it is of interest to investigate whether some of the results considered here can be generalized to Maxwell theory in more general spacetimes, in particular to analyze the generalization of the helicity hierarchy of conservation laws to certain algebraically special spacetimes. The fact that the helicity tensor is conserved in Lorenz gauge may allow us to apply the construction to interesting families of solutions of Maxwell theory, including for example Hopfions and several types of optical beams, cf. eg. [29–32] and references therein.

Appendix A. A covariant notation for 1 + 3 decomposition

Performing a time-space decomposition in any relativistic theory is crucial when considering experimental and observational measurements. It is necessary since
measurements depend on the rest space of a lab or an observer (or equivalently on
the 4-velocity of the rest space of the lab or observer). This is perhaps most often
done by introducing 3-dimensional indices which only take values which number the
coordinate axes in the observer’s rest space. While perfectly valid, this procedure
does not respect the spacetime covariance of the underlying relativistic theory. Also,
it obscures the fact that measurements depend on the world line of a lab or an
observer. In this note we will instead use a slightly different approach which is fully
spacetime covariant but in which the time-space decomposition can nevertheless be
just as clearly displayed. This approach also has the advantage of employing only
a single type of indices.

The covariant decomposition will be done by using projection tensors which can
project any tensor into either the worldline of the observer or into the observer’s
rest space. Starting with projection into the rest space, the projection tensor is
defined by
\[ h_{ab} = g_{ab} + u^a u^b \]  
(A.1)
where \( u^a \) is the 4-velocity of the observer and \( g_{ab} \) is the 4-dimensional Kronecker
delta. The latter is usually written with a “\( \delta \)” but we avoid this notation here since
it can be in conflict with \( h_{ab} \) which can serve as a Kronecker delta in the rest space.
The projection into the observer’s worldline has the form
\[ f^a_b = -u^a u^b . \]  
(A.2)
The projection property of the tensors defined in (A.1) and (A.2) is manifested by
the relations
\[ h_{ac} h_{cb} = h_{ab} , \quad f^a_c f^c_b = f^a_b . \]  
(A.3)
Taking a general vector \( V^a \) as an example, its decomposition in time and space
parts takes the form
\[ V^a = f^a_b V^b + h^a_b V^b = -u_b V^b u^a + V^a \]  
(A.4)
where the barred index is used to denote a spatially projected index
\[ V^a = h^a_b V^b . \]  
(A.5)
Note that using this notation, intrinsically spatial tensors such as \( h_{ab} \) or the electric
and magnetic fields for example can be written both with and without bars, for
example \( E_\bar{a} = E_a \). It also follows from its definition that a barred index can be
raised and lowered with respect to either \( g_{ab} \) or \( h_{ab} \), for example
\[ V^\bar{a} = g^{ab} V_\bar{b} = h^{ab} V_\bar{b} . \]  
(A.6)
Employing an observer adapted reference frame, a projection of a vector \( V^a \) into
the worldline of the observer has the single component \( V^0_\bar{a} \), which from the relation
(A.4) can be identified as the scalar \( V^0 = -u_a V^a \). In the same way, lowering
the free index in (A.4) gives \( V^0 = -V^0 = u_a V^a \). We finally note that using this
formalism allows spatial objects to be represented in a fully spacetime covariant
way. Although this has been demonstrated here for vectors, it is clear that the
formalism can be straightforwardly extended to any tensorial objects.

### Appendix B. The 1 + 3 decomposition of the helicity tensor

We start with the 1 + 3 decomposition of the trace of the helicity tensor. Performing
the 1 + 3 decomposition of \( \tilde{J}_a \) gives the following components which are
identical to the helicity vector components in transverse gauge:
\[ \tilde{J}_0 = -\frac{1}{2} (C^a \hat{A}_a - A^a \hat{C}_a) \equiv \frac{1}{2} (C \cdot E - A \cdot B) = - \mathcal{H} \]  
(B.1)
\[ \tilde{J}_a = -\frac{1}{2} (C^b A_{b,a} - A^b C_{b,a}) \equiv \frac{1}{2} (B \times A + C \times B - \frac{1}{2} (C^b A_{b,\bar{a}} - A^b C_{b,\bar{a}}) \)  
(B.2)
As the next task, we perform the $1+3$ decomposition of the real part $Y_{abc}$ of the helicity tensor. This gives

\[ Y_{000} = C_0 \dot{A}_0 - A_0 \dot{C}_0 \equiv 0 \]  
\[ Y_{00c} = C_0 (\nabla A_0)_c - A_0 (\nabla C_0)_c \equiv 0 \]  
\[ Y_{000} = \frac{1}{2} (C_0 \dot{A}_a - A_0 \dot{C}_a) \equiv 0 \]  
\[ Y_{ab0} = C_{(a} \dot{A}_{b)} - A_{(a} \dot{C}_{b)} \equiv -C_{(a} E_{b)} + A_{(a} B_{b)} = h_{ab} \mathcal{H} - \Sigma_{ab} \]  
\[ Y_{0bc} = \frac{1}{2} (C_0 A_{b,c} - A_0 C_{b,c} - A_0 C_{b,c} + C_{b,c}) \equiv 0 \]  
\[ Y_{abc} = C_{(a} A_{b)},c - A_{(a} C_{b)},c. \]

We see that only the parts $Y_{ab0}$ and $Y_{abc}$, with 6 and 18 components respectively, are non-zero in transverse gauge.

Looking at the $1+3$ decomposition of the divergences of $Y_{abc}$, the only nontrivial part is given by

\[ Y_{ab}^{c,e} = C_{(a} \Box A_{b)} - A_{(a} \Box C_{b)} \equiv 0, \]

where the weak equality “$\equiv$” refers to equality modulo the Euler-Lagrange equations of $\mathcal{L}_R$.

In the end, performing the $1+3$ decomposition of $X_{abc}$ gives the following parts which are not identically zero

\[ X_{000} = A_{[a} \dot{A}_{b]} + C_{[a} \dot{C}_{b]} \equiv 0 \]  
\[ X_{ab0} = \frac{1}{2} (A_{(a} \dot{A}_{b)} - A_{(a} \dot{C}_{b)} + C_{a} \dot{C}_{b} - C_{b} \dot{C}_{a}) \equiv E_{[a} A_{b]} + B_{[a} C_{b]} \]  
\[ X_{0bc} = A_{[a} A_{b]},c + C_{[a} C_{b]},c \equiv 0 \]  
\[ X_{abc} = A_{[a} A_{b]},c + C_{[a} C_{b]},c \]

For $X_{abc}$, the $1+3$ parts which are not identically zero are

\[ *X_{000} = \frac{1}{2} \epsilon^{bc}_a X_{bc0} \equiv \frac{1}{2} \epsilon^{bc}_a (A_{[a} \dot{A}_{c]} + A_{[a} A_{c]} + A_{b} \dot{C}_{c} - C_{b} \dot{C}_{c}) \]

\[ *X_{ab0} = \frac{1}{2} (A_{(a} \dot{A}_{b)} - A_{(a} \dot{C}_{b)} + C_{a} \dot{C}_{b} - C_{b} \dot{C}_{a}) \equiv E_{[a} A_{b]} + B_{[a} C_{b]} \]

\[ *X_{0bc} = \frac{1}{2} \epsilon^{de}_b X_{dec} = -\frac{1}{2} \epsilon^{de}_b (A_{d} A_{e,c} + C_{d} C_{e,c}) \]

\[ *X_{abc} = \frac{1}{2} \epsilon^{de}_b X_{dec} \equiv \epsilon_{ab}^e (A_{[a} A_{b]} + C_{[a} C_{b]}) \equiv 0. \]

The anti-symmetric part of $*X_{ab0}$ can be obtained as

\[ *X_{[a]b0} = \frac{1}{2} \epsilon_{ab}^e (A^d A_{e,d} + C^d C_{e,d}) - \frac{1}{4} \epsilon_{abde} (A^d A_{e,d} + C^d C_{e,d}) \]

Based on (B.18) and up to the Lorenz gauge, one can show that (B.16) can be written in the form

\[ *X_{ab0} = \Sigma_{ab} - \bar{K}_{ab}, \]

where $\Sigma_{ab}$ is the infra-zilch, Eq. (2.10) and

\[ \bar{K}_{ab} = \frac{1}{4} \left( 2 \epsilon_{a[b}^d A_{c]} A_d + 2 \epsilon_{a[b}^d C_{c]} C_d - \epsilon_{b}^d (A_d A_d + C_d C_d) \right)^c, \]

which obviously has vanishing divergence with respect to the index $\bar{b}$. 

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Taking the divergence of $\star X_{abc}$, there are two parts not identically zero given by

$$\star X_{abc,c} = \star X_{0\bar{b}c,0} + \star X_{\bar{b}0c} \quad \text{and} \quad TG = -(S_{b} + \Sigma_{bc,c}). \quad (B.21)$$

$$\star X_{\bar{a}bc,c} = \star X_{\bar{a}\bar{b}0} + \star X_{\bar{a}\bar{b}c} \quad \text{and} \quad TG = 0. \quad (B.22)$$

**References**


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